

Phys 3820, Fall 2009
Problem Set #2, Hint-o-licious Hints

1. *Griffiths, 6.11* Fairly easy algebra.
2. *Griffiths, 6.14* The relativistic perturbation to the Hamiltonian is

$$H' = -\frac{p^4}{8m^3c^2}$$

so we want to calculate

$$E_{n,\text{rel}}^1 = \left\langle n \left| -\frac{p^4}{8m^3c^2} \right| n \right\rangle$$

Use

$$p = i\sqrt{\frac{\hbar m \omega}{2}}(a_+ - a_-) \quad \implies \quad p^2 = -\frac{\hbar m \omega}{2}(a_+^2 + a_-^2 - a_+a_- - a_-a_+)$$

in

$$E_{n,\text{rel}}^1 = -\frac{1}{8m^3c^2} \langle p^2 \psi_n^0 | p^2 \psi_n^0 \rangle$$

and use orthogonality of the HO wave functions. Get

$$E_{n,\text{rel}}^1 = -\frac{3\hbar^2\omega^2}{32mc^2}(2n^2 + 2n + 1)$$

3. *Griffiths, 6.17* Take the individual terms from Eqs. (6.57) and (6.65) and consider individually the cases $j = l \pm \frac{1}{2}$. Do the algebra (which is a little messy) and show that in both cases you get simple result in terms of j .

4. *Griffiths, 6.20* Use $|\mathbf{L}| \approx \hbar$ in (6.59) With this, the critical size of the B field is around 12 T.

5. *Griffiths, 6.21* The fine structure is larger than the Zeeman contribution to the energy. The zero-field value of the energies is given by (6.67). Since it included the spin-orbit splitting, the energies depend on j (and n).

Now, for $n = 2$ we have the states $l = 0$ and $l = 1$. We must have states of “good” j , so we note that the $l = 0$ state is a $j = \frac{1}{2}$ states while the $l = 1$ state give $j = \frac{1}{2}$ and $j = \frac{3}{2}$.

Calculate the Landé g factor g_J for each state note, it depends on j and l , and then the weak-field Zeeman energy is

$$E_Z^1 = \mu_B g_J B_{\text{ext}} m_j$$

If you plot E vs. $\mu_B B_{\text{ext}}$, the slope of the line is $g_J m_j$.

For the $j = \frac{1}{2}$ states we then have a pair for the state that came from $l = 0$ (with $m_j = \pm \frac{1}{2}$) and a pair for the state that came from $l = 1$. The $j = \frac{3}{2}$ state came from $l = 1$ and with $m_j = -\frac{3}{2} \dots \frac{3}{2}$, there are four lines with their own slopes.

6. *Griffiths, 6.28* This one is trickier than it seems because you need to do some tracing to tell how to replace the various factors in the formula for the hyper-fine splitting ΔE for the

various (interesting!) two-body bound systems given. It is best to ignore the formula (6.92) and go back to (6.89), which basically says

$$\Delta E \propto \frac{g_p}{m_p m_e a^3}$$

because in this expression each symbol appears for a particular reason. (In (6.92) the two factors of m_e appear for two different reasons so they have different replacements!) Actually, this expression should also have a factor g_e in the numerator, but as long as the “orbiting” particle is electron-like (as the muon is), it’s always 2 so it won’t change.

Make replacements, noting: g_p is the “ g -factor” of the positive particle. m_p is the mass of the positive particle. m_e is the *true mass* of the orbiting negative particle, since it comes from an expression with the gyromagnetic ratio.

Most complicated is a . This factor arises from the wave function of the two-particle system and is inversely proportional to the *reduced mass* of the system:

$$a \propto \frac{1}{m(\text{red})}$$

For example the reduced mass of positronium case is $\frac{m_e}{2}$. The reduced mass for the H atom is *very close* to m_e .

7. Griffiths, 6.29 The perturbation is

$$H' = \frac{e^2}{4\pi\epsilon_0} \left(\frac{1}{r} - \frac{1}{b} \right)$$

and use this in

$$E_{\text{gs}}^1 = \int \psi^*(\mathbf{r}) H' \psi(\mathbf{r}) d^3\mathbf{r}$$

As explained in class, you can approximate the exponential as 1 to get a lowest-order answer, and that all we need to check the order of magnitude of the result. With this, show

$$\frac{E_{\text{gs}}^1}{|E_{\text{gs}}^1|} = \frac{4}{3} \left(\frac{b}{a} \right)^2$$

Compare with value with those from fine structure and the hyperfine splitting! (Use Table 6.1).