Cycle 09

Proposal for Chandra Observations

Cover Page

| Principal Investigator Dr. Christopher P. O'Dea | | |
|---|-------------------|-------------|
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| Proposal Title Why are Some Brightest Cluster Galaxies Forming Stars? | | | | | | | | | |
|---|---------------------|------------------------|--------------------|--|--|--|--|--|--|
| Subject Category CLUSTERS OF GALAXIES | | | | | | | | | |
| Proposal Type G0 | Linked Proposal | Distr. Medium WWW ONLY | Proprietary Rights | | | | | | |
| Total Requested Time 201.00 | Number of Targets 4 | | Proposed Budget | | | | | | |

| Joint Proposal? | | XMM Time | RXTE T | lime | | | | |
|-----------------|----------------|------------------|---------|------------|--|--|--|--|
| | | | | | | | | |
| HST Orbits | HST Instrument | HST Instruments: | | | | | | |
| | | | | | | | | |
| Spitzer Time | Spitzer AOTs: | | | | | | | |
| | | | | | | | | |
| NOAO Nights? | NOAO Telescop | e/Instruments: | | | | | | |
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| NRAO/VLA Hours | NRAO/VLA Ar | ray | RAO/VLA | Wavelength | | | | |
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| NRAO/VLBA Hours | NRAO/VLBA V | Vavelength | | | | | | |
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Abstract

Spitzer observations show that at least 25 percent of a sample of 63 Brightest Cluster Galaxies (BCGs) with bright Halpha emission show a mid-IR-excess (MIRE). The MIRE is found in high X-ray luminosity clusters and is consistent with star formation at the level of several solar masses per year. Why is star formation so common in BCGs? What causes it? What does it tell us about the process of AGN feedback and the "final" mass accretion rate? . We propose to study samples of ten MIRE and non-MIRE BCGs matched in X-ray luminosity. Since there is already some archival data, we are requesting 6 additional BCGs - 3 of each type. The combined dataset will establish a complete sample of the MIRE BCGs and a well matched comparison sample with which to address these key questions.

| Proposal Number | Date: | 2007-03-15 | | |
|-----------------|-------|------------|--------|----------|
| | | | Admin. | use only |

${\bf Proposal~for~\it Chandra~\bf Observations}$

General Form

| ΡI | Dr. | Christopher P. O'Dea |
|----------------|-------|--|
| \mathbf{Prc} | posal | Title |
| | Why a | are Some Brightest Cluster Galaxies Forming Stars? |

| Co-Investigator(s) | | | | | | | | | | |
|--|---|-----|--|--|--|--|--|--|--|--|
| First Name E-Mail Last Name Institute Co | | | | | | | | | | |
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| Craig Sarazin | cls7i@mail.astro.virginia.edu University of Virginia | USA | | | | | | | | |
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| Are there additional C | Co-Is listed in the science justification? N | | | | | | | | | |

Institute Endorsement

| Name of Administrator | Donald Boyd |
|--------------------------|-----------------------------------|
| Administrative Authority | Vice President for Research |
| Administrative Institute | Rochester Institute of Technology |
| Admin Signature: | Date: |
| PI Signature: | Date: |

Cycle 09

Proposal for Chandra Observations

Target Summary

PI Dr. Christopher P. O'Dea

Proposal Title

Why are Some Brightest Cluster Galaxies Forming Stars?

| | Target Name | (J200 | U) | Offsets | | | Detector | (c/s) | | Grid |
|-----|-------------------------------|---------|---------|-------------------------|---------|---------|----------|------------|-------|----------|
| | Solar System Object | (3200 | U) | Y Detector | Ontical | Observ. | Grating | Count Rate | Time- | GHU |
| Tar | Grid Name | R.A. | | Z Detector | Monitor | Time | HRC | 1st Order | | #Points |
| No | Target Description (keywords) | Dec. | | Z Detector SIM Trans | V-Mag | (ksec) | Timing | Total Fld. | | MaxDist. |
| 1 | R1442+22 | 14 42 | 19.4 | DIM Hans | N N | 42.0000 | ACIS-S | 1.182000 | N | N |
| | NONE | +22 18 | | | | | NONE | 1.102000 | N | |
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Cycle 09

Proposal for Chandra Observations

 ${\it ACIS~Parameters~(Required,~Pileup,~Telemetry~Parameters)}$

PI Dr. Christopher P. O'Dea

Proposal Title

Why are Some Brightest Cluster Galaxies Forming Stars?

