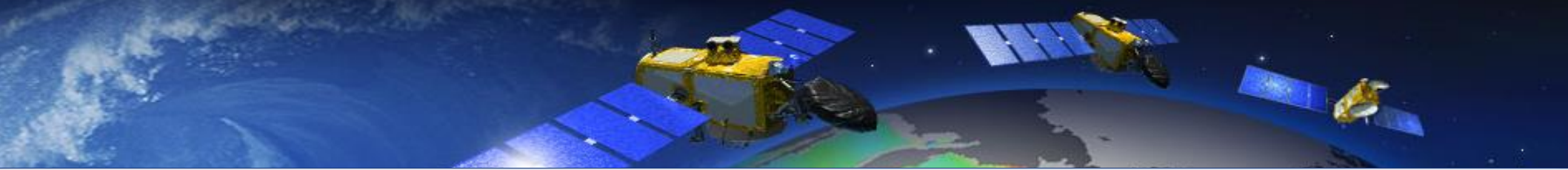


# Orbit Mechanics Tutorial

Hao Ye

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Directed by Prof. Urs Hugentobler

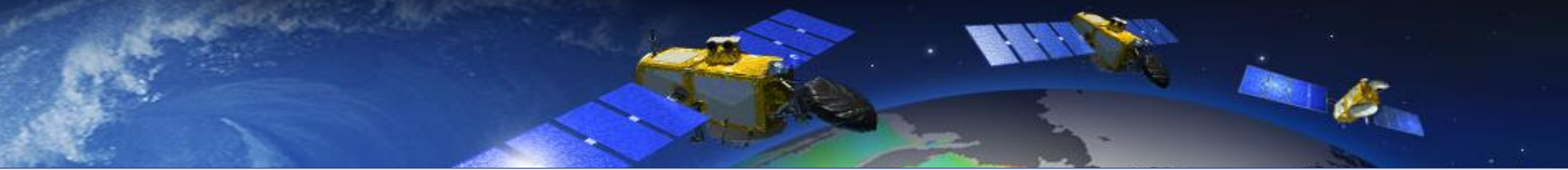


## Three tutorials in total

**1<sup>st</sup> --Keplerian Orbits (analytical way)**

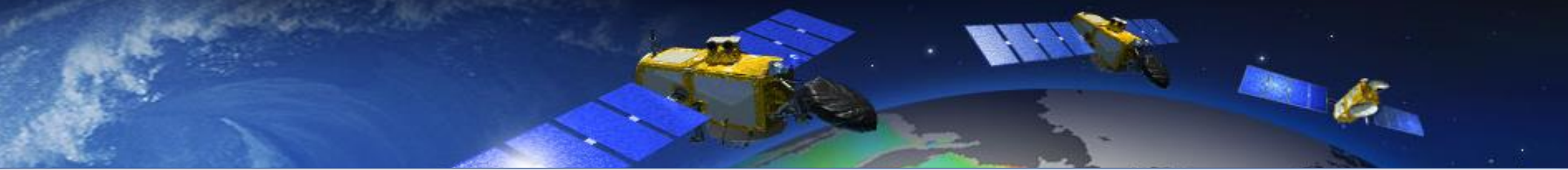
**2<sup>rd</sup> --Numerical Integration of Satellite Orbits (numerical way)**

**3<sup>rd</sup> --Orbits Integration with Different Force Models(software)**



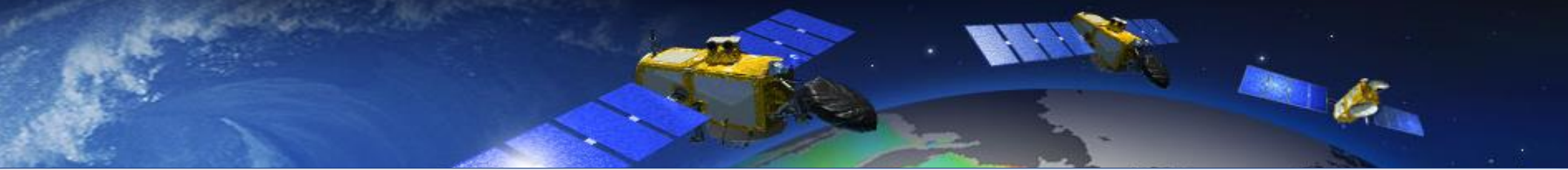
## Six Keplerian orbital elements

- semimajor axis  $a$
  - numerical eccentricity  $e$
  - Inclination  $i$
  - Right ascension  $\Omega$  of the ascending node
  - Argument of perigee  $\omega$
  - Perigee passing time  $T_0$
- } **Shape of orbit**
- } **Orientation of orbital plane**



## Six Keplerian orbital elements

Satellite	$a$ [km]	$e$	$i$ [deg]	$\Omega$ [deg]	$\omega$ [deg]
GOCE	6629	0.004	96.6	257.7	144.2
GPS	26560	0.01	55	60	0
MOLNIYA	26554	0.7	63	245	270
GEO	Geostationary	0	0	0	0
MICHIBIKI	Geosynchronous	0.075	41	195	270



