



# APODs

Day 17 (July 28th, 2010): Galaxies: Groups, Clusters, & Dark Matter

<http://apod.nasa.gov/apod/ap100529.html> - Black Holes in mergers -- **Mike B.**

<http://apod.nasa.gov/apod/ap080823.html> - The Bullet Cluster -- **Raleigh**

Day 18 (July 29th, 2010): Clusters & Large Scale Structure

<http://apod.nasa.gov/apod/ap100502.html> - Coma Cluster -- **Miles**

<http://apod.nasa.gov/apod/ap090716.html> - Hercules Cluster -- **Naveen**

Day 20 (August 2nd, 2010): Large Scale Structure & The Distance Ladder

<http://apod.nasa.gov/apod/ap091028.html> - Furthest galaxy cluster -- **Josh**

<http://apod.nasa.gov/apod/ap071211.html> - The Universe nearby -- **Emily**

Day 21 (August 3rd, 2010): Cosmology & CMB

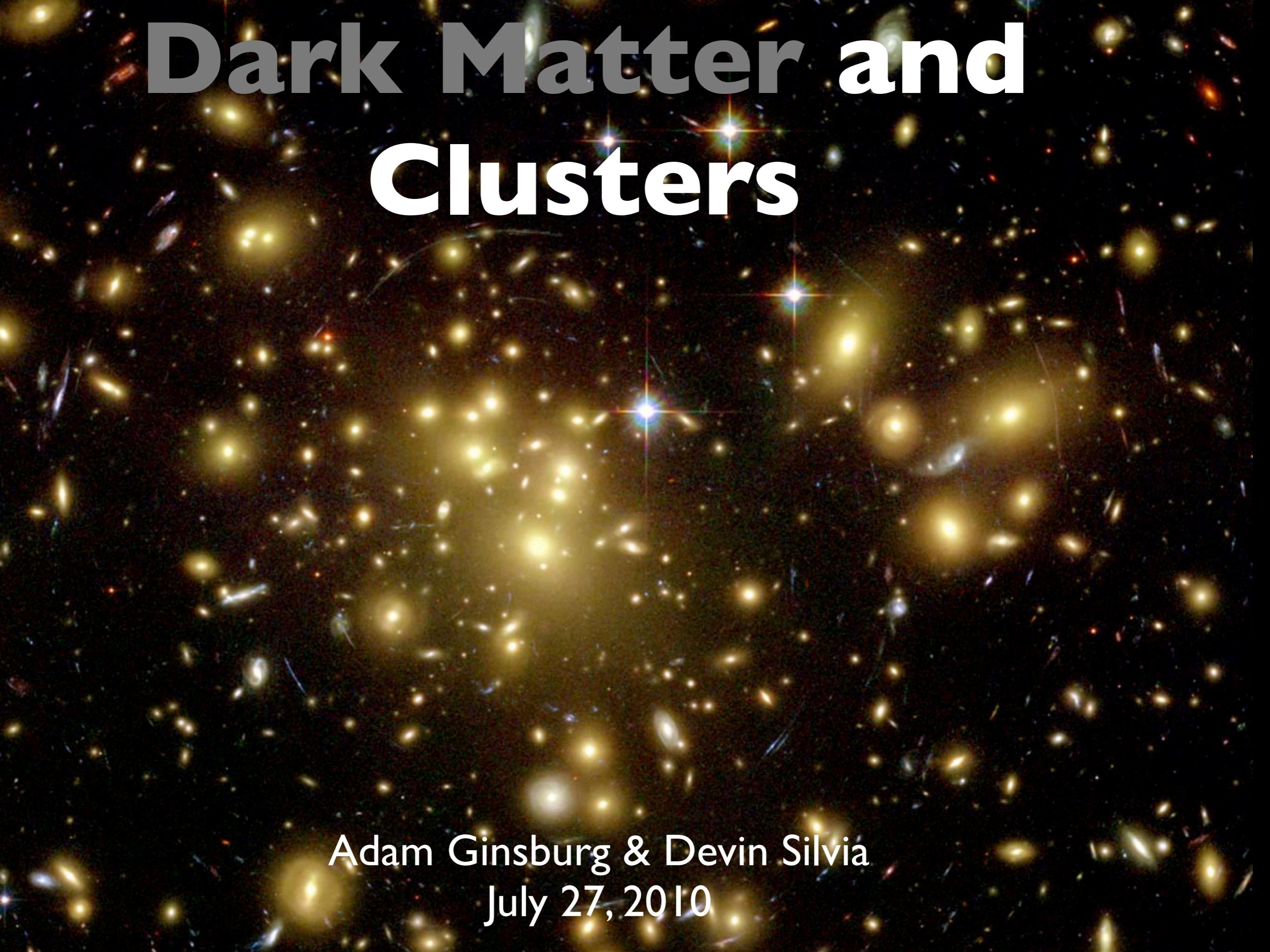
<http://apod.nasa.gov/apod/ap031231.html> - WMAP CMB, the age of the universe -- **Lisle**

<http://apod.nasa.gov/apod/ap061007.html> - COBE CMB -- **Forbes**

Day 22 (August 4th, 2010): Expansion/Fate of the Universe & Dark Energy

<http://apod.nasa.gov/apod/ap010404.html> - Type Ia and expansion -- **Tyler**

# Dark Matter and Clusters



Adam Ginsburg & Devin Silvia  
July 27, 2010

# Learning Goals

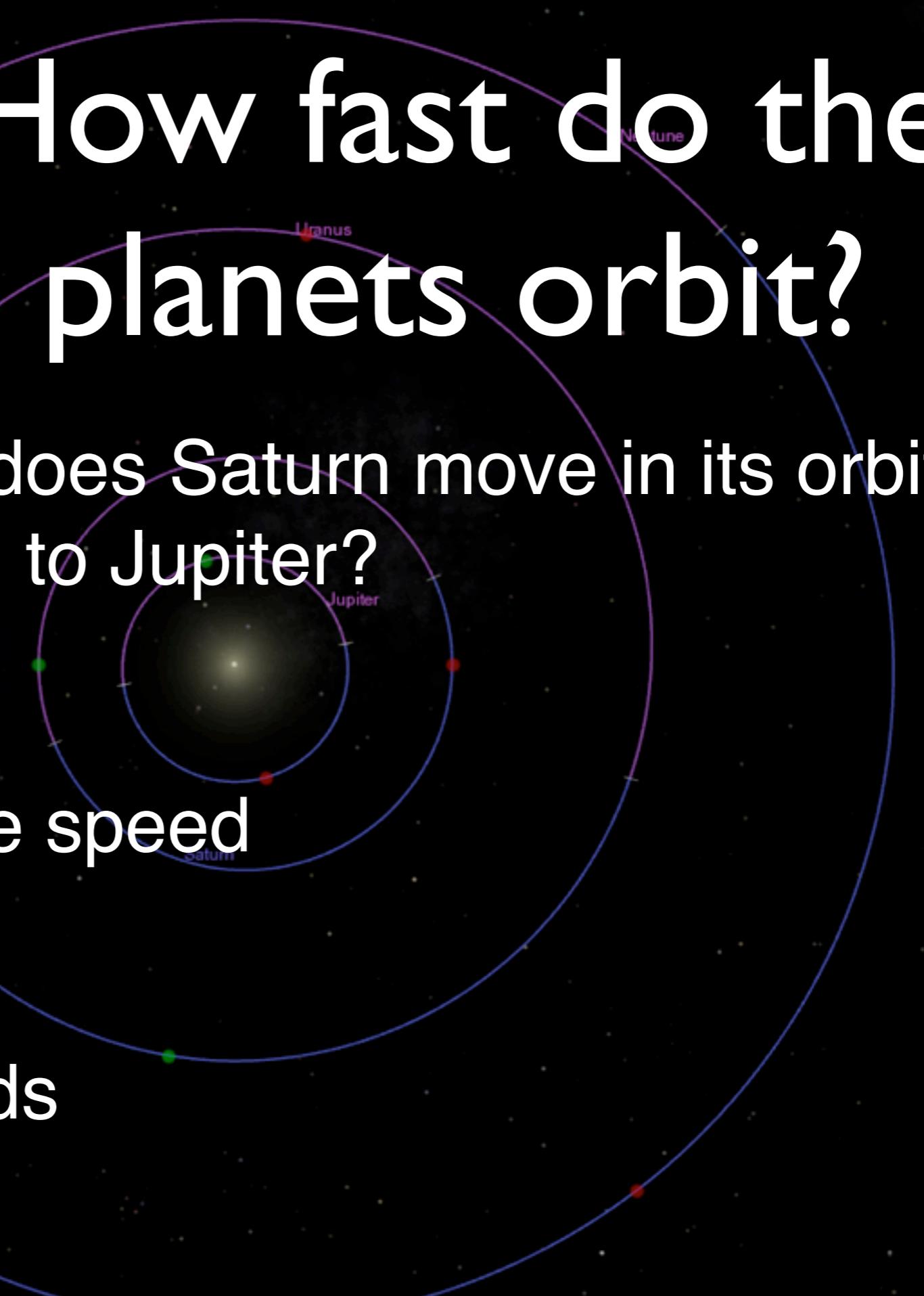
- If it's dark, how do we know it's there?
- What do you mean, dark?
- What are galaxy clusters?
- What do we use galaxy clusters for?

# Gravity, Angular Momentum

- Your old friends (or nemeses)
- $F = G M m / r^2$
- $m v r = \text{constant}$



# How fast do the planets orbit?

- How fast does Saturn move in its orbit compared to Jupiter?
- A) Faster
- B) The same speed
- C) Slower
- D) Backwards
- 



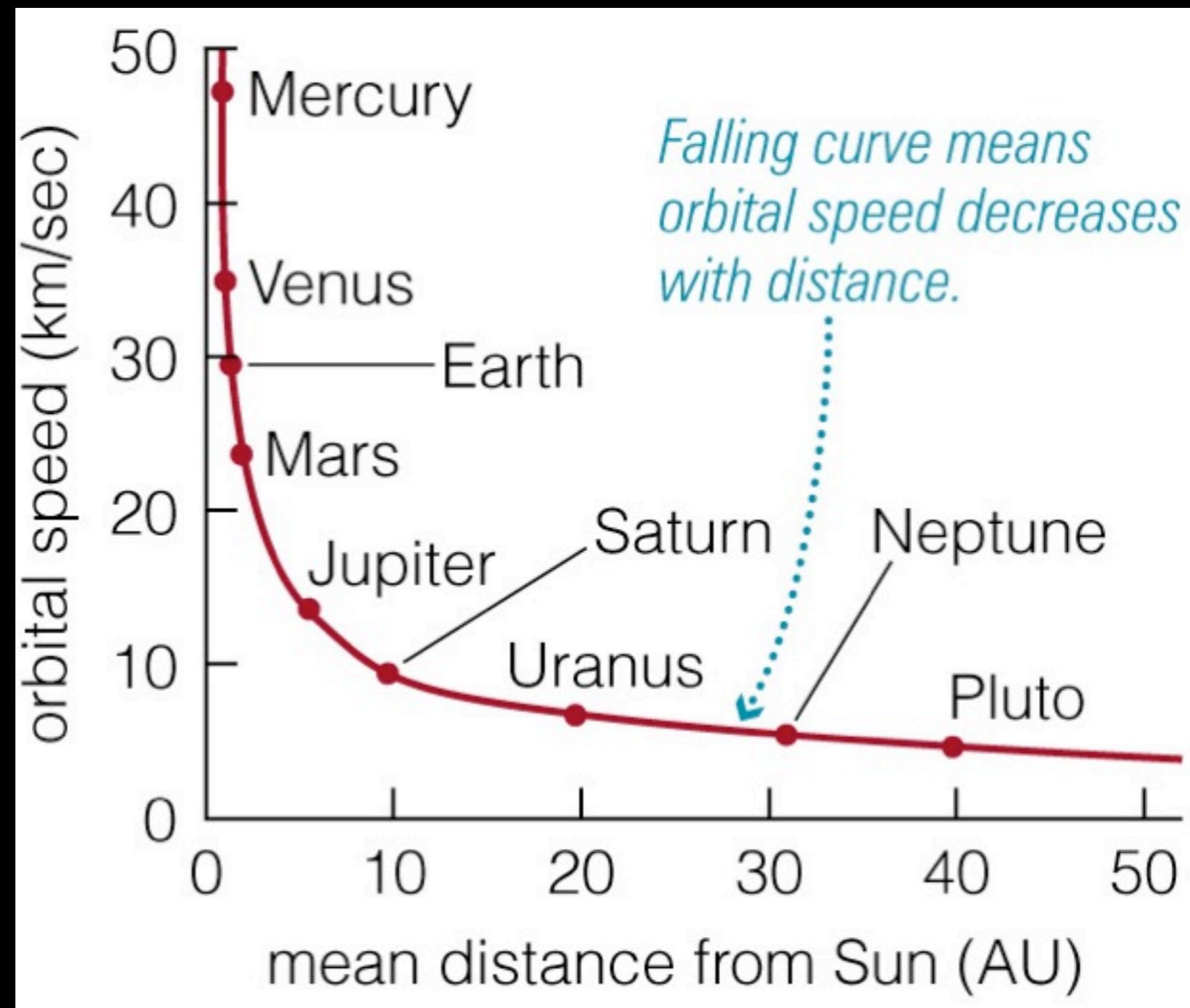
What would happen to the planets' orbits if the Sun suddenly turned into a black hole of the same mass?

