

Math Review

- $1 / 0.001 = 1000$
- $1000 \times 1000 = 1000000$
- $10^3 \times 10^3 = (10^3)^2 = 10^6$



Math

- $10^2 \times 10^5 =$

A) 10^3

B) 10^4

C) 10^{10}

D) 10^{14}

E) 10^7



Parallax and Distance

Astronomy 1120:
Stars and Galaxies
Adam Ginsburg &
Devin Silvia
July 13, 2010

Learning Goals

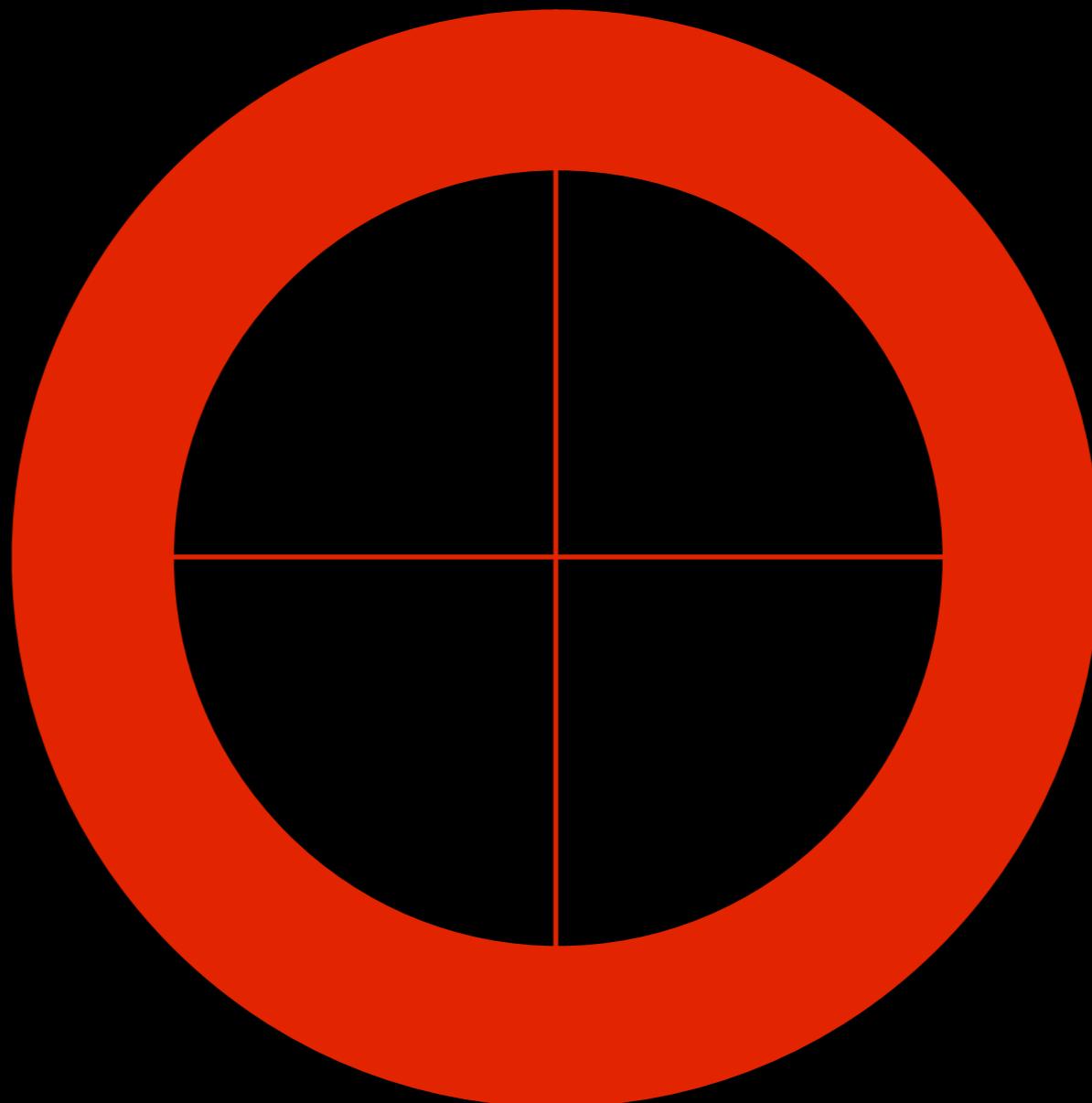
- Why is it important to know how far something is?
- How can astronomers measure distance?
- Definitions: arcsecond, parsec

A little history

- Up through the 1920's, it was not known whether Galaxies like Andromeda were nebulae within our own galaxy or independent, enormous groupings very far away
- The big uncertainty was the distance: because it is so far away, we can't resolve individual stars in Andromeda, so it just looks like fuzz
- If Andromeda had turned out to be a nebula within the Milky Way, it would be at least 100 times smaller, or $(100)^3 = 1,000,000$ times less massive!

