

LECTURE 13

$$\frac{d^2 u}{d\theta^2} + u = - \frac{f(1/u)}{m^2 L^2}$$

$$\left(\frac{du}{d\theta}\right)^2 + u = (E - V)$$

$$m(\ddot{r} - r\dot{\theta}^2) = f(r)$$

$$\dot{r}^2 + r^2 \dot{\theta}^2 = \frac{2E}{m} - V(r)$$

$$r = 1/u \quad .$$

$$\boxed{\text{II}} \quad V(r) = -\alpha/r$$

$$V_{\text{eff}}(r) = \frac{L^2}{2mr} - \frac{\alpha}{r}$$

$$\left(\frac{du}{d\theta}\right)^2 + u^2 = \frac{2m}{L^2}(E - V)$$

$$\Rightarrow \left(\frac{du}{d\theta}\right)^2 = -u^2 + \frac{2mE}{L^2} + \frac{2m\alpha}{L^2}u$$

Complete the square

$$\left(\frac{du}{d\theta}\right)^2 = -\left(u - \frac{m\alpha}{L^2}\right)^2 + B^2,$$

$$\left(\frac{dz}{d\theta}\right)^2 = -z^2 + B^2 \quad \Rightarrow$$

$$\int \frac{dz}{\sqrt{B^2 - z^2}} = \int d\theta = \theta - \theta_0$$

$$\frac{1}{r} = \frac{m\alpha}{L^2} (1 + \epsilon \cos \theta),$$

$$\epsilon = \sqrt{1 + \frac{2EL^2}{m\alpha^2}}$$

$$\epsilon = 0 \Rightarrow r = \frac{L^2}{m\alpha} = \text{const.} \rightarrow \text{circular orbit}$$

$$V_{\text{eff}}(r) = \frac{L^2}{2mr^2} - \frac{\alpha}{r}$$

$$E = - \frac{m\alpha^2}{2L^2}$$

$$\frac{dV_{\text{eff}}(r)}{dr} = -\frac{L^2}{mr^3} + \frac{\alpha}{r^2} = 0$$

$$\Rightarrow r = \frac{L^2}{m\alpha} = r_*$$

$$V_{\text{eff}}(r_*) = \frac{L^2}{2mr_*^2} - \frac{\alpha}{r_*} = \frac{m\alpha^2}{2L^2} - \frac{m\alpha^2}{L^2} = - \frac{m\alpha^2}{2L^2}$$

$$\frac{\epsilon > 1}{\downarrow} \quad \frac{0 < \epsilon < 1}{\rightarrow}$$

$$\frac{1}{r} = \frac{m\alpha}{L^2} (1 + \epsilon \cos \theta)$$

$$= \frac{m\alpha}{L^2} \epsilon \left(\cos \theta + \frac{1}{\epsilon} \right)$$

