Under Attack - Binary Classification of Hazardous Objects from Space

Katie Knight[*]
Anna-Maria Nau[*]
Christoph Metzner[*]
kknigh22@vols.utk.edu
anau@vols.utk.edu
cmetzner@vols.utk.edu
DSE 511 - Project 3
University of Tennessee, Knoxville
Knoxville, Tennessee, USA

Figure 1: Potential threat of extinction through extraterrestrial objects. (https://eparisextra.com/living/nasa-attempts-to-stop-hypothetical-asteroid-from-hitting-earth-and-fails/)

ABSTRACT

This study utilizes machine learning algorithms to identify potentially hazardous objects, such as asteroids, and therefore serve as an early warning system. To perform this binary classification task, that is, to predict whether an asteroid is hazardous or not, four supervised classifiers, including naïve Bayes, support vector machine, decision tree, and random forest were trained on data provided by the NASA API called NeoWS (Near Earth Object Web Service) and which is readily available on Kaggle (https://www.kaggle.com/shrutimehta/nasa-asteroids-classification). Specifically, each algorithm was trained and tuned on two training sets (i.e., one standardized and one reduced via principal component analysis), and then evaluated on a testing set. In addition, to identify the most important predictors, feature importance was applied. In total, eight classifiers were built and tested, and the highest performance was achieved by the decision tree with a F1-score of 0.9826, a recall of 0.9724, and a precision of 0.993.

[*]All three authors contributed equally to this research.

KEYWORDS

machine learning, supervised, classification, pca, feature importance

1 INTRODUCTION

In "A Synopsis of the Astronomy of Comets", Edmond Halley first voiced his concerns about the devastating impact of extraterrestrial objects, such as asteroids or comets, on the modern world in the 18th century [16]. A group centered around Luis Alvarez hypothesized that such objects caused several mass extinctions and, generally, has a significant influence on the well-being of our planet [1]. In the face of potential extinction, US Congress ordered NASA, in collaboration with The University of Arizona and Hawaii, to develop strategies for observation, tracking, characterization, and damage mitigation [13] for such objects. One specific strategy is the development of statistical models utilizing machine learning that are able to classify potentially hazardous objects based on their properties, such as shape, size, velocity, or shape of orbit. These extraterrestrial objects, asteroids or comets, are referred as near-earth objects (NEOs). NEOs orbit the sun, may potentially cross Earth's orbit [2], and can range in size from a dust particle to kilometers in diameter [11]. NEOs are categorized into near-Earth asteroids (NEAs) and near-Earth comets (NECs) based on their origin in space and geometric appearance. NEAs originate from the Main Belt [3], while NECs come from either the Kuiper belt or the Oort cloud [8].

There is a wide range of publications that propose different strategies for impact risk assessment of NEOs based on common features such as size or velocity. For example, a roughly twenty-year-old approach is to use scales rating the potential impact of NEOs such as the *Torino scale* or *Palermo scale* [14]. Their advantage is that the layman can easily interpret such scales. For example, the Torino scale categorizes the impact risk of a NEO from 1 to 10 by plotting the kinetic energy against the probability of impact. Objects with largest kinetic energy and highest probability of impact receive a score of 10. Another approach focuses on using visible and near-infrared spectroscopy to characterize the compositional and physical structure of NEOs to determine their hazardous potential [7].

Nugent et al. [15] were the first to apply machine learning algorithms to the classification of hazardous NEOs. Their objective was to showcase the usefulness of supervised learning algorithms in classifying hazardous objects from space. A more recent study successfully applied more complex machine learning algorithms based on neural network frameworks to infer the physical properties and impact risks of asteroids by analyzing energy deposition curves [17]. We hypothesize that using machine learning algorithms can play a vital part in the detection of hazardous objects. Specifically, we ask if standard "off-the-shelf" machine learning algorithms can be trained to help identify whether a NEOs is hazardous or not.