## Main tasks

### 1. Keplerian elements $\implies$ Orbits in 2D plane

- 2D orbit
- Mean anomaly, eccentric anomaly and true anomaly

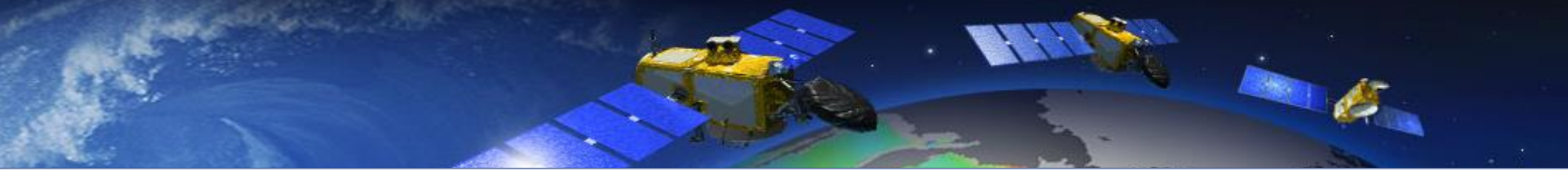
### 2. 2D plane $\implies$ 3D trajectory in space-fixed system

### 3. Orbit in space-fixed system $\implies$ Orbit in earth-fixed system

- 3D orbit
- Tracks on the ground

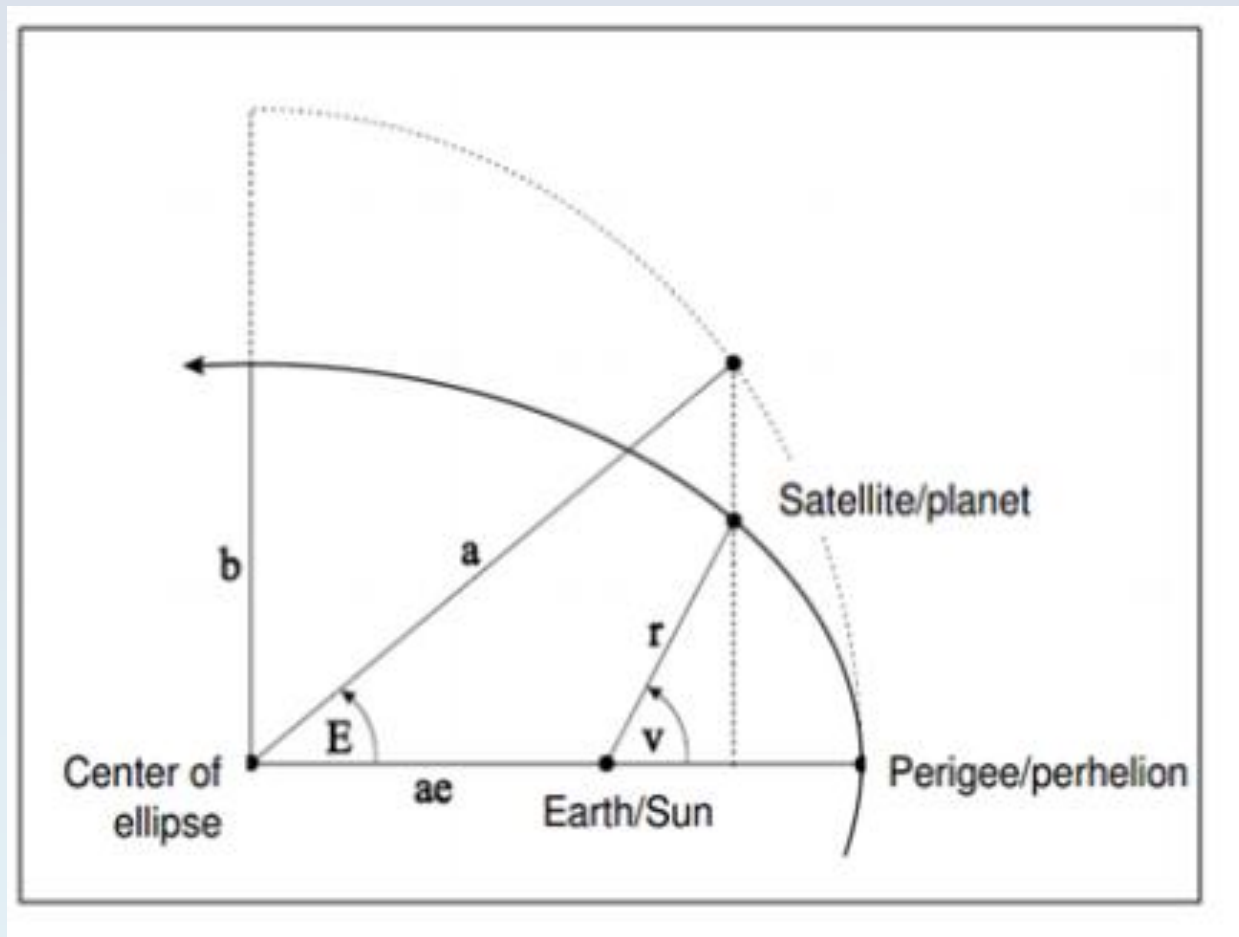
### 4. Orbit in earth-fixed system $\implies$ Orbit in topo-centric system

