

acceleration of $\vec{a}_i = \sum_{\substack{j=1 \\ j \neq i}}^N G m_j \frac{\vec{r}_j - \vec{r}_i}{(|\vec{r}_j - \vec{r}_i|^2 + \epsilon^2)^{3/2}}$

$i = \text{particle 1}$
 $j = \text{particle 2}$

$\leftarrow N\text{-body}$
 $\leftarrow \text{distance between}$
 $\leftarrow \text{gravity}$
 $\leftarrow \text{softening parameter} \rightarrow \epsilon \approx 0.01 \times \boxed{\frac{R_{\text{cluster}}}{N^{1/3}}}$ (mean inter-particle spacing)

$$r_{ij}^2 = (x_j - x_i)^2 + (y_j - y_i)^2 + (z_j - z_i)^2 \quad \leftarrow \text{distance}$$

use M_\odot , AU, yr!

↓

$$G = 4\pi^2 (\text{AU}^3 M_\odot^{-1} \text{yr}^{-2})$$

general structure:

Integration methods:

$$\begin{aligned} \dot{\vec{y}} &= \vec{f}(t, \vec{y}) = [\vec{v}, \vec{a}] \\ \vec{y} &= [\vec{r}, \vec{v}] \end{aligned}$$

- eulers method

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \cdot \boxed{\vec{f}(t_n, \vec{y}_n)}$$

same

- runge-kutta method

- RK2

$$\begin{aligned} \vec{k}_1 &= \vec{f}(t_n, \vec{y}_n) \\ \vec{k}_2 &= \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_1\right) \end{aligned}$$

same thing

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \cdot \vec{k}_2$$

same

- RK4

$$\begin{aligned} \vec{k}_1 &= \vec{f}(t_n, \vec{y}_n) \\ \vec{k}_2 &= \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_1\right) \\ \vec{k}_3 &= \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_2\right) \\ \vec{k}_4 &= \vec{f}(t_n + \Delta t, \vec{y}_n + \Delta t \cdot \vec{k}_3) \end{aligned}$$

same thing

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \frac{1}{6} (\vec{k}_1 + 2\vec{k}_2 + 2\vec{k}_3 + \vec{k}_4)$$

same

Monitor + plot:

- total KE
- total grav. PE = W
- total E = KE + W
- virial ratio

$$Q = \frac{|2KE + W|}{|W|}$$

should be < 0.01

* count each pair only once ($i < j$)

- symplectic integrators = leapfrog

*similar $\vec{v}_{n+1/2} = \vec{v}_n + \Delta t \cdot \frac{1}{2} \cdot \vec{a}(\vec{r}_n)$

similar $\vec{r}_{n+1} = \vec{r}_n + \Delta t \cdot \vec{v}_{n+1/2}$

similar $\vec{v}_{n+1} = \vec{v}_{n+1/2} + \Delta t \cdot \frac{1}{2} \cdot \vec{a}(\vec{r}_{n+1})$

Sampling:

- np.random.uniform (low, high, array shape)

- krape IMF:

$$\alpha = \begin{cases} 1.3 & \text{for } 0.08 \leq m < 0.5 \\ 2.3 & \text{for } 0.5 \leq m \end{cases}$$

$$\xi = \begin{cases} A_1 m_1^{-\alpha_1} \\ A_2 m_2^{-\alpha_2} \end{cases}$$

$$I = \begin{cases} (m_{0.5}^{1-\alpha_1} - m_{\min}^{1-\alpha_1}) / (1-\alpha_1) \\ (m_{\max}^{1-\alpha_2} - m_{0.5}^{1-\alpha_2}) / (1-\alpha_2) \end{cases}$$

$$A = \begin{cases} N / (I_1 + m_{0.5}^{\alpha_2-\alpha_1} \cdot I_2) \\ A_1 \cdot m_{0.5}^{\alpha_2-\alpha_1} \end{cases}$$

$$P = \begin{cases} A_1 \cdot I_1 / N \\ A_2 \cdot I_2 / N \end{cases}$$

* m on [a,b], $\alpha \neq 1$ ← do per segment

$$C = (m_b^{1-\alpha} - m_a^{1-\alpha}) / (1-\alpha)$$

$u \sim U(0,1)$ ← uniform distribution

$$m = \left(m_a^{1-\alpha} + u(m_b^{1-\alpha} - m_a^{1-\alpha}) \right)^{-1/(1-\alpha)}$$

