

Astr 1010
Problem Set #5, Solutions

1. For Mercury we have $P_{\text{yr}} = 0.2409$ and $a_{\text{AU}} = 0.387$ (data is from Table 4 in *Explorations*). Comparing the two sides of Kepler's Third Law, we find:

$$P_{\text{yr}}^2 = 5.803 \times 10^{-2} \quad a_{\text{AU}}^3 = 5.80 \times 10^{-2}$$

which agree to three figures.

For the four given planets, we find:

Planet	P_{yr}	a_{AU}	P_{yr}^2	a_{AU}^3
Mercury	0.2409	0.387	5.803×10^{-2}	5.80×10^{-2}
Mars	1.8809	1.524	3.5377	3.540
Saturn	29.4577	9.539	8.67756×10^2	8.680×10^2
Neptune	164.793	30.06	2.71567×10^4	2.716×10^4

The results agree (out to the permitted number of significant digits).

2. Here, $a_{\text{AU}} = 3.3$, so

$$a_{\text{AU}}^3 = (3.3)^3 = 35.9$$

By Kepler's Third Law, we also have

$$P_{\text{yr}}^2 = a_{\text{AU}}^3 = 35.9$$

and solving for P_{yr} gives

$$P_{\text{yr}} = \sqrt{35.9} = 6.0$$

The period of the planet is 6.0 yr.

3. Here, $P_{\text{yr}} = 11.18$, so

$$P_{\text{yr}}^2 = (11.18)^2 = 125.0$$

By Kepler's Third Law, we also have

$$a_{\text{AU}}^3 = P_{\text{yr}}^2 = 125.0$$

and solving for a_{AU} gives

$$a_{\text{AU}} = (125.0)^{1/3} = 5.0$$

The semi-major axis of the planet is 5.0 AU.

4. We are given the period of Halley's Comet, so Kepler's 3rd Law will give us the semi-major axis of the orbit. We get:

$$P_{\text{yr}}^2 = (76.1)^2 = 5.79 \times 10^3 = a_{\text{AU}}^3$$

so that

$$a_{\text{AU}} = (5.79 \times 10^3)^{1/3} = 18.0$$

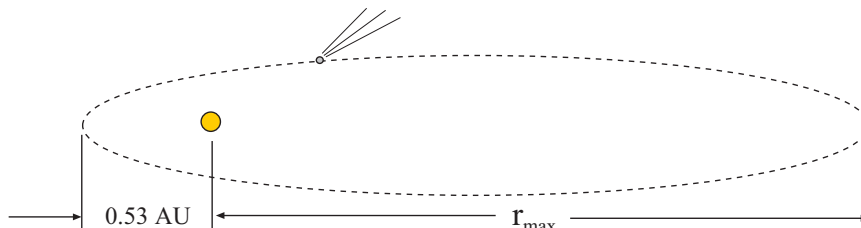
so the semi-major axis is $a = 18.0 \text{ AU}$.

Now, the full long dimension of the ellipse is equal to $2a$ but it is also equal to $r_{\min} + r_{\max}$, and we already know r_{\min} . So we can find r_{\max} :

$$r_{\min} + r_{\max} = 2a = 2(18.0 \text{ AU}) = 36.0 \text{ AU}$$

$$r_{\max} = 36.0 \text{ AU} - r_{\min} = 36.0 \text{ AU} - 0.53 \text{ AU} = 35.5 \text{ AU}$$

So Halley's Comet gets as far as 35.5 AU from the sun. (This is beyond the orbit of Neptune.)



5. If an object falls from rest in 5.0 s it has fallen a distance

$$d = \frac{1}{2}gt^2 = \frac{1}{2}9.8\left(\frac{\text{m}}{\text{s}^2}\right)(5.0 \text{ s})^2 = 122.5 \text{ m}$$

and its speed is

$$v = gt = (9.8 \frac{\text{m}}{\text{s}^2})(5.0 \text{ s}) = 49 \frac{\text{m}}{\text{s}}$$

Now suppose we are given that the distance the object has fallen is 300 m. What is the time t ? We find:

$$d = \frac{1}{2}gt^2 \quad \implies \quad t^2 = \frac{2d}{g} = \frac{2(300 \text{ m})}{(9.8 \frac{\text{m}}{\text{s}^2})} = 61.2 \text{ s}^2$$

and this gives:

$$t = \sqrt{61.2 \text{ s}^2} = 7.8 \text{ s}$$

At this time, the speed of the object is

$$v = gt = (9.8 \frac{\text{m}}{\text{s}^2})(7.8 \text{ s}) = 7.7 \frac{\text{m}}{\text{s}}$$

6. On the moon, the acceleration of gravity is $g_{\text{Moon}} = 5.3 \frac{\text{m}}{\text{s}^2}$, so in 1 s, a rock will fall a distance

$$d = \frac{1}{2}gt^2 = \frac{1}{2}(1.6 \frac{\text{m}}{\text{s}^2})(1.0 \text{ s})^2 = 0.80 \text{ m}$$

and in 3.0 s it will fall

$$d = \frac{1}{2}gt^2 = \frac{1}{2}(1.6 \frac{\text{m}}{\text{s}^2})(3.0 \text{ s})^2 = 7.2 \text{ m}$$

To find the time to fall 300 m, use $d = 300 \text{ m}$ and find t :

$$d = \frac{1}{2}g_{\text{Moon}}t^2 \quad \implies \quad t^2 = \frac{2d}{g_{\text{Moon}}} = \frac{2(300 \text{ m})}{(1.6 \frac{\text{m}}{\text{s}^2})} = 375 \text{ s}^2$$

which gives

$$t = \sqrt{375 \text{ s}^2} = 19.4 \text{ s}$$

At that time, its speed is

$$v = gt = (1.6 \frac{\text{m}}{\text{s}^2})(19.4 \text{ s}) = 31 \frac{\text{m}}{\text{s}}$$

$$P_{\text{yr}}^2 = a_{\text{AU}}^3 \quad \text{Object falling from rest:} \quad d = \frac{1}{2}gt^2 \quad v = gt \quad g = 9.8 \frac{\text{m}}{\text{s}^2}$$