2 MATERIALS AND METHODS

This work applied four supervised classification algorithms to a NEO dataset to predict whether a NEO is hazardous or not. Specifically, the classification performances of 1) naïve Bayes as the baseline, 2) support vector machine, 3) decision tree, and 4) random forest were compared. Furthermore, the impact of reducing the number of features via principal component analysis on the classification performance was analyzed. The hyperparameters of all algorithms were optimized using k-fold cross validation (k=5) on the training set via the function *GridSearchCV* from the machine learning library *sklearn*.
The performance of the developed models were evaluated on the testing set using the performance metrics *F1-Score*, *Precision*, and *Recall*. Lastly, to identify the top predictors, feature importance was applied and discussed.

2.1 Data source

The source data for this study was provided by the NASA API called NeoWS (Near Earth Object Web Service available at https://api.nasa.gov), and is readily available at https://www.kaggle.com/shrutimehta/nasa-asteroids-classification. The dataset contains information on 4687 asteroids (rows), and 40 features (columns), one being the target feature indicating the ground truth (hazardous or non-hazardous). Out of the 4687 samples, 3932 are labeled as non-hazardous, and 755 as hazardous. In addition, the data contains no missing values and all features but one are numeric with varying scales. The dataset includes typically measured features of an asteroid, such as the absolute magnitude, estimated diameter, relative velocity, distance measures of the object to the sun such as perihelion distance (minimum distance) and aphelion distance (maximum distance), or the shape of the objects orbit represented by eccentricity.

2.2 Data pre-processing

After analyzing the input features, various redundant, inexpressive, or time-related variables were removed. For example, features describing the *estimated diameter* in different scales (e.g., feet, meters, or miles) were dropped due to redundancy. Also, features containing only one unique value or row identifiers were dropped. Lastly, time-related features, such as *close approach date*, were also removed since knowing the date does not contribute to the fact whether an asteroid will be hazardous or not. After feature removal, the dataset contained 19 input features and one binary target variable representing the ground truth as shown by Table 1. The data was randomly split into an 80:20 ratio as train and test set, respectively. The train set contains 3749 samples (3139 non-hazardous and 610 hazardous), and the test set 938 samples (793 non-hazardous and 145 hazardous).

2.2.1 Standardization. Since the data contained different scales of measurement, the input features were standardized to ensure equal contribution during model fitting. Standardization re-scales all features to have a mean of 0 and a standard deviation of 1 (unit variance). Specifically, each feature was re-scaled using 1.

$$\mathbf{z}_i = \frac{x_{ij} - \overline{x}_j}{\sigma_j} \tag{1}$$

Table 1: The features after data pre-processing.

Features	Features
absolute_magnitude	inclination
est_dia_in_km_min	asc_node_longitude
relative_velocity_km_per_hr	orbital_period
miss_dist._kilometers	perihelion_distance
orbit_uncertainty	perihelion_arg
minimum_orbit_intersection	aphelion_dist
jupiter_tisserand_invariant	perihelion_time
epoch_osculation	mean_anomaly
eccentricity	mean_motion
semi_major_axis	hazardous (target)

where z_i represents the scaled value, x_{ij} the original data point for the ith sample of the jth feature, $\overline{x}_j$ is the mean of the jth feature, and σ_j the standard deviation of the jth feature.