| | Exposure Mode | | CC | Ds | On | TO. | | Most | Subar | rray | Alte | rnating Exp. | En | ergy Filter | ** |
|-------|------------------|----|---------|---------|---------|----------|----|-----------|--------------|----------|----------|--------------|----------|---|---------------------------|
| Tar | Telemetry. | ~ | | | I2 | | ~ | Eff. | | StartRow | | Nbr. Rows | | $\frac{\text{Lower Thrsh.}}{\text{Energy}}$ | Use Spatial Windows |
| No. 1 | Format TE | S0 | S1 N | S2 N | | S4 02 | S5 | Time Y | Type NONE | No.Rows | Y/N N | Exp.Time | Y/N N | Range | Windows |
| | F | N | Y | Y | Y | Y | N | | NONE | | IV. | | 14 | | 14 |
| 2 | TE F | N | N Y | N Y | 01 Y | 02 Y | N | Y | NONE | | N | | N | | N |
| 3 | TE F | N | N Y | N Y | 01 Y | 02 Y | N | Y | NONE | | N | | N | | N |
| 4 | TE F | N | N Y | N Y | 01 Y | 02 Y | N | Y | NONE | | N | | N | | N |
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Proposal for Chandra Observations

Cycle 09

ACIS Parameters (Custom:Telemetry Overflow Parameters)

| PI | Dr. | Christopher P. O'Dea |
|-----|-------|--|
| Pro | posal | Title |
| | Why a | are Some Brightest Cluster Galaxies Forming Stars? |
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| | | | | a. . | · · | Spatia | l Win | dows | | ~ . | |
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Proposal for Chandra Observations

Cycle 09

Target Constraints

| PI Dr. | Christopher P. O'Dea |
|----------|--|
| Proposal | Title |
| Why | are Some Brightest Cluster Galaxies Forming Stars? |

| | Window Constraint | | | Roll Constraints | | | | Phase Dependent Observations | | | |
|-----------|-------------------|------------|-----------|------------------|------|-----------------|---------------------|------------------------------|----------------------------|------------------------|------------------------|
| Tar No | Flag | Start Time | Stop Time | Flag | 180? | Angle (degrees) | Tolerance (degrees) | Flag | Epoch(MJD) Period(days) | Min.Phase Min.Error | Max.Phase Max.Error |
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| | N | Monitoring C | Observations | i | | Group Obser | Un- | Coord- | Add. | |
|------|------|---------------------|---|--------------------|------|------------------------------|------------------------------|---------------------------------------|---|--|
| Flag | No. | Geometric Factor | $\begin{array}{c} {\rm Interval} \\ {\rm (days)} \end{array}$ | Tolerance (%) | Flag | | Interval (days) | inter rupt? | inated Obs.? | Con- straints |
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| | Flag | | Geometric | Geometric Interval | | Geometric Interval Tolerance | Geometric Interval Tolerance | Geometric Interval Tolerance Interval | Geometric Interval Tolerance Interval inter | Geometric Interval Tolerance Interval inter inated |

Proposal for Chandra Observations

Cycle 09

TOO Details

| PI Dr. | Christopher P. O'Dea |
|-----------------|--|
| Proposal Why | Title are Some Brightest Cluster Galaxies Forming Stars? |

| | | Alternates | | Res | ponse Win | dow | | |] | Followup | Observat | ions | Obs.Params |
|-----------|---------------|--------------|--------------|-------------|-----------|------|------------------|-------------------|-------|--------------|-----------------|---------------|-------------------------|
| Tar No | Trig- ger? | Group Name | Nbr. Req. | Type (days) | Start | Stop | Prob- ability | Initial Alloc. | Order | Obs. Time | Interval (days) | Tolerance (%) | specified by Target No. |
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| TOO Trigger Criteria |
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| TOO Followup Instructions |
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If this TOO is a resubmission of a proposal approved in the previous Cycle, should this TOO be canceled if the previous Cycle TOO is triggered?

${\bf Proposal~for~\it Chandra~\bf Observations}$

Cycle 09

Target Remarks

| PI | Dr. | Christopher P. O'Dea | | | | | |
|-----|----------------|--|--|--|--|--|--|
| Pro | Proposal Title | | | | | | |
| | Why a | are Some Brightest Cluster Galaxies Forming Stars? | | | | | |

| Tar No | Remarks Coordinated Observation: Observatories |
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