$$r_{\max} = \infty$$

$$r_{\min} = \frac{L^2}{m\alpha(1+\epsilon)}$$

$$r_{\max} = \frac{L^2}{m\alpha(1-\epsilon)}$$

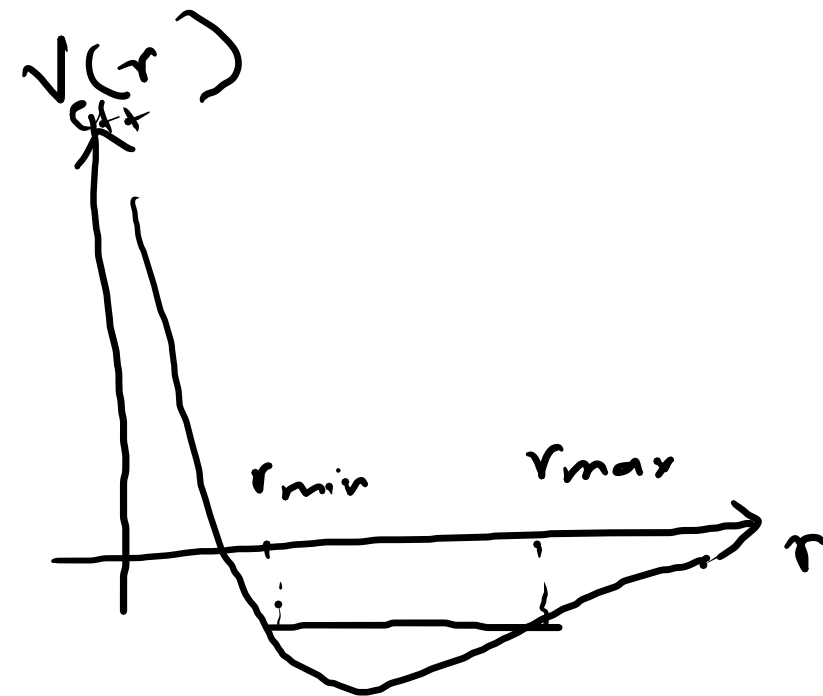
$$r_{\min} = \frac{L^2}{m\alpha(1+\epsilon)}$$

$$\underline{0 < \epsilon < 1:-}$$

$$0 < 1 + \frac{2EL^2}{m\alpha^2} < 1$$

$$\Rightarrow -1 < \frac{2EL^2}{m\alpha^2} < 0$$

$$\Rightarrow -m\alpha^2/(2L^2) < E < 0$$



Geometry of path/orbit

$$\frac{1}{r} = \left(\frac{m\alpha}{L^2} \right) (1 + \epsilon \cos \theta)$$

$$0 < \epsilon < 1$$

$$\epsilon = 1$$

$$\epsilon > 1$$

To work out geometry of orbit,
easier to use Cartesian coordinates. (x, y) .

$$k \equiv \frac{L^2}{m\alpha}$$

$$\cos \theta = x/r$$

$$\frac{1}{r} = \frac{1}{k} (1 + \epsilon \cos \theta)$$

$$\Rightarrow k = r (1 + \epsilon \cos \theta)$$

$$\Rightarrow k = r (1 + \epsilon x/r) = r + \epsilon x$$

$$k = r + \epsilon x$$

Squaring,

$$r^2 = (k - \epsilon x)^2$$

$$\Rightarrow r^2 = k^2 + \epsilon^2 x^2 - 2\epsilon kx$$

$$\Rightarrow x^2 + y^2 = k^2 + \epsilon^2 x^2 - 2\epsilon kx$$

Case :- $0 < \epsilon < 1$

$$\Rightarrow (1 - \epsilon^2)x^2 + 2\epsilon kx + y^2 = k^2$$

$$\Rightarrow x^2 + 2x \frac{k\epsilon}{1 - \epsilon^2} + \frac{y^2}{1 - \epsilon^2} = \frac{k^2}{(1 - \epsilon^2)}$$

$$\Rightarrow x^2 + 2x \frac{k\epsilon}{1 - \epsilon^2} + \left(\frac{k\epsilon}{1 - \epsilon^2} \right)^2 - \underbrace{\left(\frac{k\epsilon}{1 - \epsilon^2} \right)^2} + \frac{y^2}{1 - \epsilon^2} = \frac{k^2}{1 - \epsilon^2}$$