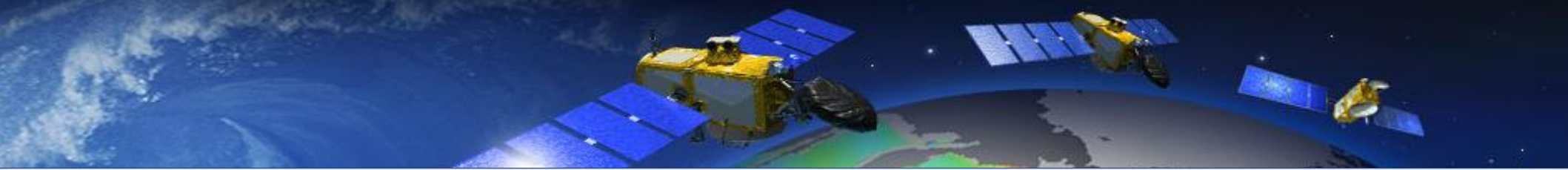
- Azimuth and elevation orbit
- Judge visibility



# Task I

Keplerian elements  $\Rightarrow$  Orbits in 2D plane





# Task I

Keplerian elements  $\Longrightarrow$  Orbits in 2D plane

$$n = \sqrt{\frac{GM}{a^3}}$$

$$M(t) = n \cdot (t - T_0)$$

Mean anomaly



