

acceleration of  
 $\vec{a}_i = \sum_{\substack{j=1 \\ j \neq i}}^N G m_j \frac{\vec{r}_j - \vec{r}_i}{(|\vec{r}_j - \vec{r}_i|^2 + \epsilon^2)^{3/2}}$ 
← gravity

$i = \text{particle 1}$   
 $j = \text{particle 2}$

$\epsilon$  softening parameter  $\rightarrow \epsilon \approx 0.01 \times \frac{R_{\text{cluster}}}{N^{1/3}}$   
mean inter-particle spacing

$$r_{ij}^2 = (x_j - x_i)^2 + (y_j - y_i)^2 + (z_j - z_i)^2 \quad \leftarrow \text{distance}$$

use  $M_0$ , AU, yr!



$$G = 4\pi^2 (\text{AU}^3 M_0^{-1} \text{yr}^{-2})$$

general structure:

Integration methods:  $\dot{\vec{y}} = \vec{f}(t, \vec{y}) = [\vec{v}, \vec{a}]$

$$\vec{v} = [\vec{r}, \vec{\dot{r}}]$$

- eulers method

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \cdot \vec{f}(t_n, \vec{y}_n)$$

same

- runge-kutta method

- RK2

$$\vec{k}_1 = \vec{f}(t_n, \vec{y}_n)$$

$$\vec{k}_2 = \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_1\right)$$

same thing

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \cdot \vec{k}_2$$

same

- RK4

$$\vec{k}_1 = \vec{f}(t_n, \vec{y}_n)$$

$$\vec{k}_2 = \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_1\right)$$

$$\vec{k}_3 = \vec{f}\left(t_n + \frac{\Delta t}{2}, \vec{y}_n + \frac{\Delta t}{2} \vec{k}_2\right)$$

$$\vec{k}_4 = \vec{f}(t_n + \Delta t, \vec{y}_n + \Delta t \cdot \vec{k}_3)$$

$$\vec{y}_{n+1} = \vec{y}_n + \Delta t \cdot \frac{1}{6} (\vec{k}_1 + 2\vec{k}_2 + 2\vec{k}_3 + \vec{k}_4)$$

same

Monitor + plot:

- total KE
- total grav. PE = W
- total E = KE + W
- virial ratio

$$G = \frac{|2KE + W|}{|W|}$$

should be  $< 0.01$

\* count each pair only once ( $i < j$ )

- symplectic integrators = leapfrog

$$\text{similar } \vec{N}_{n+\frac{1}{2}} = \vec{N}_n + \Delta t \cdot \frac{1}{2} \cdot \vec{\alpha}(\vec{r}_n)$$

$$\text{similar } \vec{r}_{n+1} = \vec{r}_n + \Delta t \cdot \vec{v}_{n+\frac{1}{2}}$$

$$\text{similar } \vec{v}_{n+\frac{1}{2}} = \vec{v}_{n+\frac{1}{2}} + \Delta t \cdot \frac{1}{2} \cdot \vec{\alpha}(\vec{r}_{n+1})$$

## Testing:

- Sun + Earth

$$r = 1 \text{ AU}, N_{\text{angular}} = \frac{\sqrt{GM}}{r}$$

10 yrs, dt = 0.01

- Sun + Earth + Jupiter

sun at origin

Earth at 1 AU ( $3 \times 10^{-6} M_\odot$ )

Jup at 5.2 AU ( $9.55 \times 10^{-4} M_\odot$ )

24 yrs

- N-body

uniform mass  $1 M_\odot$

random positions

Radius = 100 AU

$$\vec{v}_i = 0$$

then  
leapfrog  
Kraupa IMF sampling  
 $m \in [0.08, 150] M_\odot$

Plummer sphere positions  
sphere radius  $a$   
virtual equilibrium  $\sim$   
 $v_i \sim \mathcal{N}(0, \sigma_{v0}(r_i))$

test with loops of  
 $N = 10, 20, 50, 100$

then  
rewrite force eq. with  
NumPy broadcasting  
No loops!