- A) They'd fly off along a straight line, as if the string had been cut
- B) They'd continue along their orbits as if nothing had happened
- C) They would spiral in to the black hole
- D) They would get cold and tired and then just give up and forget about the whole thing

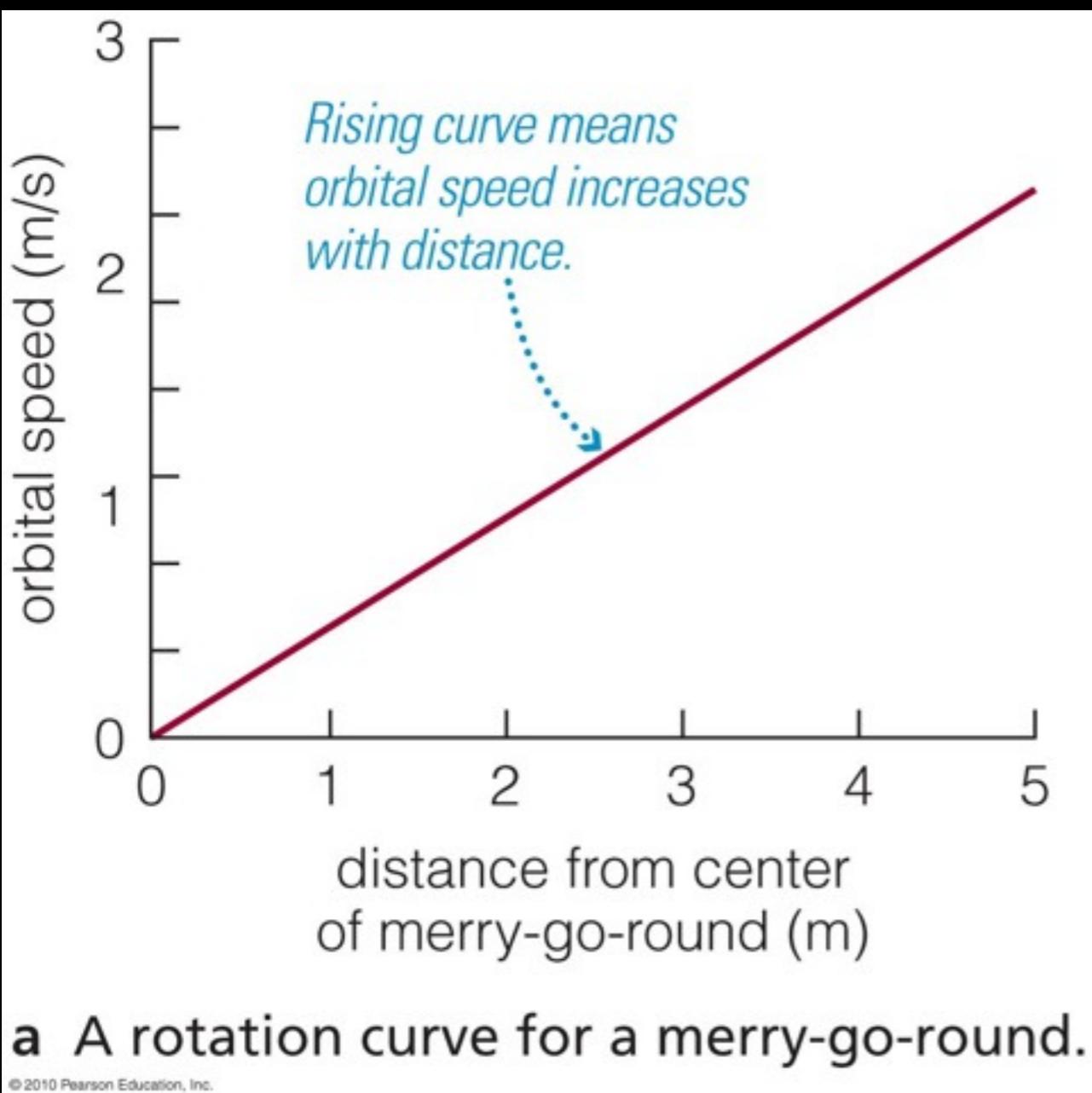


# Gravity controls orbital speed

- velocity drops as the square root of distance
  - $v \propto 1/\sqrt{r}$
- What **mass** is responsible for nearly all of the **gravitational force** in the solar system?



# What does a merry-go-round do?



- Things on the outside go faster
- Gravity won't do this, but “solid body rotation” will

# Rotation Curves

- OK, so if the **mass** is all in the center, velocity drops off with **radius**.
- Let's do the same exercise for a galaxy
  - How do we measure the rotation rate of a galaxy?

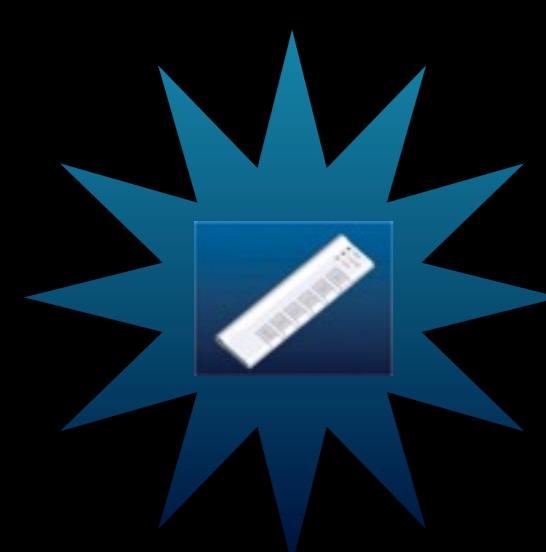


# Rotation Curves

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Spectroscopy to measure Doppler Shift

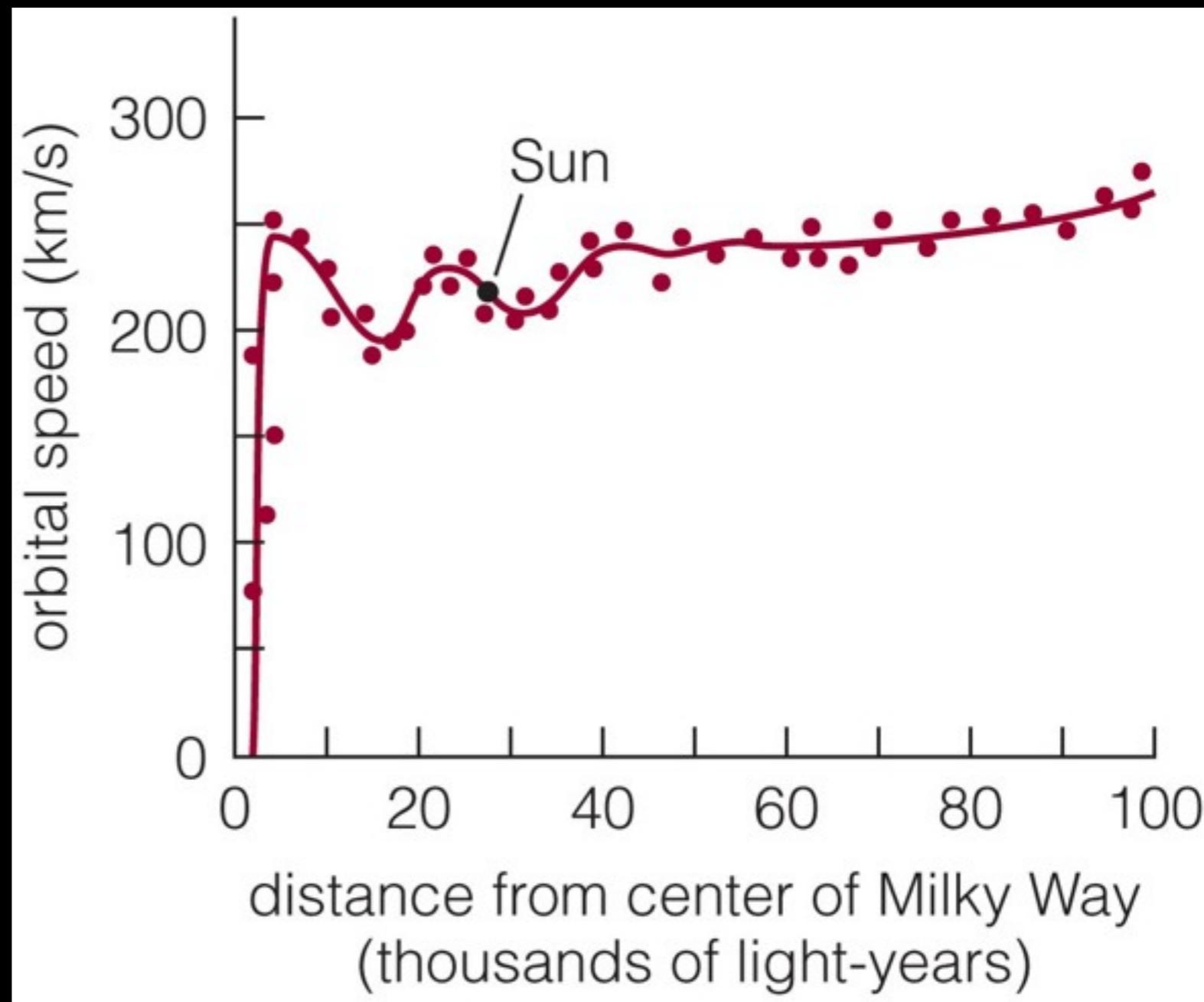


- 
- What do you expect the “velocity curve” to do based on where most of the stars are?
    - A) Decrease with distance, like in the solar system
    - B) Increase with distance, like on a merry-go-round
    - C) Stay constant
    - D) Do something wonky; increase and decrease



# What's really there?

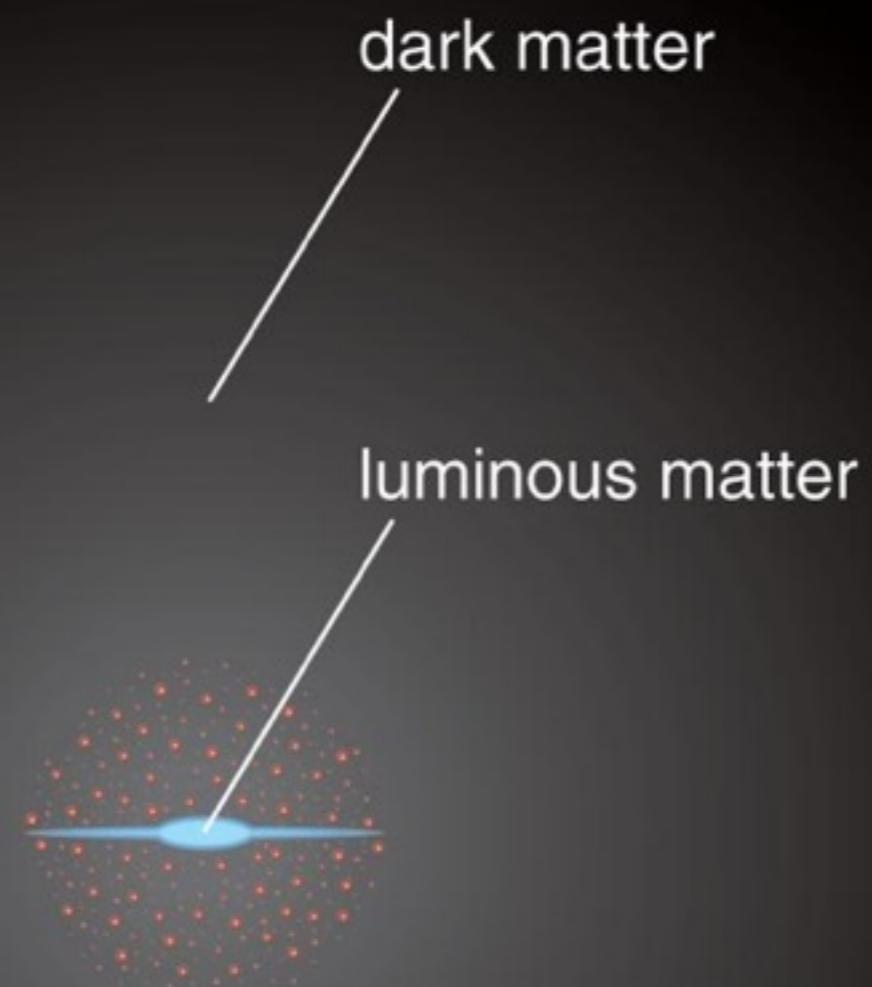
- The orbital speed stays nearly constant!
- That means there's a lot of mass that's not in the center!
- But... it *looks* like most of the mass is in the center



# Enter... *Dark Matter*

- Eh, not really that dramatic
- We know that there is extra force pulling on gas and stars in the outer galaxy
- It must be from some mass exerting gravitational pull that we can't see