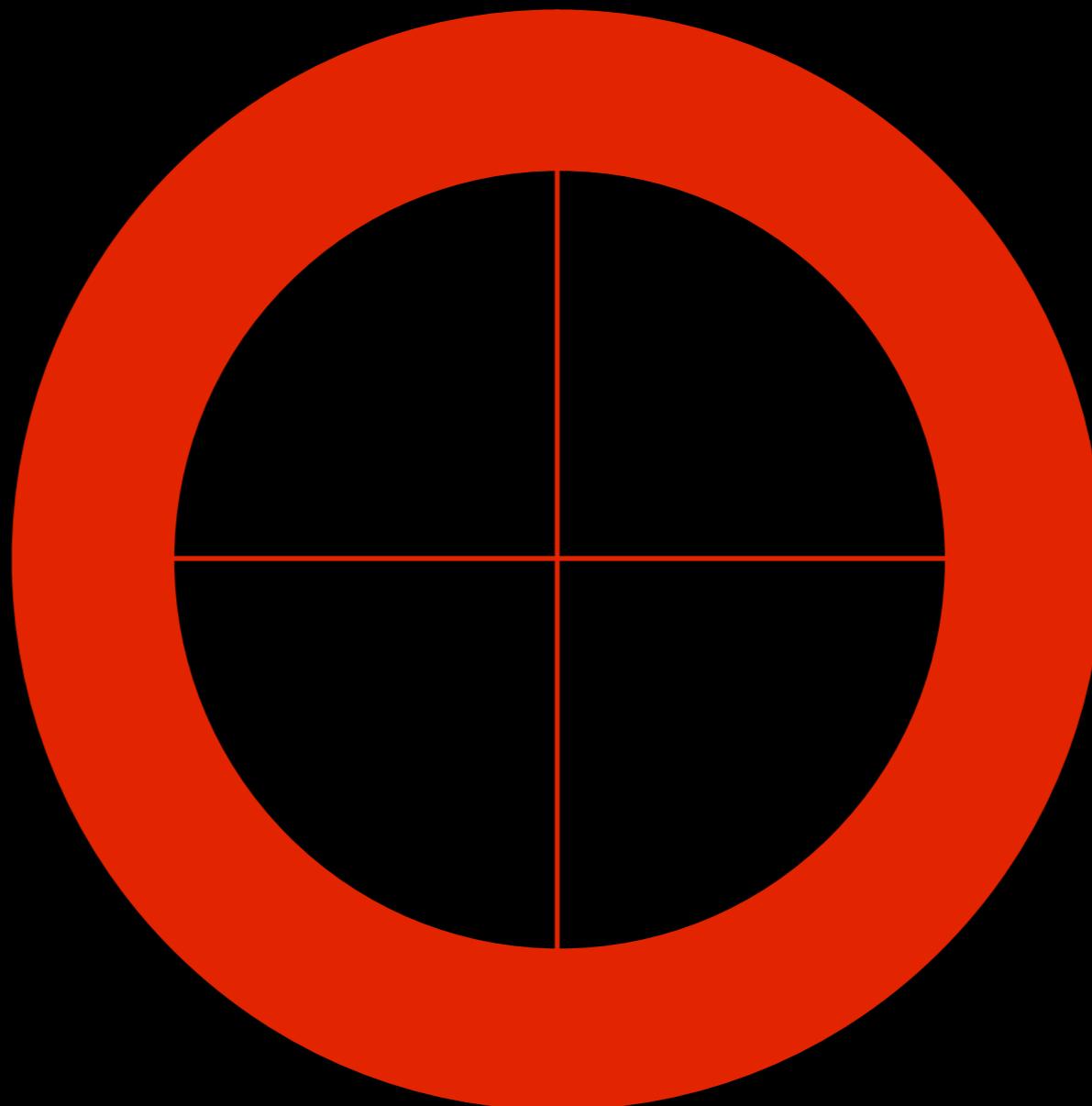
Parallax Demo

- Stick your thumb out at the target
- Close one eye, then the other



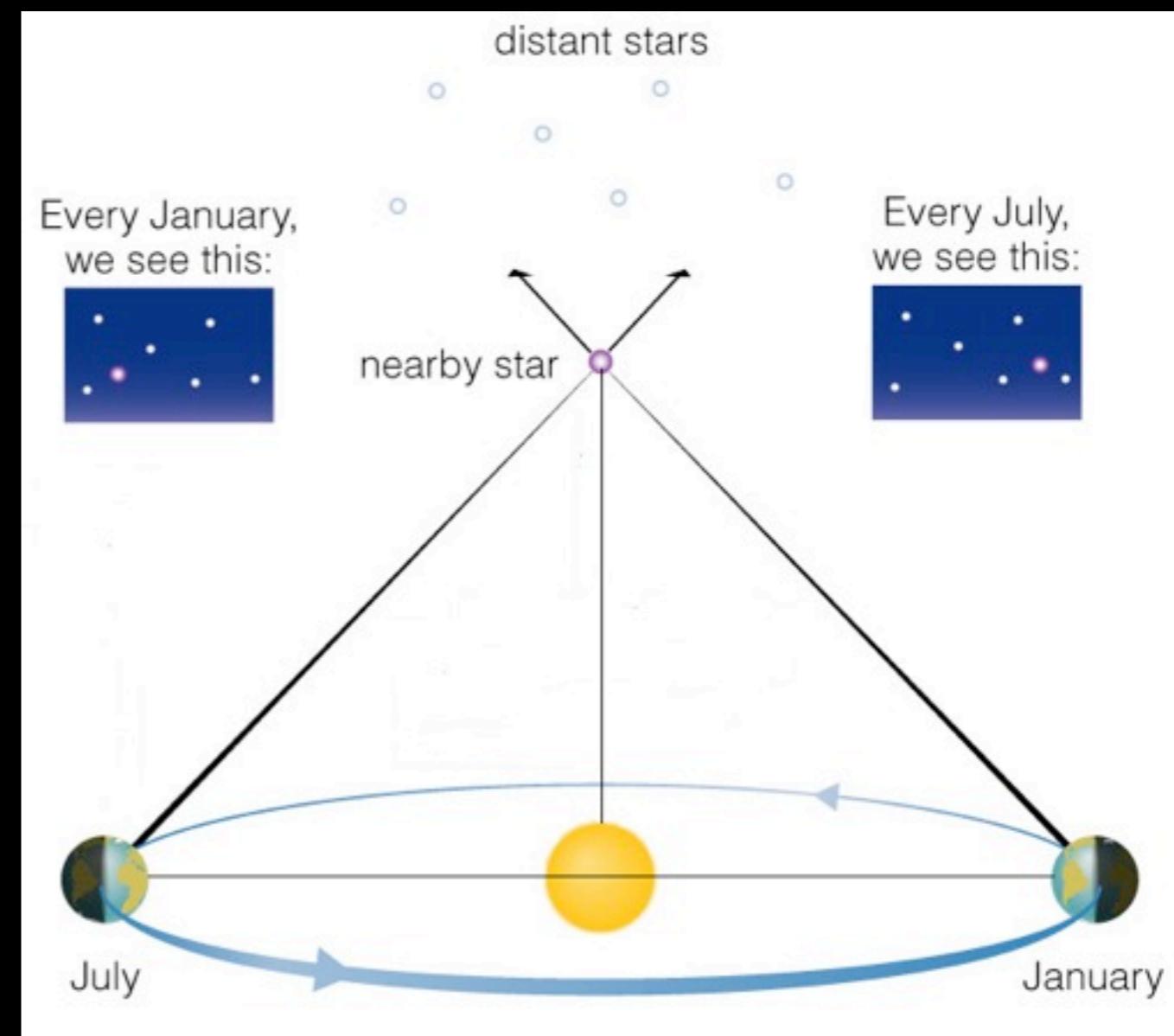
Parallax Demo

- Move your thumb closer to your eye
- Does the parallactic angle change?



Stellar Parallax

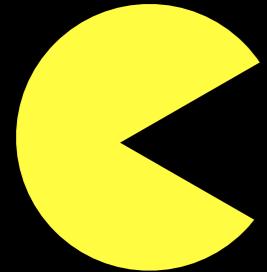
- As the Earth moves around the Sun, nearby stars will appear to move relative to “fixed” background stars





Measuring Parallax

- If an object appears to move relative to a fixed background, where is it?
 - A) closer than the background
 - B) further than the background
 - C) you can't tell

