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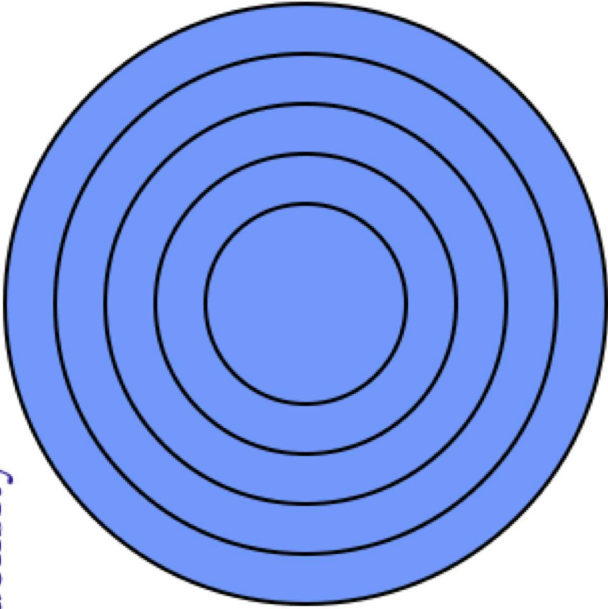
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Show code

Building a Star

This code will create a plot for a star (2-D circle) of uniform density (each shell has equal the same density) and plot how the mass and gravity change as a function of radius.

Constant density



For your Python code, you must first import the relevant packages for your code.

Here we are importing or loading 3 packages that we will use with this code:

- 1- **Numpy** is a fundamental package of Python. It allows for a wide range of data types and data manipulation capabilities. To use it you will call it simply as **np**.
- 2- **Matplotlib** is a Python 2D plotting library. We are only loading the Pyplot capabilities, which provide a collection of command that make matplotlib work like **MATLAB**.
- 3- **Math** is a module that allow for certain math functions.
- 4- **matplotlib** notebook allows for interactive plotting.

To execute a code in a cell, you can either select the cell and hit *Shift+Enter* or click on the run cell button above.

```
1 %matplotlib inline
2 import numpy as np
3 import matplotlib.pyplot as plt
4 import math
```

Once you have successfully executed a cell, a number should appear in the square brackets next to that cell.

First we will define an array `rad_star` (for radius of star) that will go from the center of the star ($r = 0$) to the star's surface ($r = 1$). We will use the `Numpy` function `arange` which generates evenly spaced values within a given interval. We specify the increment to be 0.01.

You need to specify the start and end points for the radius

Keep in mind that `Floats` or non-integer numbers (i.e. with decimal places) should be written with the decimal point. For example, if our starting point is 1 and our ending point is 10, then the code should be `np.arange(1,10,10.0)`.

```
1 start =
2 end =
3 rad_star = np.arange(start,end,0.01)
```

Mass

For this exercise you will need to determine the relation between Mass M and radius r , given that the "star" is of constant density, ρ . Use the continuity equation

$$\frac{dM_r}{dr} = 4\pi\rho r^2$$

We have worked this out in class, but work this out again on a separate sheet of paper first and **enter the expression in the cell below**. Your expression will be a normalized mass in terms of the total stellar mass, M_* .

NOTES:

- To use the constant π in Python, you use `math.pi`. For example, the circumference of the star can be expressed like `circumf = 2.0 * math.pi * rad_star`
- In your expression for mass, the radius term r is defined by the variable `rad_star`.
- Remember to hit *Shift+Enter* to execute the cell.

```
1 mass =
```

Gravity

Now that we have an expression for mass of the star in terms of radius, we will now **do the same thing but for gravity**, where

$$g = \frac{GM_r}{r^2}$$

The gravity term will also be in normalized units, in terms of the surface gravity $g_{surf,face}$.

```
1 grav =
```

Plotting

There are several ways to plot in Python. Here we will use some of the tools that are part of `matplotlib`. However, you can explore other ways to plot in [this Useful tutorial](#). Each line of code is explained below.

- Create a figure object, called `fig`
`fig = plt.figure()`
- Create an "axis object" in the figure
`ax1 = fig.add_subplot(111)`
- Creates a new set of axes (`ax2`) that shares the x-axis with `ax1`, but can have a separate y-axis.
`ax2 = ax1.twinx()`

- Tells python what to plot in the axes object, i.e. your (x,y) values. In this case ax1 will be for the mass. The last part specifies to plot a straight solid line (‘-’) in blue (‘b’).
- ax1.plot(rad_star_mass, ‘b-’)
- Plot the gravity in ax2 as a straight, solid line in red.
- ax2.plot(rad_star_grav, ‘r-’)
- When creating figures, you should always properly label your figure and include units when appropriate.
- ax1.set_title(“Stellar Interior of Star with constant density”)
- To label your axes, set_ylabel defines the y-axis labels and set_xlabel does the same for the x-axis.
- ax1.set_ylabel(“Mass (M/M_{\$})”, color=‘blue’)
- ax2.set_ylabel(“Gravity (g/g_{\$surface})”, color=‘red’)
- ax1.set_xlabel(“Radius (r/R_{\$})”)
- Each time we plot, we must show the figure plot.
- plt.show()

```

1 fig = plt.figure()
2 ax1 = fig.add_subplot(111)
3 ax2 = ax1.twinx()
4 ax1.plot(rad_star_mass, ‘b-’)
5 ax2.plot(rad_star_grav, ‘r-’)
6
7 ax1.set_title(“Stellar Interior of Star with constant density”)
8 ax1.set_ylabel(“Mass ($M/M_{$})”, color=‘blue’)
9 ax2.set_ylabel(“Gravity ($g/g_{$surface})”, color=‘red’)
10 ax1.set_xlabel(“Radius ($r/R_{$})”)
11
12 plt.show()
```