P3  
array operations  
 $r_{ij} = \text{shape}(N, N, 3)$   
test with  
 $N = 10, 20, 50, 100,$   
 $200, 500, 1000, 2000,$   
 $5000, \text{ maybe } 10000$

## Sampling:

- np.random.uniform (low, high, array shape)

- Kraupa IMF:

$$\alpha = \begin{cases} 1.3 & \text{for } 0.08 \leq m < 0.5 \\ 2.3 & \text{for } 0.5 \leq m \end{cases}$$

$$\xi = \begin{bmatrix} A_1 m_1^{-\alpha_1} \\ A_2 m_2^{-\alpha_2} \end{bmatrix}$$

$$I = \begin{bmatrix} (m_{0.5}^{1-\alpha_1} - m_{\min}^{1-\alpha_1}) / (1 - \alpha_1) \\ (m_{\max}^{1-\alpha_2} - m_{0.5}^{1-\alpha_2}) / (1 - \alpha_2) \end{bmatrix}$$

$$A = \begin{bmatrix} N / (I_1 + m_{0.5}^{\alpha_2 - \alpha_1} \cdot I_2) \\ A_1 \cdot m_{0.5}^{\alpha_2 - \alpha_1} \end{bmatrix}$$

$$P = \begin{bmatrix} A_1 \cdot I_1 / N \\ A_2 \cdot I_2 / N \end{bmatrix}$$

$$\left[ \begin{array}{l} * m \text{ on } [a, b], \alpha \neq 1 \leftarrow \text{do per segment} \\ C = (m_b^{1-\alpha} - m_a^{1-\alpha}) / (1 - \alpha) \\ U \sim U(0, 1) \leftarrow \text{uniform distribution} \\ m = \left( m_a^{1-\alpha} + U(m_b^{1-\alpha} - m_a^{1-\alpha}) \right)^{1/(1-\alpha)} \end{array} \right]$$

Normalize  $\xi_1$  and  $\xi_2$

Calculate  $P_1, P_2$  from  $\xi$

$s \sim U(0, 1) \leftarrow \text{random.uniform}$

If  $s < P_1$ , sample from seg. 1:  $\alpha_1$ ,  $m \in [\min, 0.5]$   
else, sample from seg. 2:  $\alpha_2$ ,  $m \in [0.5, \max]$

## Required Plots:

- 2 body
- 2x2 comparison of
  1. orbital trajectories (line styles)
  2. total E vs time (colors)
  3. relative E error (log)
  4. virial ratio
- 3 body
  - same as above, just only 2 methods
  - can split orbital into own plot
- N-body
  - P1
    - energy diagnostics only
      - ↳ compare RK4 to leapfrog
  - P2
    - initial cond. validation ( $N=100$ )
      - log-log mass histogram w/ kroupa slopes (-1.3, -2.3)
      - radial density profile showing plummer  $\rho \propto (1 + \frac{r^2}{a^2})^{-3/2}$
      - cluster evol. snapshots at  $\frac{1}{4}$  intervals through time
  - P3
    - verify E comp. for loop vs. vectorized ( $N=100$ )
    - performance analysis
      - log log time vs  $N$
      - overlay  $O(N^2)$  ref. line
      - speed-up factor  $\left( \frac{t_{\text{loop}}}{t_{\text{vec}}} \right)$  vs  $N$
    - large  $N$  showcase
      - 3x2 evol. snapshots at  $\frac{1}{5}$  intervals
      - color by mass
      - energy diagnostics

- plummer sphere positions

$$\rho = \frac{3 M_0}{4\pi a^3} \left(1 + \frac{r^2}{a^2}\right)^{-3/2} \quad M_0 = \text{total mass}$$

$a = \text{plummer radius}$

$$\Phi = \frac{-GM_0}{\sqrt{r^2 + a^2}}$$

$$\sigma_z = \frac{GM_0}{G\sqrt{r^2 + a^2}}$$

$$\text{velocity dispersion } f = \frac{24\sqrt{2}}{7\pi^3} \frac{a^2}{G^3 M_0^4} (-E)^{3/2} \quad \leftarrow \text{isotropic distribution function}$$

↳ If  $E < 0$  and  $f = 0$

$$\text{else } E = \frac{1}{2}N^2 + \Phi \quad \leftarrow \text{specific energy}$$

$$M( $r) = \frac{M_0 r^3}{(r^2 + a^2)^{3/2}} \quad \leftarrow \text{mass within } a \text{ radius}$$$

$$\text{core radius} \rightarrow r_c = a \sqrt{5E - 1}$$

$$\text{half-mass radius} \rightarrow r_h = \left( \frac{1}{0.5^{2/3}} - 1 \right)^{-0.5} \cdot a$$

$$\text{virial radius} \rightarrow r_v = \frac{16}{8\pi} a$$

$$\begin{aligned}x_e &\rightarrow y_s \\x_j &\rightarrow x_s \\x_j &\rightarrow x_e\end{aligned}$$

$$ps = \begin{bmatrix} x_s & y_s & z_s \\ x_e & y_e & z_e \\ x_j & y_j & z_j \end{bmatrix} = \underbrace{\text{shape}[N, d]}$$