# Revised anatomy of a galaxy



# How much dark matter?

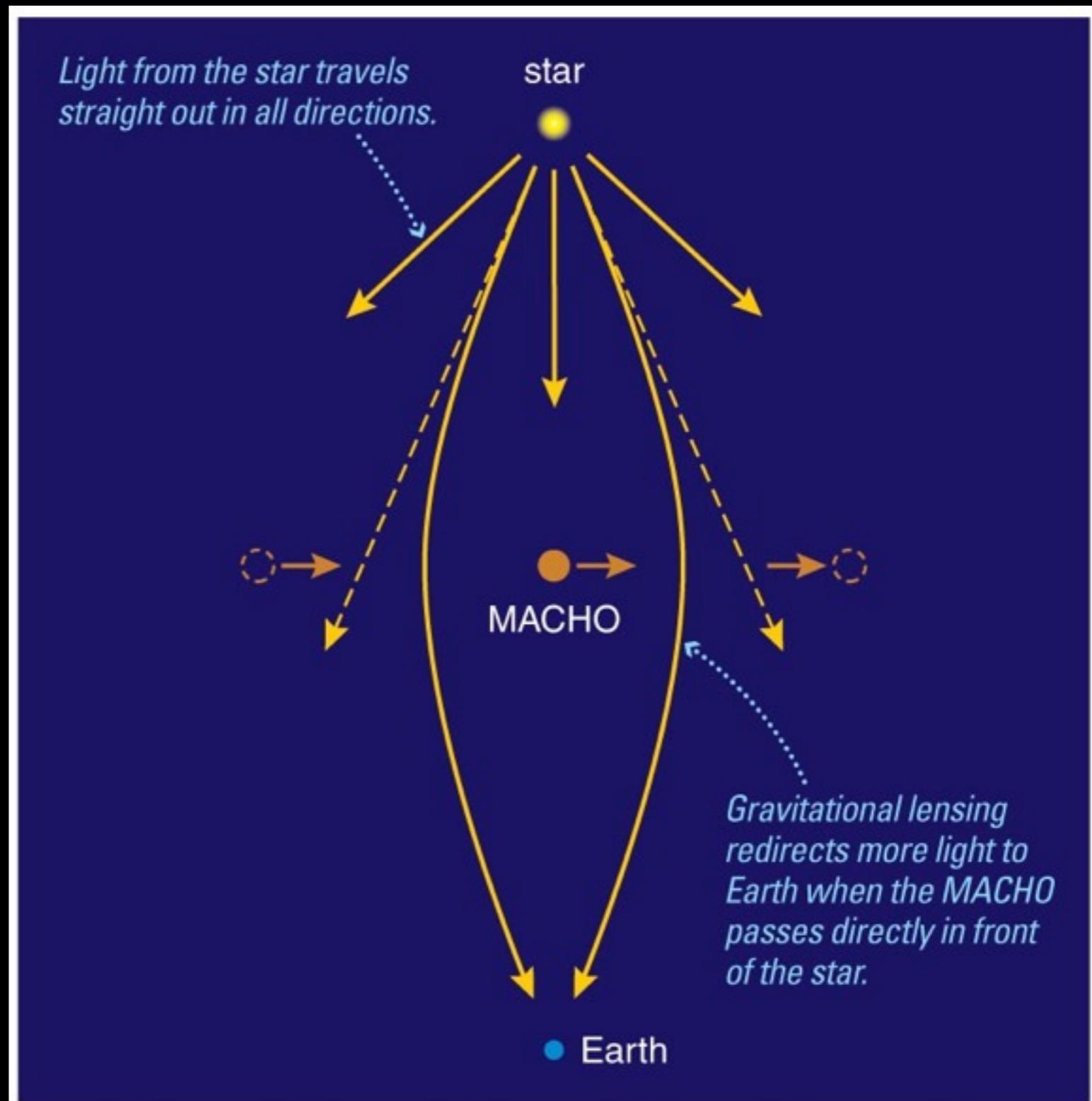
- Only 10% of the matter in a galaxy is from stars, gas, and dust (“normal” matter)
- The other 90% is from dark matter

# What is dark matter?

- Two hypotheses:
  - Dark matter is made of **normal matter**, but we can't see it because it's too faint
    - Example: black dwarfs, black holes, red dwarfs, brown dwarfs
  - Dark matter is made of **extraordinary matter** or **exotic particles**
    - For example, neutrinos
    - Or something else like neutrinos

# MACHOs

- MACHOs are Massive Compact Halo Objects (faint stars/planets)
- We can detect them if they pass in front of stars and lens them, making them briefly appear brighter



# MACHOs aren't it

- Two big surveys called MACHO and OGLE stared at the LMC and within the Milky Way
- They detected some MACHOs, but not enough (by a lot) to explain dark matter

# Exotic Particles

- WIMPs are Weakly Interacting Massive Particles
  - AKA Cold Dark Matter
  - Like neutrinos, they don't interact with normal matter most of the time
  - But while neutrinos are very very light (at most 1/500 the mass of an electron), WIMPs are heavy (maybe 10x the mass of a proton)
- These are the currently accepted explanation for dark matter. Experiments like the Cryogenic Dark Matter Search (CDMS) are trying to find WIMPs, but haven't yet





# MACHOs

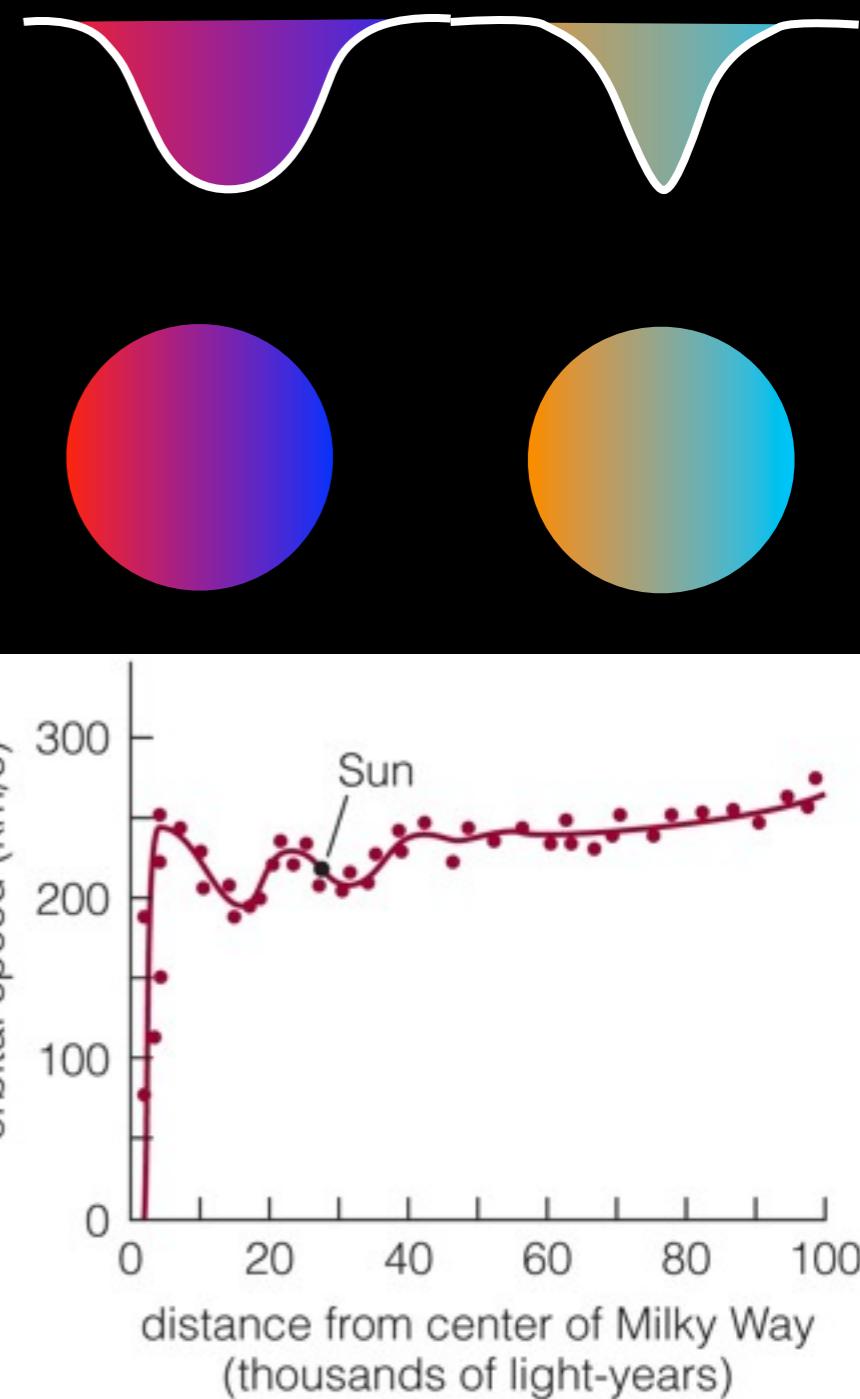
- Since 90% of the mass has to be from dark matter, how many invisible stars would there have to be in the Milky Way for MACHOs to explain dark matter? (the Milky Way has >100 billion stars)
  - A) 10 million
  - B) 10 billion
  - C) 1 trillion
  - D) 1 quadrillion
  - E) 100 bazillion

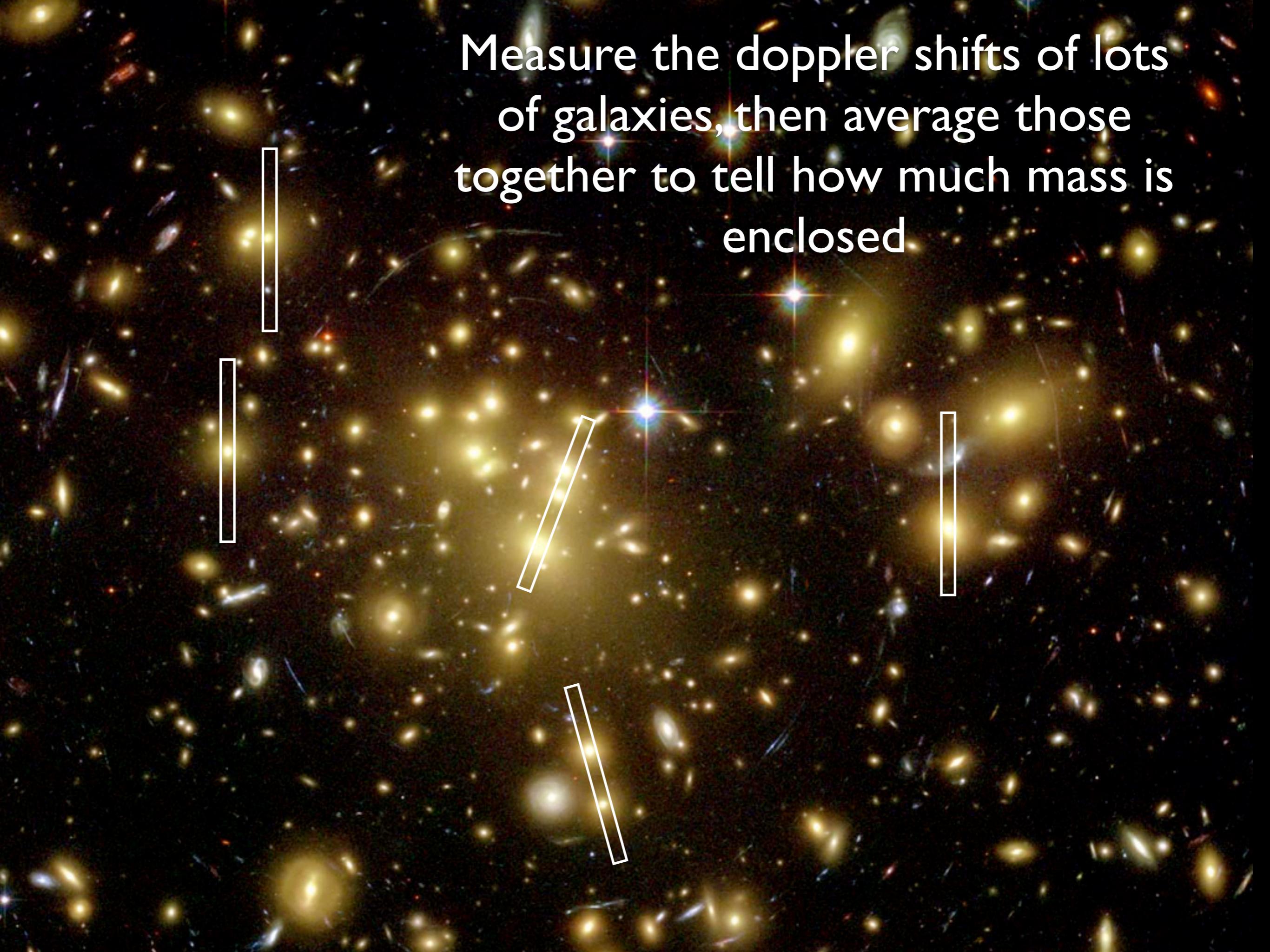
# If we haven't found it, do we *really* know it's there?

