

Phys 2112, Fall 2009
Quiz #3

1. If a spherical body with the mass of the sun (2.00×10^{30} kg) has an escape speed equal to the speed of light c , find its radius. (Assume that Newtonian physics is correct!)

Use the formula for escape speed; solve for R :

$$v_{\text{esc}} = \sqrt{\frac{2GM}{R}} \quad \Rightarrow \quad v_{\text{esc}}^2 = c^2 = \frac{2GM}{R} \quad \Rightarrow \quad R = \frac{2GM}{c^2}$$

Plug in numbers:

$$R = \frac{2(6.67 \times 10^{-11})(2.00 \times 10^{30})}{(2.998 \times 10^8)^2} \text{ m} = 2.97 \times 10^3 \text{ m} = 2.97 \text{ km}$$

The sun must squoosh down to a radius of 3 km before light is trapped at its surface!

2. In the derivation of the shape of an orbit for the gravitational force, we started with a differential equation for r as a function of time: $r(t)$. But that function does not give the *shape* of the orbit.

What did we have to do the differential equation to get a function which *does* tell us the shape of the orbit? (Summarize as much of the math as you can in a paragraph.)

What we need to tell the shape of the orbit is the function $r(\phi)$, which will give a (polar) plot of the path of the orbiting object. To get this, we had to change the time derivatives $\frac{d}{dt}$ in the differential equations to angle derivatives $\frac{d}{d\phi}$, using the chain rule; thus we got a differential equation in terms of the variable ϕ . (Solving this differential equation was quite doable but required a couple tricks!)

The relation between the two derivatives involved a factor of the constant angular momentum ℓ .

3.a) Find the speed of a proton which has a kinetic energy of 2.0 GeV.

If $T = 2.0 \text{ GeV}$ then the total energy is

$$E = m_{\text{prot}}c^2 + T = 0.9383 \text{ GeV} + 2.0 \text{ GeV} = 2.938 \text{ GeV} .$$

We can get the speed from this by

$$E = \frac{m_{\text{prot}}c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \implies \sqrt{1 - \frac{v^2}{c^2}} = \frac{m_{\text{prot}}c^2}{E} = \frac{0.938}{2.938} = 0.319$$

From this it follows that

$$\frac{v^2}{c^2} = 0.898 \implies \frac{v}{c} = 0.947 \implies v = 2.84 \times 10^8 \frac{\text{m}}{\text{s}}$$

b) What answer would you get if you used Newtonian physics?

The kinetic energy is

$$T = 2.0 \times 10^9 \text{ eV} \left(\frac{1.602 \times 10^{-19} \text{ J}}{1 \text{ eV}} \right) = 3.2 \times 10^{-10} \text{ J}$$

Using $T = \frac{1}{2}mv^2$ we get

$$v^2 = \frac{2T}{m} = \frac{2(3.2 \times 10^{-10} \text{ J})}{(1.67 \times 10^{-27} \text{ kg})} = 4.68 \times 10^{17} \frac{\text{m}^2}{\text{s}^2} \implies v = 6.2 \times 10^8 \frac{\text{m}}{\text{s}}$$

(Whoa! Faster than light!)

4. (Here we disregard the potential energy of a particle.) How does our old expression for the energy of a particle, $E = \frac{1}{2}mv^2$ relate to the new *relativistic* total energy E ? (The result was found on the last problem set; tell what you remember of it.)

On the problem set, we showed that the *relativistic* kinetic energy can be written as

$$E - mc^2 = T = \frac{1}{2}mv^2 + \text{lotsa terms with } v\text{'s and } c\text{'s}$$

where the extra terms are small when v is small compared to c . Thus the (correct) total energy is roughly equal to $mc^2 + \frac{1}{2}mv^2$. Basically Newtonian physics is ignoring all those extra terms!

Show work for all problems and include the right units!

$$\begin{aligned} m_{\text{prot}} &= 1.67 \times 10^{-27} \text{ kg} & m_{\text{prot}}c^2 &= 938.3 \text{ MeV} & 1 \text{ eV} &= 1.602 \times 10^{-19} \text{ J} \\ F &= G \frac{m_1 m_2}{r^2} & G &= 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2} & a_c &= \frac{v^2}{r} & F_c &= \frac{mv^2}{r} & c &= 2.998 \times 10^8 \frac{\text{m}}{\text{s}} \\ v_{\text{esc}} &= \sqrt{\frac{2GM}{R}} & E &= p^2 c^2 + m^2 c^4 & E &= \frac{mc^2}{\sqrt{1 - \frac{u^2}{c^2}}} & T &= E - mc^2 \end{aligned}$$