2.2.2 Principal Component Analysis (PCA). Principal Component Analysis (PCA) is an unsupervised dimensionality reduction technique. It is a way to reduce the number of features while maintaining the majority of the important information. It transforms a number of variables that may be correlated into a smaller number of uncorrelated features, known as principal components (PCs). The PCs are linear combinations of the original variables weighted by their variances (or eigenvalues) in a particular orthogonal dimension such that the first PC accounts for the largest variance in the data, the second PC accounts for the second largest variance in the data, and so on. To evaluate the usefulness of the PCs and to determine the number of components to use for model training, the total explained variance ratio metric was used which is the sum of the percentage of variance that is attributed by each of the selected components. The total explained variance ratio chosen for this study was 0.95, which resulted in 11 PCs as shown by Figure 2.
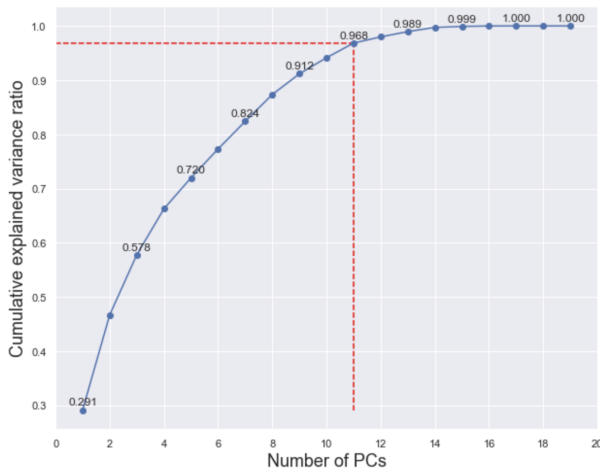

Figure 2: Line plot showing the cumulative explained variance ratio for varying number of PCs. The plot shows that 11 PCs are able to achieve a total explained variance of 0.95.

2.3 Supervised Classification Algorithms

2.3.1 Gaussian Naïve Bayes. The baseline classifier for this binary-classification problem was set to be the Gaussian Naïve Bayes algorithm. This algorithm belongs to the class of generative models, since it tries to learn the *class-conditional* density $p(x|y)$ and the *class priors* $p(y)$ for each value of y (i.e., each category). This allows to use the Bayes rule to compute the posterior probability $p(y|x)$.

$$p(y|x) = \frac{p(x|y) \times p(y)}{\sum_{y'=1}^{C} p(x|y')p(y')} \qquad (2)$$

In naïve Bayes it is assumed that all features are conditionally independent given the class label and in our case to follow a Gaussian distribution. This assumption is usually false but simplifies the computation tremendously [12].

2.3.2 Support Vector Machine. This algorithm belongs to the class of discriminant models that separate the samples based on a boundary (i.e., hyperplane). Vladimir Vapnik [18] developed support vector machine (SVM) applying the idea of structural risk minimization. Specifically, SVM tries to identify a decision boundary that maximizes the margin, distance from the closest positive to the closest negative data sample [5]. The main advantage of SVM is its ability to generalize and properly classify inseparable categories by projecting the data in lower space into higher dimensions using the *kernel trick* [18].

2.3.3 Decision Tree. The decision tree algorithm [10] is a supervised learning algorithm, most useful for either solving problems related to regression or classification. This algorithm creates a training model that predicts a variable class via learning decision rules inferred from the training data. For each class label, the decision tree begins at the "root" attribute, with each value of the root attribute compared to the target variable. This comparison dictates what "branch" corresponds to the attribute value, and creates a set of "if-then-else" decision rules.

2.3.4 Random Forest. The random forest classifier consists of multiple decision trees. Each tree in the algorithm produces a prediction, with the most common prediction among each tree becoming the overall model's prediction [4]. Since individual decision trees tend to overfit and exhibit high variance, the advantage of randomness is the production of decision trees with decoupled prediction errors; averaging the prediction can cancel out some of these errors.

2.4 Performance Evaluation

Performance evaluation is addressed using the harmonic mean of the precision and recall (also known as the *true positive rate*), the so-called F1-score performance metric. The F1-score is defined as follows:

$$F1 = \frac{2 \times precision \times recall}{precision + recall} \qquad (3)$$

$$precision = \frac{tp}{tp + fp} \qquad (4)$$

$$recall = \frac{tp}{tp + fn} \qquad (5)$$

where tp being the true positive events, fp the false positive events, tn true negative events, and fn the false negative events. The positives events are samples indicating hazardous objects. Also, to further evaluate the accuracy of specific models, the receiver operating characteristic curve is used. This plot provides an accuracy measure by plotting the recall (y-axis) vs. 1 - specificity (x-axis). The closer the curve is to the upper left corner, the better the discriminatory ability of the model [6]. The outcomes of the confusion matrix for the test data are also reported in the results section.