- We see it in **galaxies**, where else?
- Clusters
- There are 3 different ways we can tell there is dark matter in clusters
  - **Velocities of galaxies**
  - **Hot Gas x-rays**
  - **Gravitational Lensing**

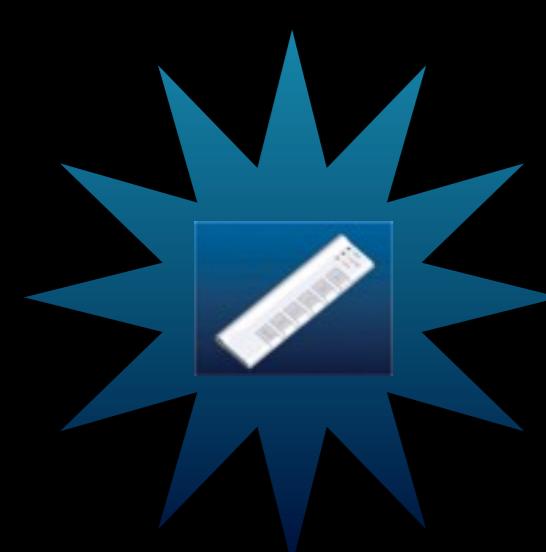
# Dopplerizing Galaxy Clusters

- Remember, a faster-spinning planet would have a *wider* emission or absorption line because one side is coming *towards*, one side is going *away* from you
- Also, if there is **more mass** in the center of a system, orbits are *faster*
- Put these together: **wider lines mean more mass!**
  - This is true even for clusters of galaxies and stars, in which they can be orbiting in opposite directions. Then, the width is a measurement of “**velocity dispersion**”





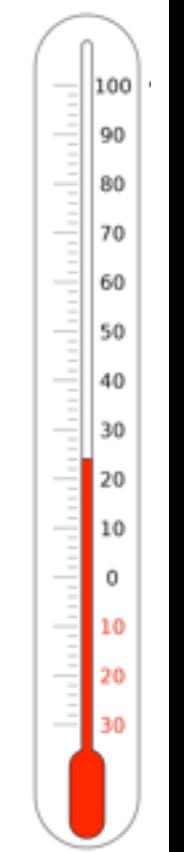
Measure the doppler shifts of lots  
of galaxies, then average those  
together to tell how much mass is  
enclosed



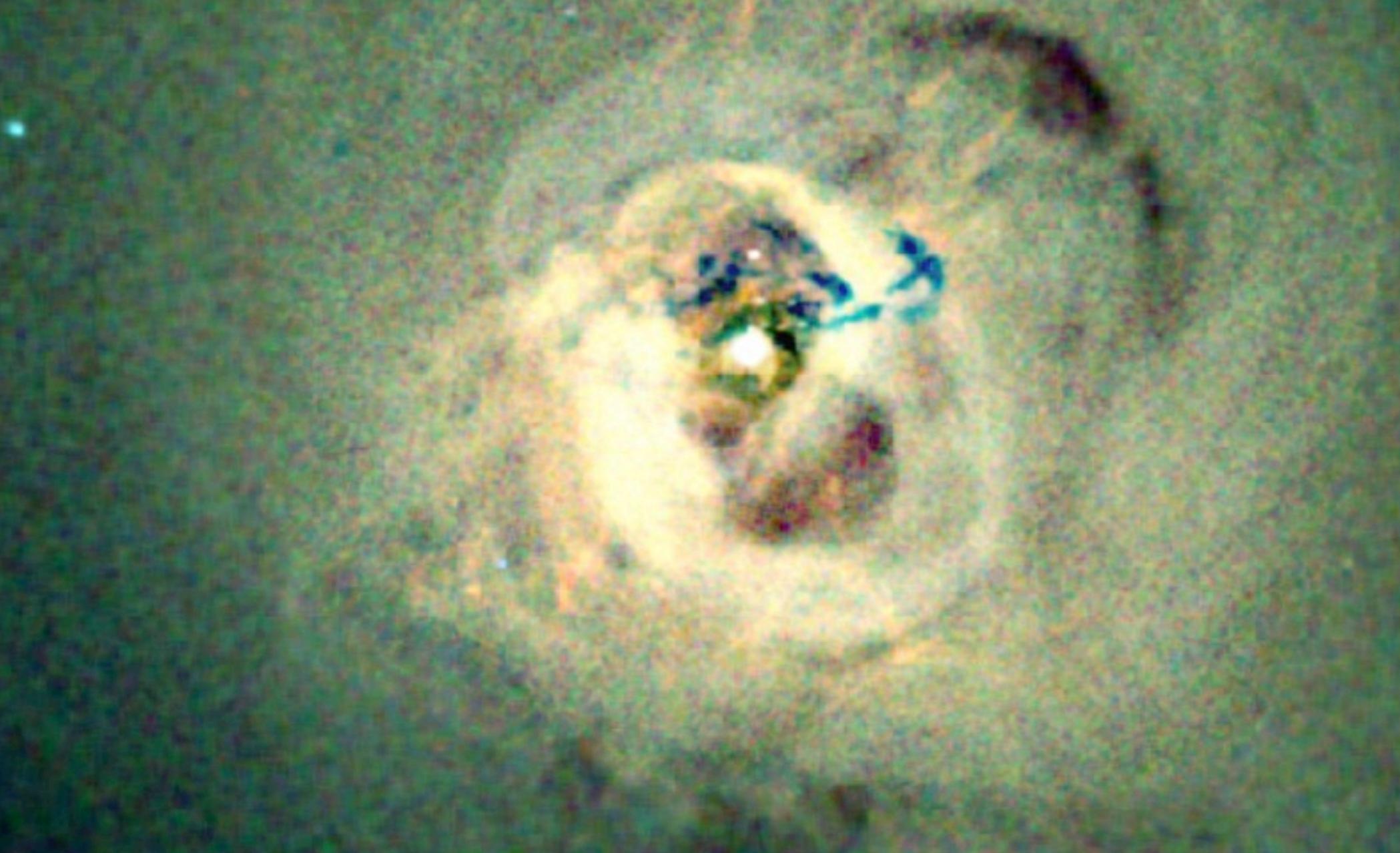
- We infer the presence of Dark Matter by assuming that galaxy clusters are *gravitationally bound*. If there was no dark matter in these **clusters** and galaxies were moving at the **same speed**, what would happen?
  - A) nothing special, there would just be less mass in the cluster
  - B) the galaxies would all fall in to the center
  - C) the gas would act like glue and hold the cluster together
  - D) the cluster would “evaporate” as galaxies escaped from the weaker gravity

# X-ray the clusters

- Clusters are filled with HOT gas
  - about 1 million K
- The gas has very, very low density, but clusters are BIG
- Let's look at a picture...



# BAM! Galaxy Cluster!

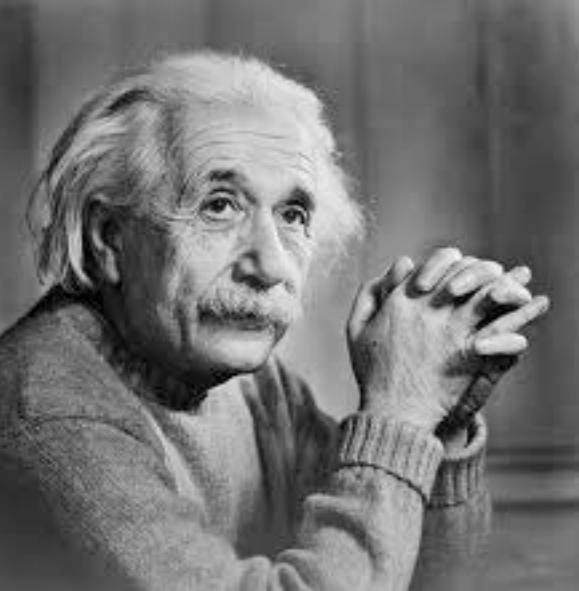
- 
- Notice anything weird about this cluster of galaxies?

# X-rays and hot gas

- We can only see the really hot gas
  - galaxies don't even show up in X-rays
- The hot gas outweighs the galaxies by about a factor of 2
- We know there is dark matter because we measure the *temperature* of the gas, and determine that it is so hot, without dark matter, it would evaporate too quickly to be seen

# X-ray the clusters

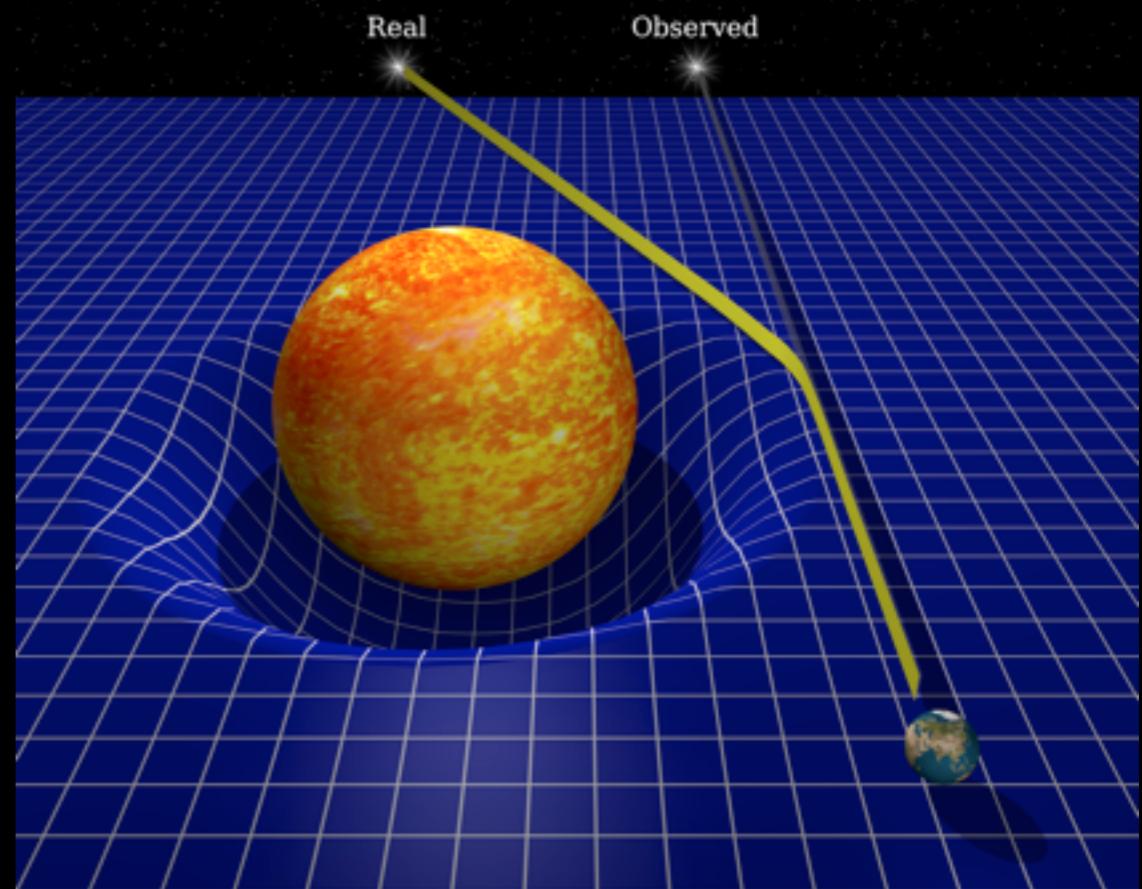
- Lots of galaxies in one big, diffuse, blob of X-ray gas