$$\underline{\epsilon = 1} \quad 1 + \frac{2EL^2}{m\alpha^2} = 1$$

$$\Rightarrow E = 0$$

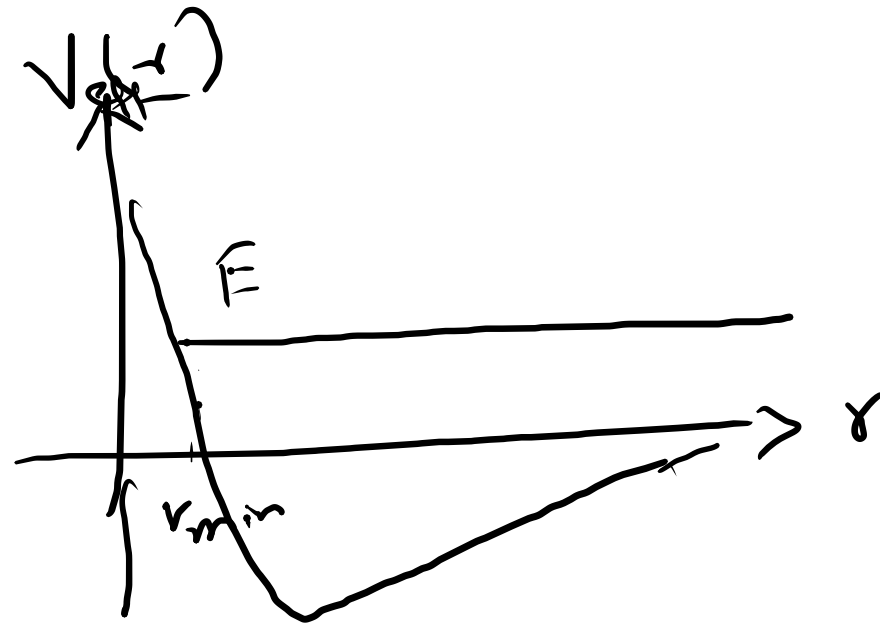
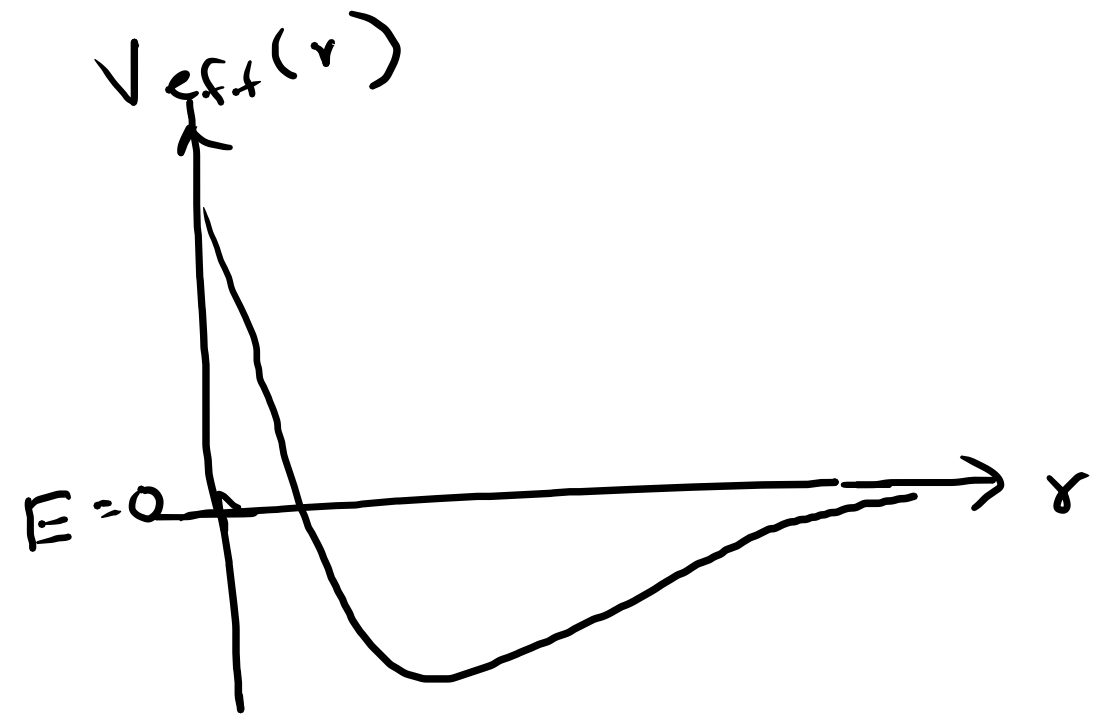
and $V_{\text{eff}} \rightarrow 0$ for large r .

