

# Proposal for *Chandra* Observations

Cycle 10

Cover Page

<b>Principal Investigator</b> Prof. Megan Donahue		
<b>Department/Mailstop</b> Physics & Astronomy Department		
<b>Institute</b> MICHIGAN STATE UNIVERSITY		
<b>Address / Street</b>		<b>City / Town</b> East Lansing
<b>State / Province</b> MI	<b>Zip / Postal Code</b> 48824-2320	<b>Country</b> USA
<b>Telephone</b> 517 355 9200	<b>Fax</b> 517 432 8802	
<b>E-Mail Address:</b> donahue@pa.msu.edu		

<b>Proposal Title</b> Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei			
<b>Subject Category</b> CLUSTERS OF GALAXIES			
<b>Proposal Type</b> GO	<b>Linked Proposal</b> N	<b>Distr. Medium</b> WWW ONLY	<b>Proprietary Rights</b> S
<b>Total Requested Time</b> 65.00	<b>Number of Targets</b> 1		<b>Proposed Budget</b>

<b>Joint Proposal?</b>	
<b>HST Orbits</b>	<b>HST Instruments:</b>
<b>XMM Time</b>	<b>RXTE Time:</b>
<b>NOAO Nights?</b>	<b>NOAO Telescope/Instruments:</b>
<b>NRAO Hours</b>	<b>NRAO Telescopes</b>

<b>Abstract</b>		
<p>Low-redshift brightest cluster galaxies (BCGs) having luminous radio sources are almost always located in the centers of galaxy clusters whose core gas is particularly cool and dense. Abell 193 seems to be an exception to that general trend. However, the existing 17 ksec Chandra observation reveals an extended (4 kpc) low-entropy corona centered on the BCG and perhaps a second corona associated with another of the three optical nuclei detected in the BCG with HST imaging. We propose a 65 ksec observation of the core of Abell 193 in order to characterize these nuclear X-ray sources, including the corona, and to search for cavities excavated by the radio source.</p>		
Proposal Number 10800668	Date: 2008-03-20	Admin. use only

## Cycle 10

<b>PI</b> Prof.   Megan Donahue		
<b>Proposal Title</b> Feeding the Feedback in Brightest Cluster Galaxies:   Radio Sources, Coronae, and Multiple Nuclei		
<div style="text-align: center;">Co-Investigator(s)</div>		
<b>First Name</b> <b>Last Name</b>	<b>E-Mail</b> <b>Institute</b>	<b>Country</b>
Ken Cavagnolo	cavagnolo@pa.msu.edu MICHIGAN STATE UNIVERSITY	USA
Ming Sun	sunm@pa.msu.edu MICHIGAN STATE UNIVERSITY	USA
G. Mark Voit	voit@msu.edu MICHIGAN STATE UNIVERSITY	USA
Are there additional Co-Is listed in the science justification?      N		
Is the first Co-I doing observing, rather than the PI?   N   Telephone:		

<b>Name of Administrator</b>	Katie Cook
<b>Administrative Authority</b>	Senior Contract and Grant Administrator
<b>Administrative Institute</b>	Contracts and Grants, Michigan State University
<b>Admin Signature:</b>	<b>Date:</b>
<b>PI Signature:</b>	<b>Date:</b>

# Proposal for *Chandra* Observations

Cycle 10

## Target Summary

<b>PI</b>	Prof. Megan Donahue
<b>Proposal Title</b>	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

Tar No	Target Name	(J2000)	Offsets Y Detector Z Detector SIM Trans	Optical Monitor V-Mag	Observ. Time (ksec)	Detector	(c/s)	Time- Constr? Ext.Src?	Grid #Points MaxDist.
	Solar System Object					Grating	Count Rate		
	Grid Name					HRC	1st Order		
	Target Description (keywords)					Timing	Total Fld.		
1	Abell 193 NONE  RICH CLUSTER (LOW-REDSHIFT)	01 25 07.5 +08 41 57.0		N	65.0000	ACIS-S NONE N	2.000000  2.000000	N Y	N

# Proposal for *Chandra* Observations

Cycle 10

ACIS Parameters (Required, Pileup, Telemetry Parameters)

<b>PI</b>	Prof. Megan Donahue
<b>Proposal Title</b>	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

Tar No.	Exposure Mode	CCDs On						Most Eff.	Subarray		Alternating Exposures		Energy Filter		Spectra	
	Telemetry. Format	I0	I1	I2	I3			CCD Time	Type	StartRow No.Rows	Y/N	Nbr. Rows Exp.Time	Y/N	Lower Thresh. Range	Max Count	Mult. Lines
1	TE VF		N	N	01	02		Y	NONE		N		N			
		N	Y	Y	Y	Y	N									

## Cycle 10

<b>PI</b>	Prof. Megan Donahue
<b>Proposal Title</b>	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

[illegible]

Target Constraints

PI	Prof. Megan Donahue
Proposal Title	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

Tar No	Window Constraint			Roll Constraints			Phase Dependent Observations				
	Flag	Start Time	Stop Time	Flag	180?	Angle (degrees)	Tolerance (degrees)	Flag	Epoch(MJD) Period(days)	Min.Phase Min.Error	Max.Phase Max.Error

Tar No	Monitoring Observations				Group Observations			Un-inter rupt?	Coordinated		Add. Con- straints
	Flag	No.	Geometric Factor	Interval (days)	Tolerance (%)	Flag	Group ID	Interval (days)	Flag	Interval (days)	

# Proposal for *Chandra* Observations

Cycle 10

## TOO Details

<b>PI</b>	Prof. Megan Donahue
<b>Proposal Title</b>	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

Tar No	Trig-ger?	Alternates		Response Window			Prob-ability	Initial Alloc.	Followup Observations				Obs.Params specified by Target No.
		Group Name	Nbr. Req.	Type (days)	Start	Stop			Order	Obs. Time	Interval (days)	Tolerance (%)	
									1				
									2				
									3				
									4				
									5				
									6				
									7				
									8				

## TOO Trigger Criteria

--

## TOO Followup Instructions

--

If this TOO is a resubmission of a proposal approved in the previous Cycle, should this TOO be canceled if the previous Cycle TOO is triggered?

Proposal for *Chandra* Observations

Cycle 10

Target Remarks

<b>PI</b>	Prof. Megan Donahue
<b>Proposal Title</b>	Feeding the Feedback in Brightest Cluster Galaxies: Radio Sources, Coronae, and Multiple Nuclei

Tar No	Remarks Coordinated Observation: Observatories