2.5 Feature Importance

Feature importance analysis lets us look at what features were most useful for each algorithm. Since Gaussian Naive Bayes and SVM Classifier using RBF-Kernel do not have an intrinsic way of determining feature importance, we used SciKit's permutation importance as a means to inspect which features were significant for classifying NEOs as hazardous or not. Permutation feature importance is when a model's score decreases after one feature value is randomly shuffled. Because this severs the feature and target relationship the score drop indicates how much the model relies on the feature.

3 RESULTS

This section presents the performance results and the confusion matrix of each evaluated model. In addition, a receiver-operating characteristic curve is plotted illustrating the diagnostic ability of each binary classifier under different thresholds. Four algorithms were trained on either the standardized or standardized-pca data, which resulted in a total of eight models. These eight models were evaluated on the respective standardized or standardized-pca test data.

3.1 Performance Evaluation

Table 2 presents the performance results of the models for the F1-score, recall, and precision evaluation metrics. The main takeaway of this study is that models trained on only standardized data outperformed their respective counterpart trained on the standardized-pca data. Decision tree, random forest, and support vector machine outperformed the baseline algorithm, Naïve Bayes, for both datasets. The highest performing model across all metrics was the decision tree trained on standardized data with a F1-score of 0.9826, a recall of 0.9724 and, 0.993 for the precision. Random forest performed similarly as the decision tree, were both performed 6-8 points better than the support vector machine algorithm. Interestingly, support vector machine significantly outperformed the decision tree and random forest when trained on the standardized-pca data. In addition, support vector machine seems to be the most robust algorithm based on the type of pre-processing of the input data given the present results. The fastest algorithm was Naïve Bayes given its simplicity. However, decision tree significantly outperformed Naïve Bayes with similar computing time.

The predicted conditions for each test sample are presented in Table 3. Decision tree was not able to correctly identify four and random forest six hazardous objects on the standardized data. But both only mislabeled one non-hazardous NEO as being hazardous. In

other words, both models were certain in discriminating hazardous from non-hazardous objects.

Table 2: Performance evaluation of the trained models for the F1-score, recall, and precision. The algorithms naïve bayes (NB), support vector machine (SVM), decision tree (DT), and random forest (RF) were trained on standardized data and/or on features created by principal component analysis (PCA). The average computing time per model training during hyperparameter tuning is reported.

Algorithm	Data	F1-Score	Recall	Precision	Time [s]
NB	Scaled	0.8542	0.869	0.84	0.01173
NB	PCA	0.6345	0.5448	0.7596	**0.009791**
SVM	Scaled	0.9181	0.8897	0.9485	9.1910
SVM	PCA	0.8531	0.8414	0.8652	17.5618
DT	Scaled	**0.9826**	**0.9724**	**0.993**	0.0281
DT	PCA	0.6823	0.7034	0.6623	0.0555
RF	Scaled	0.9754	0.9586	0.9929	3.867
RF	PCA	0.741	0.6414	0.8774	5.6026

Table 3: Confusion matrix of the trained models shows the prediction conditions true positive (TP), false positive (FP), false negative (FN), and true negative (TN) of each test sample. The algorithms naïve bayes (NB), support vector machine (SVM), decision tree (DT), and random forest (RF) were trained on the standardized data or on features created by principal component analysis (PCA).

Algorithm	Data	TP	FP	FN	TN
NB	Scaled	126	24	19	769
NB	PCA	79	25	66	768
SVM	Scaled	129	7	16	786
SVM	PCA	122	19	23	774
DT	Scaled	141	1	4	792
DT	PCA	102	52	43	741
RF	Scaled	139	1	6	792
RF	PCA	93	13	52	780

The individual model performances are further analyzed using the receiver operating characteristic curve (ROC-curve) that visualizes the diagnostic ability of a binary classifier with varying threshold boundaries (Figure 3). The ROC-curve supports the statement that model performance depends on the pre-processing method used for the input dataset. The best model performance is arguably the random forest classifier with standardized data (yellow-green line) showing superior performance across almost all thresholds. Both curves for the support vector machine algorithm performed similarly. Decision tree trained standardized-pca data performed worse across all models.