$$v|_{r \rightarrow \infty} = 0$$

$$\underline{\epsilon > 1} !-$$

$$\frac{1}{2} m v_{\infty}^2 = E.$$

$$\Rightarrow v_{\infty} = \sqrt{\frac{2E}{m}}$$



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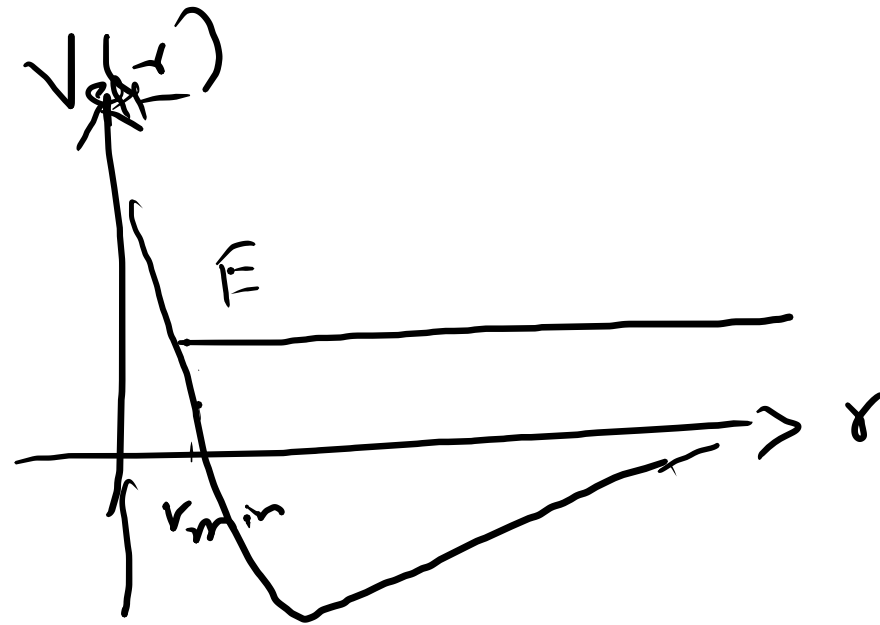