Kepler's equation:  $M = E - e \sin E \quad \Delta E_i < 10^{-6}$

Eccentric anomaly



$$r = a(1 - e \cos E)$$

$$\tan \frac{\nu}{2} = \sqrt{\frac{1+e}{1-e}} \tan \frac{E}{2}$$

Radius  
True anomaly

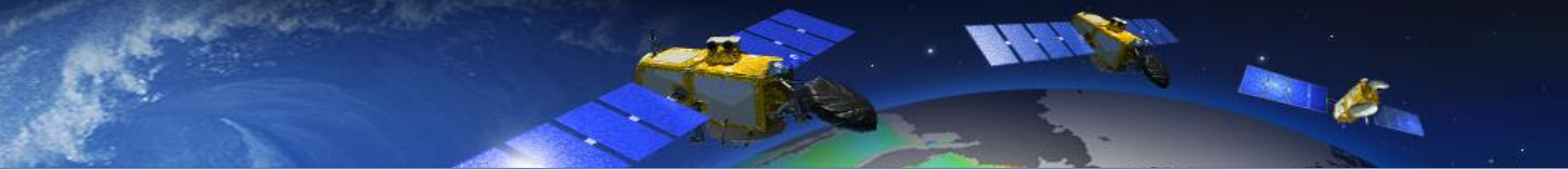


$$x = r \cos \nu$$

$$y = r \sin \nu$$

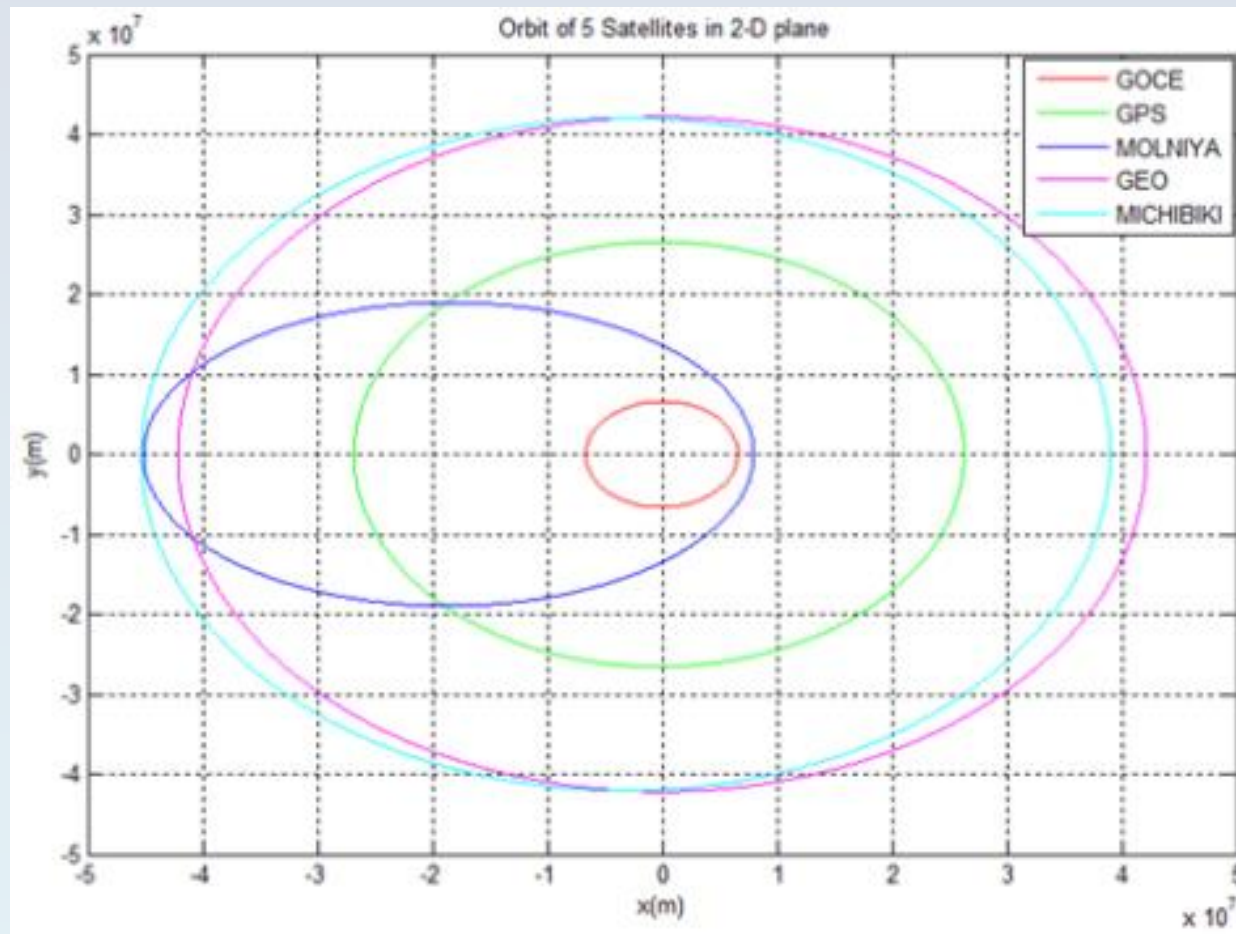
2D coordinates