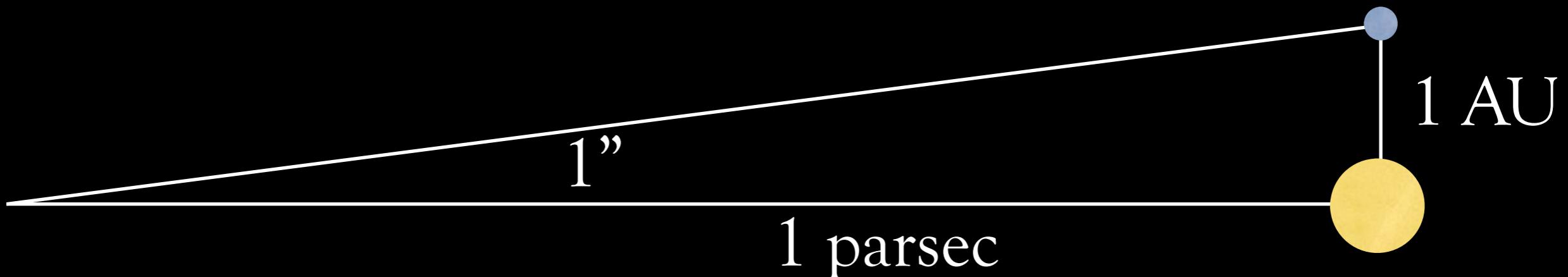


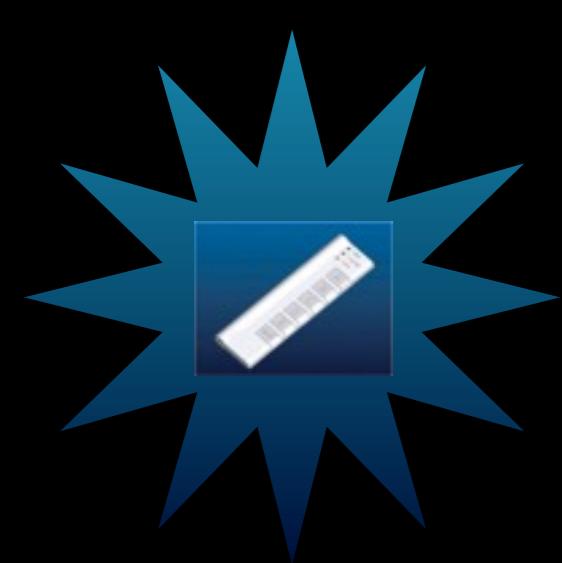
Measuring Angle

- There are 360 **degrees** in a circle
- 60 **arcminutes** ('') in a **degree**
- 60 **arcseconds** (") in an **arcminute**
 - or 3600 arcseconds in a degree
- Your thumb held out at arm's length is about 2 **degrees** (it should cover the sun or moon completely)

Measuring Parallax

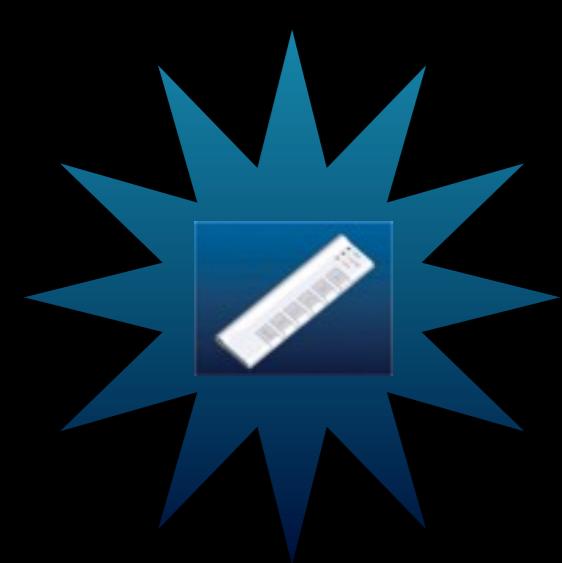
- The “Parsec” is a unit of **distance**
 - Parallax Arcsecond
 - If the apparent motion of a star is 1 arcsecond, its distance is 1 parsec (3.26 light-years, or $19 \text{ trillion} = 19 \times 10^{12}$ miles)
 - A star 10 parsecs away has a parallax of 0.1”





Parallax

- A star has a parallax of 0.01" (10 milli-arcseconds). How far away is it?
 - A) 0.01 parsecs
 - B) 0.1 parsecs
 - C) 1 parsec
 - D) 10 parsecs
 - E) 100 parsecs



Parallax

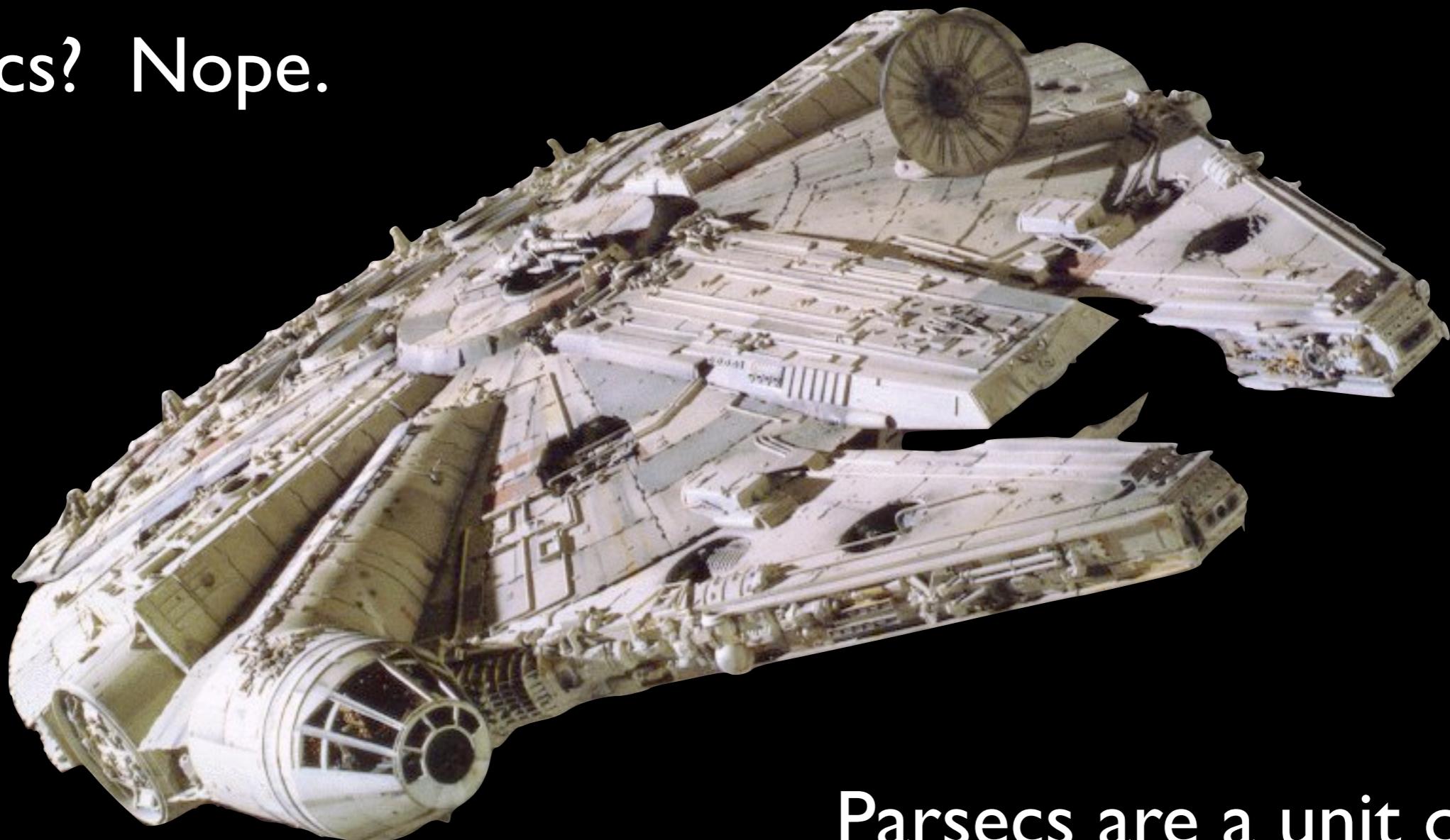
- A star has a parallax of 0.01" (10 milli-arcseconds). How far away is it?
E) 100 parsecs
- $100 \text{ pc} = 1/(0.01 \text{ arcseconds})$
- The nearest 100 stars are within about 7 pc, or 22 light-years. There are 380 known stars within 10 parsecs (32.6 ly)