Normalize ξ_1 and ξ_2

Calculate P_1, P_2 from ξ

$s \sim U(0,1)$ ← random uniform

If $s < P_1$, sample from seg. 1: $\alpha_1, m \in [\min, 0.5]$
else, sample from seg. 2: $\alpha_2, m \in [0.5, \max]$

Testing:

all 4 methods

- Sun + Earth
 $r = 1 \text{ AU}$, $v_{\text{circular}} = \sqrt{\frac{GM}{r}}$
10 yrs, $dt = 0.01$

RK4
leapfrog

- Sun + Earth + Jupiter
sun at origin
Earth at 1 AU ($3 \times 10^{-6} M_\odot$)
Jup at 5.2 AU ($9.55 \times 10^{-4} M_\odot$)
24 yrs

- N-body

uniform mass $1 M_\odot$

random positions

Radius = 100 AU

$$\vec{v}_i = 0$$

P1

then
leapfrog

Krape IMF sampling
 $m \in [0.08, 150] M_\odot$

Plummer sphere positions
sphere radius a

virial equilibrium \sim
 $v_i \sim N(0, \sigma_{10}(r_i))$

test with loops of
 $N = 10, 20, 50, 100$

then

rewrite force eq. with
Numpy broadcasting

No loops!

array operations

$r_{ij} = \text{shape}(N, N, 3)$

test with

$N = 10, 20, 50, 100,$

200, 500, 1000, 2000,

5000, maybe 10_000

P3

Required Plots:

- 2 body

- 2x2 comparison of

1. orbital trajectories (line styles)
2. total E vs time (colors)
3. relative E error (log)
4. virial ratio

- 3 body

- same as above, just only 2 methods
- can split orbital into own plot

- N-body

- P1

- energy diagnostics only
 - ↳ compare RK4 to leapfrog

- P2

- initial cond. validation (N=100)

- log-log mass histogram w/ kroupa slopes (-1.3, -2.3)

- radial density profile showing plummer $\rho \propto (1 + \frac{r^2}{a^2})^{-5/2}$

- cluster evol. snapshots at $\frac{1}{4}$ intervals through time

- P3

- verify E comp. for loop vs. vectorized (N=100)

- performance analysis

- log log time vs N

- overlay $O(N^2)$ ref. line

- speed-up factor $\left(\frac{t_{\text{loop}}}{t_{\text{vec}}} \right)$ vs N

- large N showcase

- 3x2 evol. snapshots at $\frac{1}{5}$ intervals

- color by mass

- energy diagnostics

- plummer sphere positions

$$\rho = \frac{3M_0}{4\pi a^3} \left(1 + \frac{r^2}{a^2}\right)^{-3/2}$$

M_0 = total mass
 a = plummer radius

$$\Phi = \frac{-GM_0}{\sqrt{r^2 + a^2}}$$

$$\sigma^2 = \frac{GM_0}{6\sqrt{r^2 + a^2}}$$

velocity dispersion $f = \frac{24\sqrt{2}}{7\pi^3} \frac{a^2}{G^3 M_0^4} (-E)^{3/2} \leftarrow \text{isotropic distribution function}$

↳ If $E < 0$ and $f = 0$

else $E = \frac{1}{2} v^2 + \Phi \leftarrow \text{specific energy}$

$$M(<r) = \frac{M_0 r^3}{(r^2 + a^2)^{3/2}} \leftarrow \text{mass within } a \text{ radius}$$

core radius $\rightarrow r_c = a \sqrt{\sqrt{2} - 1}$

half-mass radius $\rightarrow r_h = \left(\frac{1}{0.5^{2/3}} - 1 \right)^{-0.5} \cdot a$

virial radius $\rightarrow r_v = \frac{16}{8\pi} a$

$$\begin{aligned}x_e &\rightarrow y_s \\x_s &\rightarrow x_s \\x_s &\rightarrow x_e\end{aligned}$$

$$ps = \begin{bmatrix} x_s & y_s & z_s \\ x_e & y_e & z_e \\ x_j & y_j & z_j \end{bmatrix} = \text{shape}[N, d]$$