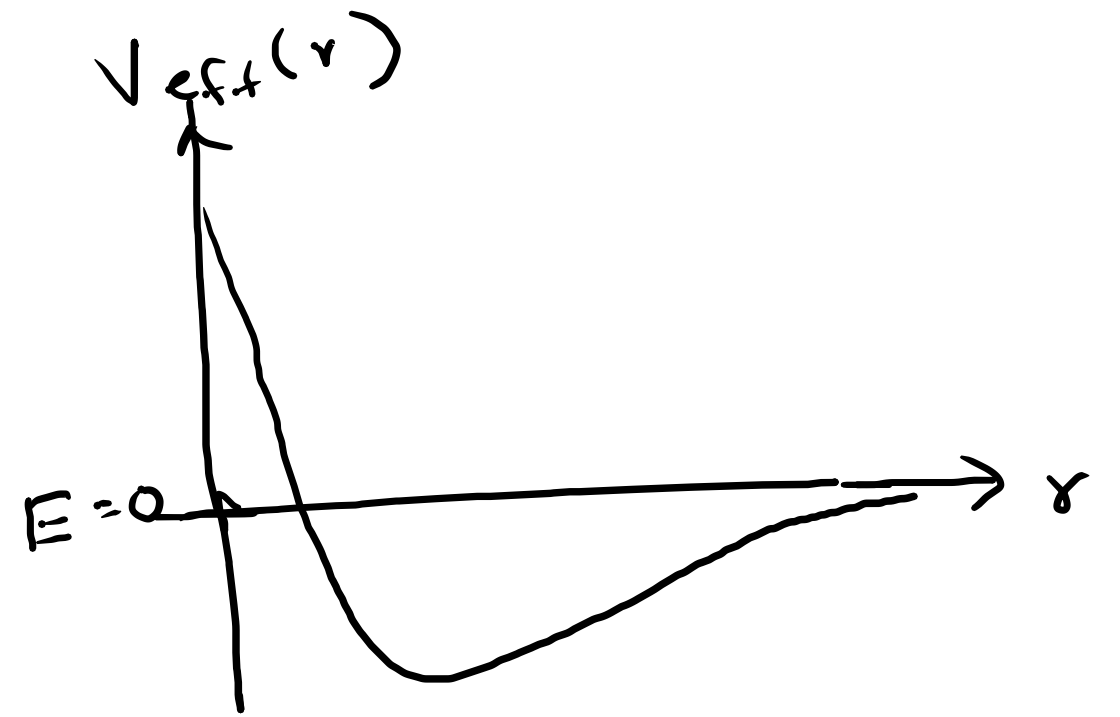
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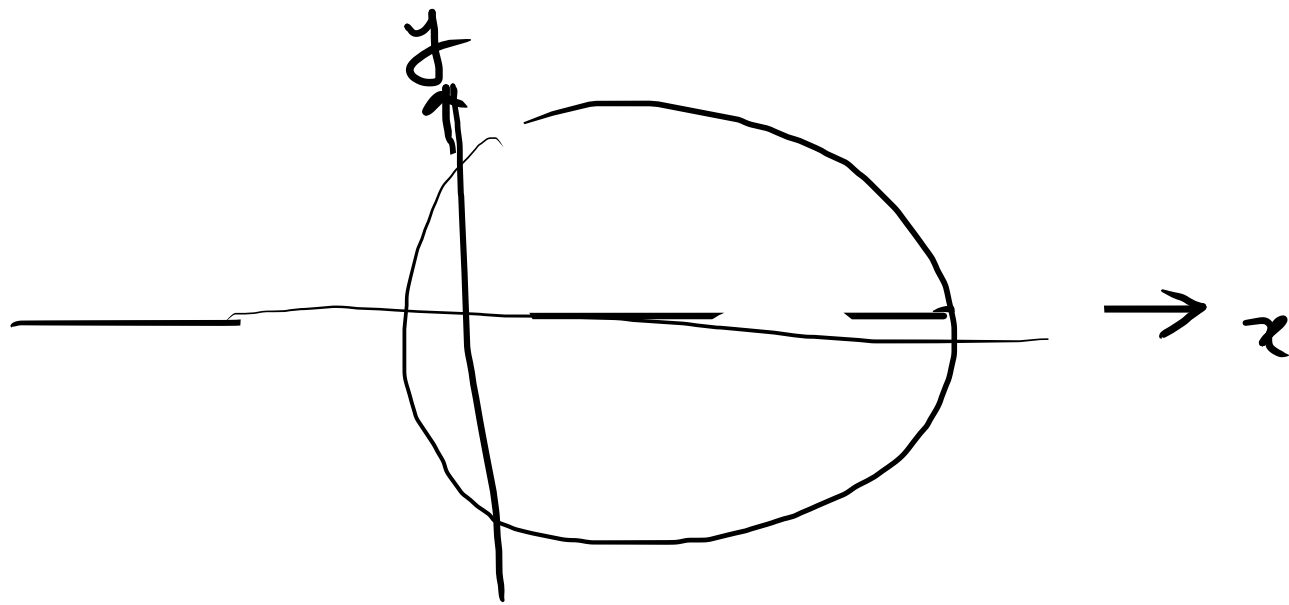
$$\Rightarrow \left(x + \frac{ke}{1-e^2}\right)^2 + \frac{y^2}{1-e^2} = \frac{k^2}{1-e^2} \left[1 + \frac{e^2}{1-e^2}\right] = \frac{k^2}{(1-e^2)^2}$$