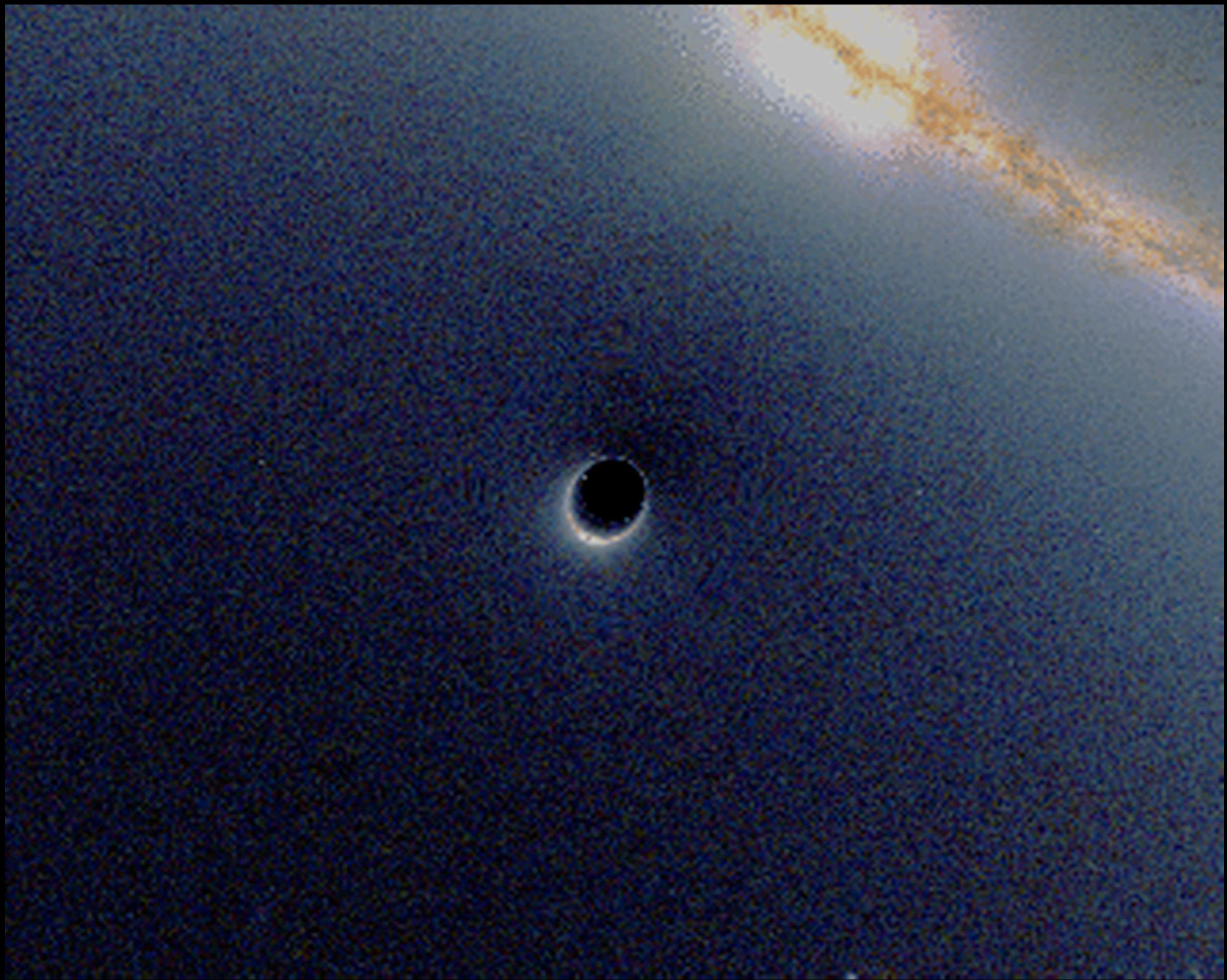


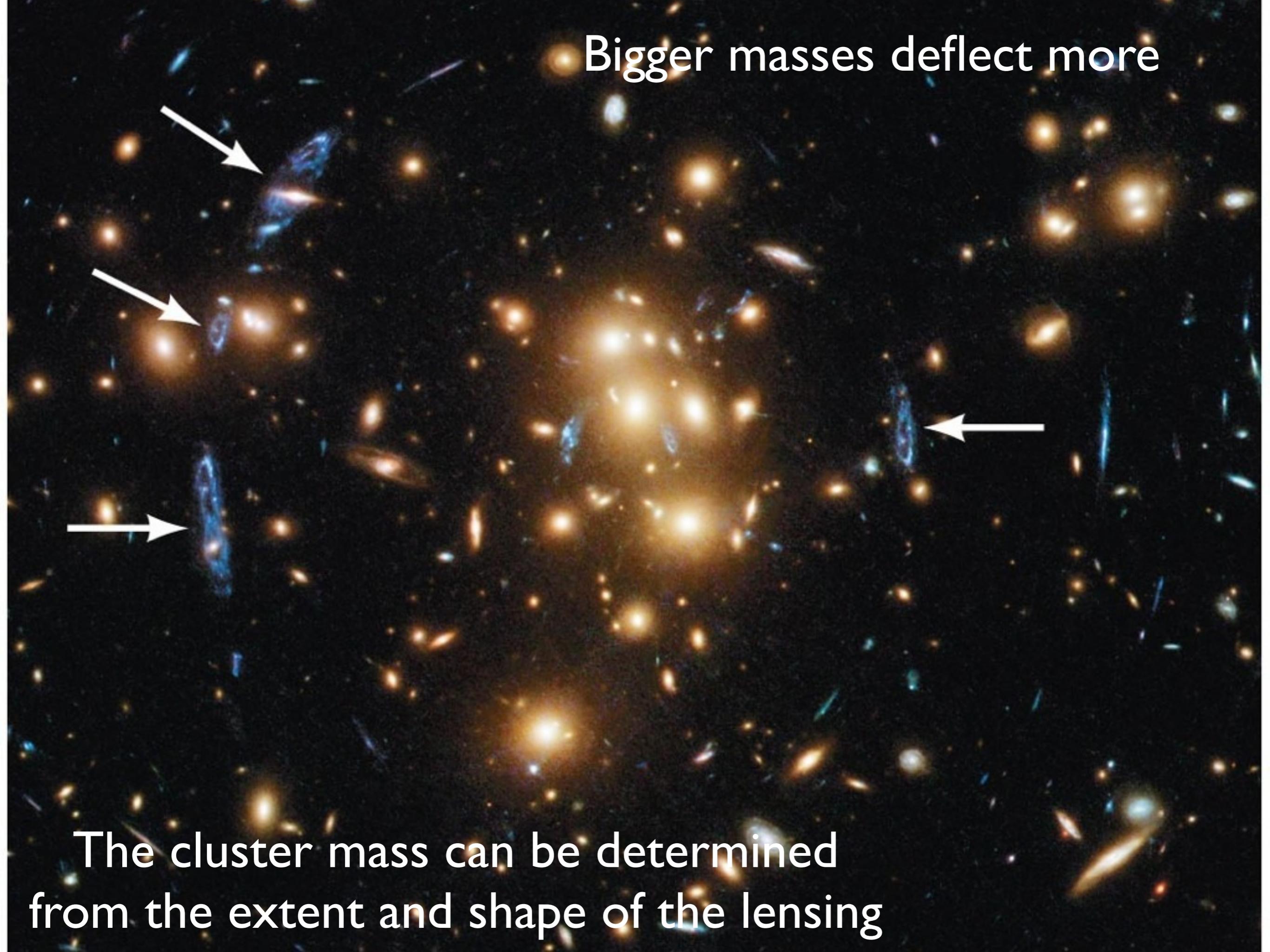
# Gravitational Lensing



- General Relativity says that light curves around massive objects
- Has been proved in the solar system: stars behind the sun are shifted by a small amount
- Made Einstein famous in 1919: an expedition by Arthur Eddington used a solar eclipse to prove general relativity







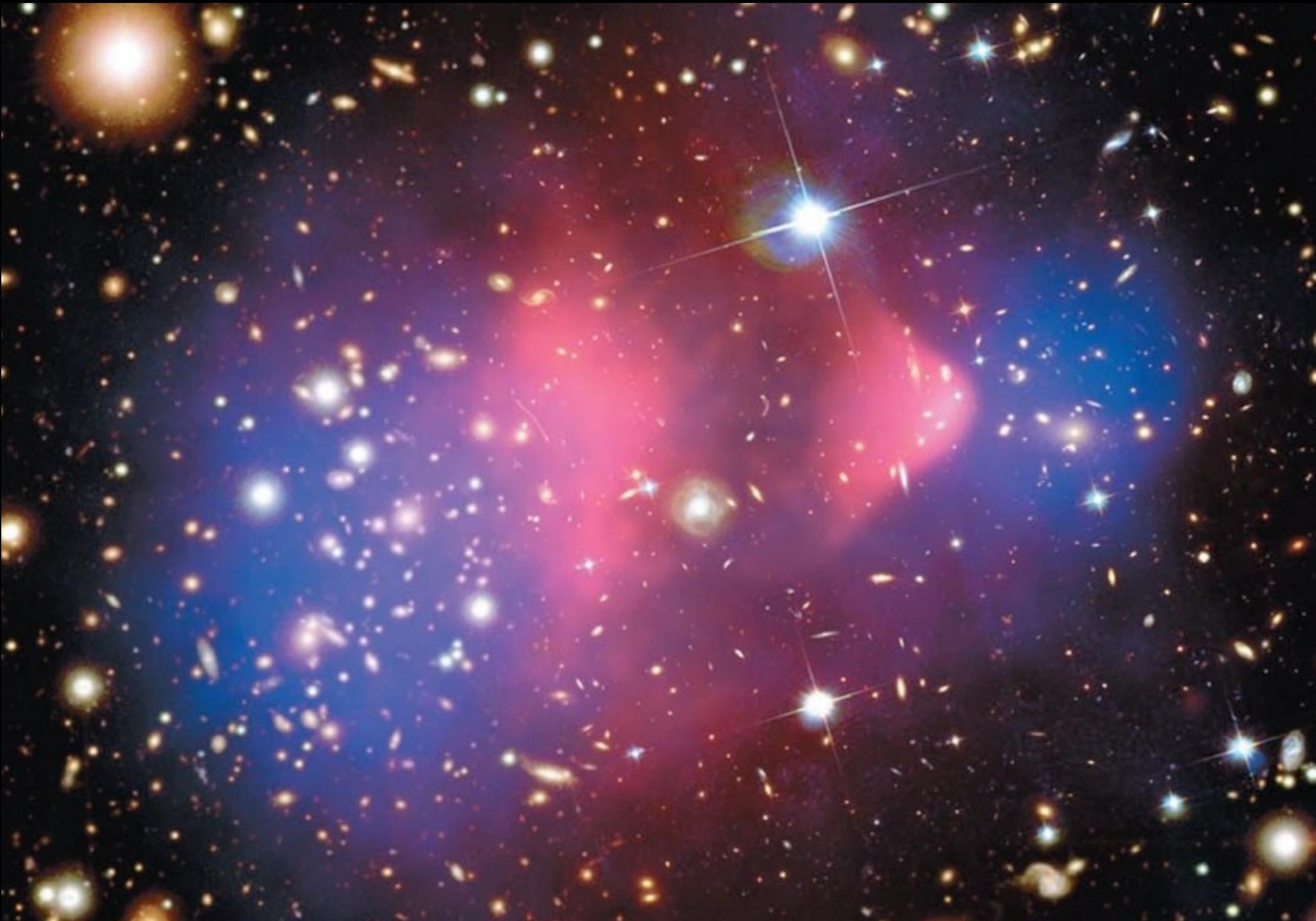
Bigger masses deflect more

The cluster mass can be determined  
from the extent and shape of the lensing

# Mass, mass, mass

- 3 ways to measure the **mass!**
- What do they tell us?
  - X-rays say that there's a lot more **hot gas** than there are stars - but that's still not dark matter
  - Doppler shift and gravitational lensing say there's a lot more mass than we can see