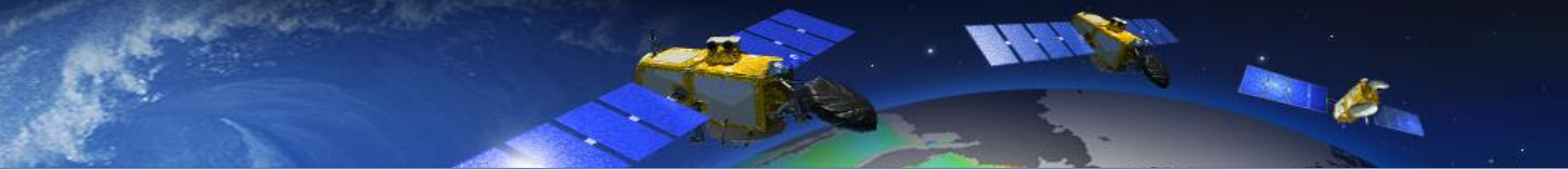


# Task I

Keplerian elements  $\Rightarrow$  Orbits in 2D plane

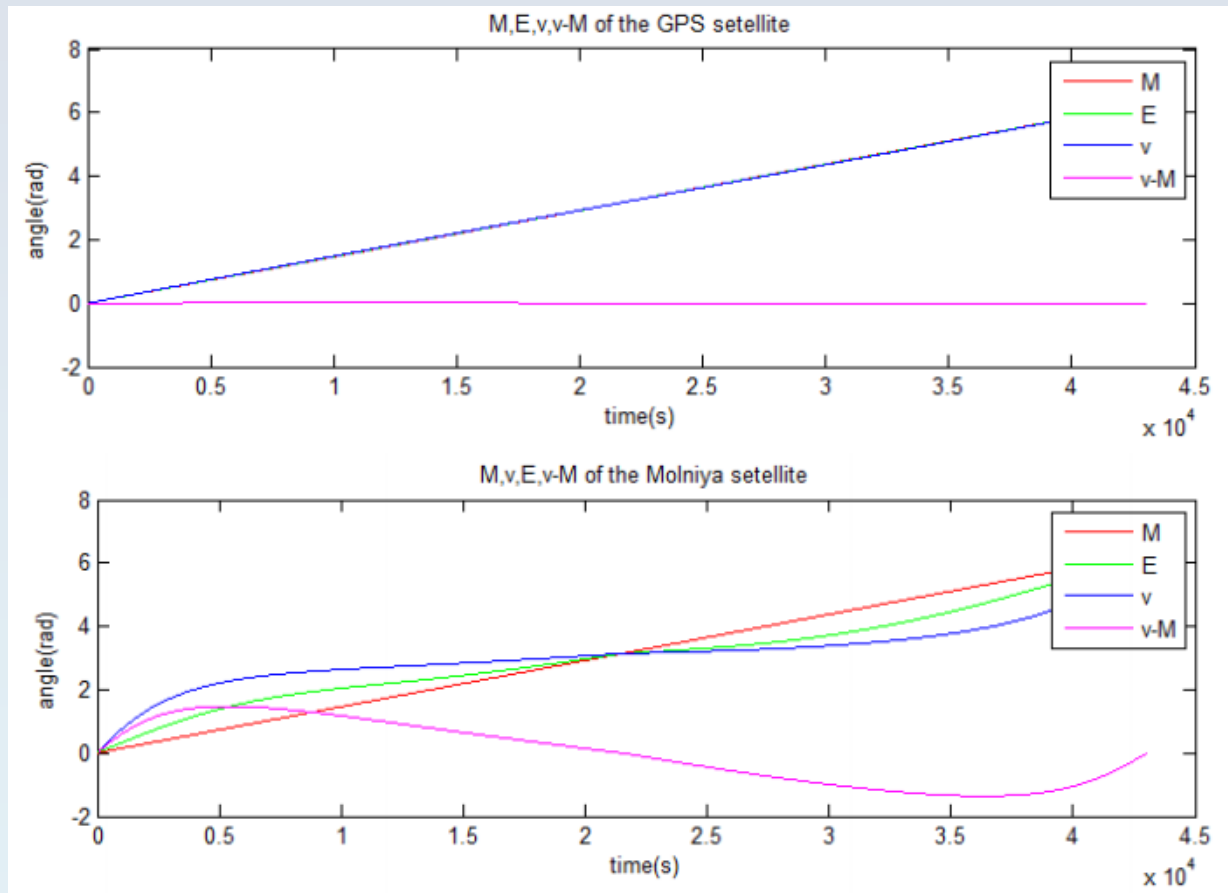




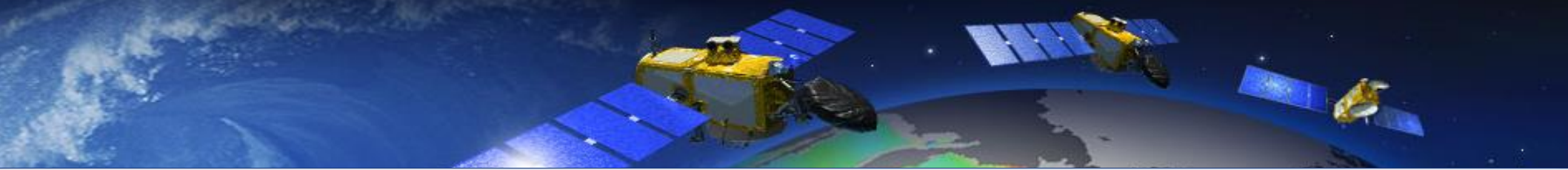


# Task I

Keplerian elements  $\Rightarrow$  Orbits in 2D plane

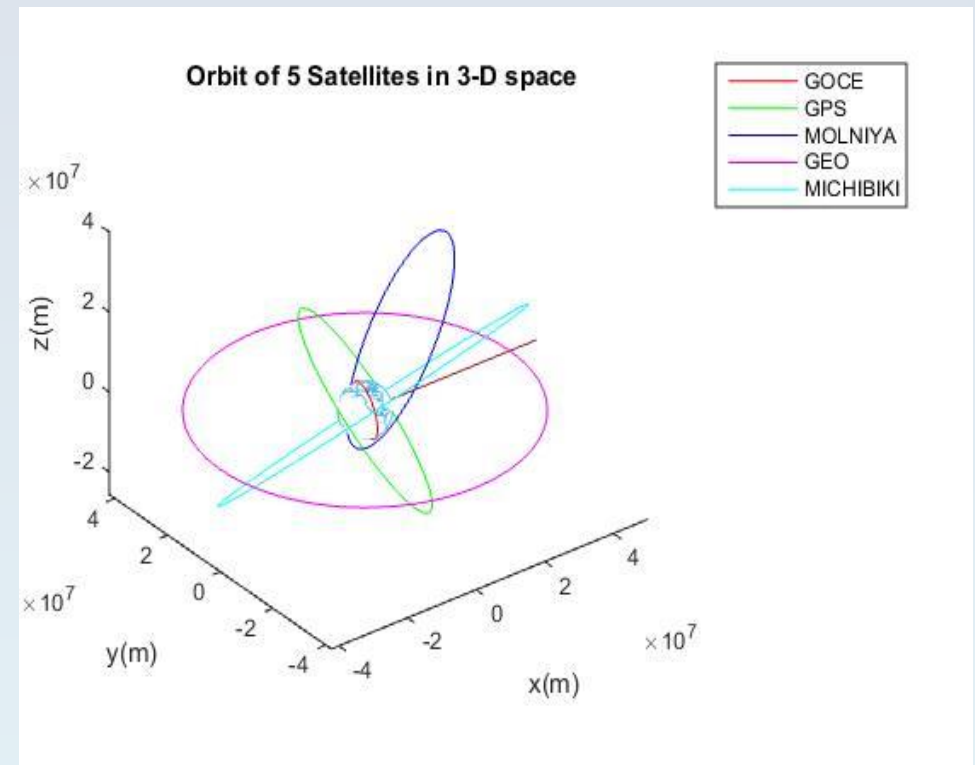
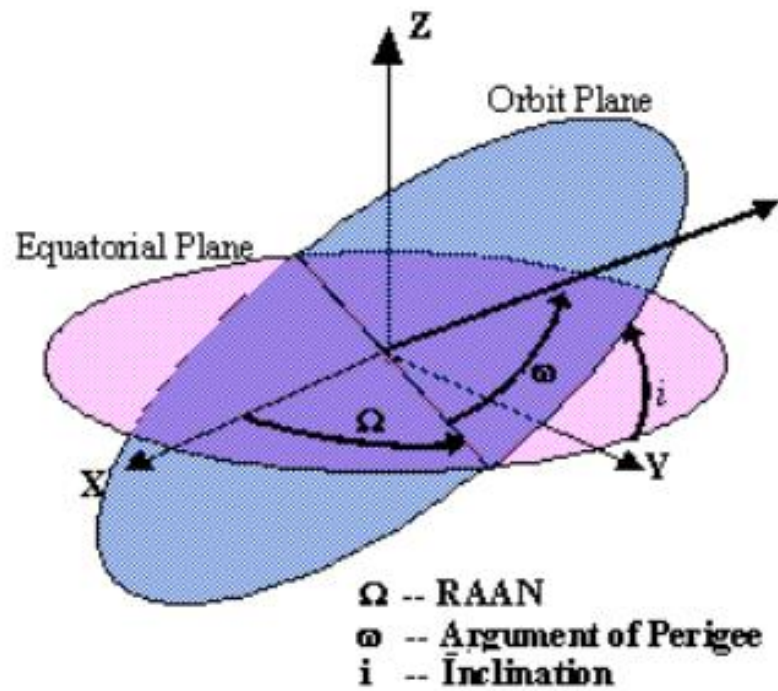