Parsecs and Light-Years

- The parsec is a unit based on the **geometry** of the Solar System
- The Light-Year is based on the **constant speed of light**: it is the **distance** light travels in 1 year
 - A light-second is 186,000 miles
 - Light can travel around the Earth 7.5 times per second

The Millenium Falcon

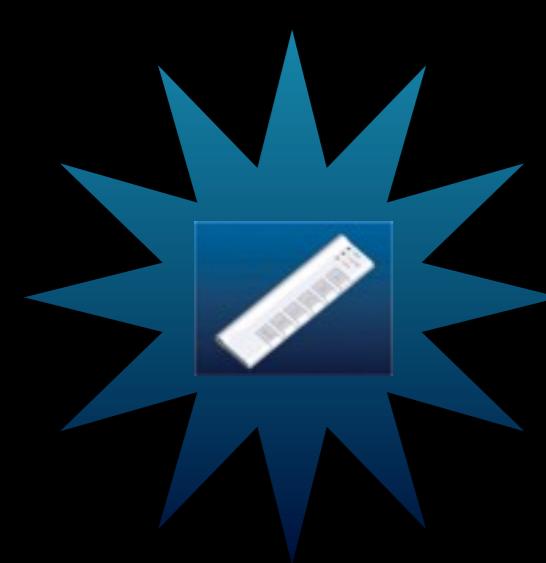
Kessel run in 12
parsecs? Nope.



Parsecs are a unit of
distance, not time!

The closest stars

- After the sun, the closest stars are Proxima and Alpha Centauri, which are at a distance of 1.3 pc (4 ly)
- Their parallax is <1" (**arcsecond**)
 - recall that ground-based telescopes have a **resolution** of ~1" because of the atmosphere
 - *Parallax is hard to measure!*

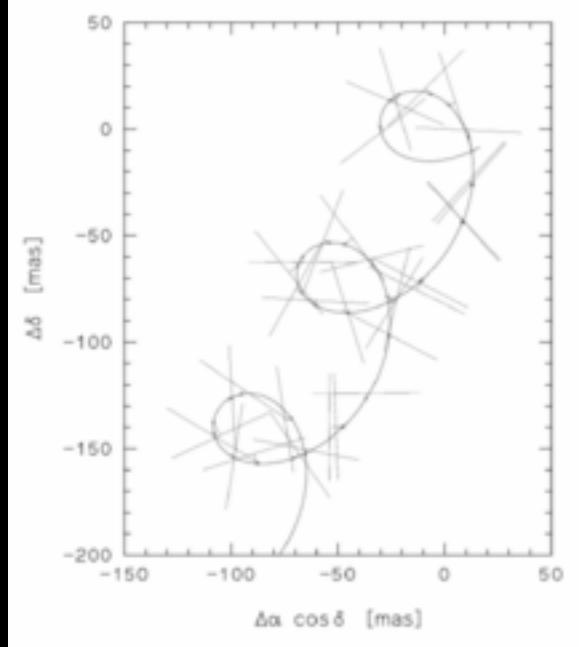


Travel Time

- Spacecraft launched from Earth using chemical rockets travel around 5 km/s (12,000 mph). How long would it take a satellite to reach Alpha Centauri, which is 4 light-years away, at that speed? (don't calculate, just guess what you think sounds reasonable. Hint: the speed of light is 300,000 km/s)
 - A) 4 years
 - B) 400 years
 - C) 240,000 years
 - D) 24 million years
 - E) 1 billion years

Astrometry & Parallax

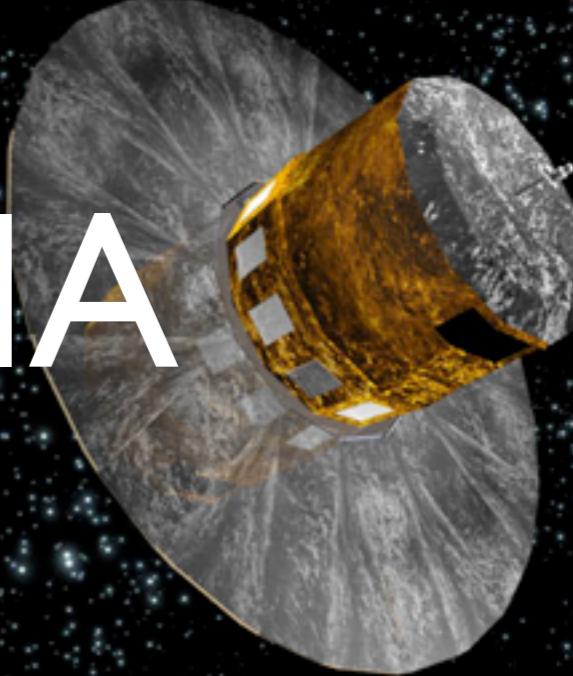
- The measurement of positions of objects is known as **astrometry**
- Precise astrometry is needed to get parallax measurements, but is also useful for measuring **proper motion**, which is the tangential motion of an object on the sky
- **Astrometry + Doppler Shift** gives full 3D velocity information
- **Astrometry + Parallax** gives full 3D location information



Hipparcos

- Since astrometric precision from the ground is limited, we used a satellite called Hipparcos to measure parallaxes towards stars within 3300 light-years of the sun
- It observed 120,000 stars with 0.001" accuracy