# Bullet Cluster

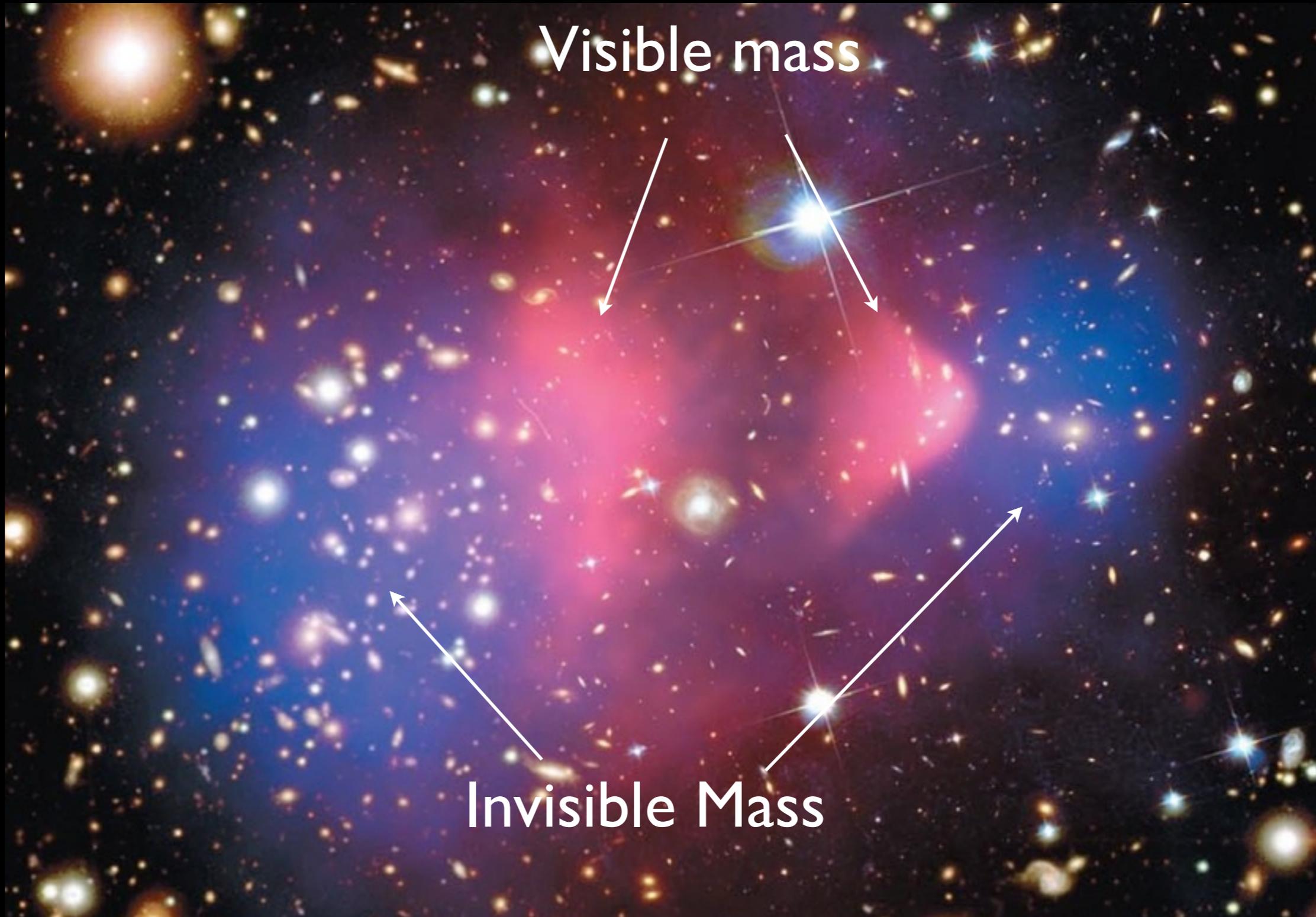


Dark Matter

X-rays

Optical

# Bullet Cluster



Dark Matter

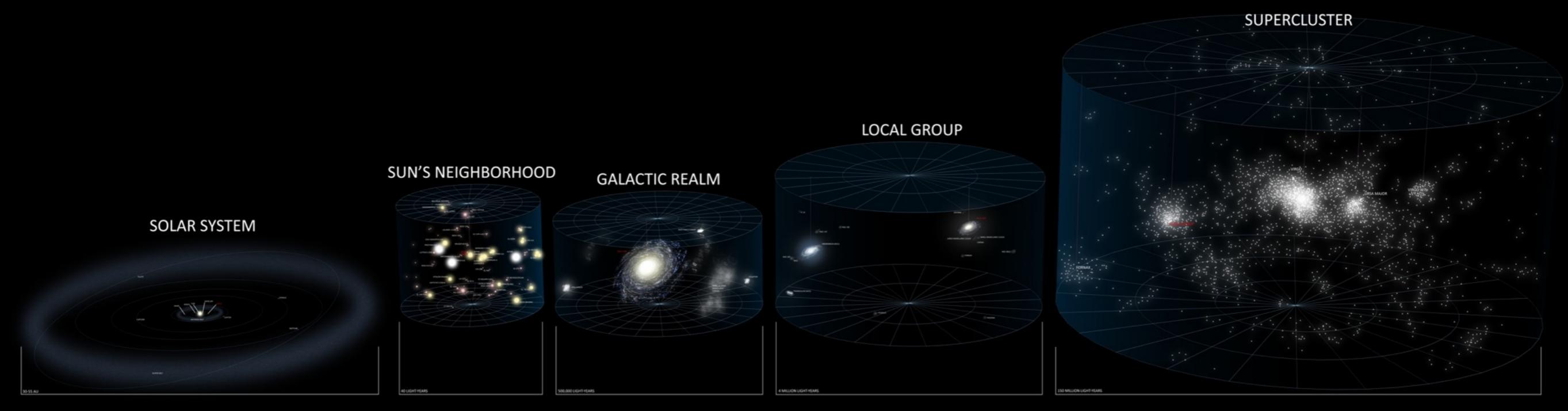
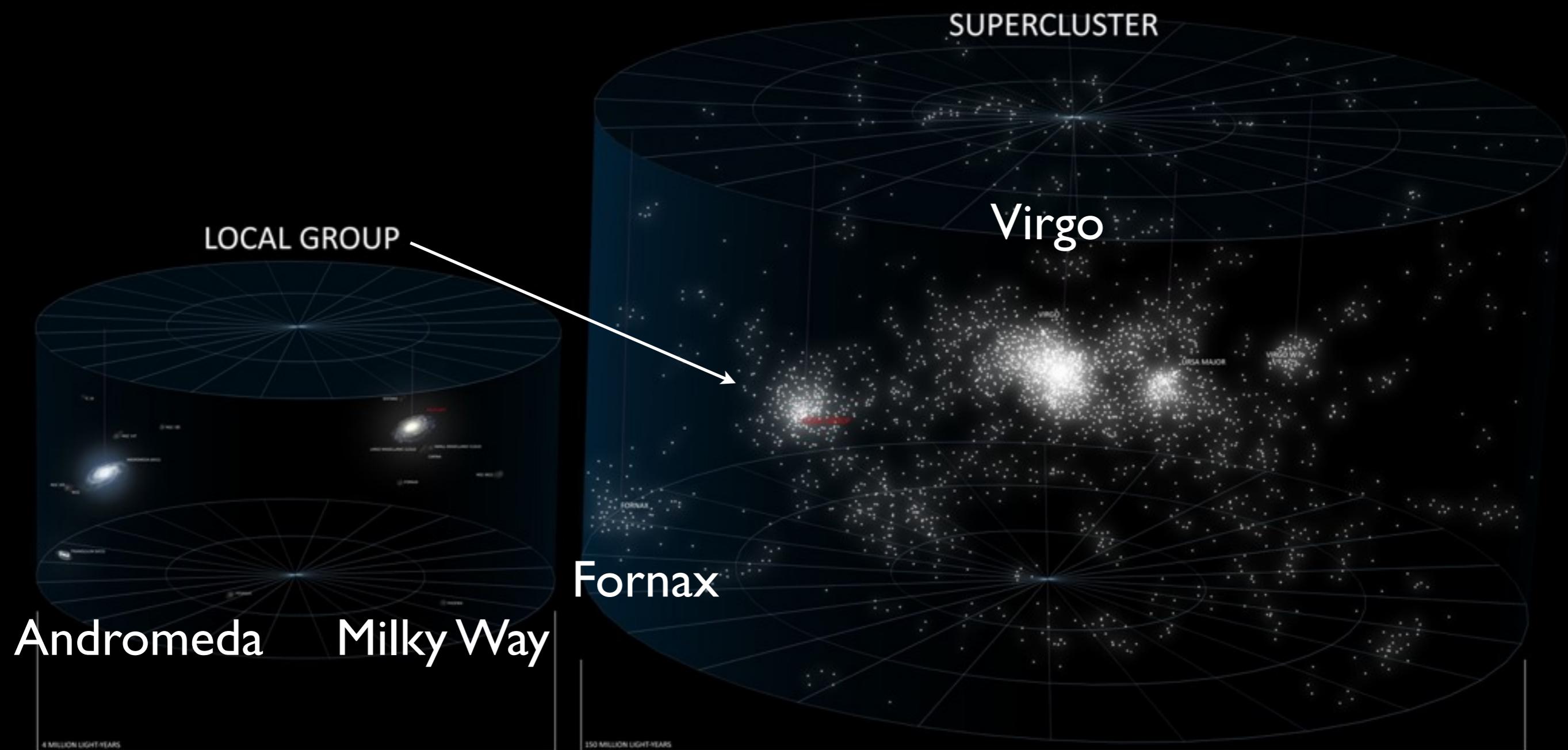
X-rays