$$ps[:, \text{None}, :, :, :] = \underbrace{\text{shape}[1, N, d]}$$

$$ps[:, :, \text{None}, :, :] = \underbrace{\text{shape}[N, 1, d]}$$

↖

$$diff = \underbrace{\text{shape}[N, N, d]}$$

$$m = \text{shape}[1, N, 1] = \underbrace{\begin{bmatrix} 0 & 0 & 0 \\ x_{es} & y_{es} & z_{es} \\ x_{js} & y_{js} & z_{js} \end{bmatrix} \cdot m_e}_{\rightarrow D3} \cdot \underbrace{\begin{bmatrix} x_{se} & y_{se} & z_{se} \\ 0 & 0 & 0 \\ x_{je} & y_{je} & z_{je} \end{bmatrix} \cdot m_s}_{\rightarrow D2} \cdot \underbrace{\begin{bmatrix} x_{sj} & y_{sj} & z_{sj} \\ x_{ej} & y_{ej} & z_{ej} \\ 0 & 0 & 0 \end{bmatrix} \cdot m_j}_{\rightarrow D1} \cdot m_i$$

receiving force ↓      acting force ↓      dimension ↓

$$\vec{a}_i = \sum_{\substack{j=1 \\ j \neq i}}^N G m_j \frac{\vec{r}_j - \vec{r}_i}{(|\vec{r}_j - \vec{r}_i|^2 + \epsilon^2)^{3/2}}$$

$m_j \cdot \vec{r}_j - \vec{r}_i$

$$diff_{sx} = diff[:, 0, 0]$$

$M_0$  = total mass

$a$  = scale radius (given)

$$\rho(r) = \frac{3M}{4\pi a^3} \left(1 + \frac{r^2}{a^2}\right)^{-5/2}$$

$$M( $r)$  =  $M_0 \frac{r^3}{(r^2 + a^2)^{5/2}}$$$

$$F(r) = M( $r$ ) / M \in [0,1]$$

$$F(r) = U , U \sim U(10^{-12}, 1-10^{-12})$$

$$r = a(U^{-2/3} - 1)^{-1/2}$$

$$\phi = 2\pi U , U \sim U(0,1)$$

$$\cos \theta = 1 - 2\omega , \omega \sim U(0,1)$$

$$x = r \sin \theta \cos \phi$$

$$y = r \sin \theta \sin \phi$$

$$z = r \cos \theta$$

$$r_h = \frac{a}{\sqrt{U^{2/3} - 1}}$$

$$a = r_h \sqrt{U^{2/3} - 1}$$

$$\vec{y} = [\vec{r}, \vec{v}]$$

## Modules

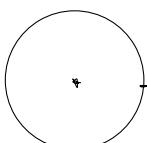
constants ✓  
 integration methods  
 sampling methods  
 plotting  
 testing  
 acting forces

$\text{pos} = [x, y, z] \downarrow^N$   
 $\text{vel} = [vx, vy, vz] \downarrow^N$   
 $\text{acc} = [ax, ay, az] \downarrow^N$   
 $\text{mass} = [m] \downarrow^N$   
 $\text{system\_t} = [\text{system}, t] \downarrow^{\text{steps}}$   
 $\text{system} = \text{N-body}$