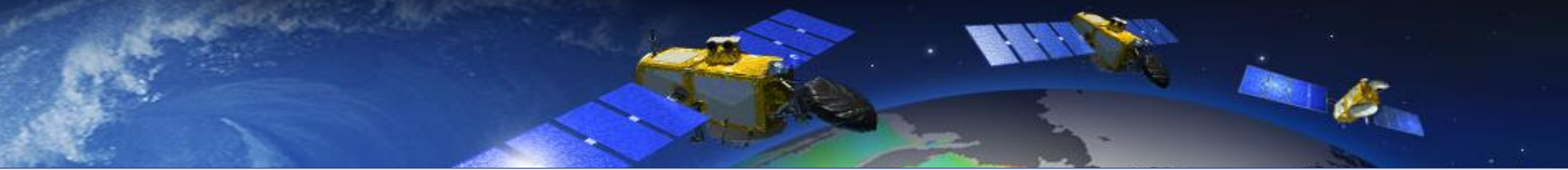
Mean anomaly  
Eccentric anomaly  
True anomaly



## Task II

2D plane  $\Rightarrow$  3D plane in space-fixed system





## Task II

2D plane  $\Rightarrow$  3D plane in space-fixed system

$$\vec{r}_b = r(\cos \nu, \sin \nu, 0)$$

$$\dot{\vec{r}}_b = \sqrt{\frac{GM}{a(1-e^2)}}(-\sin \nu, e + \cos \nu, 0)$$

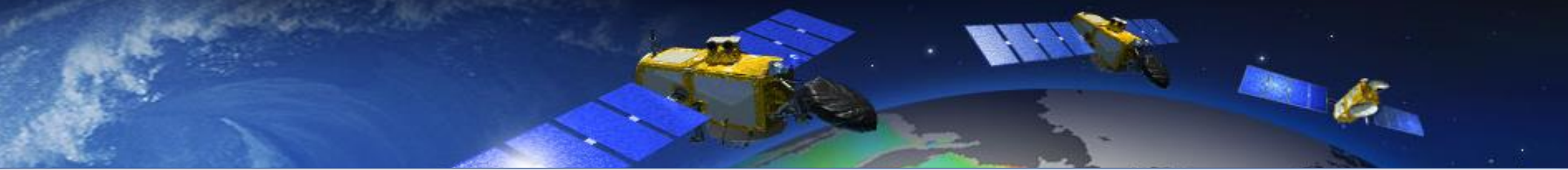
rotation



$$\vec{r} = R_3(-\Omega)R_1(-i)R_3(-\omega)\vec{r}_b$$

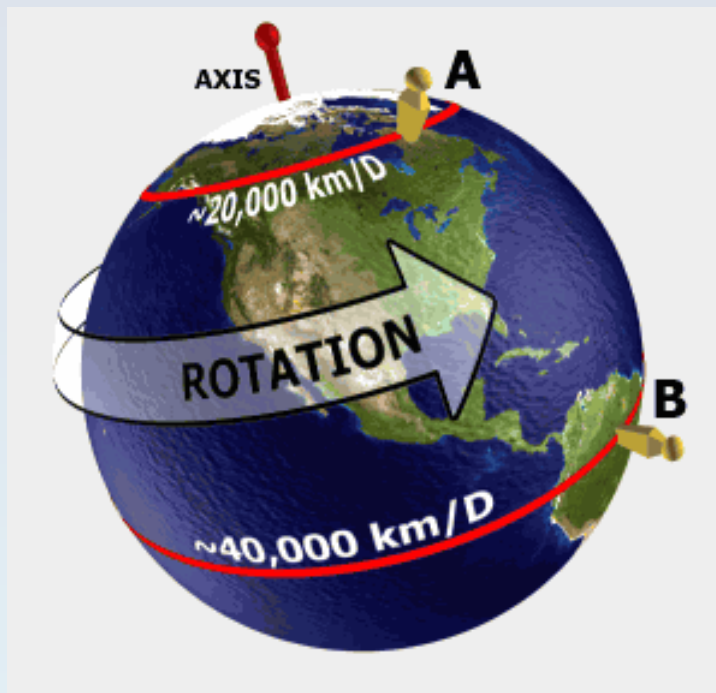
$$\dot{\vec{r}} = R_3(-\Omega)R_1(-i)R_3(-\omega)\dot{\vec{r}}_b$$

$$R_1(\theta) = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & \sin \theta \\ 0 & -\sin \theta & \cos \theta \end{bmatrix} \quad R_2(\theta) = \begin{bmatrix} \cos \theta & 0 & -\sin \theta \\ 0 & 1 & 0 \\ \sin \theta & 0 & \cos \theta \end{bmatrix} \quad R_3(\theta) = \begin{bmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix}$$



