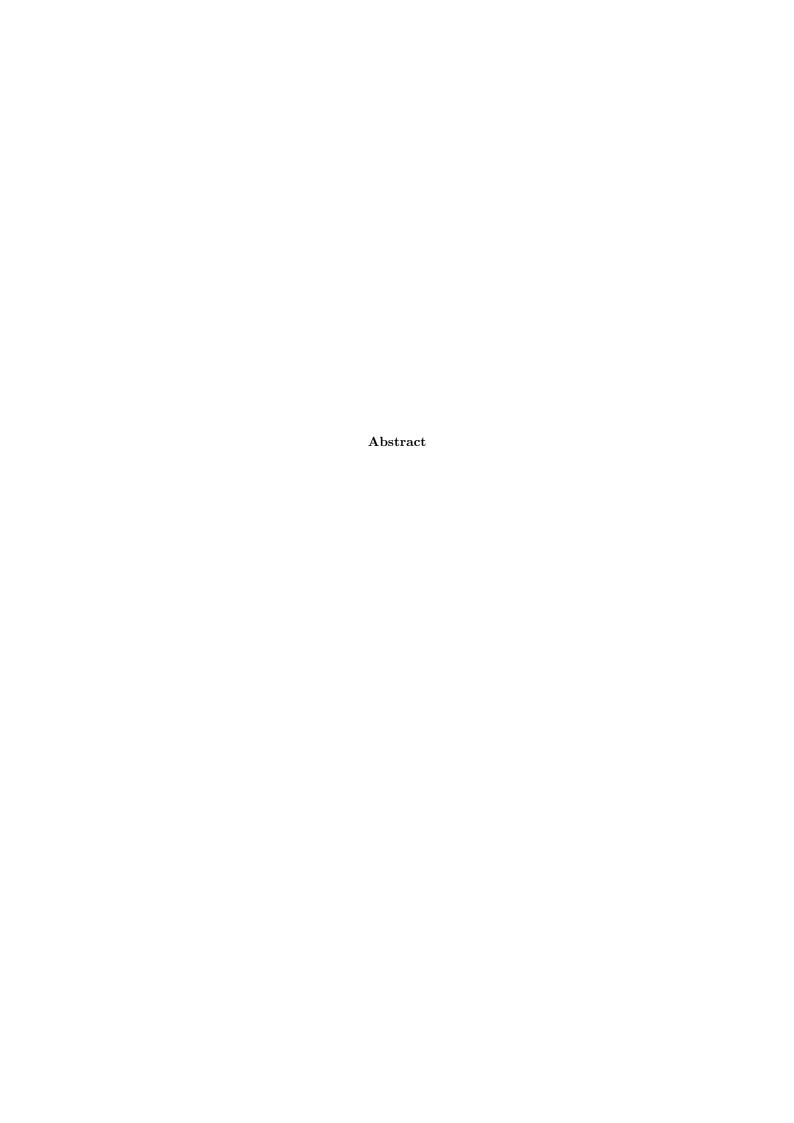
## Mitschrift KOMA



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## Chapter 1

## Charge Order

### 1.1 Peierl Transition

## 1.2 From Causality to Kramer-Kronig relation

Looking at a causal function  $\tilde{\chi}(t)$ , we can split it, like every analytical function, in an even  $\chi_{even}(t)$  and an odd  $\chi_{odd}(t)$  part.

$$\tilde{\chi}(t) = \begin{cases} 0 & t < t_0 \\ \chi(t) & t > t_0 \end{cases} \qquad \text{University of Zurich}$$

$$\tilde{\chi}(t) = \frac{\tilde{\chi}(t) + \tilde{\chi}(-t)}{2} + \frac{\tilde{\chi}(t) - \tilde{\chi}(-t)}{2} = \chi_{even}(t) + \chi_{odd}(t) \qquad (1.1)$$

Multiplying the even part of this function with the signum function yields,

$$\operatorname{sign}(t) \cdot \chi_{even} = \operatorname{sign}(t) \cdot \left\{ \frac{\tilde{\chi}(t)}{2} + \frac{\tilde{\chi}(-t)}{2} \right\} = \frac{\tilde{\chi}(t)}{2} - \frac{\tilde{\chi}(t)}{2} = \chi_{odd}(t) \quad (1.2)$$

Using this relation to replace  $\chi_{odd}(t)$  in Eq 1.1.

$$\tilde{\chi}(t) = \chi_{eve} + \chi_{odd} = (1 + \text{sign}(t)) \cdot \chi_{even}(t) = \sigma(t) \cdot \chi_{even}(t)$$
 (1.3)

### **Onsager Relation**

$$S \cdot \left(\frac{1}{B_{n+1}} - \frac{1}{B_n}\right) = \frac{2\pi e}{\hbar} \tag{1.4}$$

S is the Fermi surface (  $S=\pi k_F^2.$ 

$$F \equiv \left(\frac{1}{B_{n+1}} - \frac{1}{B_n}\right)^{-1} \tag{1.5}$$

$$F = \frac{\Phi_0}{2\pi^2} \cdot S \tag{1.6}$$

where  $\Phi_0 = \frac{h}{2e} = \text{Flux Quantum}$ 

## Chapter 2

## Second Quantization

### 2.1 1D-Chain

To have a fitst insight into second quantization, we look at a phonons emerging in a one dimensional chain of atoms

#### skecht 1D lattice chain

In the sketch above  $u_i$  referes to the distance from equilibrium of the *i*-th atom,  $p_i$  stands for the momentum of the *i*-th atom, K is the spring constant and the lattice constant a. The Hamiltonian of this system is given as

$$H_{\rm ph} = \sum_{j_1}^{N} \left\{ \frac{1}{2M} p_j^2 + \frac{1}{2} K (u_j - u_{j-1})^2 \right\}$$
 (2.1)

Since our problem has one dimension in real space it has also one dimension in the reciprocal space. We can introduce two length scales in the reciprocal space

- 1. The size of the one dim. unit cell defines the length of the reciprocal lattice vector  $\vec{G}$  which is equal to the size of the Brillouin zone:  $G = 2\pi/a$
- 2. The length L of the chain of atoms in real space defines the lowest possible distance two points in reciprocal space can have  $\Delta k_{\rm min} = 2\pi/L$ .

#### sketch 1D k-space

We now want to rewritte the Hamiltonian Eq 2.1 by replacing the distance from equilibrium and the momentum of the atoms by its fourer transformations

$$p_j = \frac{1}{\sqrt{N}} \sum_{k \in BZ} P_k e^{ikR_j^0} \qquad \Longleftrightarrow \qquad P_k = \frac{1}{\sqrt{N}} \sum_j p_j e^{-ikR_j^0}$$
 (2.2)

$$u_j = \frac{1}{\sqrt{N}} \sum_{k \in BZ} U_k e^{ikR_j^0} \qquad \Longleftrightarrow \qquad U_k = \frac{1}{\sqrt{N}} \sum_j u_j e^{-ikR_j^0}$$
 (2.3)

Since it is a cumbersome procedure we look at the calculate the different terms from Eq 2.1 separatly. We start with  $\sum_i p_j^2$ :

$$p_{j} = \frac{1}{\sqrt{N}} \sum_{k} P_{k} e^{ikR_{j}^{0}} = \frac{1}{2\sqrt{N}} \left\{ \underbrace{\sum_{k} P_{k} e^{ikR_{j}^{0}}}_{A} + \underbrace{\sum_{-k} P_{-k} e^{-ikR_{j}^{0}}}_{B} \right\}$$
(2.4)

The second equality sign holds because summing two times and sum and dividing by two gives the same result. Additionally in the second sum we replaced  $k \to -k$ . No we can write

$$p_j^2 = \frac{1}{4N} (A^2 + B^2 + 2AB) = \frac{2}{4N} \sum_j AB$$

$$= \frac{1}{2N} \sum_j \sum_k \sum_{-k} P_k P_{-k} = \sum_k P_k P_{-k}$$
 (2.5)

#### is this correct?

To get to this result  $\sum_j = e^{ikR_j^0} = \delta(k)$  was used. For evaluating now  $\sum_j (u_j - u_{j-1})^2$  we use

$$u_{j} = \left\{ \underbrace{\sum_{k} u_{k} e^{ikR_{j}^{0}}}_{\mathbf{A}^{'}} + \underbrace{\sum_{-k} u_{-k} e^{-ikR_{j}^{0}}}_{\mathbf{B}^{'}} \right\}$$
 (2.6)

$$u_{j-1} = \frac{1}{\sqrt{N}} \sum_{k} u_{k} e^{ikR_{j}^{0}} e^{-ika}$$

$$= \frac{1}{2\sqrt{N}} \left\{ \sum_{k} u_{k} e^{ikR_{j}^{0}} e^{-ika} + \sum_{-k} u_{-k} e^{-ikR_{j}^{0}} e^{ika} \right\}$$
(2.7)

To get the final expression for  $u_{j-1}$  we used  $R_{j-1}^0 = r_j^0 - a$ .

$$\sum_{j} (u_{j} - u_{j-i})^{2} = \sum_{j} (u_{j}^{2} + u_{j-1}^{2} + 2u_{j}u_{j-1})$$

$$= \sum_{k} \left\{ u_{k}u_{-k} + u_{k}u_{-k} - 2u_{k}u_{-k} \frac{e^{-ika} + e^{ika}}{2} \right\}$$

$$= \sum_{k} 2u_{k}u_{-k} \left\{ 1 - \cos(ka) \right\}$$

$$= \sum_{k} 2u_{k}u_{-k} \sin^{2}(ka/2)$$
(2.8)