$$ps[\text{none}, :, :] = \text{shape}[1, N, d]$$

$$ps[:, \text{none}, :] = \text{shape}[N, 1, d]$$

$$\text{diff} = \text{shape}[N, N, d]$$

$$m = \text{shape}[1, N, 1] = \begin{matrix} \downarrow \text{D2} \\ \begin{bmatrix} 0 & 0 & 0 \\ x_{es} & y_{es} & z_{es} \\ x_{js} & y_{js} & z_{js} \end{bmatrix} \cdot m_e \quad \dots \quad \begin{bmatrix} x_{se} & y_{se} & z_{se} \\ 0 & 0 & 0 \\ x_{je} & y_{je} & z_{je} \end{bmatrix} \cdot m_s \quad \dots \quad \begin{bmatrix} x_{sj} & y_{sj} & z_{sj} \\ x_{ej} & y_{ej} & z_{ej} \\ 0 & 0 & 0 \end{bmatrix} \cdot m_e \end{matrix}$$

$\rightarrow \text{D3} \quad \quad \quad \rightarrow \text{D1}$

receiving force \downarrow acting force \downarrow dimension \downarrow

$$\vec{a}_i = \sum_{\substack{j=1 \\ j \neq i}}^N G m_j \frac{\vec{r}_j - \vec{r}_i}{(|\vec{r}_j - \vec{r}_i|^2 + e^2)^{3/2}}$$

\uparrow

$$m_j \cdot \vec{r}_j - \vec{r}_i$$

$$\text{diff}_{sx} = \text{diff}[:, 0, 0]$$

$M_0 = \text{total mass}$

$a = \text{scale radius (given)}$

$$\rho(r) = \frac{3M}{4\pi a^3} \left(1 + \frac{r^2}{a^2}\right)^{-5/2}$$

$$M(<r) = M_0 \frac{r^3}{(r^2 + a^2)^{3/2}}$$

$$F(r) = M(<r) / M \in [0, 1)$$

$$F(r) = u, \quad u \sim U(10^{-12}, 1 - 10^{-12})$$

$$r = a(u^{-2/3} - 1)^{-1/2}$$

$$\phi = 2\pi u, \quad u \sim U(0, 1)$$

$$\cos \Theta = 1 - 2\omega, \quad \omega \sim U(0, 1)$$

$$x = r \sin \Theta \cos \phi$$

$$y = r \sin \Theta \sin \phi$$

$$z = r \cos \Theta$$

$$r_h = \frac{a}{\sqrt{z^{2/3} - 1}}$$

$$a = r_h \sqrt{z^{2/3} - 1}$$

$$\vec{y} = [\vec{r}, \vec{v}]$$

Modules

constants ✓

integration methods

sampling methods

plotting

testing

acting forces

eulers ($\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$)

RK ($\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$)

leapfrog ($\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$)

random (cluster radius, N \rightarrow pos)

kroupa IMF (N, m_min, m_max \rightarrow mass)

plummer sphere (cluster radius, N, mass-total \rightarrow pos)

orbits (pos, t, N, ax, fig \rightarrow ax, fig)

— vs time (—, t, N, ax, fig \rightarrow ax, fig)

— vs N (—, t, N, ax, fig \rightarrow ax, fig)

evol. screenshots (pos, t_i, ax, fig \rightarrow ax, fig)

2 body

3 body

N body

P1

P2

P3

gravity (m_i , separation, $\epsilon \rightarrow \vec{a}_i$)

ϵ (cluster radius, N $\rightarrow \epsilon$)