$$\Rightarrow \frac{(x-c)^2}{a^2} + \frac{y^2}{b^2} = 1$$

$$a = \frac{k}{1-e^2}$$

$$b = \frac{k}{\sqrt{1-e^2}}$$

$$a = \frac{b}{\sqrt{1-e^2}}$$




Case:- $\epsilon = 1$

$$\cancel{x}^2 + y^2 = k^2 + \cancel{x}^2 - 2kx$$

$$\Rightarrow y^2 = k^2 - 2kx \rightarrow \text{parabola.}$$

Case:- $\epsilon > 1$

$$\frac{\left(x - \frac{k\epsilon}{\epsilon^2 - 1}\right)^2}{a^2} - \frac{y^2}{b^2} = 1 \rightarrow \text{hyperbola.}$$


$$\boxed{V(r) = -\alpha/r}$$

$$F = -\frac{dV}{dr} \propto \frac{1}{r^2}$$

Gravitational potential



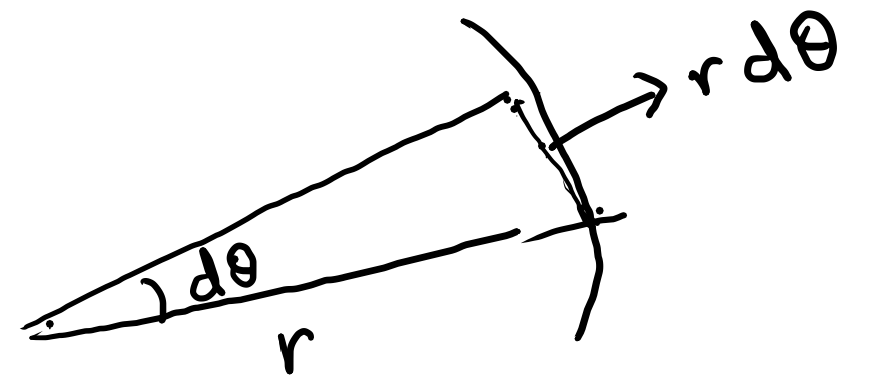
Kepler's law.

Kepler's law

- (i) Planets move around sun in elliptic orbits.
- (ii) Areal velocity of planets is const.

$$dA = \frac{1}{2} r (r d\theta)$$

$$\frac{dA}{dt} = \frac{1}{2} r^2 \dot{\theta} = \frac{L}{2m} = \text{constant} = \text{areal velocity.}$$



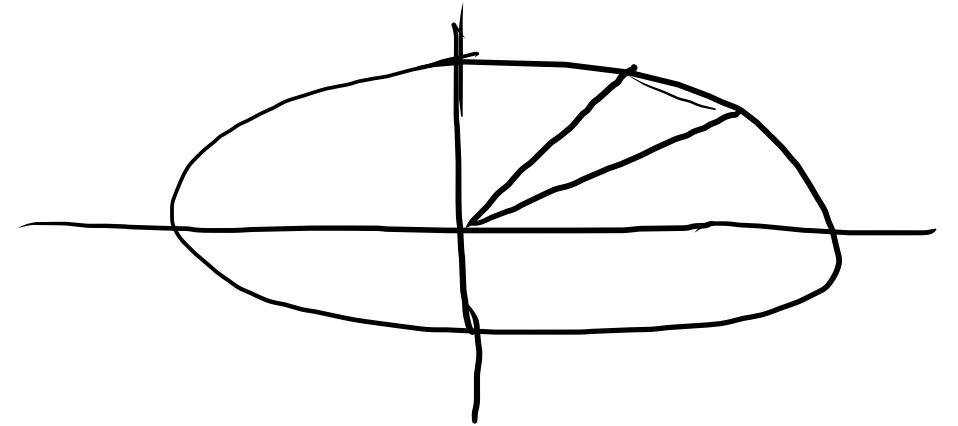
(iii) Time period T of planets around the sun,