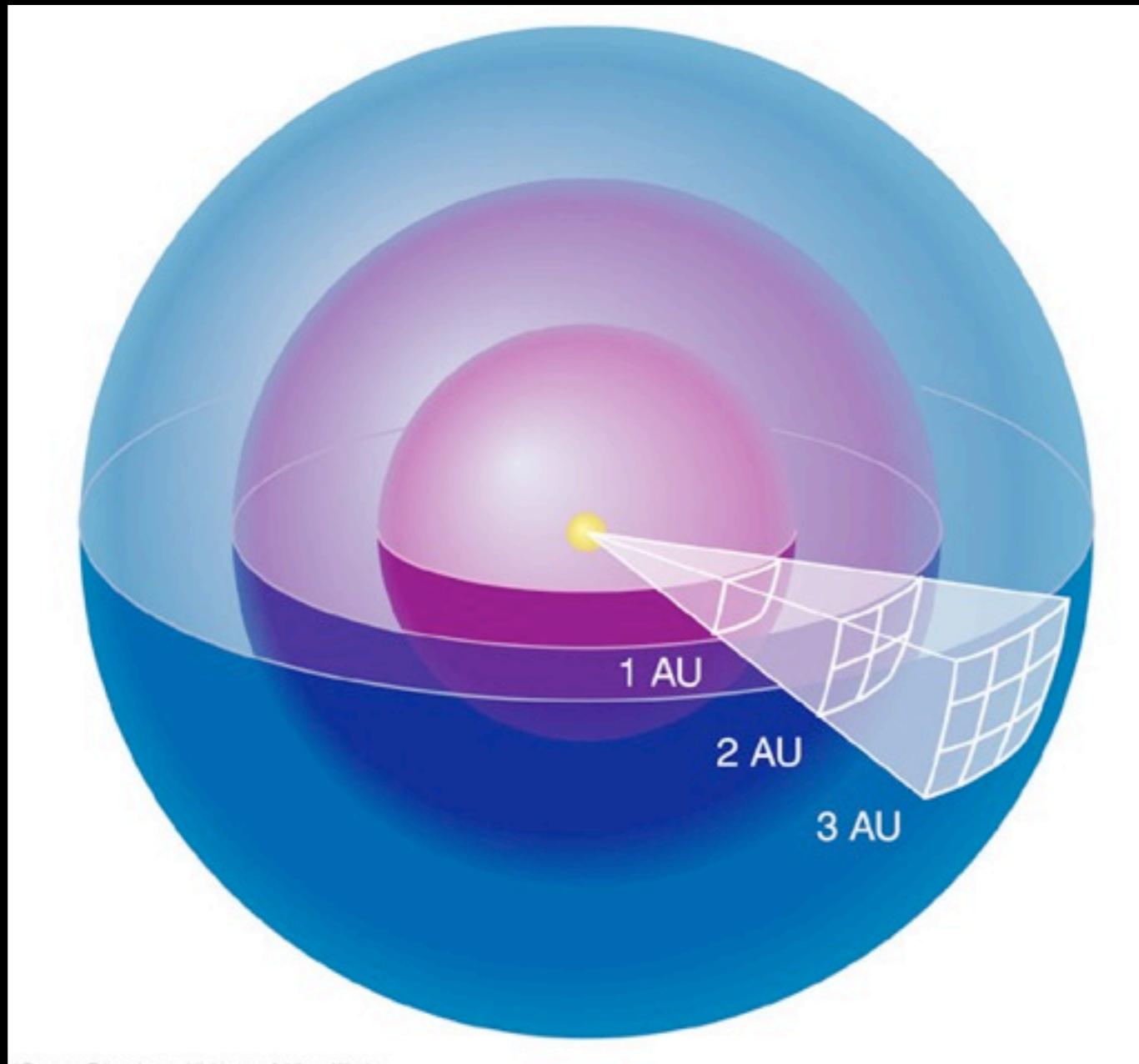
VLBI and GAIA

- Very Long Baseline Interferometry allows resolutions of ~ 10 micro-arcseconds ($0.00001''$), so distances to some special objects called masers can be measured throughout the galaxy
- GAIA is a new astrometric satellite that will launch in 2012 and measure stellar parallaxes to millions of stars with ~ 20 micro-arcsecond accuracy

Stop: Tutorial Time

- Page 35: The Parsec

Distance past Parallax



$$\text{Apparent Brightness} = \frac{\text{Luminosity}}{4\pi \cdot \text{distance}^2}$$

Luminosity → Distance

- If you know exactly how bright a source is and how bright it appears, you can say how far away it is

$$F = L / 4 \pi d^2$$



Brightness & Distance

- Spica is about 3x further away than Canopus, but is the same luminosity. Which star is brighter, and by how much?
 - A) Spica, 3x
 - B) Spica, 9x
 - C) Canopus, 3x
 - D) Canopus, 9x
 - E) They are the same brightness



Brightness & Distance

- Spica is about 3x further away than Canopus, but is the same luminosity. Which star is brighter, and by how much?

D) Canopus, 9x

$$L_{\text{Spica}} = L_{\text{Canopus}} \quad d_{\text{Spica}} = 3 d_{\text{Canopus}}$$

$$F_{\text{Spica}} = L_{\text{Spica}} / 4 \pi d_{\text{Spica}}^2$$

$$F_{\text{Canopus}} = L_{\text{Canopus}} / 4 \pi d_{\text{Canopus}}^2$$

Substitution:

$$F_{\text{Spica}} = L_{\text{Canopus}} / 4 \pi (3 d_{\text{Canopus}})^2 = 1/9 F_{\text{Canopus}}$$

$$\text{or } F_{\text{Canopus}} = 9 F_{\text{Spica}}$$

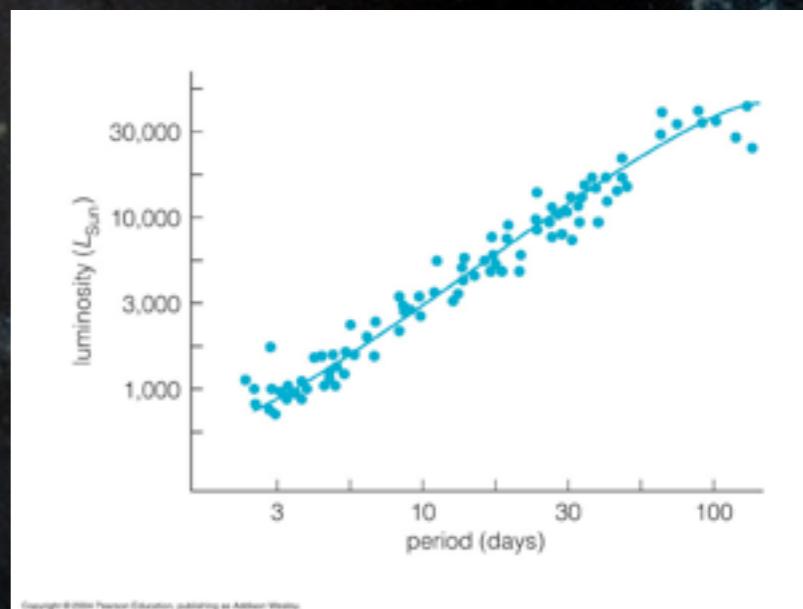


Brightness & Distance

- Becrux and Deneb have the same brightness, but Deneb is 9x more luminous. Which is closer, and by how much?
 - A) Deneb, 3x
 - B) Deneb, 9x
 - C) Becrux, 3x
 - D) Becrux, 9x
 - E) They are at the same distance

Measuring Luminosity

- More details in stellar classification, but...
- There are special stars called Cepheid Variables that follow a period-luminosity relationship
- Luckily, these stars are very large and very bright - we can use them to measure the distance to other galaxies
- The trick is that we use timing measurements to get the luminosity and imaging to get the brightness