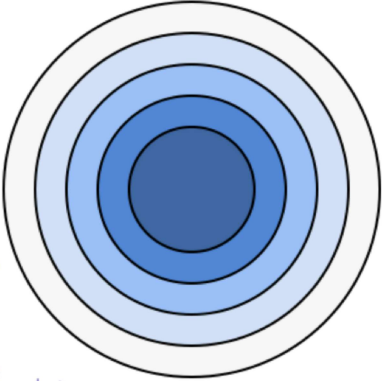
⇒ **Now see if you can edit the code above and add density to your plot.**

Decreasing density Star

You will now adapt the code from the constant density case to create a plot for a star (2-D circle) of decreasing density and plot how the mass, gravity, density, pressure and potential energy change as a function of radius.

Decreasing density

$$\frac{\rho}{\rho_o} = 1 - \frac{r}{R_{\star}}$$



First we will define an array rad_star_dec (for radius of star with decreasing density) that will go from the center of the star (r = 0) to the star's surface (r = 1). We will use the NumPy function arange which generates evenly spaced values within a given interval. We specify the

increment to be 0.01. **As above, you need to specify the start and end points for the radius.**

Keep in mind that *Floats* or non-integer numbers (i.e. with decimal places) should be written with the decimal point. For example, if our starting point is 1 and our ending point is 10, then the code should be np.arange(1.0,10.0) .

```

1 start_dec =
2 end_dec =
3 rad_star_dec = np.arange(start,end,0.01)
```

Mass

For this exercise you will need to determine the relation between Mass M and radius r , given that the ‘star’ is of constant density, ρ . Use the continuity equation

$$\frac{dM_r}{dr} = 4\pi\rho r^2$$

We have worked this out in class, but work this out again on a separate sheet of paper first and **enter the expression in the cell below**. Your expression will be a normalized mass in terms of the total stellar mass, M_{\star} .

NOTES:

- To use the constant π in Python, you use math.pi. For example, the circumference of the star can be expressed like
circumf = 2.0 * math.pi * rad_star
- In your expression for mass, the radius term r is defined by the variable rad_star.
- Remember to hit Shift+Enter to execute the cell.

```
1 mass_dec =
```

Gravity

Now that we have an expression for mass of the star in terms of radius, we will now **do the same thing but for gravity**, where

$$g = \frac{GM_r}{r^2}$$

The gravity term will also be in normalized units, in terms of the surface gravity $g_{surf,free}$.

```
1 grav_dec =
```

Continue to build your model.

Density

Express the normalized density ρ/ρ_0 .

```
1 rho_dec =
```

Pressure

Now enter the expression for normalized pressure in terms of the central pressure P_c .

```
1 press_dec =
```

Potential Energy

Enter the expression for the normalized potential energy Ω_r/Ω_{total} , where Ω_r is the enclosed potential energy. The potential energy of a small shell is

$$d\Omega_r = -\frac{GM_r}{r} dr$$

```
1 Epot_dec =
```

Plotting

There are several ways to plot in Python. Here we will use some of the tools that are part of *matplotlib*. However, you can explore other ways to plot in [this useful tutorial](#). Each line of code is explained below.

- Create a figure object, called fig

```
fig = plt.figure()
```
- Create an 'axis object' in the figure

```
axes = fig.add_subplot(111)
```
- Tells python what to plot in the axes object, i.e. your (x,y) values. The last part specifies to plot a line in blue ('b').

```
axes.plot(rad_star, mass, 'b-')
```
- When creating figures, you should always properly label your figure and include units when appropriate.

```
axes.set_title("Stellar Interior of Star with decreasing density")
```
- To label your axes, `set_ylabel` defines the y-axis labels and `set_xlabel` does the same for the x-axis.

```
axes.set_xlabel('r', color='blue')
axes.set_ylabel('')
```
- Since we are plotting more than two lines, and all the parameters are normalized, we will use a legend to label each line. You can specify where you want the legend to appear. Best lets Python place it where it thinks it is best. `axes.legend(loc='best')`
- Each time we plot, we must show the figure plot.

```
plt.show()
```

```
1 fig = plt.figure()
2 axes = fig.add_subplot(111)
3
4 axes.plot(rad_star_dec, rho_dec, 'r', label='Density')
5 axes.plot(rad_star_dec, mass_dec, 'b', label='Mass')
6 axes.plot(rad_star_dec, grav_dec, 'g', label='Gravity')
7 axes.plot(rad_star_dec, press_dec, 'm', label='Pressure')
8
9 axes.set_title('')#<--Enter appropriate plot title in the quotations
10 axes.set_xlabel('r')#<--Enter appropriate x-axis title in the quotations
11 axes.set_ylabel('')#<--Enter appropriate y-axis title in the quotations
12
13 axes.legend(loc='best')
14 plt.show()
```

Adapt the plotting routine to include the potential energy.

You can use the following Python color chart to add more colors to your plots.