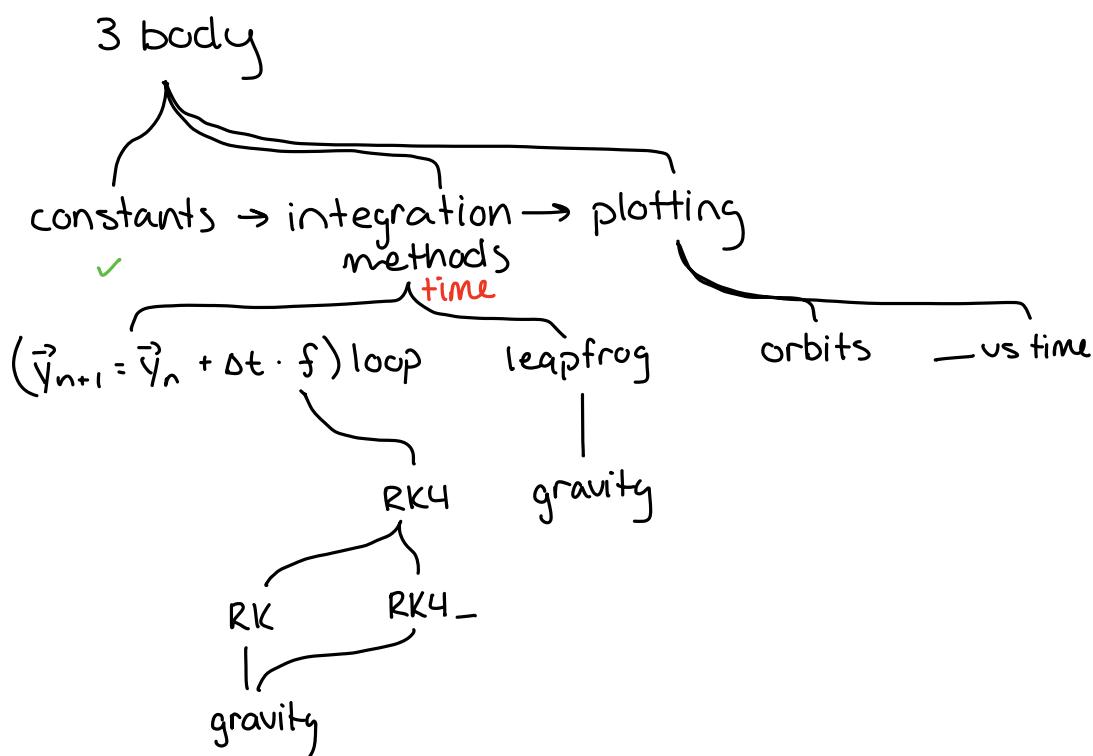
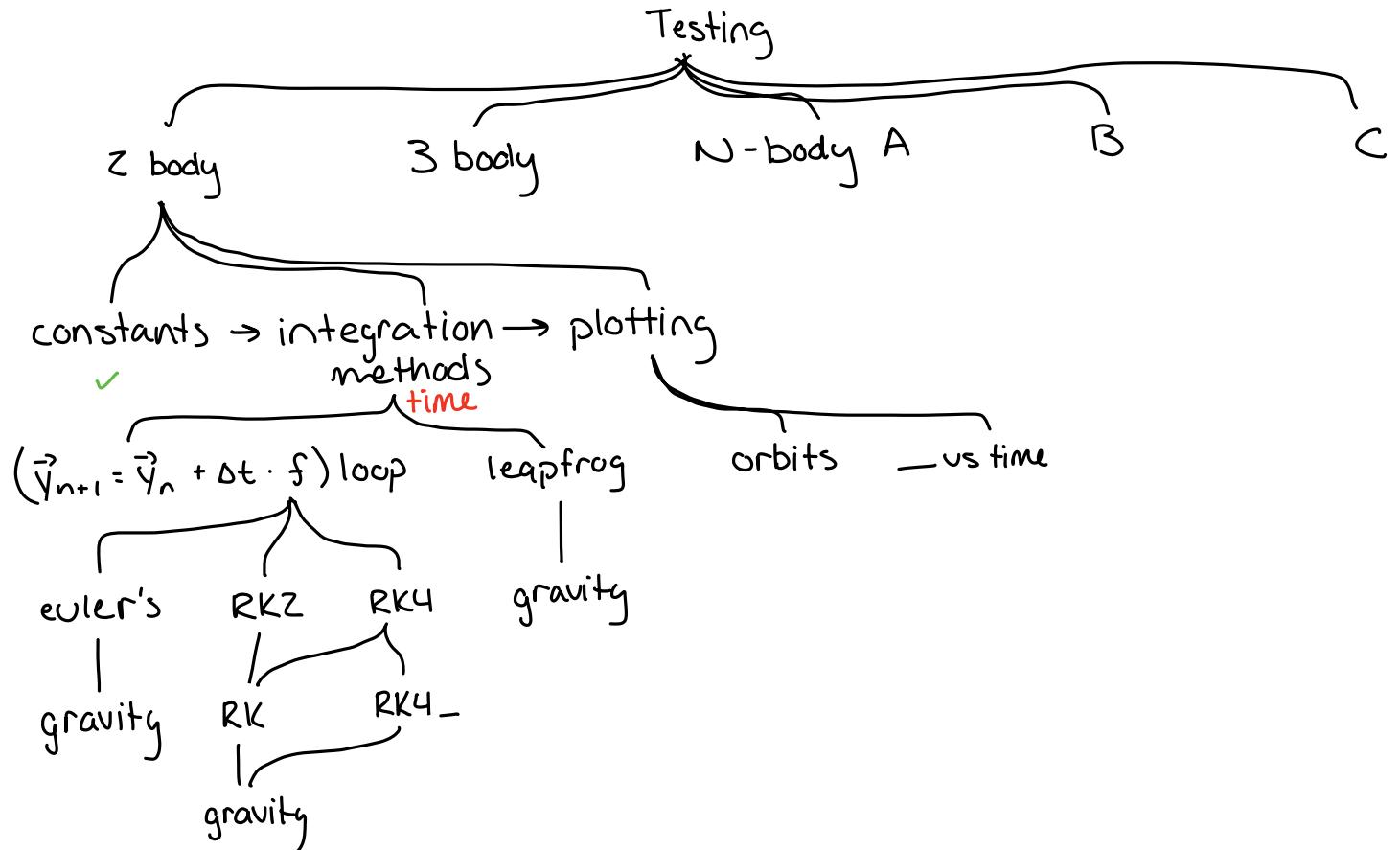
eulers ( $\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$ )  
 RK ( $\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$ )  
 leapfrog ( $\vec{y}_n, \Delta t, t_n, \vec{f} \rightarrow \vec{y}_{n+1}$ )  
 random (cluster radius, N → pos)  
 kroupa IMF (N, m\_min, m\_max → mass)  
 plummer sphere (cluster radius, N, mass\_total → pos)  
 orbits (pos, t, N, ax, fig → ax, fig)  
 — vs time (—, t, N, ax, fig → ax, fig)  
 — vs N (—, t, N, ax, fig → ax, fig)  
 evol. screenshots (pos, t\_i, ax, fig → ax, fig)