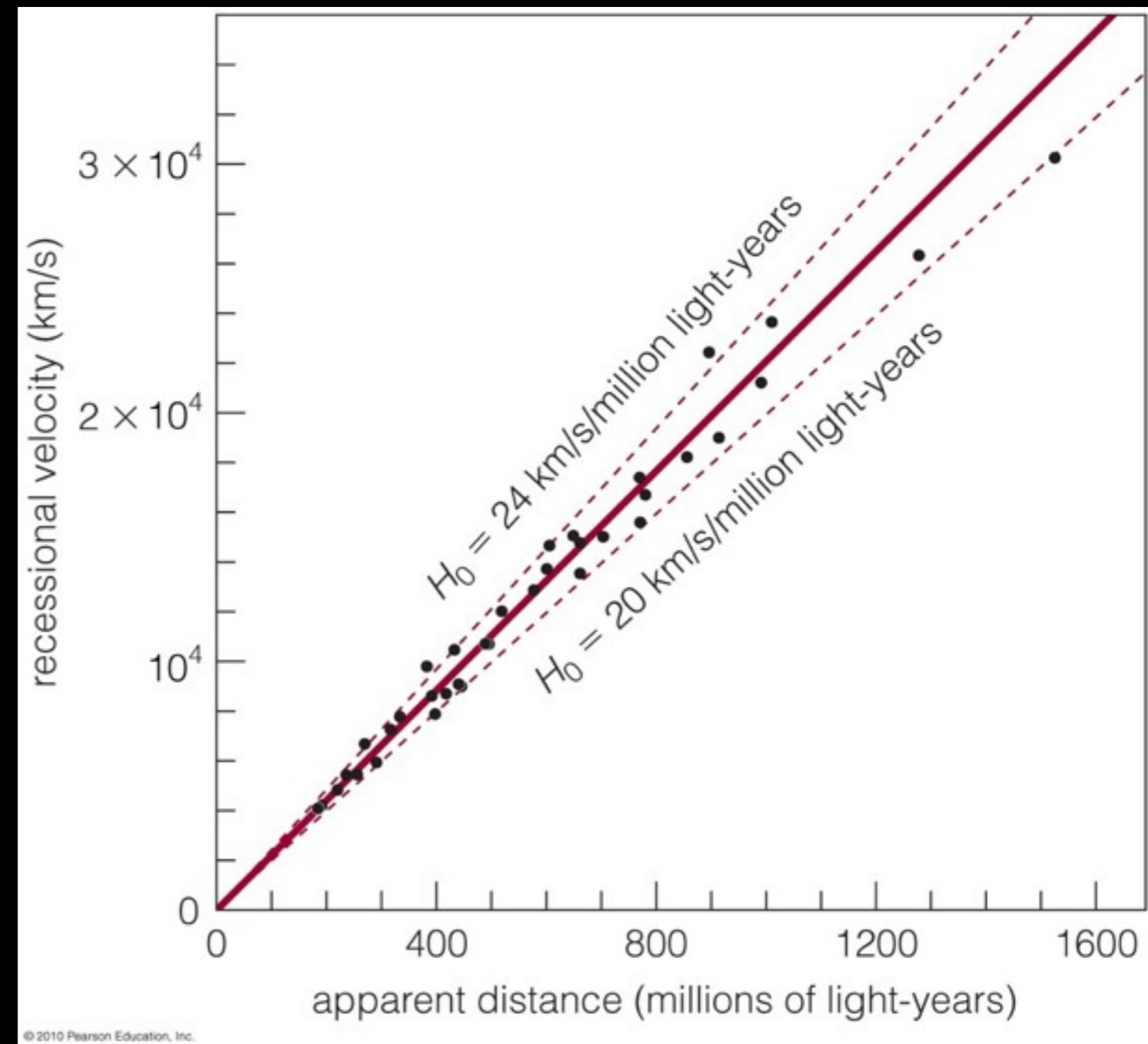


Cepheid Period-Luminosity Relation

- The luminosity of Cepheids is directly measured using parallax distances:
- $F = L / 4 \pi d^2$
- Timing measurements show that luminosity and pulsation period are related

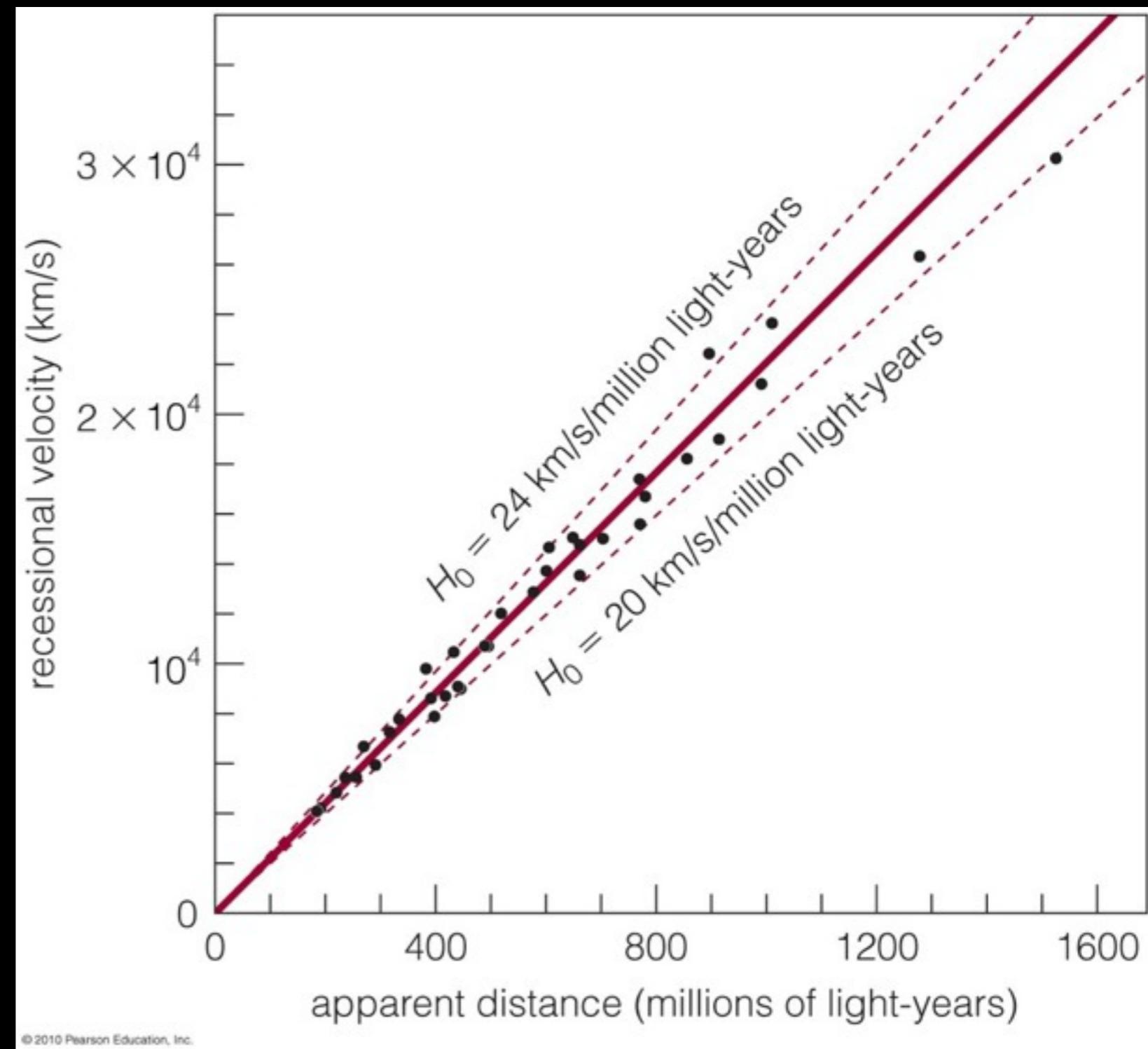
Distance past Cepheids

- Hubble's Law:
More distant galaxies are moving away faster
- Doppler Shift and Distance are correlated!