In the step were the summing over j changed to a summing over k the same arguments as for Eq 2.5 were used. As we already know from the *introduction to solid state physics* class, we can assign the argument of the sum in the final expression as the square of the phonon dispersion relation  $\omega(k)$ 

$$\omega(k) = \omega_k = \sqrt{\frac{2K}{M}(1 - \cos(ka))} = \sqrt{\frac{4K}{M}}|\sin(ka/2)| \qquad (2.9)$$

Pluggin our solutions for  $\sum_j p_j^2$  and  $\sum_j (u_j - u_{j-1})^2$  into Eq 2.1 yields

$$H_{\rm ph} = \sum_{j=1}^{N} \left\{ \frac{1}{2M} P_k P_{-k} + \frac{M\omega_k^2}{2} u_k u_{-k} \right\}$$
 (2.10)

This expression similar to the Hamilatonian from the harmonic oscillator. We therefore define the latter operators  $b_k$  and  $b_k^{\dagger}$ 

$$b_k = \frac{1}{\sqrt{2}} \left( \frac{u_k}{l_k} + i \frac{p_k}{\hbar/l_k} \right) \tag{2.11}$$

$$b_k^{\dagger} = \frac{1}{\sqrt{2}} \left( \frac{u_k}{l_k} - i \frac{p_k}{\hbar/l_k} \right) \tag{2.12}$$

and rewrite  $H_{\rm ph}$  as

$$H_{\rm ph} = \sum_{k} \hbar \omega_k (b_k^{\dagger} b_k + 1) \tag{2.13}$$

This analogy between the Hamiltonian from the harmonic oscillations and Eq 2.10 emphasises that we can describe the state if excited phonons in a solid as a combination of different modes similar to the different modes of oscillations of an excited harmonic oscillator. Further we can use the latter operators applied to the vacuum state to describe the excited states mathematically.

picture single atom phonon dispersion relation picture two atom basis phonon dispersion relation ...if going down in temperature peierl transition occurs -; kohn anomaly

why is neutron scattering good for probing accustic bands? - check lecture notes for explanation

#### ...neutrons only probe lattice and neglect electronic structure

This way of decribing the physical processes in terms of latter operators can be extended much more. It turns out that it is a very convenient way of describing physical problems and is used broadly in all fields in condensed matter physics. We refere to this notation as **Second quantization**. In the next sections we will learn how to apply this formalism more generally.

## 2.2 Second quantization: Free electron gas

Hamiltonian of a free electron gas can be written as.

$$H = \sum_{k\sigma} \frac{(\hbar \hat{k})^2}{2m} c_{k\sigma}^{\dagger} c_{k\sigma} \tag{2.14}$$

The ground state at T = 0K is create by applying the electron-creation operators  $c_{k_i\uparrow}$  of the i – th electron  $e_{i\uparrow}$  to the vacuum state  $|0\rangle$  for every electron that is present in the electron gas.

$$|FS\rangle = c_{k_{N/2}\uparrow}c_{k_{N/2}\downarrow}...c_{k_1\uparrow}c_{k_1\downarrow}|0\rangle =$$
 (2.15)

To get the total number of electrons N the number operator  $\hat{N} = \sum_{k\sigma} c_{k\sigma}^{\dagger} c_{k\sigma}$  can be applied on the ground state to the ground state  $|FS\rangle$ 

$$N = \langle FS|\hat{N}|FS\rangle = \langle FS|\sum_{k\sigma} c_{k\sigma}^{\dagger} c_{k\sigma}|FS\rangle \qquad (2.16)$$

For the ground state at T=0 the

Since we consider a free electron gas in its ground state at T=0K the momentum values k are limited to lay inside a sphere with radius  $k_F$ . Furthermore we have to take into account that in the system under consideration the orientation of the spin  $\sigma$  doesn't matter. Therefore we can replace the sum  $\sum_{\sigma}$  by a factor two.

$$N = \sum_{\mathbf{k}\sigma} \langle \mathrm{FS} | c_{k\sigma}^{\dagger} c_{k\sigma} | \mathrm{FS} \rangle = 2 \sum_{|\mathbf{k}| \le k_F} \langle \mathrm{FS} | \mathrm{FS} \rangle = 2V \int \frac{d^3k}{(2\pi)^3} \theta(k_F - |\mathbf{k}|) \langle \mathrm{FS} | \mathrm{FS} \rangle$$
$$= \frac{2V}{(2\pi)^3} \int_{|\mathbf{k}| \le k_F} d^3k$$
(2.17)

Solving this integral in spherical coordinates gives

$$N = \frac{2V}{(2\pi)^3} \int_0^{k_F} dk k^2 \int_{-1}^1 d\cos(\theta) \int_0^{2\pi} d\phi = \frac{1}{3} \frac{V}{\pi^2} k_F^3$$
 (2.18)

From that equation we can determine the electron density n for a three dimensional free electron gas at zero Temperature

$$n = N/V = k_F^3/(3\pi^2) \tag{2.19}$$

As another example of the usage of second quantization we want to determine now the ground state energy  $E_0$  of the free electron gas

$$E_{0} = \langle FS|H|FS \rangle = \sum_{\mathbf{k}\sigma} \frac{\hbar \mathbf{k})^{2}}{2m} \langle FS|c_{\mathbf{k}\sigma}^{\dagger}c_{\mathbf{k}\sigma}|FS \rangle = \frac{\hbar^{2}}{2m} \sum_{\mathbf{k}\sigma} k^{2}\theta(|\mathbf{k}| - k_{F})$$

$$= 2\frac{4\pi V}{(2\pi)^{3}} \frac{\hbar^{2}}{2m} \int dk\theta(|\mathbf{k}| - k_{F})k^{4}$$

$$= \frac{4\pi V}{(2\pi)^{3}} \frac{\hbar}{m} \frac{k_{F}^{5}}{5} = \frac{3}{5}NE_{F}$$
(2.20)

To get to the final expression we used the electron density from Eq 2.19 and the Fermi Energy  $E_F = (\hbar k_F)^2/(2m)$ .

#### **Electron-Electron Interaction term**

The electron-electron interaction term in a material is given as

$$V_{\text{el-el}} = \frac{1}{2V} \sum_{\sigma_1, \sigma_2} \sum_{k_1, k_2, q} V_q c_{k_1, \sigma_1}^{\dagger} c_{k_2, \sigma_2}^{\dagger} c_{k_1 + q, \sigma_1}, c_{k_2 - q, \sigma_2}$$
 (2.21)

To understand this expression better we look first at one single interaction between two electrons with initial momentum and spin  $k_1, \sigma_1$  and  $k_2, \sigma_2$ . The spin doesn't change. When colliding the exchanged momentum due to the interaction term  $V_q$ , were we assume  $q \neq 0$ . The electron momenta in the final state are related by  $k_1 + q = k_2$ .

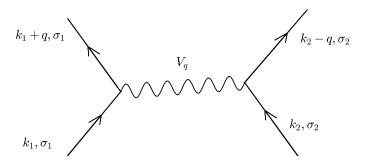


Figure 2.1: illustration of an e-e Interaction were the incoming electrons have initial momentum  $k_1$ ,  $k_2$  and spin  $\sigma_1$ ,  $\sigma_2$ . The interaction  $V_q$  causes them to exchange a momentum q.

The interaction is due to the coulomb repulsion, therefore we can write the interaction term in reciprocal space as

$$V_q = \int d^3r e^{i\mathbf{q}\cdot\mathbf{r}} V(\mathbf{r}) = \int d^3r \frac{e^2}{r} e^{i\mathbf{q}\cdot\mathbf{r}} = \frac{4\pi e^2}{q^2}$$
 (2.22)

Calculating the change in energy resulting from the coulomb interaction between the electrons gives

$$\frac{E_1}{N} = \frac{\langle FS|V_{\text{el-el}}|FS\rangle}{N} = \frac{1}{2VN} \sum_{q} \sum_{\mathbf{k}_1, \mathbf{k}_2} \sum_{\sigma_1, \sigma_2} \frac{4\pi^2}{q^2} \langle FS|c_{\mathbf{k}_1 + q}^{\dagger} c_{\mathbf{k}_2 - q}^{\dagger} c_{\mathbf{k}_2} c_{\mathbf{k}_1} |FS\rangle \quad (2.23)$$

$$\langle \mathrm{FS} | c_{\mathbf{k}_{1}+q,\sigma_{1}}^{\dagger} c_{\mathbf{k}_{2}-q,\sigma_{2}}^{\dagger} c_{\mathbf{k}_{2},\sigma_{2}} c_{\mathbf{k}_{1},\sigma_{2}} | \mathrm{FS} \rangle = \delta_{\sigma_{1}\sigma_{2}} \delta_{\mathbf{k}_{1}+q,\mathbf{k}_{2}} \langle \mathrm{FS} | c_{\mathbf{k}_{1}+q}^{\dagger} c_{\mathbf{k}_{2}-q}^{\dagger} c_{\mathbf{k}_{2}} c_{\mathbf{k}_{1}} | \mathrm{FS} \rangle$$

$$= -\delta_{\sigma_{1}\sigma_{2}} \delta_{\mathbf{k}_{1}+q,\mathbf{k}_{2}} \langle \mathrm{FS} | c_{\mathbf{k}_{1}+q}^{\dagger} c_{\mathbf{k}_{1}+q} c_{\mathbf{k}_{1}+q}^{\dagger} c_{\mathbf{k}_{1}} | \mathrm{FS} \rangle$$

$$= -\delta_{\sigma_{1}\sigma_{2}} \delta_{\mathbf{k}_{1}+q,\mathbf{k}_{2}} \theta(k_{F} - |\mathbf{k}_{1}+q|) \theta(k_{F} - |\mathbf{k}_{1}|)$$

$$(2.24)$$

The two delta functions impose the fact that the spin does not change during the interaction and that the momentum is conserved in the process. Furthermore since we can relate  $\mathbf{k}_1$  and  $\mathbf{k}_2$  we replaced the latter one in the latter operator index.