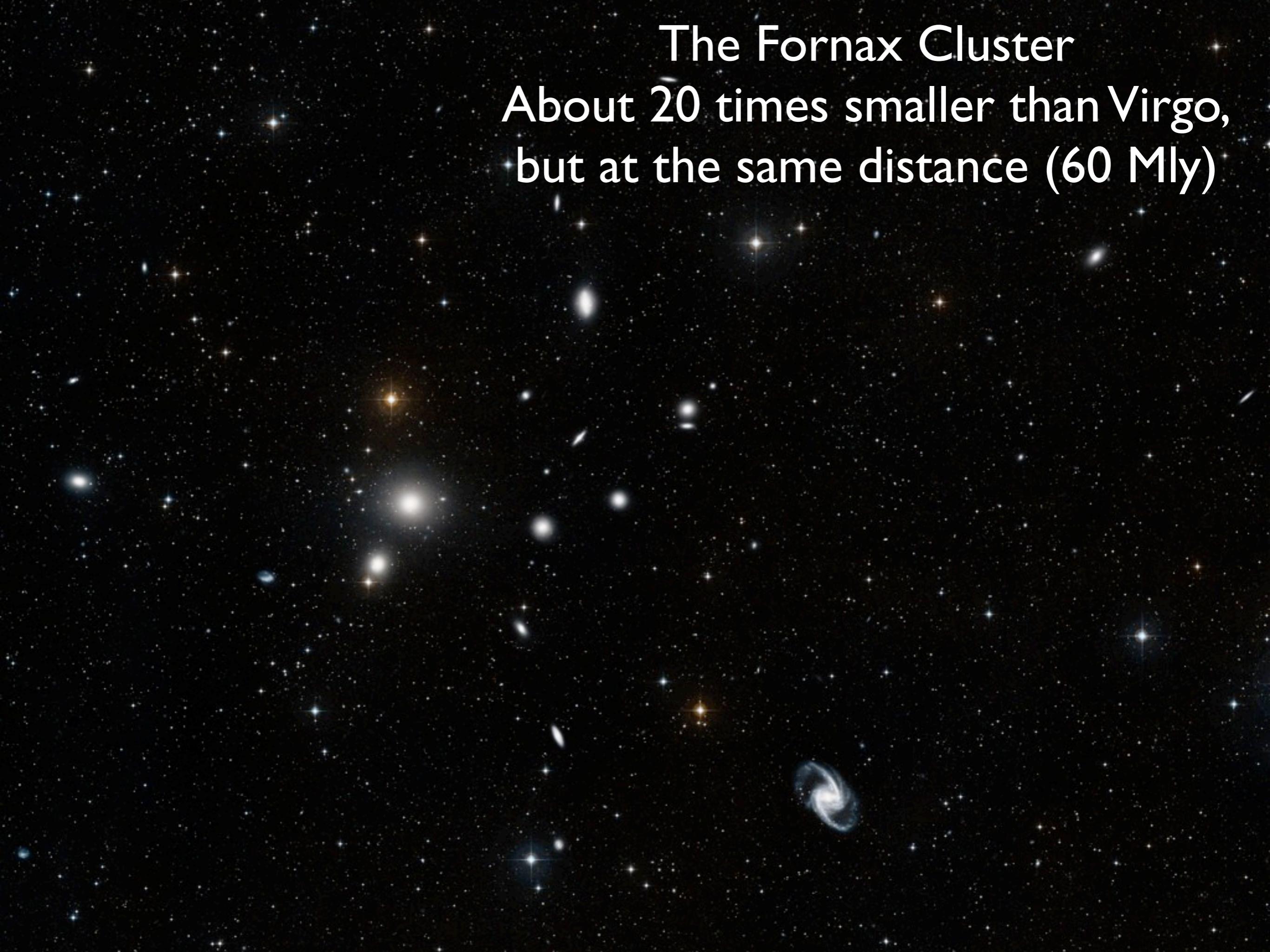
Optical



# Break

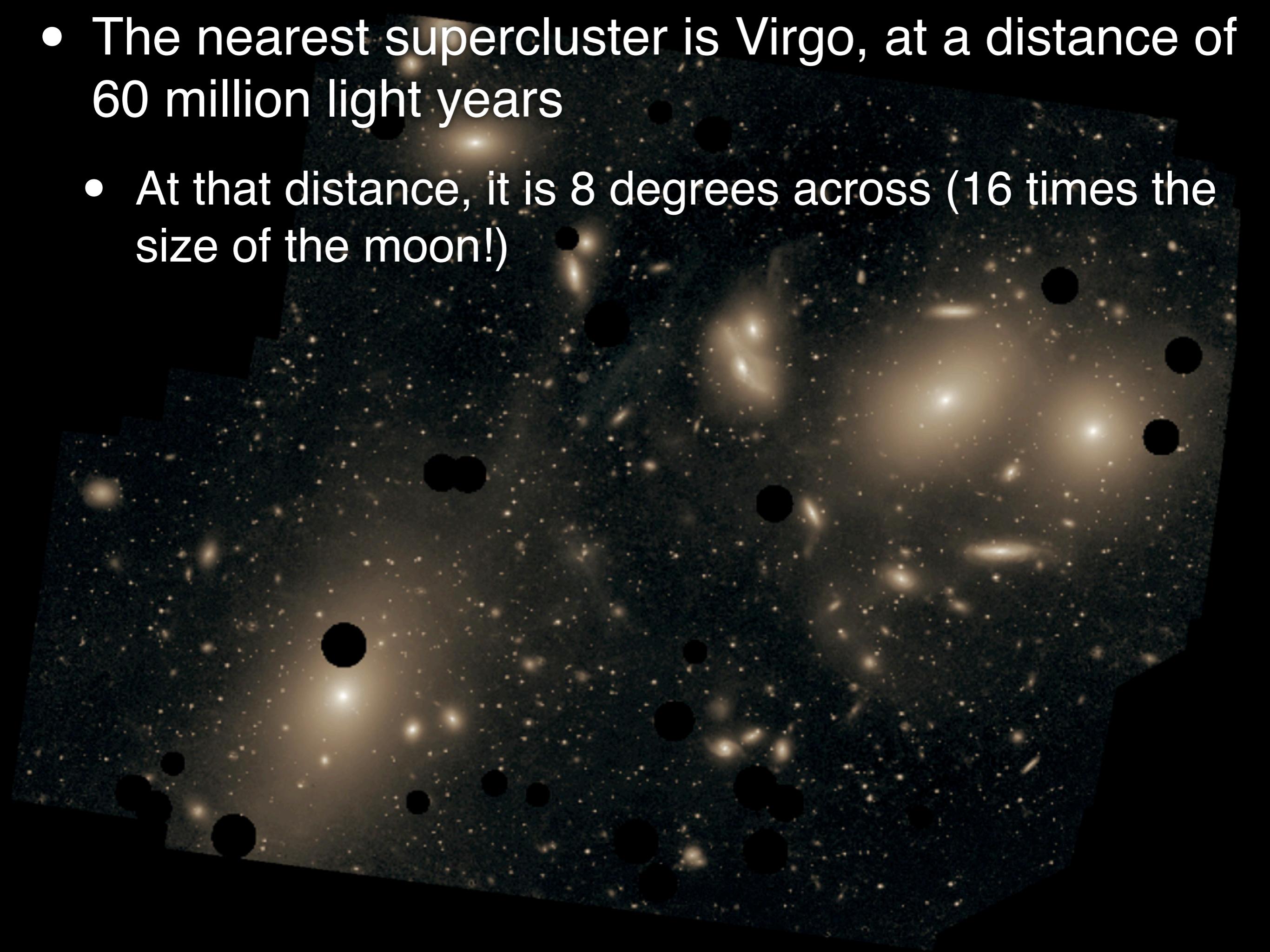
- Come back and do Dark Matter LT



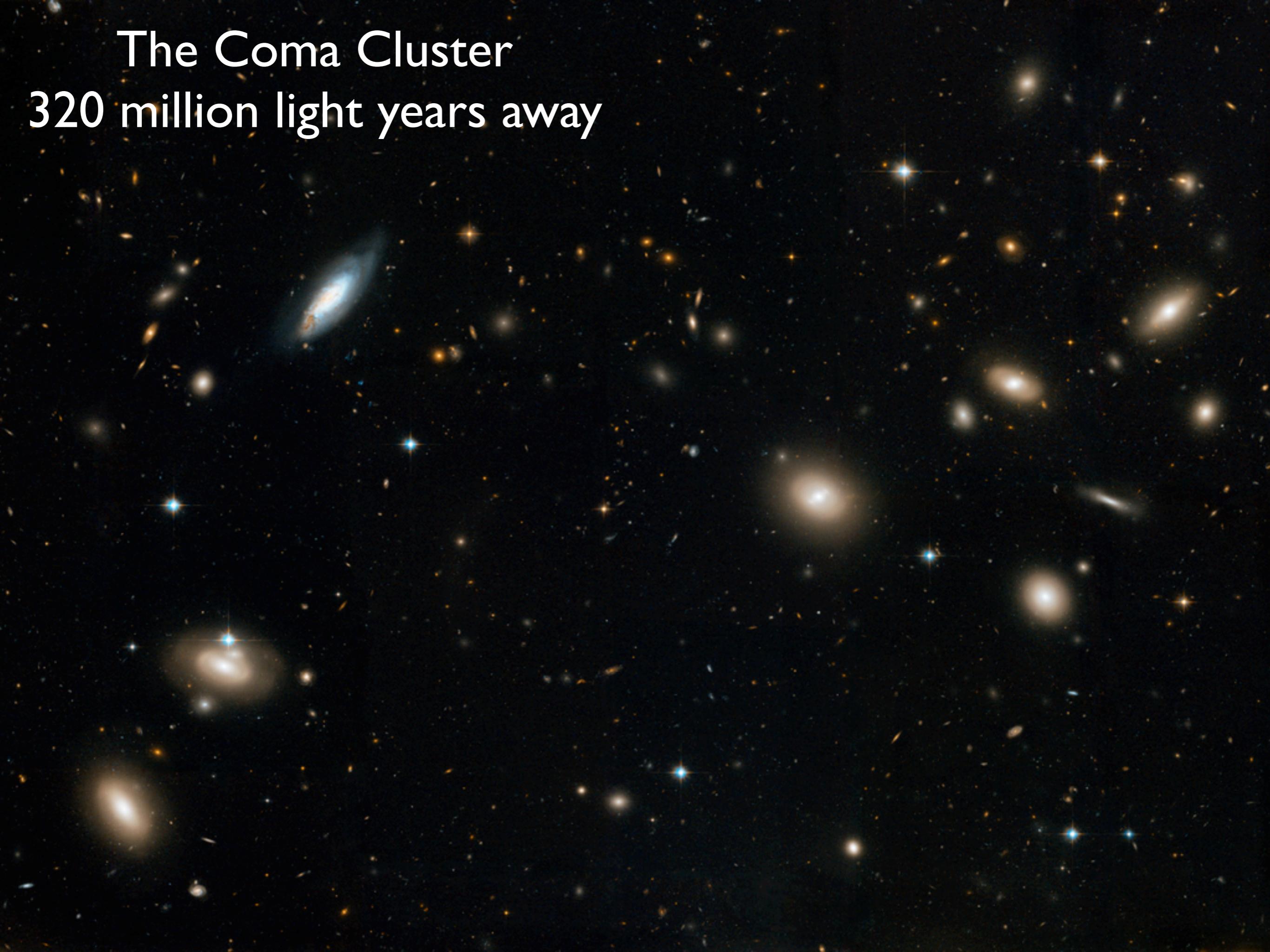


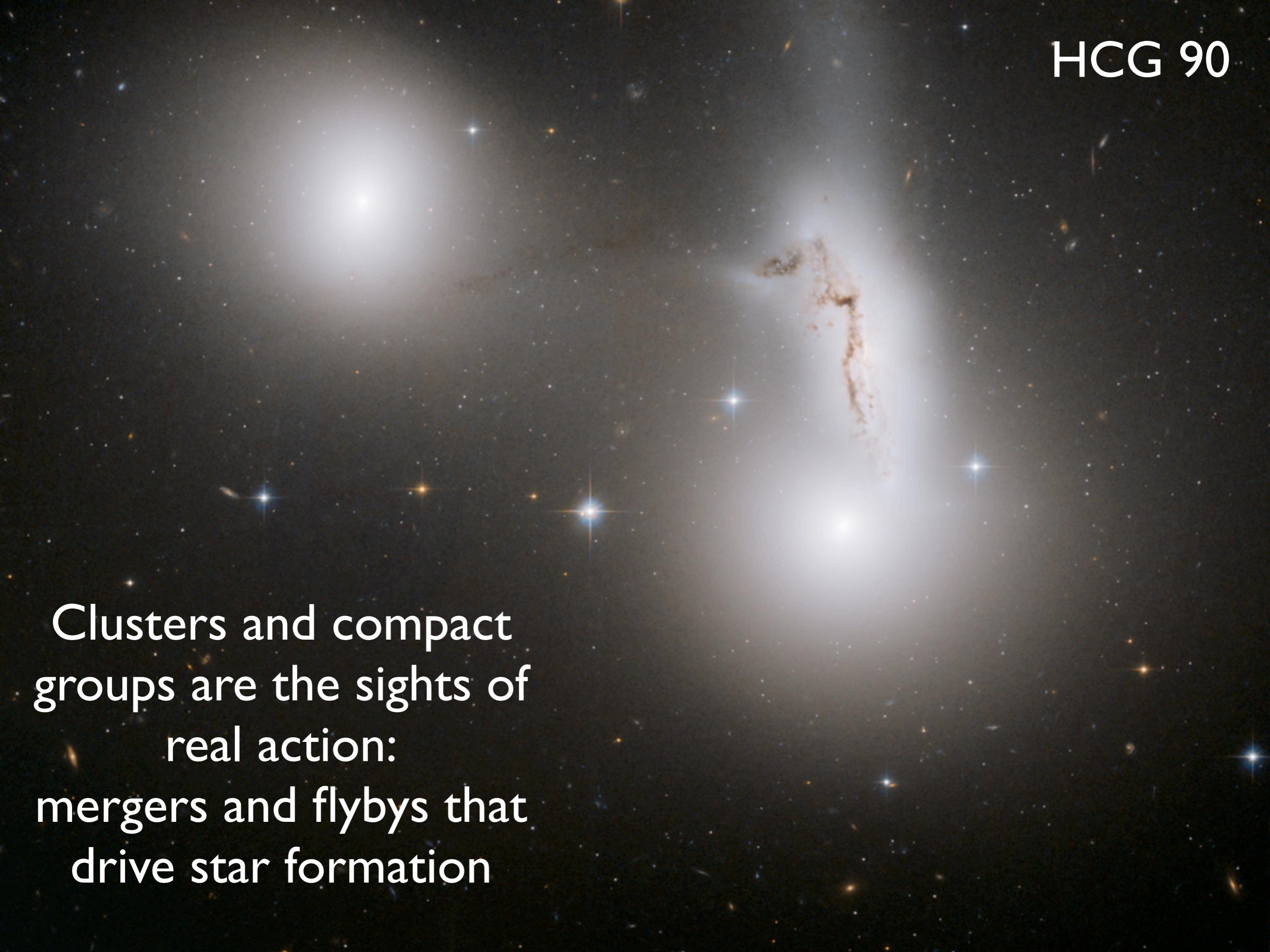
The Fornax Cluster  
About 20 times smaller than Virgo,  
but at the same distance (60 Mly)

- The nearest supercluster is Virgo, at a distance of 60 million light years
  - At that distance, it is 8 degrees across (16 times the size of the moon!)



The Coma Cluster  
320 million light years away





HCG 90

Clusters and compact  
groups are the sights of  
real action:  
mergers and flybys that  
drive star formation



HCG 31

# Stephan's Quintet



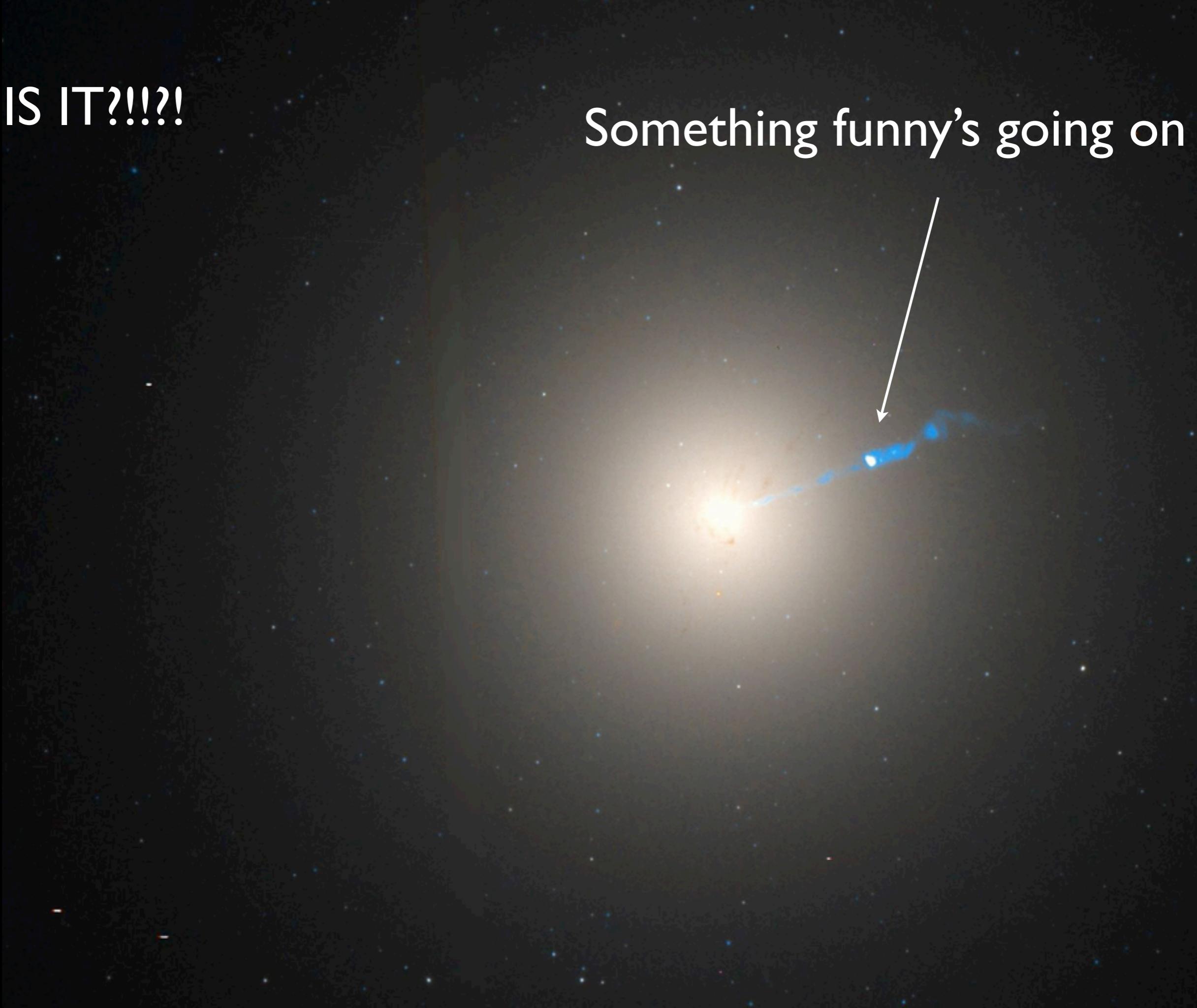
Back to Virgo...



M87, a boring elliptical...

OR IS IT?!!?!

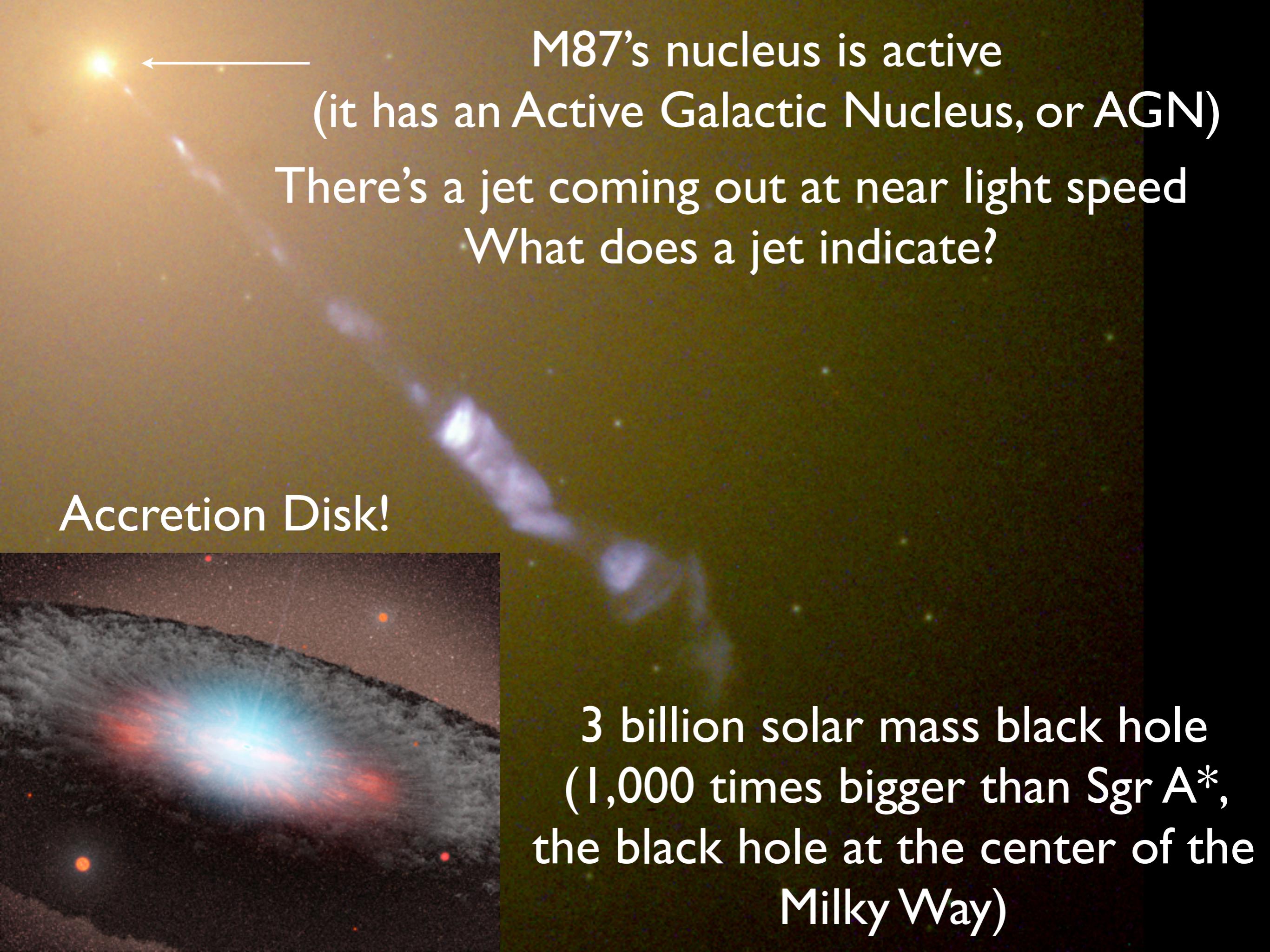
Something funny's going on





M87's nucleus is active  
(it has an Active Galactic Nucleus, or AGN)

There's a jet coming out at near light speed  
What does a jet indicate?