3.2 Feature Importance results

Figures 3, 4, 5, and 6 show the resulting significant features for each classifier.
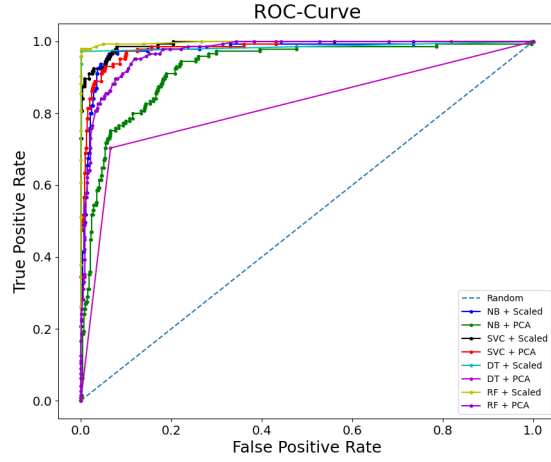
Figure 3: Receiver operating characteristic curve for the eight different models. All algorithms performed better on the standardized data than their respective counterpart using the components explaining at least 95% of the variance contained by the data of the principal component analysis.
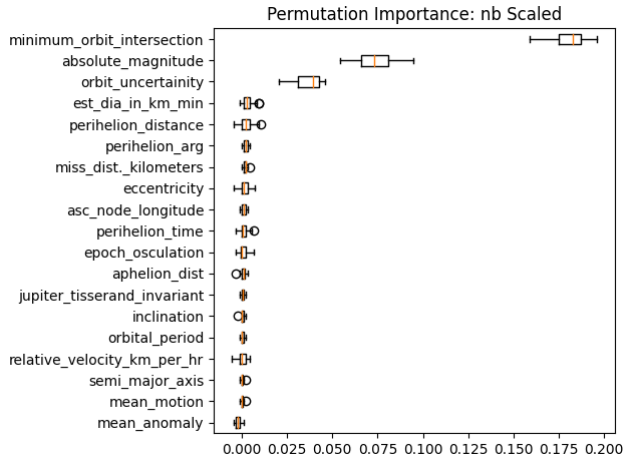

Figure 4: Naive Bayes feature importance.

4 DISCUSSION

Decision tree and random forest algorithm showed similar performances due to the nature of the random forest algorithm and the already stable, high-performing results of the individual decision tree. Given the already well performing decision tree it is unlikely that an ensemble of decision trees will perform significantly better. It is more likely, that the random forest performs slightly worse due to the inherent variation of the algorithm. Models trained on standardized data performed better than the models trained on PCA-transformed data. This could stem from the fact that PCA does not consider the relationship between the independent variables with the target variable. In fact, PCA will treat components
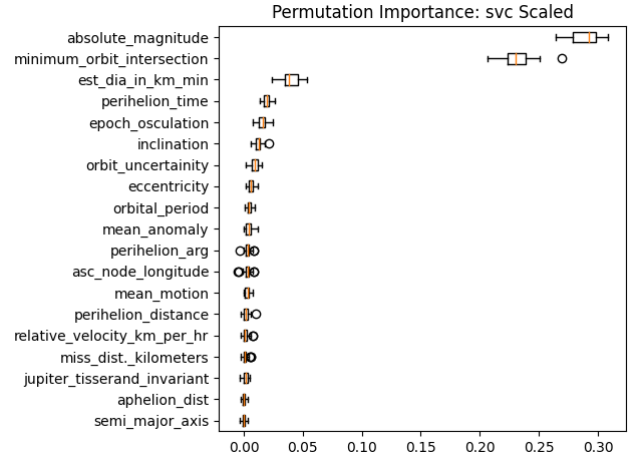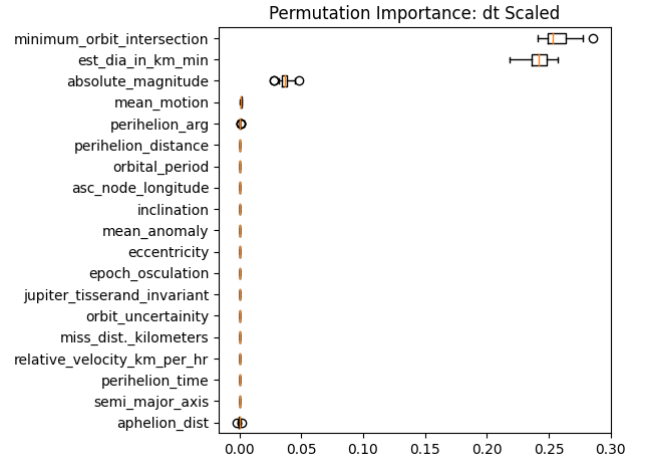