Here we see the benefit which comes along with this notation. By only using commutation relation for fermion-latter operators we were able to change the rathre complicated expression to one that can easily be integrated.

The way to perform this integration can be illustrated graphically Fig.2.2.

How to choose the boundaries of integration can be illustrated.

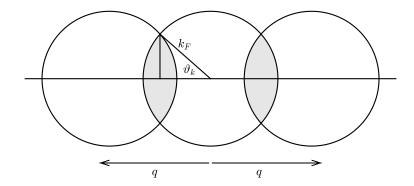


Figure 2.2: illustration of an e-e Interaction were the incoming electrons have initial momentum  $k_1$ ,  $k_2$  and spin  $\sigma_1$ ,  $\sigma_2$ . The interaction  $V_q$  causes them to exchange a momentum q.

#### Second Quantization: Free Electron Gas

$$k_F^3 = 3\pi^2 n$$

$$E_0 = \frac{3}{5}N\epsilon_F$$

Bohr Radius:  $a_0 = \frac{\hbar}{me^2}$ 

$$\frac{3\pi^2}{k_F^3} = \frac{1}{n} = \frac{V}{N} = \frac{4\pi}{3} (r_S a_0)^3 \Rightarrow r_S = \left(\frac{9\pi}{4}\right)^{1/3} \frac{1}{a_0 k_F}$$

$$\frac{E_0}{N} = \frac{2.21}{r_S^2} \frac{e^2}{2a_0}$$

#### **Electron Interaction**

$$\frac{E_1}{N} = \frac{\langle FS | V_{el-el} | FS \rangle}{N} = -\frac{e^2}{2} \frac{V}{N} \frac{k_F^4}{2\pi^3} = -\frac{0.916}{r_S} \frac{e^2}{2a_0}$$

plot - energy minimum at  $r_S$ 

$$V_{\text{el-el}} = \frac{1}{2V} \sum_{\sigma_1 \sigma_2} \sum_{k_1 k_2 q} V_q c_{k_1 + q}^{\dagger} c_{k_2 - q}^{\dagger} c_{k_2} c_{k_1}$$
 (2.25)

#### Fermi Liquid Theory

#### **Electron Tight Binding Model**

An approximation for an electron wavefunction over the whole crystal can be given as

$$\Psi_k = \sum_j c_{kj} \varphi(\mathbf{r} - \mathbf{r}_j) \tag{2.26}$$

Where  $\varphi(\mathbf{r})$  is the ground state wavefunction (atomic orbital) of an electron in the potential  $V(\mathbf{r})$  of an isolated atom. This approximation is only valid if the interaction between neighborign atoms is small. The sum goes over all lattice points.

Using

$$c_{kj} = N^{-1/2}e^{i\mathbf{k}\mathbf{r}_j} \tag{2.27}$$

The factor  $N^{-1/2}$  ensures the wavefunction to meet the normalization criterion We can express the over all wavefunction of an electron in the crystal

$$\Psi_k(\mathbf{r}) = N^{-1/2} \sum_j e^{i\mathbf{k}\mathbf{r}_j} \varphi(\mathbf{r} - \mathbf{r}_j)$$
 (2.28)

Recalling the **Bloch Theorem** which states, that the wavefunction of an electron in a periodic potential  $(V(\mathbf{r}) = V(\mathbf{r} + \mathbf{a}))$  with the periodicity a is given by

$$\phi_k(\mathbf{r}) = u_k(\mathbf{r}) \cdot e^{i\mathbf{k}\mathbf{r}} \tag{2.29}$$

where  $u_k(\mathbf{r})$  has the periodicity of the crystal  $(u_k(\mathbf{r} + \mathbf{a}) = u_k(\mathbf{r}))$ . Since the tightbinding ansatz is meant to describe the electrons in a periodic potential, so it also have to be a bloch wave.

We can show, that our electron wavefunction Eq 2.28 fulfills the Bloch criterion

$$\Psi_{k}(\mathbf{r} + \mathbf{a}) = N^{-1/2} \sum_{j} e^{i\mathbf{k}\mathbf{r}_{j}} \varphi(\mathbf{r} + \mathbf{a} - \mathbf{r}_{j})$$

$$= N^{-1/2} e^{i\mathbf{k}\mathbf{a}} \sum_{j} e^{i\mathbf{k}(\mathbf{r}_{j} - \mathbf{a})} \varphi(\mathbf{r} - (\mathbf{r}_{j} - \mathbf{a}))$$

$$= N^{-1/2} e^{i\mathbf{k}\mathbf{a}} \sum_{n} e^{i\mathbf{k}\mathbf{r}_{n}} \varphi(\mathbf{r} - \mathbf{r}_{n})$$

$$= e^{i\mathbf{k}\mathbf{a}} \Psi_{k}(\mathbf{r}) \tag{2.30}$$

To get to this result we used the fact, that the summation goes over all lattice points and we therefore can replace the dummy variables j with n by writting  $\mathbf{r}_j - \mathbf{a}$  as  $\mathbf{r}_n$ . This is allowed since  $\mathbf{a}$  describes a lattice vector connecting two lattice points.

We can compute now for the energy levels of the tight-binding ansatz

$$\epsilon_k = \langle \mathbf{k} | H | \mathbf{k} \rangle = N^{-1} \sum_j \sum_m e^{i\mathbf{k}(\mathbf{r}_j - \mathbf{r}_m)} \langle \varphi_m | H | \varphi_j \rangle$$
 (2.31)

The double sum can be understood in the following sense: to get the total energy, we have to sum over all the electrons  $(\sum_j)$ , whereas every electron feels a contribution from the potential of every atom in the crystal  $(\sum_m)$ . Since we assum, that the electrons on an atom site mainly feel interactions on the atom site itself and the nearest neighbors, we can neglect all terms with  $j \neq m$  of the first sum. We also define  $\rho_m = \mathbf{r}_m - \mathbf{r}_j$ 

$$\epsilon_{k} = \sum_{m} e^{-i\mathbf{k}\rho_{m}} \langle \varphi(\mathbf{r} - \rho_{m}) | H | \varphi(\mathbf{r}) \rangle$$

$$= \langle \varphi(\mathbf{r}) | H | \varphi(\mathbf{r}) \rangle + \sum_{\rho_{m} \in \text{NN}} \langle \varphi(\mathbf{r} - \rho_{m}) | H | \varphi(\mathbf{r}) \rangle$$
(2.32)

Defining the coefficients

$$\epsilon_0 = \langle \varphi(\mathbf{r}) | H | \varphi(\mathbf{r}) \rangle = \int d^3 \mathbf{r} \ \varphi^*(\mathbf{r}) H \varphi(\mathbf{r})$$
 (2.33)

$$t = \langle \varphi(\mathbf{r} - \rho_{\mathbf{m}}) | H | \varphi(\mathbf{r}) \rangle = \int d^3 \mathbf{r} \ \varphi^*(\mathbf{r} - \rho_{\mathbf{m}}) H \varphi(\mathbf{r})$$
 (2.34)

allows us to write Eq 2.32 in the abbreviated way

$$\epsilon_k = \epsilon_0 - t \sum_{\rho_m} e^{-i\mathbf{k}\rho_m} \tag{2.35}$$

where  $\epsilon_0$ , the energy level of the singel atom, defines the position of the energy band and t, defined through the overlap of neighboring atoms, defines the width of the energy band.

#### Example: cubic lattice structure

For a cubic lattice structure, taking only the nearest neighbors into account we have to sum over the following lattice sites in the sum of Eq 2.35

$$\rho_m \in \{(\pm a, 0, 0), (0, \pm a, 0), (0, 0, \pm a)\}$$
(2.36)

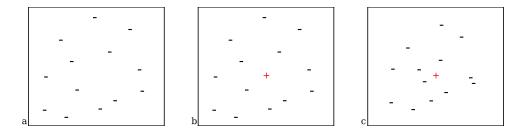


Figure 2.3: Sketch of boxes filled with charged particles. a. equally distributed negatively charged particles. b. negative charges with one positive charge without interactions. c. negative charged particles are attracted by positive charge.

This leads to the following expression for the energy bands

$$\epsilon_k = E(\mathbf{k}) = \epsilon_0 + 2t \left[\cos(k_x a) + \cos(k_y a) + \cos(k_z a)\right] \tag{2.37}$$

The position  $(\epsilon_0)$  and width W = 12t is determined by the particular single atom orbital wavefunction.