## Task III

Space-fixed system  $\Longrightarrow$  Earth-fixed system

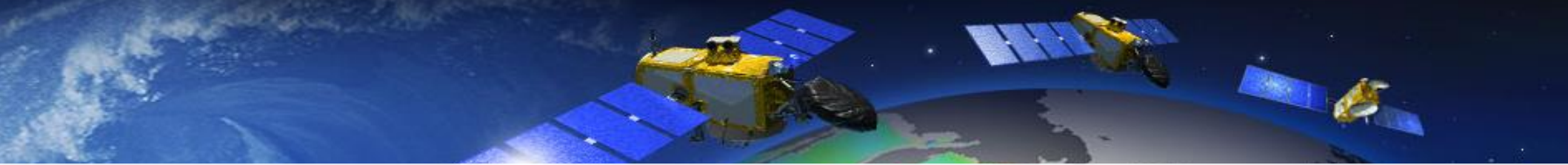


Rotating rate:  $\dot{\Omega}_E = \frac{2\pi}{86164s}$

Rotating angle:  $\theta_0(t) = \dot{\Omega}_E t$

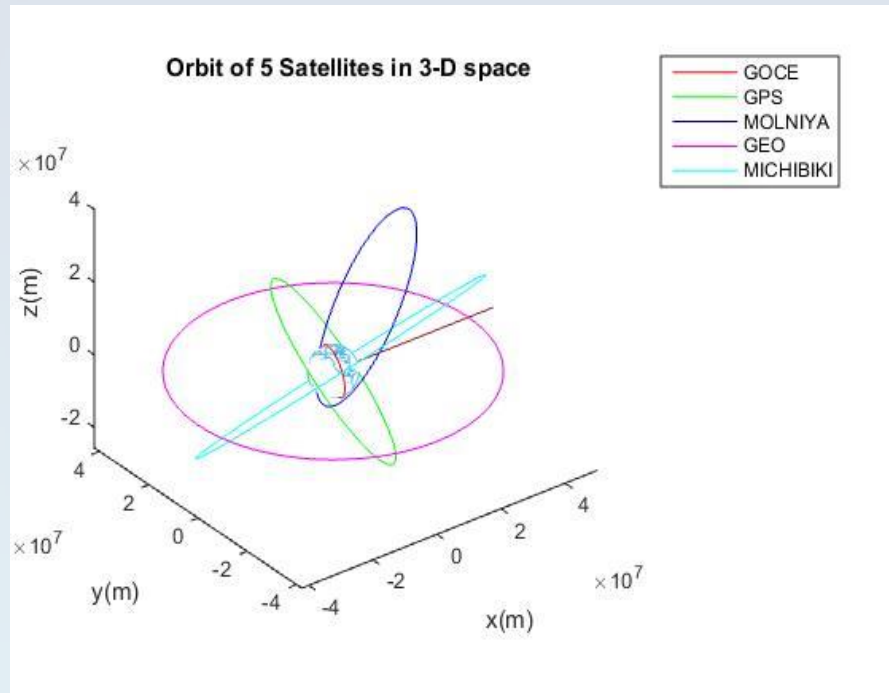
$$\vec{r}_{Earth-fixed}(t) = R_3(\theta_0(t)) \cdot \vec{r}_{Space-fixed}(t)$$

$$R_3(\theta) = \begin{bmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

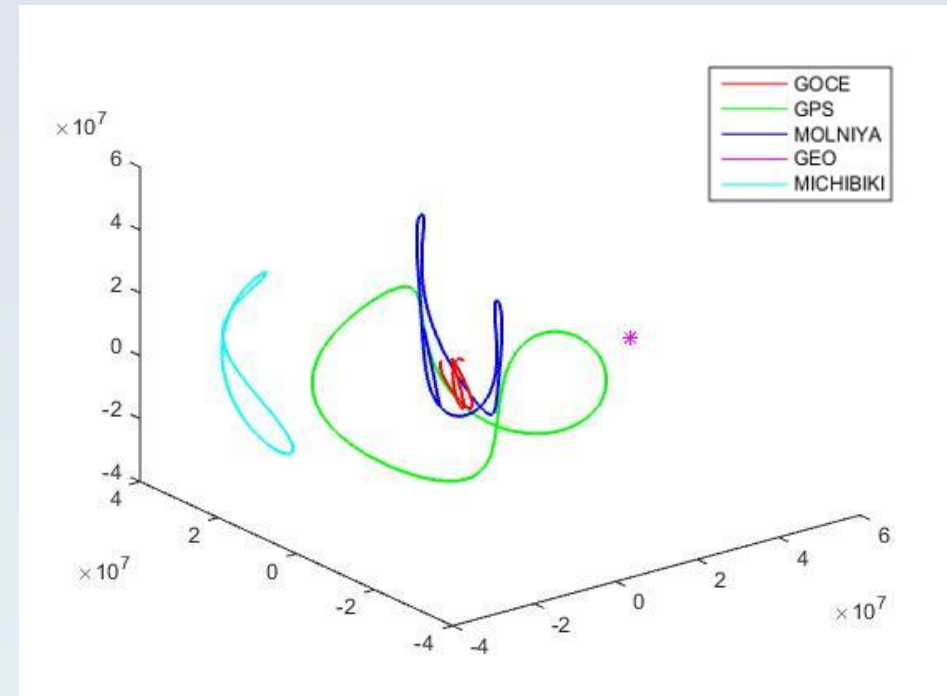
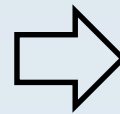