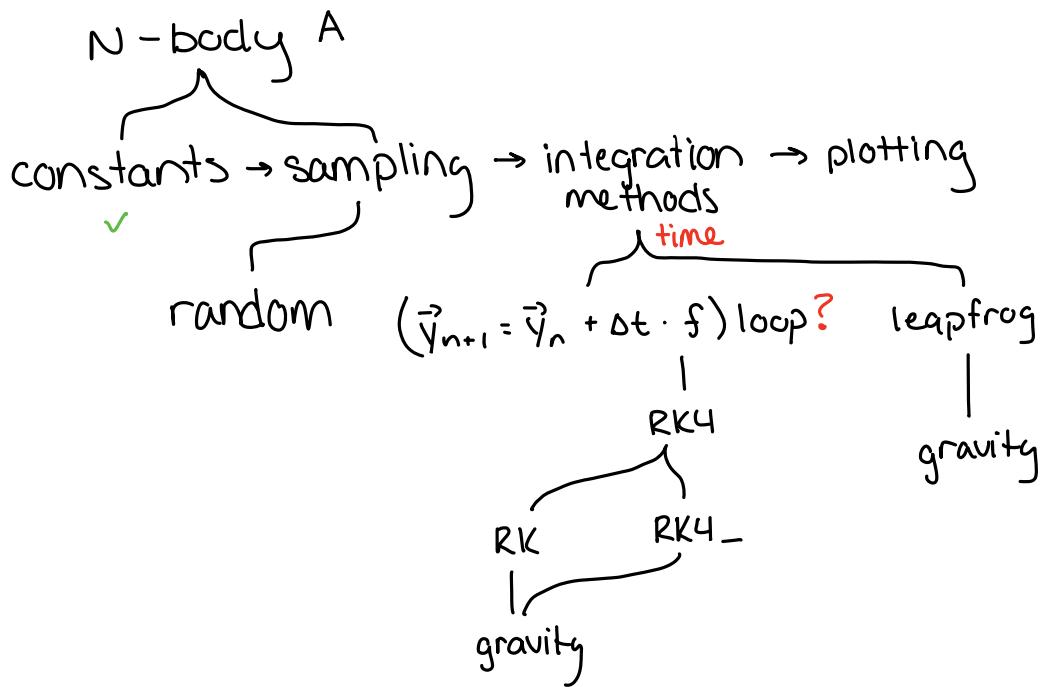
2 body  
 3 body  
 N body  
 P1  
 P2  
 P3

gravity (m\_i, separation, ε →  $\vec{a}_i$ )  
 ε (cluster radius, N → ε)



$$v = \frac{d}{t} = \frac{2\pi(1 \text{ AU})}{1 \text{ yr}} = \frac{2\pi(5.2 \text{ AU})}{12 \text{ yr}}$$





class NbodyIntegrator

pos =    x    y    z  
 $x^2 \quad y^2 \quad z^2$

masses  
 positions  
 velocities

$x^3 \quad y^3 \quad z^3$

$\epsilon$   
 method

functions:

init  
 acceleration  
 step  
 energy