#### 2.2.1 **Charge Ordering**

$$\rho(r) = \rho^{\text{Total}}(\mathbf{r}) = \rho^{\text{Ext}}(\mathbf{r}) + \rho^{\text{Ind}}(\mathbf{r})$$
(2.38)

With the poisson equation we can determine the electric potential  $\phi$  produced by the charge density

$$\nabla^2 \phi^{\text{Ext}} = -\frac{1}{\epsilon_0} \rho^{\text{Ext}}(\mathbf{r})$$
 (2.39)

$$\nabla^2 \phi = -\frac{1}{\epsilon_0} \rho(\mathbf{r}) \tag{2.40}$$

This differential equation can be solved by transforming it to the Fourier space

$$q^{2}\phi^{\text{Ext}}(\mathbf{q}) = \frac{1}{\epsilon_{0}}\rho^{\text{Ext}}(\mathbf{q})$$

$$q^{2}\phi(\mathbf{q}) = \frac{1}{\epsilon_{0}}\rho(\mathbf{q})$$
(2.41)

$$q^2 \phi(\mathbf{q}) = \frac{1}{\epsilon_0} \rho(\mathbf{q}) \tag{2.42}$$

Using the definitions for the dielectric function  $\epsilon$  and the electric susceptibility  $\chi$ 

$$\frac{1}{\epsilon(\mathbf{q})} = \frac{\phi(\mathbf{q})}{\phi^{\text{Ext}}(\mathbf{q})} \tag{2.43}$$

$$\chi(\mathbf{q}) = \frac{\rho^{\text{Ind}}(\mathbf{q})}{\phi(\mathbf{q})} \tag{2.44}$$

We can relate the dielectric function and the susceptebility by using the formulas for the electric potential from above

$$\epsilon(\mathbf{q}) = 1 - \frac{\chi(\mathbf{q})}{\epsilon_0 q^2} \tag{2.45}$$

#### Thomas-Fermi Screening

Calculate  $\chi(\mathbf{q})$ 

Charge density of an electron gas can be expressed in terms of the electron density  $n(\mathbf{r})$ 

$$\rho_0 = -en(\mu) \quad \text{with} \quad n(\mu) = \int \frac{d\mathbf{k}}{4\pi^3} f_k(\mu)$$
 (2.46)

The electron density itself is then given as the integral over the Fermi-Dirac distribution. We can write the Schrodinger Equation as

$$-\frac{\hbar^2}{2m} \nabla^2 \Psi_k - e \phi^{\text{Total}} \Psi_k = \varepsilon_k \Psi_k \tag{2.47}$$

Using the Thomas-Fermi approximation in which we assume for the total electric potential to be constance on the length scale of our problem. Then we can write

$$\varepsilon_k = \frac{(\hbar \mathbf{k})^2}{2m} + \varepsilon' \quad \text{where} \quad \varepsilon' = e \, \phi(\mathbf{r})$$
 (2.48)

With this new expression for the electron energy  $\varepsilon_k$  we are able to write an expression for the electron density  $n(\mathbf{r})$  of the whole system

$$n^{\text{Total}} = \int \frac{d\mathbf{k}}{4\pi^3} \left( e^{(E_k + \mu)/k_B T} + 1 \right)^{-1}$$

$$= \int \frac{d\mathbf{k}}{4\pi^3} \left( e^{(\frac{(\hbar \mathbf{k})^2}{2m} + \varepsilon' + \mu)/k_B T} + 1 \right)^{-1} = n(\mu + \varepsilon')$$
(2.49)

As we know from Eq 2.38 the induced charge density is given as the difference between the total charge density  $n(\mathbf{r})$  and the external charge density  $n_0$ . Therefore we can write

$$\rho^{\text{Ind}}(\mathbf{r}) = -e \left[ n(\mu + \varepsilon') - n(\mu) \right] = -e \cdot \varepsilon' \frac{dn}{d\mu}$$
 (2.50)

The difference of the two electron density was replaced by de derivative by using the defintion of derivatives  $df/dx = (f(x + \Delta x) - f(x))/\Delta x$ . Eq 2.50 is known as the basic equation of the Thomas-Fermi Theory. Replacing  $\varepsilon'$  by the electric potential again and Fourier transforming Eq 2.50 gives

$$\rho^{\text{Ind}}(\mathbf{q}) = -e^2 \frac{dn}{d\mu} \phi(\mathbf{q}) \tag{2.51}$$

Comparing this to Eq 2.44 we found for the susceptibility and the dielectric function

$$\chi = -e^2 \frac{dn}{d\mu}$$
 and  $\epsilon(\mathbf{q}) = 1 + \frac{e^2 dn/d\mu}{\epsilon_0 q^2}$  (2.52)

It is important to note that the susceptibility does here not depend on the wavevector  $\mathbf{q}$ . This is a result of our approximation that  $\phi(\mathbf{r})$  does not vary much over  $\mathbf{r}$ .

#### Thomas-Fermi Wavevector

Defining the Thomas-Fermi Wavevector  $\mathbf{k}_0 = \frac{e^2 dn/d\mu}{\epsilon_0}$  allows us to write Eq 2.52 as

$$\epsilon(\mathbf{q}) = 1 + \frac{\mathbf{k}_0^2}{\mathbf{q}^2} \tag{2.53}$$

From this expression it is visible that we can define a characteristic lengthscale of the electron screening. This lengthscale is given as  $1/\mathbf{k}_0$  and we refer to it as **Thomas-Fermi Screening Length**. For copper we have a screening length of  $1/\mathbf{k}_0 = 0.55 \,\text{Å}$ .

#### **Example: Coulomb Potential**

Looking at a point charge Q placed in a sea of conduction eletrons, as illustrated in Fig.2.3 .b we can express the externel potential generated from the point charge as  $\phi^{\rm Ext}(\mathbf{r}) = Q/r$  with its Fouriere transform  $\phi^{\rm Ext}(\mathbf{q}) = 4\pi \ Q/q^2$ . Rearanging Eq ?? and replacing the dielectric function by Fig.2.53 , we obtain

$$\phi(\mathbf{q}) = \frac{\phi^{\text{Ext}}(\mathbf{q})}{\epsilon(\mathbf{q})} = \frac{4\pi Q}{q^2 + k_0^2}$$
(2.54)

Fourier transform this back to real space leads to the well known Yukawa Potential

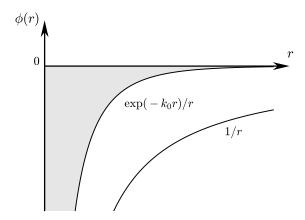


Figure 2.4: Illustration of bare coulomb potential  $(r^{-1})$  and the Yukawa potential.

$$\phi(\mathbf{r}) = Q \frac{e^{-ik_0r}}{r} \tag{2.55}$$

The difference between the Coulomb- and Yukawa potential are displayed in Fig.2.4. The screening effect added point charge by the surrounding negative charges is clearly visible in the faster convergence to zero of the Yukawa Potential. The screening also weakens the longe range interaction of the charges in the solid.

#### 2.2.2 Lindhard Potential

A more sophisticated approach is provided by the Lindhard Theory from which the Thomas-Fermi screening is a special case ( $\mathbf{q} \ll \mathbf{k}_0$ ).

As we know, the charge density for one electron with momentum  $\mathbf{k}$  is given as  $\rho_{\mathbf{l}}(\mathbf{r}) = e|\psi_{\mathbf{k}}(\mathbf{r})|^2$ . Describing the the overall charge density of a crystal, assuming it is in its groundstate we have to sum the one electron charge density multiplied with the fermi function  $f_{\mathbf{k}}$  over all momenta  $\mathbf{k}$ 

$$\rho(\mathbf{r}) = -e \sum_{\mathbf{k}} f_{\mathbf{k}} |\Psi_{\mathbf{k}}(\mathbf{r})|^2$$
 (2.56)

Assuming the additional term  $e\phi$  in our Hamiltonian ( Eq 2.47 , introduced throught the external charge, can be treated as a pertubation, we can express the one electron wavevector  $\psi_{\bf k}$  as

$$\psi_{\mathbf{k}} = |\mathbf{k}_0\rangle + \sum_{\mathbf{k}}' \frac{\langle \mathbf{k}' e \phi | \mathbf{k}_0 \rangle}{\varepsilon_{\mathbf{k}_0} - \varepsilon_{\mathbf{k}'}}$$
 (2.57)

Plugging now our ansatz for  $\psi$  into  $\mbox{ Eq }\ref{eq:property}$  leads to

## Chapter 3

## Magnetism

## 3.1 Paramagnetism

#### Magnetic Moment

#### insert picture - magnetic moment

If  $\vec{A}$  is the area inside the loop and I the current, the magnetic moment can be written as

$$\vec{\mu} = I\vec{A} \tag{3.1}$$

#### Example: Hydrogen Atom

#### insert picture - hydrogen atom and orbiting electron

The magnetic moment of a hydrogen atom can be described semi-classically by assuming the electron to be on a fixed trajectory orbiting the hydrogen nucleus with constant radius r and velocity v. The current I produced by the moving electron can be written as  $I=-e/\tau$  with the orbital period  $\tau=2\pi r/v$ . For the loop area we use the formula  $A=\pi r^2$ . This leads the magnetic moment to be

$$\mu = \frac{-ev\pi r^2}{2\pi r} = \frac{-evr}{2} = \frac{-emvr}{2m} \tag{3.2}$$

We can identify the classical definition of the angular momentum  $\vec{l}=\vec{r}\times m\vec{v}$  in the numerator. Taking the quantisation of angular momentum  $|\hat{l}|=n\hbar$  in quantum mechanics into account we can rewritte expression 3.2 for n=1

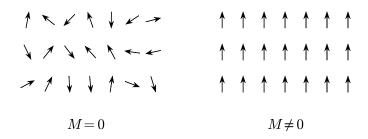


Figure 3.1: hohohoh

$$\mu \simeq \frac{-e\hbar}{2m} \equiv -\mu_B \tag{3.3}$$

Which we define as the Bohr magneton  $\mu_B$ .

#### Magnetic Moment of Atoms

The total angular momentum in an atom is given as the sum of total orbital angular Momentum  $\hat{L}$  and total spin angular momentum  $\hat{S}$ 

$$\hat{J} \simeq \hat{L} + \hat{S} \tag{3.4}$$

#### Magnetisation

Considering a solid with N atoms each having a magnetic moment  $\vec{\mu}$ . We define the magnetization as

$$\vec{M} \equiv \frac{\text{Magnetic moment}}{\text{Volume}}$$
 (3.5)

### Magnetic Moment in a $\vec{B}$ -Field

From classical electrodynamics we know that the potential energy E of a magnetic moment  $\vec{\mu}$  in a magnetic field  $\vec{B}$  is described as

$$E = -\vec{\mu} \cdot \vec{B} \tag{3.6}$$

To get a feeling for the order of magnitudes of magnetic energy of atomic scale we

$$\mu_B \times 1 \text{Tesla} \simeq 0.05 \,\text{meV}$$

This is comparable to the energy one has to put into a system to increase its temperature by 1 K ( $k_B \cdot 1$  K  $\simeq 0.084$  meV).