Figure 5: SVC feature importance.

Figure 6: Decision tree feature importance.

with the highest variances as its features but said features may be insignificant for the target variable. As a result, PCA could create many meaningless features and purge actually useful ones from the feature space.

Feature importance analysis showed that 3 of our 4 algorithms favored minimum orbit intersection as the most important feature. Minimum orbit intersection distance measures any possible near approaches and collision risks between astronomical objects [9] and is considered an important calculation to use when determining if any object is at risk for collision with the Earth. Absolute magnitude also featured prominently with one algorithm (SVC), and appeared among the top three for the others. NASA defines this as "the visual magnitude an observer would record if the asteroid were placed 1 Astronomical Unit (au) away", and assists in measuring the diameter of any NEO. Interestingly, two algorithms (DT and RF) also favored the estimated diameter in kilometers as important; though it may be that this feature is essentially the same as absolute magnitude.
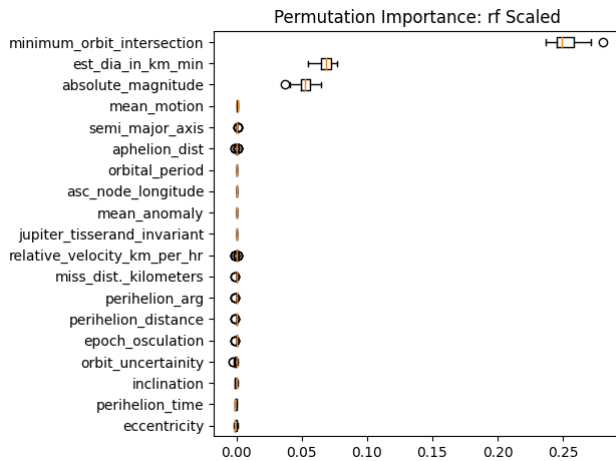
Figure 7: Random forest feature importance.

Also of note is that only two classifiers (SVC and NB) favored orbit uncertainty as significant.

Permutation scores do indicate the relative predictive power of a feature for a model, but the scores are only useful in this context, and features with high scores should only be considered in that context. Additionally, feature importance is not the same as statistical inference, and give us no information about the nature of the relationship of that feature (e.g., linear, etc.).

When we first began this experiment, we aimed to also include the distributed gradient boosting library, XGBoost[1], among the classifiers to test. However, after running into significant difficulty with getting this library to work properly (namely, memory seemed to be a significant issue), we chose to test the Random Forest classifier instead. Still, future research may explore XGBoost as another possibility for this type of classification.

5 CONCLUSION

In this study, we demonstrated that "off-the-shelf" machine learning algorithms can be used to identify the hazardous nature of extraterrestrial objects. Based on the performance metrics F1-score, precision, and recall, the decision tree slightly outperforms the random forest algorithm. However, in cases that requires to have a flexible threshold for binary classification one may want to use the random forest. Support vector machine algorithm seems to be the most robust among the four evaluated algorithm and may be considered in the future if the input data is varying frequently. The use of principal component analysis to accomplish feature reduction, reduced the performance of all algorithms, and is therefore not recommended as a pre-processing step.