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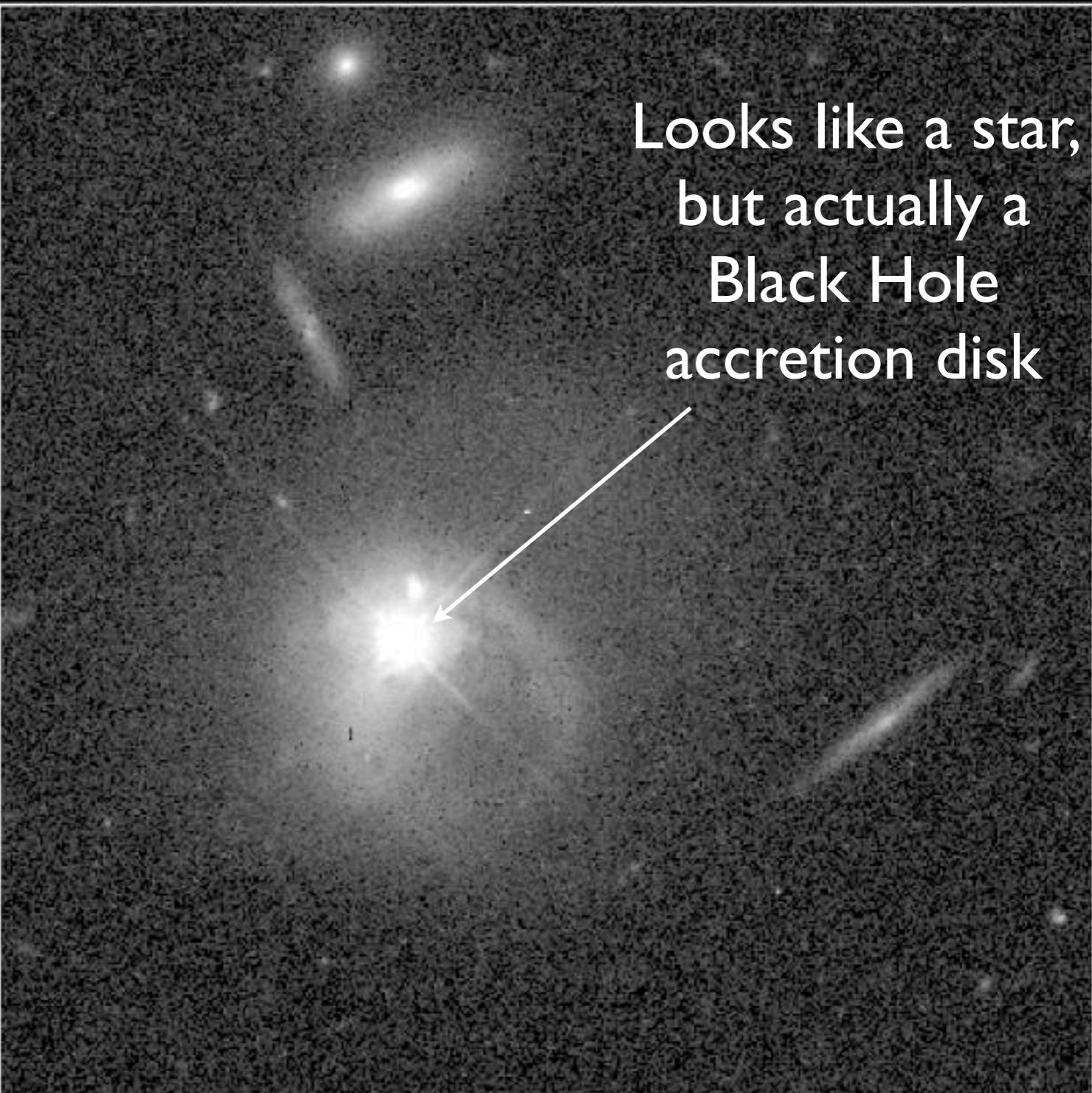
Accretion Disk!



3 billion solar mass black hole  
(1,000 times bigger than Sgr A\*,  
the black hole at the center of the  
Milky Way)

In the early universe,  
AGN were much  
brighter.

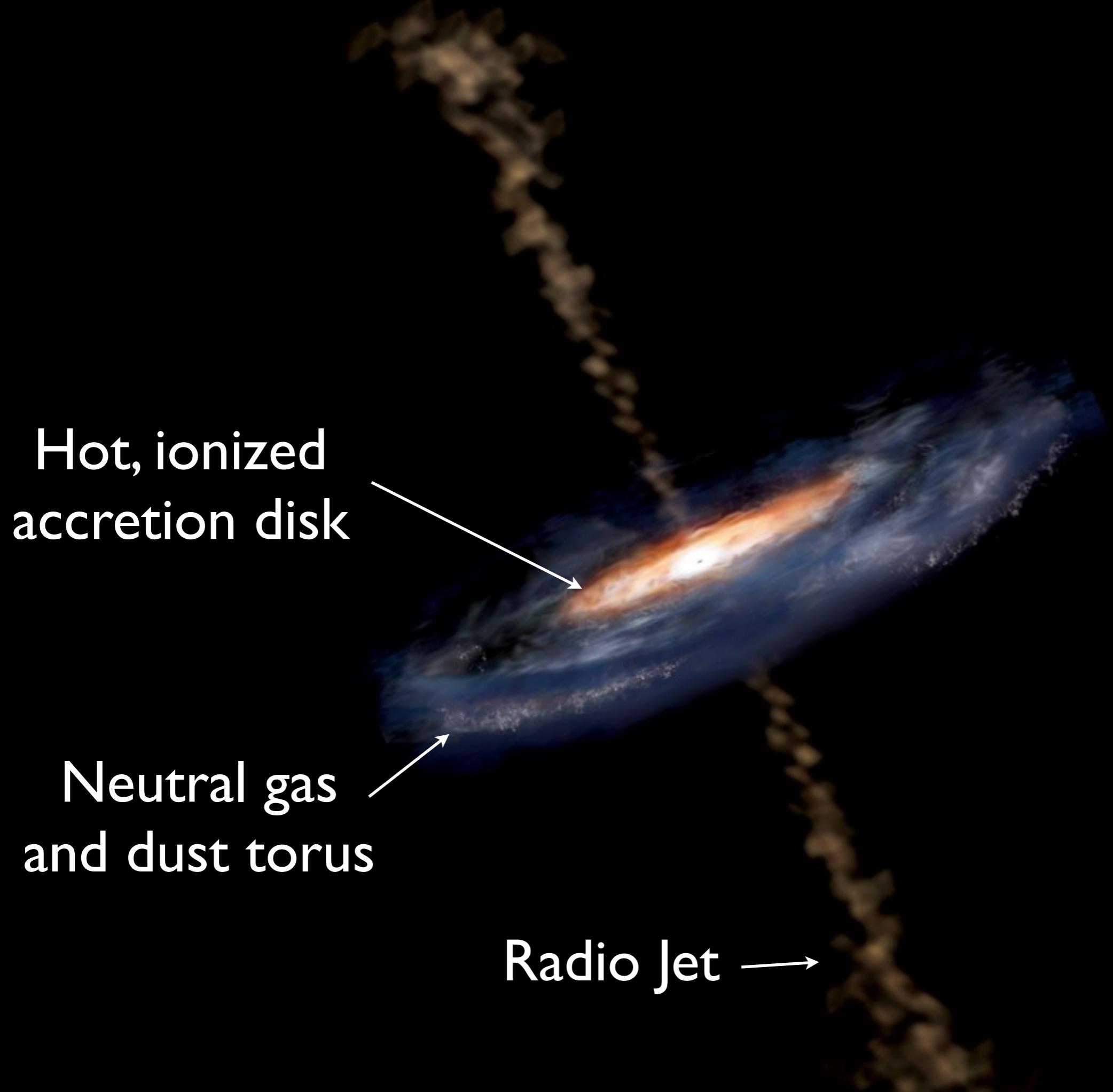
They are called  
“Quasars” because  
they are QUAsi-  
StellAr Radio  
Sources



**Quasar PKS 2349**

ST Scl OPO · January 1995 · J. Bahcall (Princeton), NASA

HST · WFPC2



Hot, ionized  
accretion disk

Neutral gas  
and dust torus

Radio Jet →

Looking down  
the jet:  
Blazar

Hot, ionized  
accretion disk

Neutral gas  
and dust torus

Radio Jet →



Looking down  
the jet:  
Blazar

From the side:  
Seyfert I

Hot, ionized  
accretion disk

Neutral gas  
and dust torus

Radio Jet →

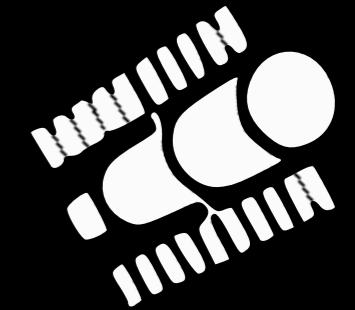
Looking down  
the jet:  
Blazar

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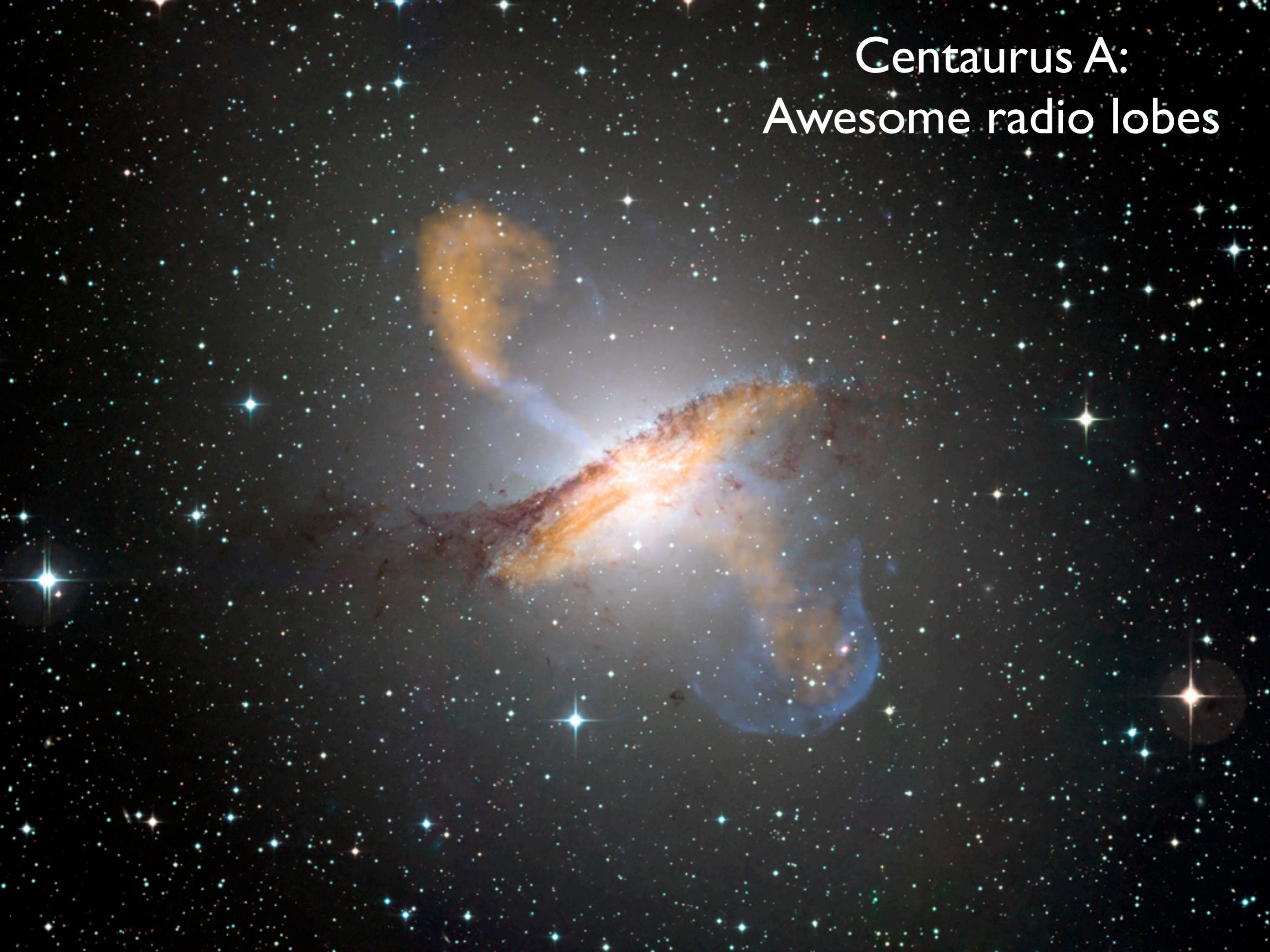
Neutral gas  
and dust torus

Radio Jet →

From the side:  
Seyfert I



Edge on:  
“Radio Galaxy”  
or  
Seyfert 2



Centaurus A:  
Awesome radio lobes

# Summary

- If it's dark, how do we know it's there?
  - Rotation curves, X-ray clusters, gravitational lensing
- What do you mean, dark?
  - Doesn't interact with normal matter; doesn't emit light
- What are galaxy clusters?
  - What do we use galaxy clusters for?
  - Just what they sound like! But they also have a boatload of hot gas in them, and they're convenient places to find dark matter
- And maybe we talked about Active Galactic Nuclei, the big black holes at the centers of galaxies