#### Magnetic susceptibility

$$\chi \simeq \frac{\mu_0 \vec{M}}{\vec{B}} \tag{3.7}$$

We can look at the magnetic susceptibility  $\chi$  as response function.  $\chi$  describes the response of the system  $\vec{M}$  when exposed to an external changing field  $\vec{B}$ . To simplify things in the calculations we will make the following constraints on  $\vec{B}$ .

 $\vec{B}$  is static  $\Rightarrow$  no time dependence

 $\vec{B}$  is homogeneous  $\Rightarrow$  no dependence on  $\vec{r}$ 

The definition of the susceptibility allows us now to classify materials depending on how the respond to an external magnetic field. We calc materials to be *paramagnetic* if they align there spin parallel to the applied  $\vec{B}$ -Field. This is the case if  $\chi > 0$ . On the contrary we refer to materials as *diamagnetic* if their inner magnetic moments align anti-parallel in respect to an applied field  $\vec{B}$ . This results in a negative value for the susceptibility  $\chi < 0$ .

In reality, the response of systems is composed of multiple differenent responses which can have different origins.

$$\chi_{total} \simeq \chi_{paramagnetic} + \chi_{diamagnetic} + \dots$$
 (3.8)

Paramagnetism and diamagnetism can also have different origins. For example:

$$\chi_{\text{paramagnetic}} \simeq \chi_{\text{Langevin}} + \chi_{\text{Van Vleck}} + \chi_{\text{Pauli}} + \dots$$

$$\chi_{\text{diamagnetic}} \simeq \chi_{\text{electronic}} + \chi_{\text{superconductivity}} + \dots$$

#### Diamagnetism

#### insert picture - loop current to illustrate lens law

To symbolize diamagnetic behaviour we recall lens law. It states, that a changing magnetic field induces a current which creates for itself a magnetic field that points in the opposit

direction as the initial magnetic field change. In that way the system tries to compensate the externaly induced field changes.

To describe that, we use 3.1, with the area  $\vec{A} = \pi < r^2 >$  and the current  $I = -Ze/\tau$  with the number of electrons Z per atom and the orbit period  $\tau$ . expression the orbit period in terms of the cyclotron frequency  $\omega = eB/2m$  shows the dependency on the magnetic field.

$$\tau = \frac{2\pi}{\omega} = \frac{4\pi m}{eB} \tag{3.9}$$

Putting everything together we get the following expression to approximate the magnetic moment of a diamagnet

$$\vec{\mu} = -\frac{Ze^2B}{4\pi m}\pi < r^2 > = -\frac{Ze^2B}{4m} < r^2 >$$
 (3.10)

The magnetisation of a macroscopic material can be written as the sum over all the magnetic moments of its atoms resulting in

$$\vec{M} = \mu_0 N \vec{\mu} \tag{3.11}$$

where N denotes the number of atoms per volume. Recalling 3.7, we find for the diamagnetic susceptibility caused by electrons

compare formula with wikipedia. probably wrong by a prefactor

$$\chi_{\text{Diamagnetic}} = -\frac{\mu_0 N e^2 < r^2 > Z}{4m}$$
 (3.12)

Simplest Case:  $e^-$  only

Consider a system consisting of atoms with only one electron

$$\rightarrow \quad \hat{L} = 0, \quad \hat{S} \simeq 1/2 \qquad \Rightarrow \hat{J} = 1/2 \tag{3.13}$$

Further more we asume for the spin g-factor  $g \simeq 2$  (not proved). This system has only to states: spin-up and spin-down.

Consider the energy of these two states leads to

$$E = -\vec{\mu} \cdot \vec{B} = \pm \mu_B \cdot |\vec{B}| \tag{3.14}$$

vector diagram nergy consideration of magnetic moment in B-field

As stated before, the Energy resulting from magnetic moments of the order of magnituted of a bohr magneton  $\mu_B$  to magentic fields to several Tesla are in the same the energy range of some Kelvin (some  $k_BT$ ). Therefore we use Boltzman statistics to describe the **population** distribution of the system. The probability of energy level  $E_1$  to be populated can be written as

$$p_{\alpha} = \frac{e^{\beta E_{\alpha}}}{\mathcal{Z}} = \frac{e^{\beta E_{\alpha}}}{e^{\beta E_{1}} + e^{\beta E_{2}}}, \quad \text{with } \alpha = \{1, 2\}$$
 (3.15)

where  $\beta = 1/k_BT$  and the partition function  $\mathcal{Z} = \sum_i e^{\beta E_i}$ .

$$p_1 = \frac{e^{\beta E_1}}{e^{\beta E_1} + e^{-\beta E_1}}, \quad p_2 = \frac{e^{-\beta E_1}}{e^{\beta E_1} + e^{-\beta E_1}}$$
 (3.16)

This has to be equal to  $p_1 = N_1/N$  and  $p_2 = N_2/N$  with  $N_1$  ( $N_2$ ) being the number of electrons in the energy state  $E_1$  ( $E_2$ ) and  $N = N_1 + N_2$  being the total number of electrons in the system. Furthermore we subtitute  $\bar{x} = \beta E_1 = \mu_B b/k_B T$ .

### plots energy level polulation

The magnetisation M can now be written as the difference in the number of electrons with opposite magnetic moment

$$M = (N_1 - N_2)\mu_B = N\mu_B \left(\frac{N_1}{N} - \frac{N_2}{N}\right) = N\mu_B \left(\frac{e^{\bar{x}} - e^{-\bar{x}}}{e^{\bar{x}} + e^{-\bar{x}}}\right)$$

$$M = N\mu_B \tanh(\bar{x})$$
(3.17)

Looking at the case  $\bar{x} \ll 1$ , which stands for having large Temperatures and/or low magnetic fields. We get for the magnetisation

$$M \simeq N\mu_B \bar{x} = N \frac{\mu_B^2 B}{k_B T} \tag{3.18}$$

With the help of this equation we can now write down an expression for the paramagnetic susceptibility for (large temperature or low magnetic fields)

$$\chi_{\text{paramagnet}} = \frac{N\mu_B^2}{k_B T}$$
(3.19)

We refer to this formula as Curie law and we call the introduced constant  $C = N\mu_B^2/k_B$ Curie-constant. 3.1.1 How to fill the valence shell

Hund's rules

The hund's rules provide simply rules which give an idea about the total spin-, total orbital-

and total angular momentum of an atom in its ground state

1. Maximize S to minimize the Coulomb energy

2. Maximize L to minimize Coulomb repulsion

3. J = |L - S| for less than half filling and J = |L + S| for more than half filling.

Crystal field effects

Atoms arranged in a lattice feel the presence of the atoms around. This happens in a way,

that the valence electrons of the atom under consideration are experiencing the electric field from the ligands surrounding the atom. That causes the energy degeneracy, apparent in a

free atom, to be lifted. This splitting of energy levels has impact on the occupation of the

different orbitals when "filling" the atom with electrons. For studying this effect now we

will consider the valence electrons to partially fill the 3d-orbital because this is also the case

in a variety of materials in nature. Partially filled 3d-bands can especially be found in the

Transition metals. We now exemplify that at at systems with ligands arranged octahedraly

and tetragonal.

Octahedral crystal environment

Tetrahedral crystal environment

Variation of crystal field environment

Jahn-Teller Distortion

complete this subsubsection with infos from lecture notes

22

Resonant inelastic x-ray scattering (RIXS)

### 3.2 Ferromagnetism

### 3.2.1 $H_2$ Molecule

#### **Wave Function Considerations**

$$\Psi^{Total}(2 \text{ Electrons}) \rightarrow \text{Antisymmetric}$$
(3.20)

$$\Psi^{Total}(\vec{r}_1, \vec{r}_2) = -\Psi^{Total}(\vec{r}_2, \vec{r}_1) \tag{3.21}$$

$$\Psi_A = \psi_{\alpha}(\vec{r}_1)\psi_{\beta}(\vec{r}_2) - \psi_{\alpha}(\vec{r}_2)\psi_{\beta}(\vec{r}_1) \tag{3.22}$$

$$\Psi_S = \psi_{\alpha}(\vec{r}_1)\psi_{\beta}(\vec{r}_2) + \psi_{\alpha}(\vec{r}_2)\psi_{\beta}(\vec{r}_1)$$
(3.23)

Consider now spin wave function:

$$\chi_{S} = \chi_{Symmetric} = \begin{cases} |\uparrow_{\alpha}\downarrow_{\beta}\rangle & |1,1\rangle \\ (|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle)/\sqrt{2} & |1,0\rangle \\ |\downarrow_{\alpha}\uparrow_{\beta}\rangle & |1,-1\rangle \end{cases}$$
(3.24)

$$\chi_A = \chi_{Antisymmetric} = (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle)/\sqrt{2} \quad |0,0\rangle$$
(3.25)

Where to the  $\chi_S$  is referred to as **Triplet state** and to the wave function  $\chi_A$  is referred to as textbfSinglet state.

#### Quantum mechanical Spin-Operators

Recalling that  $\hat{S}^2|S,m\rangle=S(S+1)|S,m\rangle$  we get for the eigenvalues of  $\hat{S}^2_{\alpha}$  and  $\hat{S}^2_{\beta}$ 

$$\hat{S}_{\alpha}^{2}|S_{\alpha},m\rangle = S_{\alpha}(S_{\alpha}+1)|S_{\alpha},m\rangle = 3/4$$
 (3.26)

$$\hat{S}_{\beta}^{2}|S_{\beta},m\rangle = S_{\beta}(S_{\beta}+1)|S_{\beta},m\rangle = 3/4 \tag{3.27}$$

$$\hat{S} = \hat{S}_{\alpha} + \hat{S}_{\beta} \implies \hat{S}^{2} = \hat{S}_{\alpha}^{2} + \hat{S}_{\beta}^{2} + 2\hat{S}_{\alpha}\hat{S}_{\beta} \implies \hat{S}_{\alpha} \cdot \hat{S}_{\beta} = \frac{\hat{S}^{2} - \hat{S}_{\alpha}^{2} - \hat{S}_{\beta}^{2}}{2}$$
(3.28)

Calculating  $\langle \hat{S}_{\alpha} \cdot \hat{S}_{\beta} \rangle$  leads to 1/4 for  $\chi_S$  and -3/4 for the  $\chi_A$  case.