Distance past Cepheids

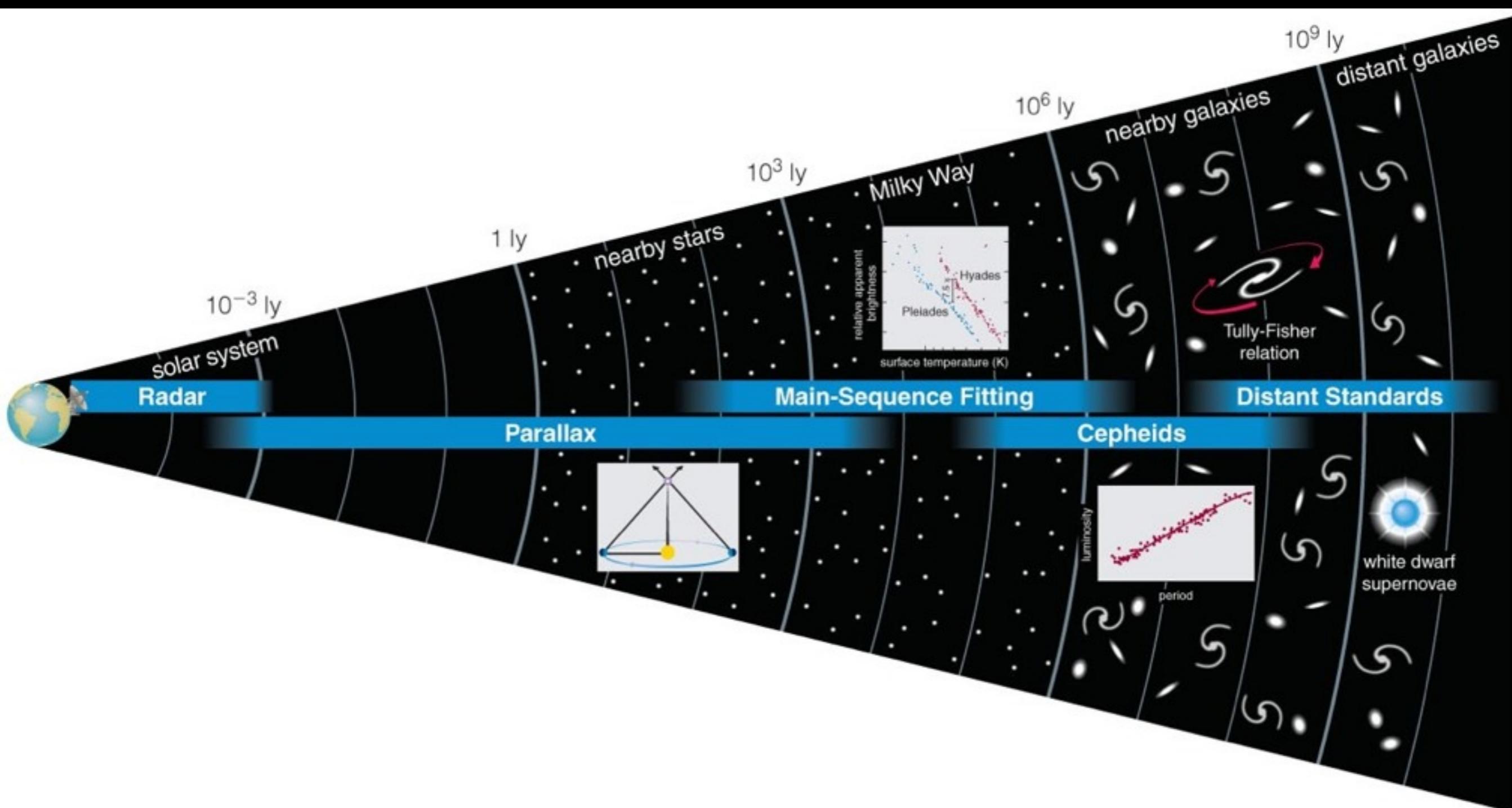
- Cepheids are used to determine the **distance** to galaxies, and **spectroscopy** is used to measure the **doppler shift**
- We use multiple observational methods to infer distances we can't directly measure