$$\boxed{T^2 \propto a^3}$$

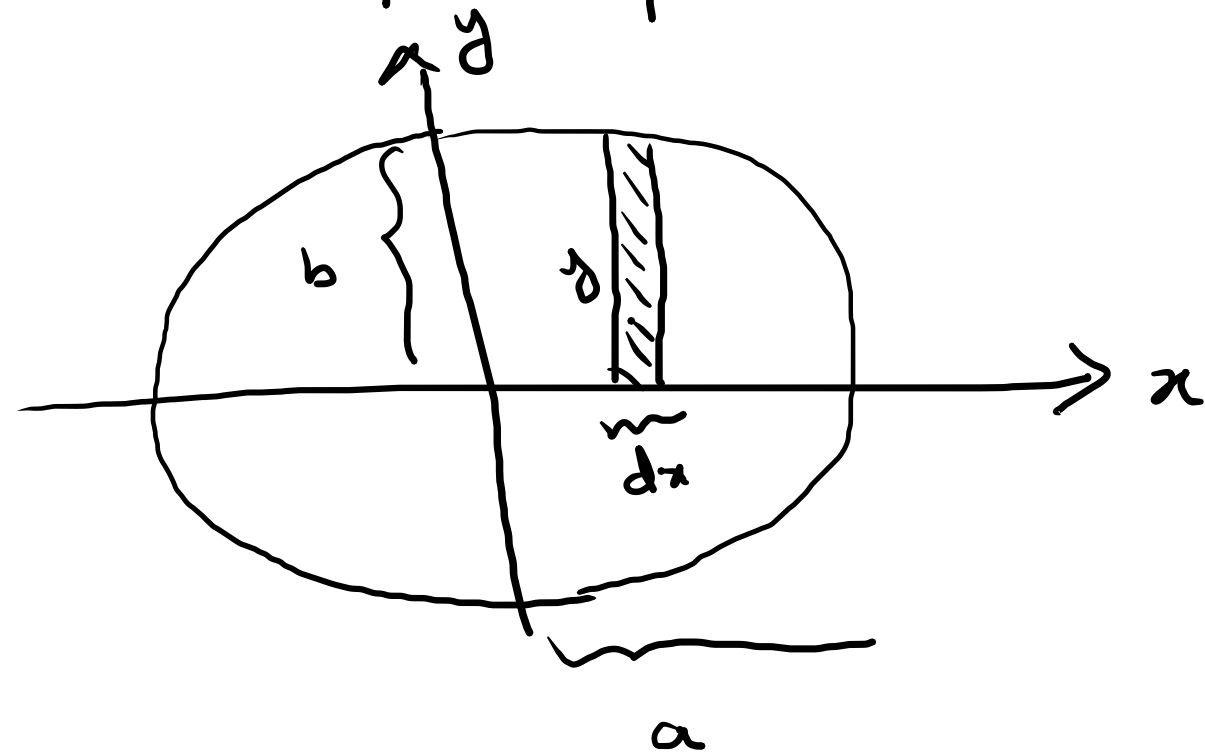
Proof:- Path/orbit is an area

$$T = \frac{A}{\text{areal velocity}}$$

$$= \frac{A}{(L/2m)}$$



— Area of ellipse :



Area of shaded region

$$= \int_0^a dx \, y(x)$$

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

$$\frac{A}{4} = \int_0^a dx \, y(x) = ab \int_0^a dx \sqrt{a^2 - x^2}$$

Sub:- $x = a \sin \alpha$

$$\frac{A}{4} = ab \int_0^{\pi/2} d\alpha \cos^2 \alpha = \frac{ab\pi}{4} \Rightarrow A = \pi ab$$