REFERENCES

[1] Luis W Alvarez, Walter Alvarez, Frank Asaro, and Helen V Michel. 1980. Extraterrestrial cause for the Cretaceous-Tertiary extinction. *Science* 208, 4448 (1980), 1095–1108.
[2] Space Studies Board, National Research Council, et al. 2010. *Defending planet earth: Near-Earth-Object surveys and hazard mitigation strategies.* National Academies Press.
[3] William F Bottke Jr, Alessandro Morbidelli, Robert Jedicke, Jean-Marc Petit, Harold F Levison, Patrick Michel, and Travis S Metcalfe. 2002. Debiased orbital and absolute magnitude distribution of the near-Earth objects. *Icarus* 156, 2 (2002), 399–433.
[4] L. Breiman. 2001. Random Forests. *Machine Learning* 45, 1 (2001), 5–32.
[5] Ronan Collobert and Samy Bengio. 2004. Links between perceptrons, MLPs and SVMs. In *Proceedings of the twenty-first international conference on Machine learning.* ACM, 23.
[6] Jerome Fan, Suneel Upadhye, and Andrew Worster. 2006. Understanding receiver operating characteristic (ROC) curves. *Canadian Journal of Emergency Medicine* 8, 1 (2006), 19–20.
[7] Michael J Gaffey, Paul A Abell, and Paul S Hardersen. 2006. COMPOSITIONAL AND PHYSICAL CHARACTERIZATIONS OF NEOS FROM VNIR SPECTROSCOPY. (2006).
[8] Sarah Greenstreet, Henry Ngo, and Brett Gladman. 2012. The orbital distribution of near-Earth objects inside Earth's orbit. *Icarus* 217, 1 (2012), 355–366.
[9] José M Hedo, Manuel Ruíz, and Jesús Peláez. 2018. On the minimum orbital intersection distance computation: a new effective method. *Monthly Notices of the Royal Astronomical Society* 479, 3 (2018), 3288–3299.
[10] R. Olshen L. Breiman, J. Friedman and C. Stone. 1984. *Classification and Regression Trees.* Wadsworth.
[11] Amy Mainzer, T Grav, J Bauer, J Masiero, RS McMillan, RM Cutri, R Walker, E Wright, P Eisenhardt, DJ Tholen, et al. 2011. NEOWISE observations of near-Earth objects: preliminary results. *The Astrophysical Journal* 743, 2 (2011), 156.
[12] Kevin P Murphy et al. 2006. Naive bayes classifiers. *University of British Columbia* 18, 60 (2006), 1–8.
[13] NASA. [n. d.]. Near-Earth Object Observations Program. https://www.nasa.gov/planetarydefense/neoo.
[14] NASA. [n. d.]. Torino Impact Scale. https://web.archive.org/web/20070224184143/http://impact.arc.nasa.gov/torino.cfm.
[15] Carrie R Nugent, John Dailey, Roc M Cutri, Frank J Masci, and Amy K Mainzer. 2017. Machine learning and next-generation asteroid surveys. In *AAS/Division for Planetary Sciences Meeting Abstracts# 49*, Vol. 49.
[16] Alfred Romer. 1984. Halley's comet. *The Physics Teacher* 22, 8 (1984), 488–493.
[17] A. M. Tarano, J. Gee, L. Wheeler, S. Close, and D. Mathias. 2020. Automating the Inference of Asteroid Physical Properties and Motion. In *AGU Fall Meeting Abstracts*, Vol. 2020. Article P008-02, P008-02 pages.
[18] Vladimir Vapnik. 2013. *The nature of statistical learning theory.* Springer science & business media.

---

[1]https://xgboost.readthedocs.io/en/stable/