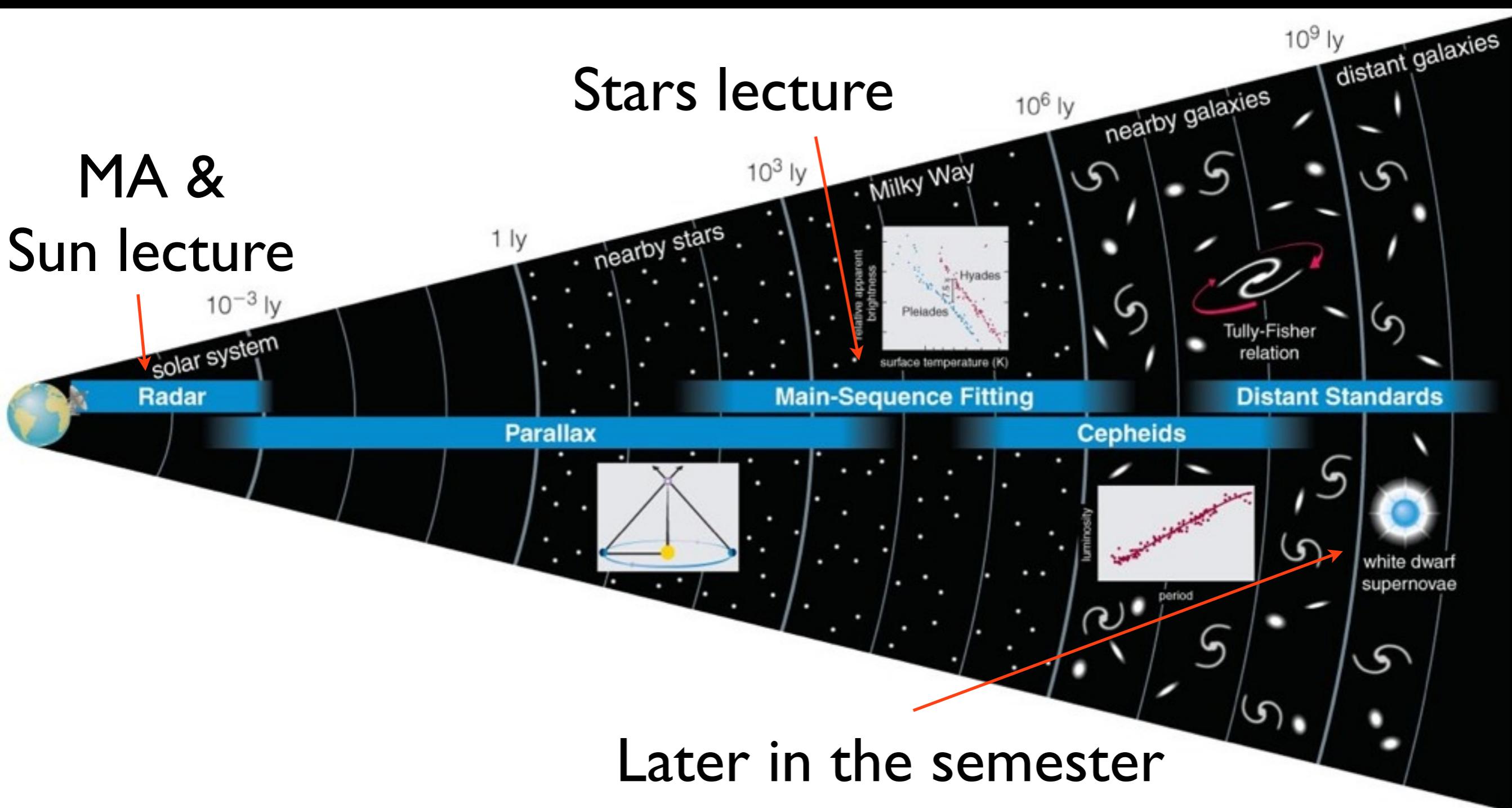
Process of Science: Empirical Law

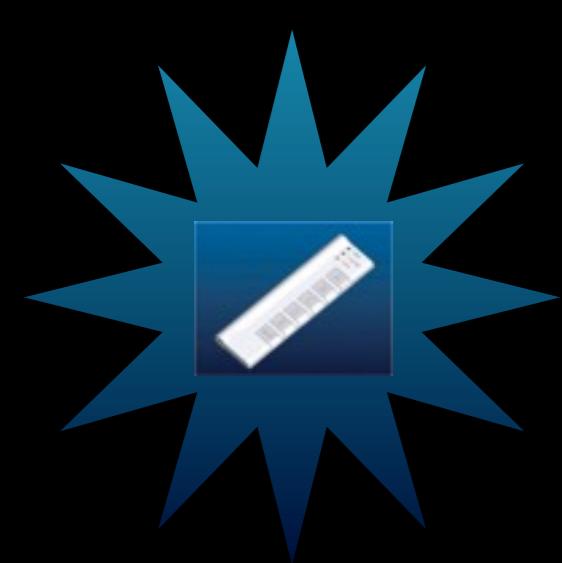
- Parallax is simply geometry; it can easily be verified by experiments in every-day life
- The Cepheid Period-Luminosity relation and Hubble's Law are “Empirical Laws”: we observe that they are true for many objects, and assume they are true for all similar objects
- Empirical Laws can easily be disproved if counterexamples are discovered

The Distance Ladder



The Distance Ladder





Distance

Can we use parallax to measure the distance to galaxies?

- A) Yes, Very Long Baseline Interferometry can do it
- B) No, they are too faint
- C) No, they are too far away



Distance

Can we use parallax to measure the distance to galaxies?

C) No, they are too far away

The closest galaxy, Andromeda, is at a distance of 720 kiloparsecs, or a **parallax** of 1.4 micro-arcseconds, which is about 10x smaller than modern telescopes can measure

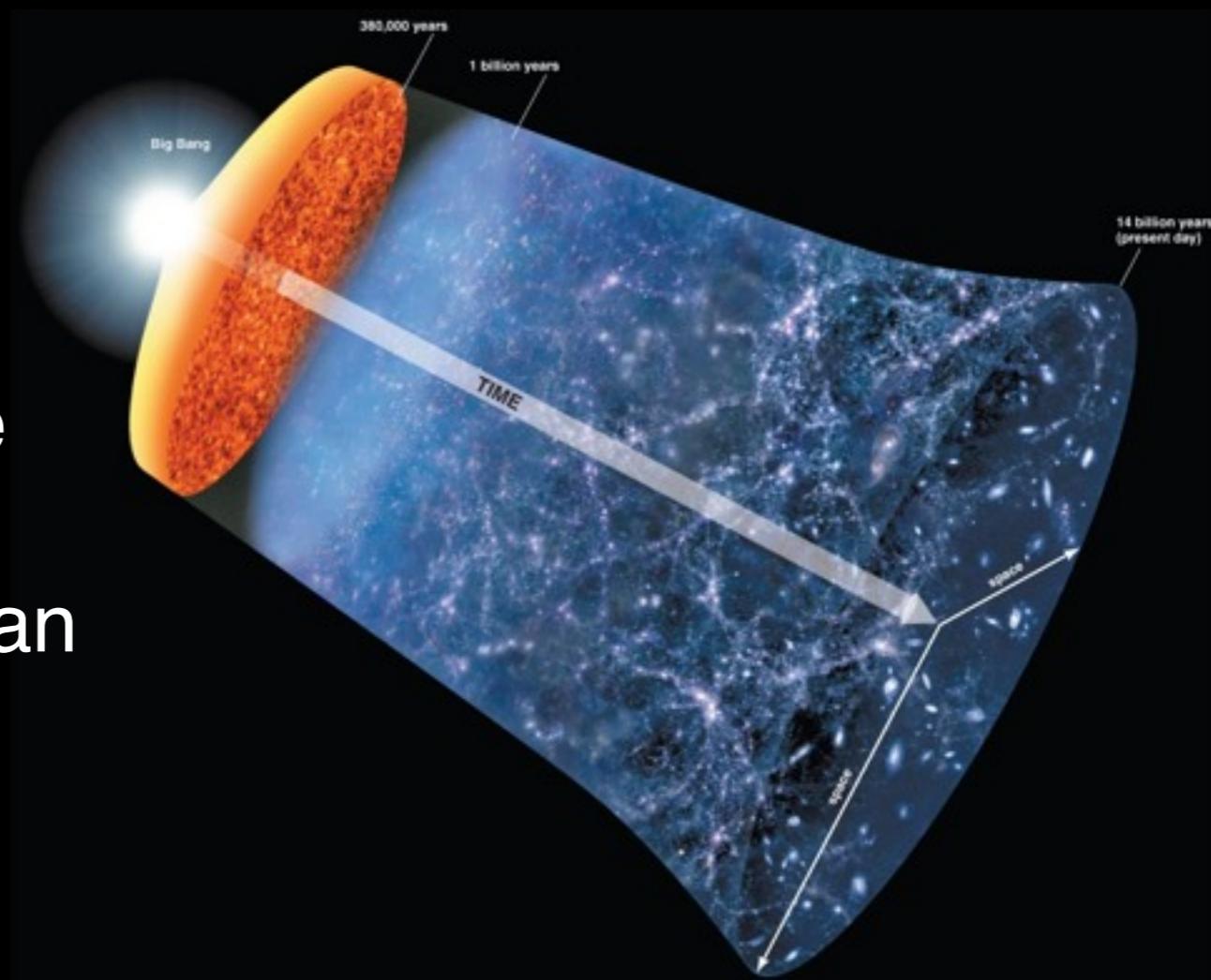
Distance and Time

- A light-year is a measure of distance
- But because it takes light 1 year to travel a light-year, we are looking *back in time* when we look at distant objects
- The sun's light we see on Earth escaped its surface 8 minutes ago



Look-back time

- The center of our Galaxy is ~24,000 light-years away, so any light from the galactic center was emitted before the written human history began
- The nearest galaxy, Andromeda, is 2.5 million light-years away
- Millions of galaxies have been observed at ages >5 billion years, older than the Sun





What time is it?

- Even though the light we see was emitted a long time ago, we talk about objects in space as they appear **now** - we talk about things in terms of the time we **receive** their light



Look-back time

- What light can we see **now** that was emitted around the time **dinosaurs ruled the earth**?
 - A) Light from nearby planets
 - B) Light from nearby stars
 - C) Light from far away stars in the center of the Milky Way
 - D) Light from moderate-distance galaxies
 - E) Light from the beginning of the universe

Learning Goals Revisited

- Why is it important to know how far something is?
 - Need distance to determine size and other properties
- How can astronomers measure distance?
 - Parallax and many bootstrap methods

Lecture Tutorial

- Page 39: Parallax and Distance



A detailed, high-resolution image of the Sun's surface dominates the background. The Sun is shown in a greenish-yellow hue, with its intricate patterns of solar flares and prominences visible. The flares are bright yellow and orange, contrasting with the darker, purple-tinted areas of the solar atmosphere.

Tomorrow:
THE SUN