$$\Rightarrow \pi ab = \frac{LT}{2m}$$

$$\Rightarrow \pi^2 a^2 b^2 = \frac{L^2}{4m^2} T^2$$

$$\Rightarrow \pi^2 a^4 = \frac{m \alpha k}{m(1-\epsilon^2)} \frac{T^2}{4m}$$

$$\Rightarrow \pi^2 a^4 \propto a T^2$$

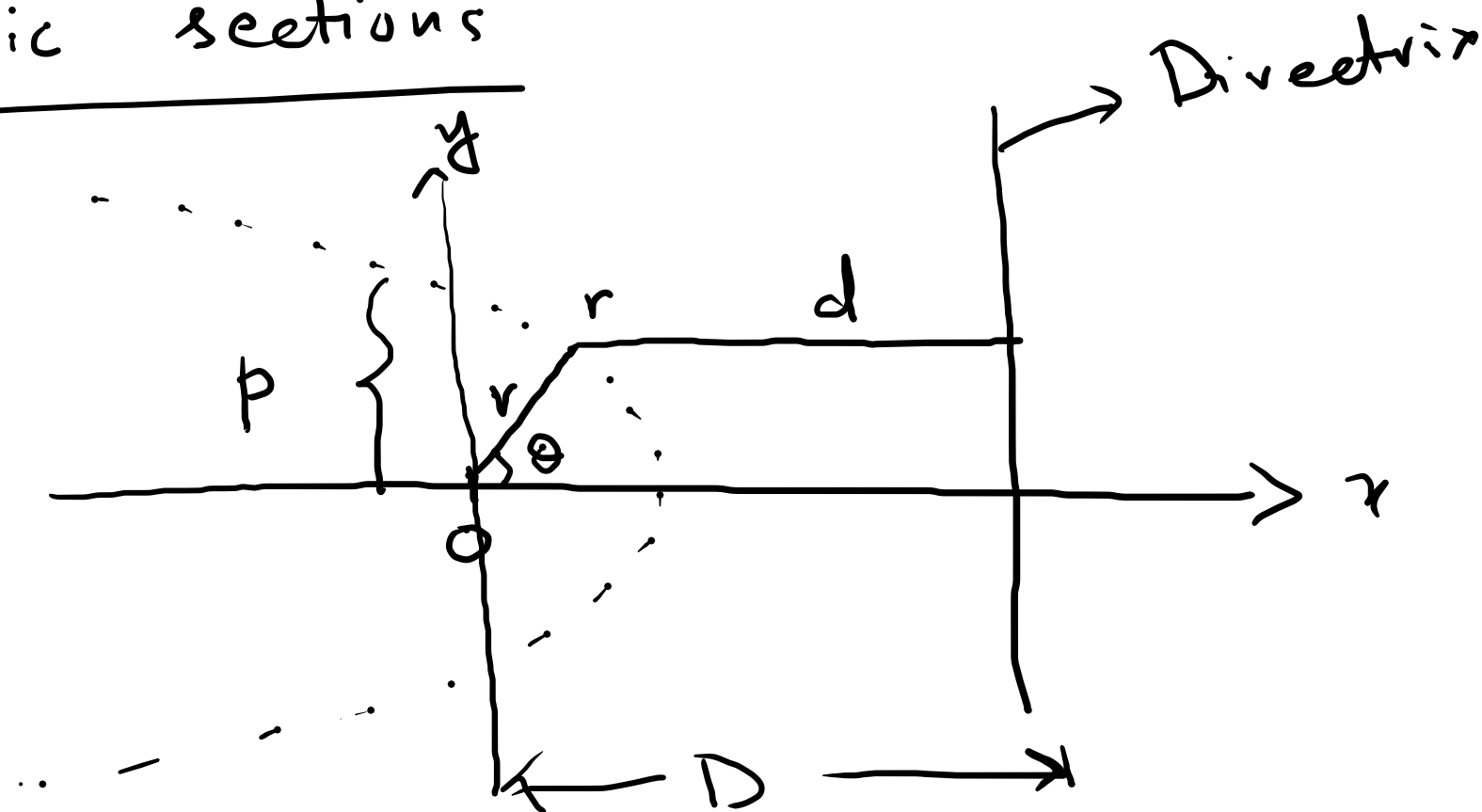
$$\Rightarrow T^2 \propto a^3$$

$$\Rightarrow \boxed{T^2 \propto a^3}$$

$$a = \frac{b}{\sqrt{1-\epsilon^2}} \quad .$$



Conic sections



$O \equiv$ fixed pt.

$$\frac{r}{d} = \boxed{\epsilon = \text{const.}} = \boxed{\frac{p}{D}}$$

$$D = d + r \cos \theta$$

$$\Rightarrow p = \epsilon (d + r \cos \theta) = \epsilon d + \epsilon r \cos \theta$$

$$\Rightarrow \boxed{r = \frac{p}{1 + \epsilon \cos \theta}}$$

\rightarrow general eqn. for conic section.