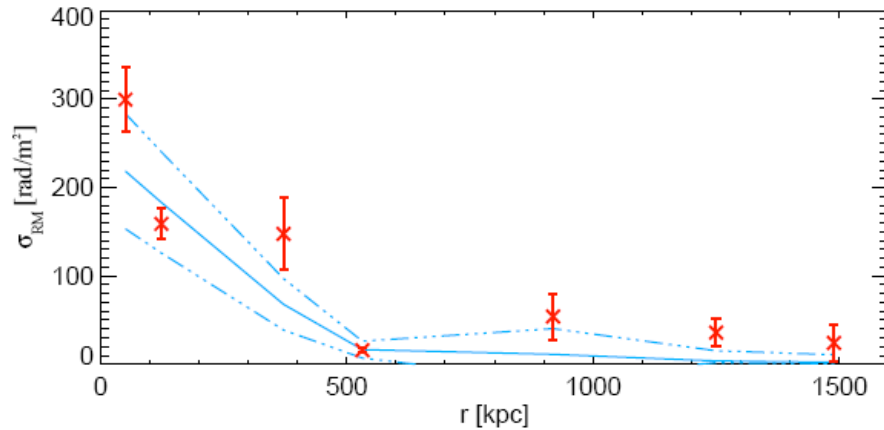


# MAGNETIC FIELDS IN THE ICM

## THE COMA CLUSTER (BONAFEDE ET AL. 2010)

*RM observations of 7 sources in the field - magnetic field numerical simulations*



Best model:  
 $B_0 = 4.7 \mu G$ ,  $\eta = 0.5$

$$\langle B \rangle (r) = B_0 \left( \frac{n_e}{n_0} \right)^\eta$$

$B_0 > 7 \mu G$  or  $< 3 \mu G$   
 $\eta > 1$  or  $< 0.2$   
excluded at 99%  
confidence level

*→ Implications for radio halos formation models*

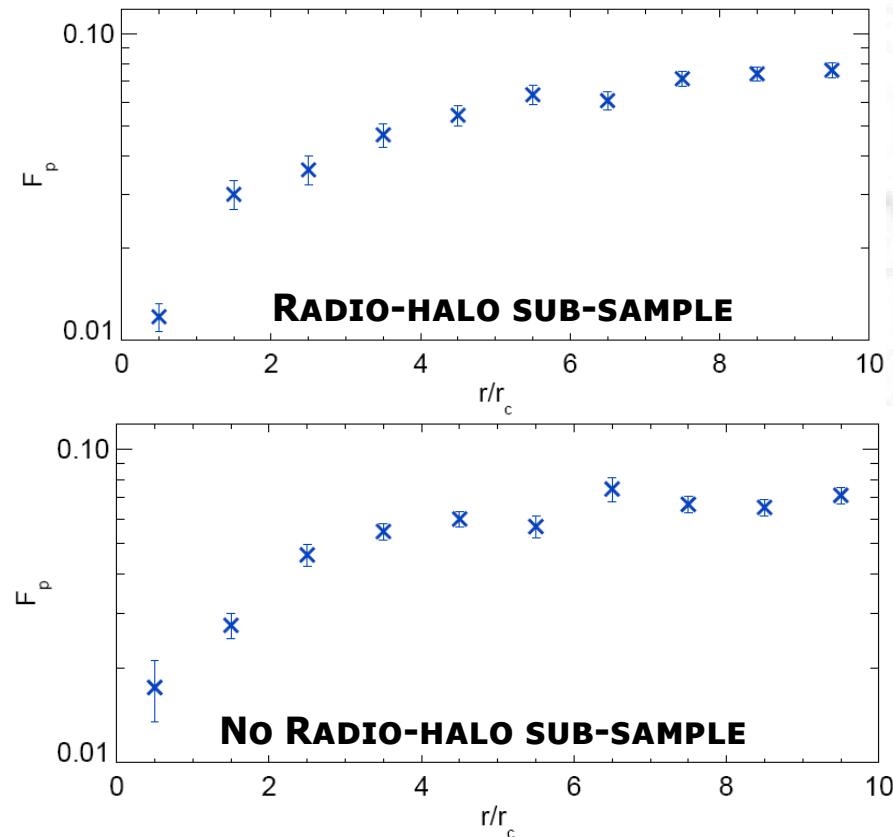
# MAGNETIC FIELDS IN THE ICM

## FROM DEPOLARIZATION OF RADIO SOURCES

33 clusters  
from HIFLUGCS catalog

8 host radio halos

Cluster center  
→ higher B and gas  
density  
→ higher RM  
→ lower fractional  
polarization



K S test:  $P = 0.97$  No significant difference from this analysis for clusters with and without radio halo

→ *Hadronic (secondary) models disfavored*