#### Consider weak Coulomb interaction

$$H = H_{\text{signel-H}} + H_{\text{int}} = H_0 + H_{\text{int}} \tag{3.29}$$

Where the interaction Hamiltonian  $H_{\text{int}}$  includes the proton-proton, electron-electron, proton 1 - electron 2 and electron 1 - proton 2 interactions.

$$H_{int} = \frac{e^2}{d_{pp}} + \frac{e^2}{d_{ee}} - \frac{e^2}{d_{ep}} - \frac{e^2}{d_{pe}}$$
 (3.30)

Here  $d_{pp}$  stands for the proton-proton distance,  $d_{ee}$  for the electron-electron distance. Furthermore contains  $H_{\text{single-H}}$  both Hamiltonians of the single hydrogen atoms

$$H_{\text{single-H}} = H_{H_1} + H_{H_2} = \frac{\hbar}{2m} \left( \nabla_{\alpha}^2 + \nabla_{\beta}^2 \right) - \left( \frac{e^2}{d_{p_{\alpha}e_{\alpha}}} + \frac{e^2}{d_{p_{\beta}e_{\beta}}} \right)$$
 (3.31)

$$E_{+} = E_{S} = \langle \Psi_{S} | H_{\text{int}} | \Psi_{S} \rangle = \int (\psi_{\alpha} \psi_{\beta} + \psi_{\beta} \overline{\psi_{\alpha}})^{*} H_{\text{int}} (\psi_{\alpha} \psi_{\beta} + \psi_{\beta} \psi_{\alpha}) d^{3}r \quad (3.32)$$

$$E_{-} = E_{A} = \langle \Psi_{S} | H_{\text{int}} | \Psi_{S} \rangle = \int (\psi_{\alpha} \psi_{\beta} - \psi_{\beta} \psi_{\alpha})^{*} H_{\text{int}} (\psi_{\alpha} \psi_{\beta} - \psi_{\beta} \psi_{\alpha}) d^{3}r \quad (3.33)$$

By defining  $C \equiv C_1 + C_2$  and  $J \equiv J_1 + J_2$  one can write the two energies as

$$E_{+} = C \pm J \tag{3.34}$$

Furthermore for the difference of the singlett- and triplett energy we get

$$E_{+} - E_{-} = 2J \int \psi_{\alpha}^{*} \psi_{\beta}^{*} H_{\text{int}} \psi_{\alpha} \psi_{\beta} d^{3}r \qquad (3.35)$$

From this equation we can associate the introduced variable J as the **Exchange Integral**.

$$J = \frac{E_{+} - E_{-}}{2} = \int \psi_{\alpha}^{*} \psi_{\beta}^{*} H_{\text{int}} \psi_{\alpha} \psi_{\beta} d^{3} r$$
 (3.36)

#### What is H<sup>Spin</sup>int

$$E_{\pm} = C \pm J = C + J/2 + 2J + \cdot \rangle \hat{S}_{\alpha} \cdot \hat{S}_{\beta} \langle = \text{constant} + 2J \rangle \hat{S}_{\alpha} \cdot \hat{S}_{\beta} \rangle$$
 (3.37)

The constant contribution to the energies  $E_{\pm}$  can be neglected since absolute energies values arbitrary. The interesting term for us is the second one on the right-most side. It gives us a quantitative measure of how large the energy difference between the two spin configurations is

$$\Rightarrow H_{\text{int}}^{\text{Spin}} = -2J\hat{S}_{\alpha} \cdot \hat{S}_{\beta} \quad \begin{cases} J > 0 & \Rightarrow E_S > E_A \\ J < 0 & \Rightarrow E_A < E_S \end{cases}$$
 (3.38)

From the upper formula we see that J us if  $\chi_S$  or  $\chi_A$  is preferred. Therefore it can be seen as an indication if ferro- or antiferromagnetism is present in a material.

#### Ferromagnetism

$$H = -\sum_{ij} J_{ij} \vec{S}_i \cdot \vec{S}_j + g\mu_B \cdot \sum_j \vec{S}_j \cdot \vec{B}$$

$$= -\sum_j \sum_i j_{ij} \vec{S}_i \cdot \vec{S}_j + g\mu_B \sum_j \vec{S}_j \cdot \vec{B}$$

$$= g\mu_B \sum_j \vec{S}_j \cdot (\vec{B}_{mf} + \vec{B}_j)$$
(3.39)

Using a mean field approximation we rewrite the interaction from all spins on  $\vec{S}_j$  from the first term as with a mean magnetic field  $\vec{B}_{mf}$  wheresa we dfined  $\vec{B}_{mf} \equiv -2/g\mu_B \sum_i J_{ij} \vec{S}_i$ .

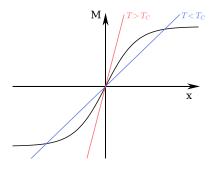
#### Conjecture

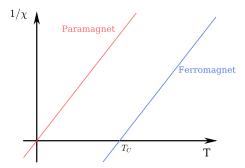
By making an educated guess one could assume, that the mean magnetic field  $\vec{B}_{mf}$  can be approximated macroscopically with the following expression.

$$\vec{B}_{mf} \simeq \lambda \cdot \vec{M}$$
 (3.40)

#### Solution

Solution can be adapted from the results about paramagnetism we gained last week. By focusing on the case  $\vec{B}=0$  we get





**Figure 3.2:** Illustration of the grafical solution of 3.41. the straight lines refer to different values of of T.

Figure 3.3: Illustration of susceptibility  $\chi$  of a Para- and Ferromagnet.

$$M \simeq N\mu_B \tanh(x)$$
 with  $x = \frac{\mu_B}{k_B T} (\vec{B} + \lambda \vec{M})$  (3.41)

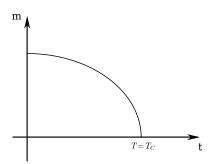
Since the argument of the tangent hyperbolicus depends also on the magnetisation  $\vec{M}$  we have a implicit equation. A solution of this equation is illustrated in ?? as the crossing point between the hyperbolic tangent and the straight line. In this graph it is also visible, that above a certain Temperature  $T > T_C$  there only exists one solution for of the implicit equation which can be associated with the paramagnetic phase of the material. On the other hand for  $T < T_C$  we see that there are exists multiple solution of 3.41 which is in accordance with the magnetisation curve of a ferromagnet. We can determine the Transition temperature  $T_C$  by comparing the slopes of the two curves at the origin

$$\frac{d}{dx}N\mu_B \tanh(x) = \frac{d}{dx}\frac{k_B T}{\mu_B \lambda}x \quad \Rightarrow \quad T_C = \frac{\lambda N \mu_B^2}{k_B} = \lambda \cdot C \quad (3.42)$$

Looking at the limit  $x \ll 1$ 

Checking the magnetisation M(T) at zero field  $\vec{B}=0$ . Using 3.41 with zero field and the definitions  $m=M/N\mu_B$  and  $t=k_BT/N\mu_B^2\lambda=T/T_C$  we get

$$m = \tanh\left(\frac{m}{t}\right) \tag{3.43}$$



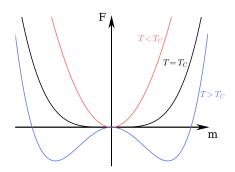


Figure 3.4: Temperature dependence of magnetisation.

Figure 3.5: Free Energy dependence on the order parameter m for the 3 cases  $T < T_C$ ,  $T = T_C$  and  $T > T_C$ .

#### Landau Theory

According to the Landau theory of phase transisiton the free energy F can be expressed

$$F = F_0 + a(T)m^2 + bm^4 + \dots (3.44)$$

Were the parameter a and b has to meet the conditions

$$a(T) = a_0(T - T_C)$$
 and  $b > 0$  (3.45)

We find the thermodynamical state of our system by minimizing the free energy

$$\frac{dF}{dm} = m(2a(T) + 4bm^2) = 0 \quad \Rightarrow \quad m = \begin{cases} 0 \\ \pm \sqrt{\frac{a_0(T - T_C)}{2b}} \end{cases}$$
 (3.46)

#### 3.2.2 Exchange Interaction J

Analog to the susceptibility of ferromagnets  $\chi_{FM} = C/(T - T_C)$ , we can define the susceptibility for anti-ferromagnets

$$\chi_{AFM} = \frac{C}{T + T_N} \tag{3.47}$$

Where we refer to the Transition Temperature  $\mathcal{T}_N$  as Neel-Temperature.

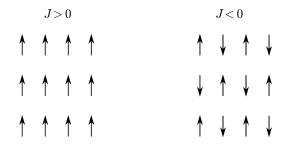
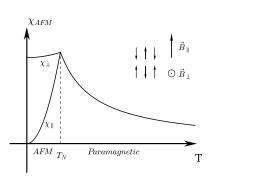


Figure 3.6: hohoho



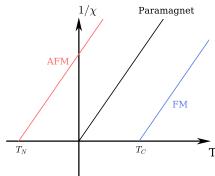


Figure 3.7: Temperature dependence of magneti- Figure 3.8: Free Energy dependence on the order sation. parameter m for the 3 cases  $T < T_C$ ,  $T = T_C$  and  $T > T_C$ .