pos = $[x, y, z] \downarrow_N$

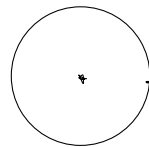
vel = $[v_x, v_y, v_z] \downarrow_N$

acc = $[a_x, a_y, a_z] \downarrow_N$

mass = $[m] \downarrow_N$

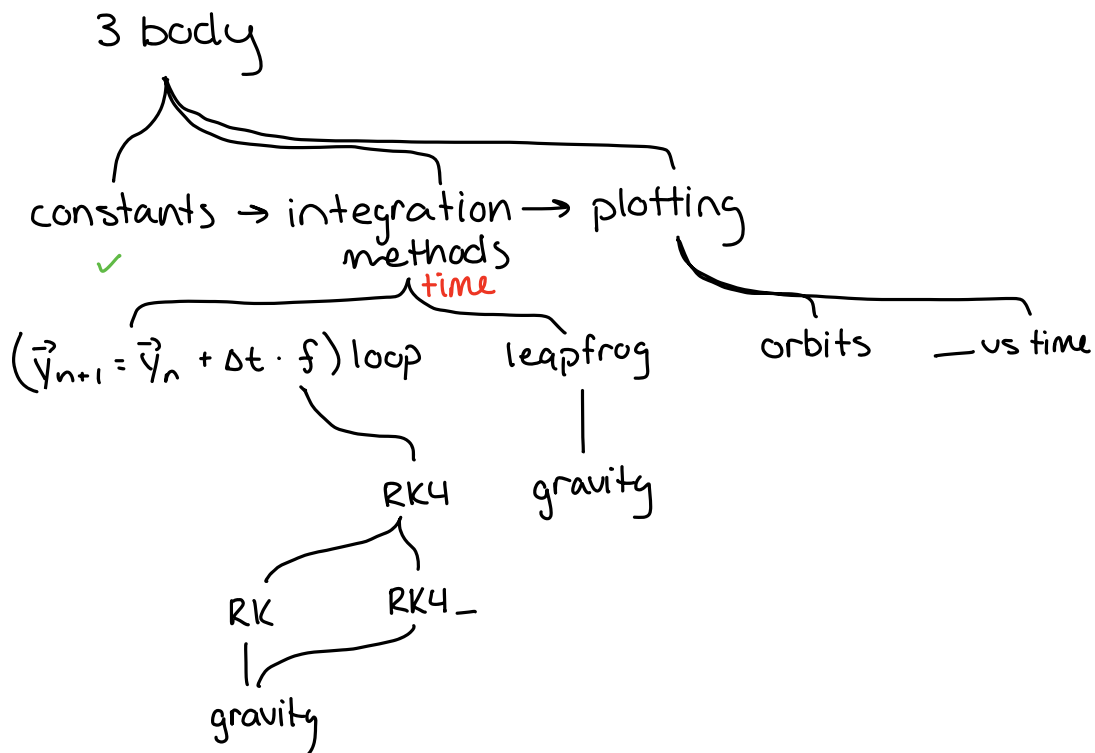
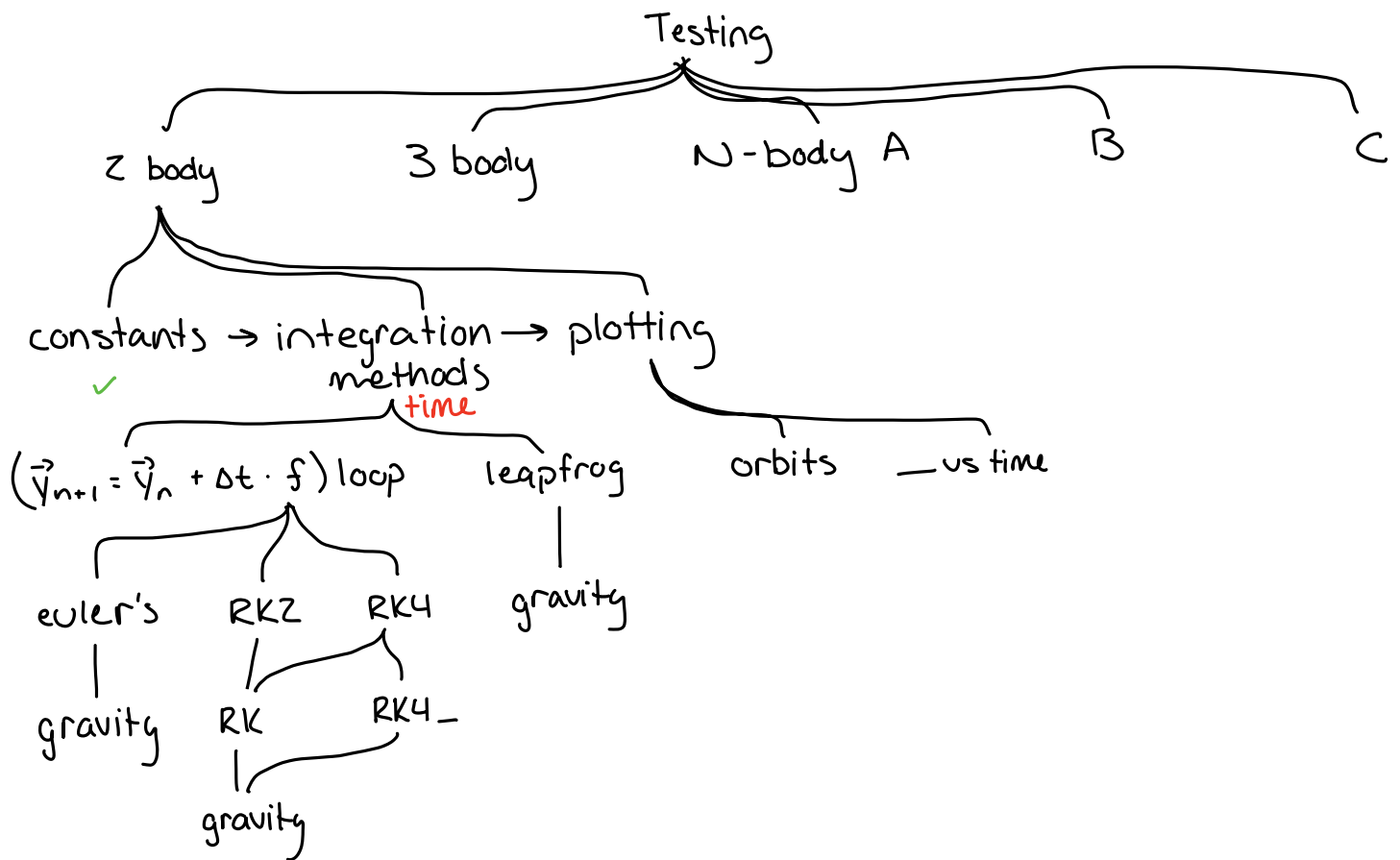
system_t = [system, t] $\downarrow_{\text{steps}}$