### 3.2.3 Ferromagnetic Magnons

Consider a linear FM chain  $|\text{FM}\rangle=|\uparrow\uparrow\uparrow\dots\rangle$ . Applying the latter operator  $s_j^-$  onto this expression leads to

$$|j\rangle = s_j^-|\text{FM}\rangle = |\uparrow\uparrow\dots\uparrow\downarrow\downarrow\uparrow\dots\rangle$$
 (3.48)

Defining

$$|q\rangle = \frac{1}{\sqrt{N}} \sum_{j} e^{iqR_j} |j\rangle \tag{3.49}$$

The Hamilton is given as

$$H = -\sum_{ij} J_{ij} \hat{S}_i \cdot \hat{S}_j = -2J \sum_i \hat{S}_i \cdot \hat{S}_{i+1}$$
$$= -2J \sum_i \left\{ \hat{S}_i^z \hat{S}_{i+1}^z + \frac{1}{2} \left[ \hat{S}_i^+ \hat{S}_{i+1}^- + \hat{S}_i^- \hat{S}_{i+1}^+ \right] \right\}$$
(3.50)

By taking only nearest neighbour interactions into account we get the The second equality sign holds if we only take neares neighbour interactions into account. To get the final expression we used the substitution

$$\hat{S}^2 = \hat{S}^{z^2} + \frac{1}{2} \left[ \hat{S}_i^+ \hat{S}_{i+1}^- + \hat{S}_i^- \hat{S}_{i+1}^+ \right]$$
 (3.51)

$$H|FM\rangle = -2JNS^2|FM\rangle = E_0|FM\rangle$$
 (3.52)

$$H|j\rangle = -2J\{(N-4)S^{2}|j\rangle + S[|j+1\rangle + |j-1\rangle]\}$$
 (3.53)

$$H|q\rangle = \frac{1}{\sqrt{N}} \sum_{j} e^{iqR_{j}} \left\{ NS^{2}|j\rangle - 2S^{2}|j\rangle + S|j+1\rangle + S|j-1\rangle \right\}$$

$$= -2JNS^{2}|q\rangle - 2J\left\{ -2S^{2} + \left( e^{iqa} + e^{-iqa} \right) \right\} |q\rangle$$

$$= E_{0}|q\rangle + 2JS\left\{ 1 - \cos(qa) \right\} |q\rangle$$
(3.54)

Since we are consindering an infinit long 1D chain of spin states we can rewrite the expression

$$H|q\rangle = E_0|q\rangle + 2JS\left\{2 - 2\cos(qa)\right\}|q\rangle \tag{3.55}$$

which leads to

$$H|q\rangle = E(q)|q\rangle$$
 with  $E(q) \simeq E_0 + 2JS(2 - 2\cos(qa))$  (3.56)

This is the dispersion relation for ferromagnets. For anti-ferromagnets we have a similar relation (not derived)

$$\hbar\omega = 2J|\sin(qa)| \tag{3.57}$$

## Chapter 4

## Superconductivity

## 4.1 Conventional superconducters

#### 4.1.1 Link $\rho$ & Meissner Effect

As we know, superconductors are characterised by vanishing resist vity ( $\rho=0$ ) for undergoing a critical temperature  $T_C$  ( Fig.4.1 ). Recalling Ohm's law,

$$\vec{j} = \sigma \vec{E} \tag{4.1}$$

which relates the electron current density  $\vec{j}$  and the electric field inside a conductor  $\vec{E}$  by introductin the material dependent conductivity  $\sigma$ .

Looking at the case of superconductors, we can emphasize, by rearranging Ohm's law, that the electric field vanishes.

$$\vec{E} = \frac{1}{\sigma} \vec{j} = \rho \vec{j} \simeq 0 \tag{4.2}$$

By using one of the Maxwell equation  $\nabla \times \vec{E} = -d\vec{B}/dt$  we can conclude, that for a vanishing electric field  $\vec{E} = 0$  there can't be a change of the magnetic field over time  $d\vec{B}/dt = 0$ .

diagram vanishing B-field

## 4.2 London Equation

The lonodon equation is a macroscopic theory that was one of the first theories which allowed to describe effects like the Meissner-Ochsenfeld effect quantitatively. We know want

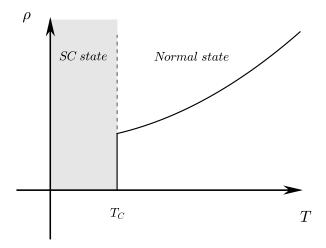


Figure 4.1: hohoho

to qualitiatively derive the london equation. For that consider a current oscillating if the frequency  $\omega$ 

$$\vec{j}e^{-i\omega t} = \sigma(\omega)\vec{E}e^{-i\omega t} \tag{4.3}$$

Using the expression for the conductivity  $\sigma(\omega)$  from the Drude model for a time-dependent electric field

$$\sigma(\omega) = \frac{\vec{j}}{\vec{E}} = \frac{ne^2\tau}{m} \frac{1}{1 - i\omega\tau}$$
 (4.4)

From this we get, as already proven in a previous excercise, that the real part of the conductivity can be written in the limit of  $\tau \to \infty$  as

$$\operatorname{Re}(\sigma(\omega)) = \frac{\pi n e^2}{m} \delta(\omega)$$
 (4.5)

This can also be made clear, when looking at the limits

$$\operatorname{Re}(\sigma(\omega)) = \frac{ne^2}{m} \frac{\tau}{1+\omega\tau} \xrightarrow{\omega\to 0} \frac{ne^2\tau}{m}$$
 (4.6)

$$\sigma(\omega) \xrightarrow{\tau \to \infty} -\frac{ne^2}{im\omega} = \operatorname{Im}(\sigma(\omega)) \implies \operatorname{Re}(\sigma(\omega)) = 0 \quad \text{for } \omega \neq 0$$
 (4.7)

From Eq.4.5 , we see that in the regime of very large mean free times  $\tau$  between ionic collisions, the real part of the conductivity  $\sigma(\omega)$  is only different from zero for vanishing oscillations of the electric field.

To derive the London equation, the following two assumptions have to be made

- 1.  $\tau \to \infty$ , which is given since there are no collisions between electrons and lattice ions in the superconducting state.
- 2. We have to split up the conductivity in a part  $\sigma_N(\omega)$  of normal state electrons and a conductivity contribution due to super electrons  $\sigma_S$ :  $\sigma(\omega) = \sigma_N(\omega) + \sigma_S(\omega)$ , were only  $\sigma_S(\omega) \neq 0$  in the superconducting state

Taking the curl of the oscillating current

$$\nabla \times (\vec{j}e^{-i\omega t}) = -\sigma(\omega)\nabla \times (\vec{E} \cdot e^{-i\omega t}) = -\sigma(\omega)\frac{d}{dt}(\vec{B} \cdot e^{-i\omega t})$$
$$= -\frac{n_S e^2}{m}\vec{B} \cdot e^{-i\omega t}$$
(4.8)

Here we used Eq.4.3 in the first and farradays law  $(\nabla \times \vec{E} = d/dt\vec{B})$  in the second step. Furthermore is only the *super electron* density  $n_S$  contributing to the conductivity  $\sigma(\omega)$ . From this equation we can deduce that

$$\nabla \times \vec{j} \propto \vec{B}$$
 and thereof  $\vec{j} = -\frac{n_S e^2}{m} \vec{A}$  (4.9)

with  $\vec{A}$  being the vector potential. The equation describing this proprtionality between the current density  $\vec{j}$  and the vector potential  $\vec{A}$  is the London equation.

Together with Ampere's law  $(\nabla \times \vec{B} = \mu_0 \vec{j})$  allows us the London equation (Eq.4.9) now to set up a differential equation for the magentic field  $\vec{B}$ .

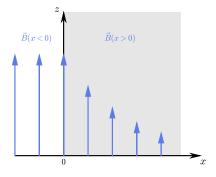
$$\nabla \times \nabla \times \vec{B} = -\frac{1}{\lambda^2} \vec{B}, \quad \text{where } \lambda = \sqrt{\frac{m}{\mu_0 n_S e^2}}$$
 (4.10)

This equation can also be written as

$$\nabla^2 \cdot \vec{B} = \frac{1}{\lambda^2} \vec{B} \tag{4.11}$$

To exemplify the solution of this differential equation we look at it in one dimension. Then Eq.4.11 becomes  $d^2/dx^2B_z(x) = 1/\lambda^2B_x$ .

To exemplify the solution of this differential equation we look at a situation were we have a superconductor filling the space for all values  $x \ge 0$  correspinding to that we have vacuum



 $B_z$   $\lambda_1$   $\lambda_2$ 

Figure 4.2: Illustration of the magnetic field decrease inside the superconductor according to Eq.4.13 . The magnetic field is pictorially drawn as blue vectors. Adapted from redKittel - include citation

**Figure 4.3:** Illustration of Eq.4.13 for two different values for the london penetration depth.

at x < 0. Applyling now a magnetic field in the vacuum region aligned along the z-axis  $\vec{B} = (0, 0, B_0)$ . in this case the differential equation Eq.4.11 simplifies to

$$\frac{d^2}{dx^2}B_z(x) = \frac{1}{\lambda^2}B_z(x) \tag{4.12}$$

Using the boundary conditions  $B_z(x) \xrightarrow{x \to \infty} 0$  (Meissner Effect) and  $B_z(x \to 0) = B_0$  leads to the solution

$$B_z(x) = B_0 e^{-x/\lambda} (4.13)$$

This solution describes an exponentially decreasing magnetic field inside the the superconductor close to the surface. From this formula one can also recognize that  $\lambda = \sqrt{m/\mu_0 n_S e^2}$  introduced a characteristic length scale over which the magnetic penetrates the superconductor. For that reason  $\lambda$  is referred to as *London penetration depth*. In 4.7 we see Eq.4.13 plotted for two different values of lambda  $\lambda_1 > \lambda_2$ . Looking at the definition of the London penetration depth we see that  $n_{S_1} < n_{S_2}$ , which means that the higher the charge density the more is the penetrating magnetic field  $B_z(x)$  diminished.