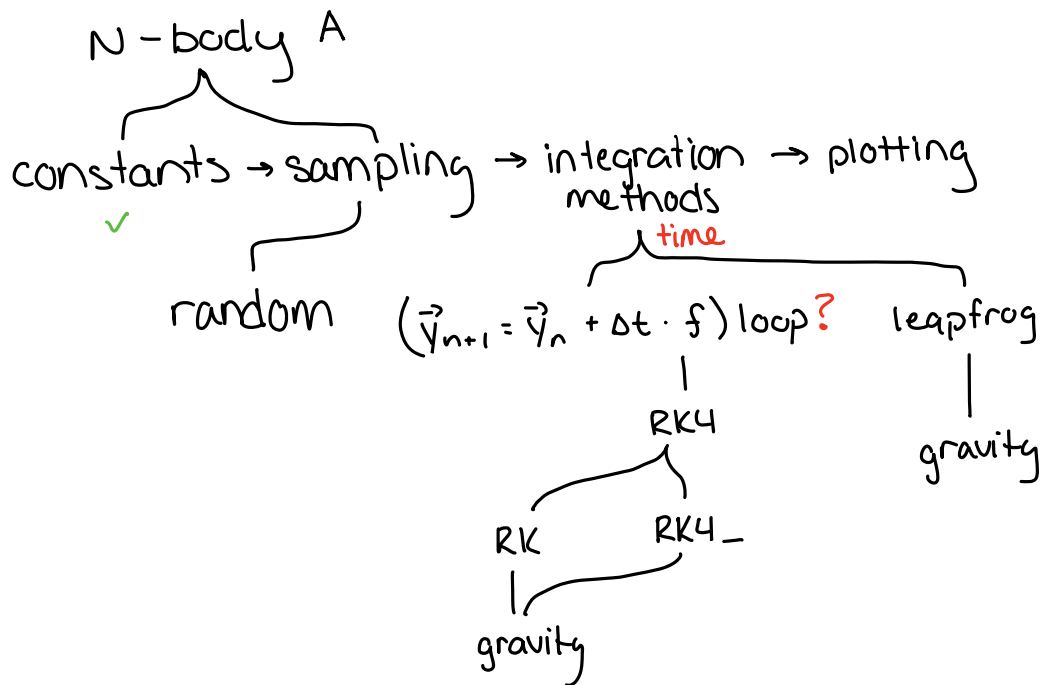
system = N-body



$$v = \frac{d}{t} = \frac{2\pi(1 \text{ AU})}{1 \text{ yr}} = \frac{2\pi(5.2 \text{ AU})}{12 \text{ yr}}$$

2π





class NbodyIntegrator

pos = x y z
x2 y2 z2
x3 y3 z3

masses
positions
velocities

€
method

functions:

init

acceleration

step

energy