## 4.3 Coherence Length

Another lengthscale that helps us to classify the different types of superconductors is the coherence length  $\xi$ . The coherence length is a measure of the distance within which the superconducting electron concentration cannot change drastically in a spatially-varying magnetic field.

$$\xi = \frac{\hbar v_F}{\pi \Delta} \tag{4.14}$$

Where  $v_F$  stands for the fermi velocity and  $\Delta$  for superconducting gap being present at the fermi surface in the SC state.  $\xi$  is a parameter in the Ginzburg-Landau Theory and can be derived there.

The two length scales, penetration depth and coherence length, allow us now, to categorize the superconductors in the groups *Type 1* and *Type 2*. This goes as follows:

By defining  $\kappa \equiv \xi/\lambda$  we say a Superconductor contains to the group of Type 1 (Type 2) if  $\kappa < (>) 1/\sqrt{2} \simeq 0.707$ .

Values of  $\kappa$  for some materials are given in length scale table. The classification of SC into Type 1 and Type 2 does not fully coincide with the classification of conventional and unconvential. Although a lot of the conventional (elementary) supperconductors are of Type 1, there are also some expetions like Niobium (Nb) which is a conventional Type 2 Superconductor.

include table from PPP

## 4.4 Ginzburg-Landau Theory

The Landau Theory of Phase Transition can also be used in the case of the transition from normal to superconducting state. The free energy per volume of the system in the superconducting state is given as

$$f_{SC} = f_{NS} + a(T)|\psi(T)|^2 + \frac{1}{2}b(T)|\psi(T)|^4 + \dots$$
 (4.15)

 $|\psi|$  is treated as the order parameter. To find the temperature dependence of the free energy per volumne we make the following assumptions: near the transition temperature  $T_C$  has a to be linear in temperature  $(a(T) = a_0(T - T_C) + ...)$  and b given as  $b(T) = b_0 + ...$  with  $a_0, b_0$  being constant and larger than zero.

Looking for the extrema of  $(f_{SC} - f_{NS})$ 

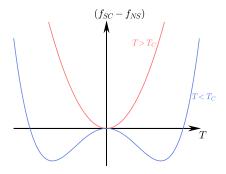
$$\frac{d(f_{SC} - f_{NS})}{d|\psi|} = 2|\psi| \{a(T) + b(T)|\psi|^2\} \stackrel{!}{=} 0$$
 (4.16)

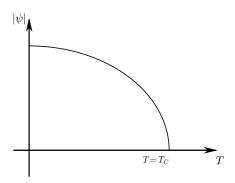
We find the following condition for  $|\psi|$  to minimize  $(f_{\rm SC} - f_{\rm NS})$ 

$$|\psi| = \begin{cases} 0, & T > T_C \\ \pm \sqrt{-a(T)/b(T)}, & T < T_C \end{cases}$$

$$(4.17)$$

Plugging Eq.4.17 into Eq.4.16 leads to





**Figure 4.4:** Difference of the free energy per volume between the normal- and superconducting state.

**Figure 4.5:** Temperature dependence of the order parameter  $|\psi|$ .

$$(f_{SC} - f_{NS}) = -\frac{a^2(T)}{b(T)} + \frac{1}{2} \frac{a^2(T)}{b(T)} = -\frac{1}{2} \frac{a_0^2(T - T_C)^2}{b_0}$$
(4.18)

Since  $f_{SC}$  is the free energy, we can write

$$S \equiv -\frac{df}{dT}$$
 and  $C \equiv T \cdot \frac{dS}{dT}$  (4.19)

This gives for the difference in the heat capacity

$$\Delta C \equiv C_{\rm SC} - C_{\rm NS} = -T \frac{d^2 (f_{\rm SC} - f_{\rm NS})}{dT^2} = T \frac{a_0^2}{b_0}$$
 (4.20)

Evaluating Eq.4.20 at the transition Temperature gives us the jump in specific heat at the phase transition

$$\Delta C(T_C) = \frac{a_0^2}{b_0} T_C (4.21)$$

#### 4.4.1 Orderparameter in the Ginzburg-Landau Model

To point out the connection between the penetration depth resulting from the London Equation and the Ginzburg-Landau Theory, we look a bit closer at the order parameter  $|\psi|$ . Before we defined the term of the super electrons  $n_s$  in the London's theory. In BCS-Theory this quantity is referred to as density of cooper pairs.

$$n_S = \begin{cases} \# \text{super electrons} & \leftarrow \text{London Theory} \\ \# \text{cooper pairs} & \leftarrow \text{BCS Theory} \end{cases}$$

Here we assign

$$n_S = 2|\psi|^2 = 2\frac{a_0(T_C - T)}{b_0}$$
 (4.22)

This allows us now to express the penetretion depth  $\lambda$  from Eq.4.10 as explicitly temperature dependent

$$\lambda(T) = \sqrt{\frac{b_0 m}{2\mu_0 e^2 a_0 (T_C - T)}} \tag{4.23}$$

### 4.5 Classification of different superconductors

#### 4.6 Vorteces

To define the appearance of vorteces from Type 2 Superconductors into into our formalism we define the free energy per volumne as

$$f_{\rm SC} - f_{\rm NS} \simeq a(T)|\psi(r)|^2 + b(T)|\psi(r)|^4 + \frac{\hbar^2}{2m}\nabla|\psi|^2$$
 (4.24)

To get the total Free energy we integrate Eq.4.24 over all space

$$F_{\rm SC} - F_{\rm NS} = \int d^3r (f_{\rm SC} - f_{\rm NS})$$
 (4.25)

We are interested now in finding the wavefunction  $\psi(r)$  for which the Free energy is minimized. Since  $\psi(r)$  is a function, this condition is formulated in a functional of the following form

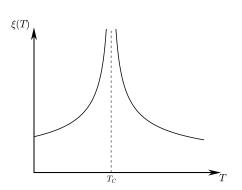
$$\frac{\delta F_{\rm SC}}{\delta \psi(r)} \stackrel{!}{=} 0 \tag{4.26}$$

This is a differential equation for  $\psi(r)$ . In one dimnesion this DEG looks like the following

$$-\frac{\hbar^2}{2m}\frac{d^2\psi(x)}{dx^2} + a(T)\psi(x) + b(T)\psi^3(x) = 0 (4.27)$$

A solution to this differential equation is

$$\psi(x) = \psi_0 \tanh\left(\frac{x}{\sqrt{2}\xi(T)}\right) \tag{4.28}$$



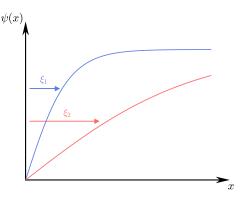


Figure 4.6: Coherence length  $\xi(T)$  plotted over Temperature according to Eq.?? . Noteworthy is the diverging of  $\xi$  when going close to  $T_C$ .

Figure 4.7: Wavefunction in one dimension according to Eq.4.28. Comparing the the blue and red curve shows that the wave function gets *pushed* down for increasing coherence length  $(\xi_2 > \xi_1)$ .

The coherence length  $\xi(T)$  introduces a temperature dependent length scale and is given as

$$\xi(T) = \left(\frac{\hbar}{2m|a(T)|}\right)^{1/2} = \xi_0 t^{-1/2} \quad \text{with} \quad t = \frac{T - T_C}{T_C}$$
 (4.29)

Here we used the same expression for a(T) as in the ginzburg-Landau Theory.

We can now use our expressions for the penetration depth ( Eq.4.23 ) and the coherence length ( Eq.4.29 ) to express the Ginzburg-Landau paramater  $\kappa$  as

$$\kappa \equiv \frac{\lambda(T)}{\xi(t)} \tag{4.30}$$

Since the temperature dependence of both characteristic length scales cancels out,  $\kappa$  is independent of T.

Each vortex has one flux quantum  $\Phi_0$ . The diameter of a vortex is given by the coherence length  $\xi$ . Looking at a cut through a superconductor with edge length L, we can define the total surface of the material covered by vorteces as

$$2\pi \xi^2 N = A_{\text{vortex}} \tag{4.31}$$

where N is the number of vorteces in the material an. Looking for the situation where  $A_{\rm vortex} = L$  we can say, that the superconducting state is fully repressed by the vortex inherent magnetic field. The magnetic field inside the superconductor can then be written as

$$H = \frac{N\Phi_0}{L^2} \tag{4.32}$$

This allows us to set up a relation between the upper critical field  $H_{c_2}$  for which the superconducting state is entirely suppressed and the coherence length, which defines the vortex diameter

$$H_{c_2} \simeq \frac{\Phi_0}{2\pi\xi^2} \tag{4.33}$$

Since one vortex has exactly one flux quantum, we can set up a relation between the penetration depth  $\lambda$  and a lower critical field  $H_{c_1}$ . This lower critical field marks the point when vorteces start to be *produced* at the materials surface. We assume, that the magnetic field penetration the superconductor has at least to be as strong as one flux quantum to be able to create a vortex

$$\pi \lambda^2 H \stackrel{!}{=} \Phi_0 \quad \Rightarrow \quad H_{c_1} = \frac{\Phi_0}{\pi \lambda}$$
 (4.34)

sketch - vorteces in the material

sketch - condition for creating one vortex

$$\mathscr{F}\{p,u\}\tag{4.35}$$

$$\mathcal{F}\{p, u\} \tag{4.36}$$

$$\mathscr{F}^{pu} \tag{4.37}$$

$$F^{p}_{u} \tag{4.38}$$

$$f_{pu} (4.39)$$

$$\psi\pi\nu$$
 (4.40)

# Bibliography