# Task III

Space-fixed system  $\Rightarrow$  Earth-fixed system

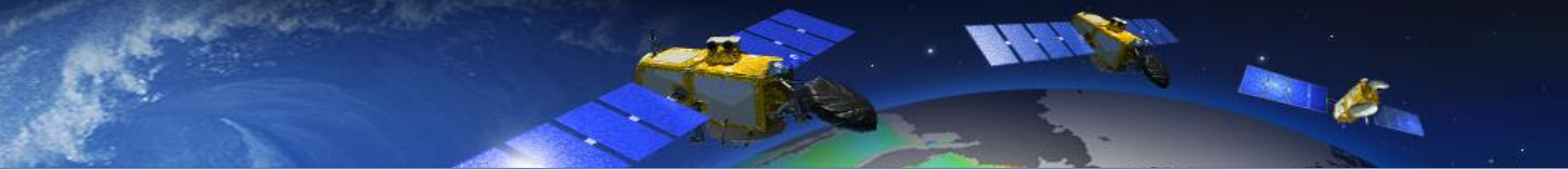


Space-fixed system



Earth-fixed system

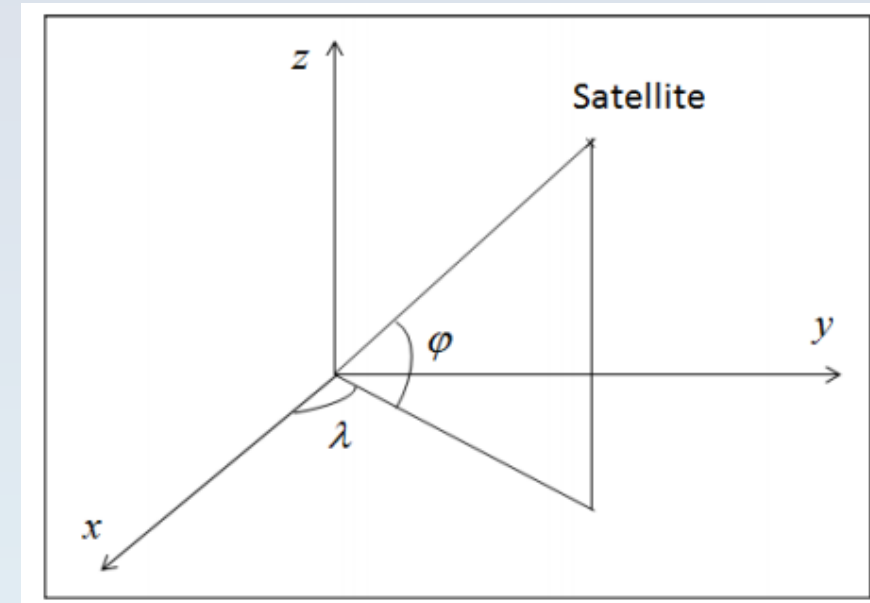




## Task III

Space-fixed system  $\Rightarrow$  Earth-fixed system

Ground-tracks on the Earth surface



Latitude



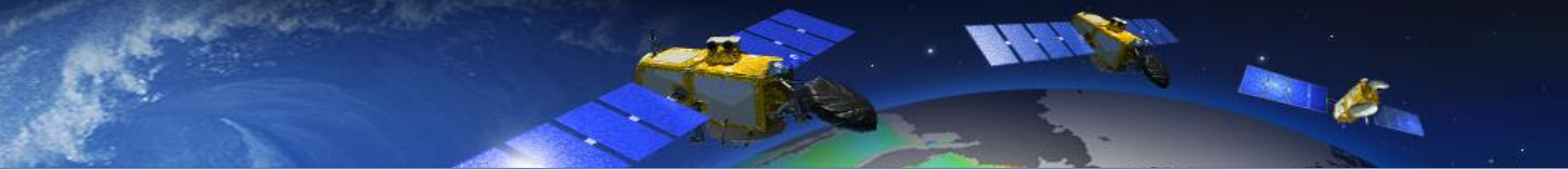
$$\tan \lambda = \frac{y_{Earth-fixed}}{x_{Earth-fixed}}$$

Longitude



$$\tan \psi = \frac{z_{Earth-fixed}}{\sqrt{x_{Earth-fixed}^2 + y_{Earth-fixed}^2}}$$

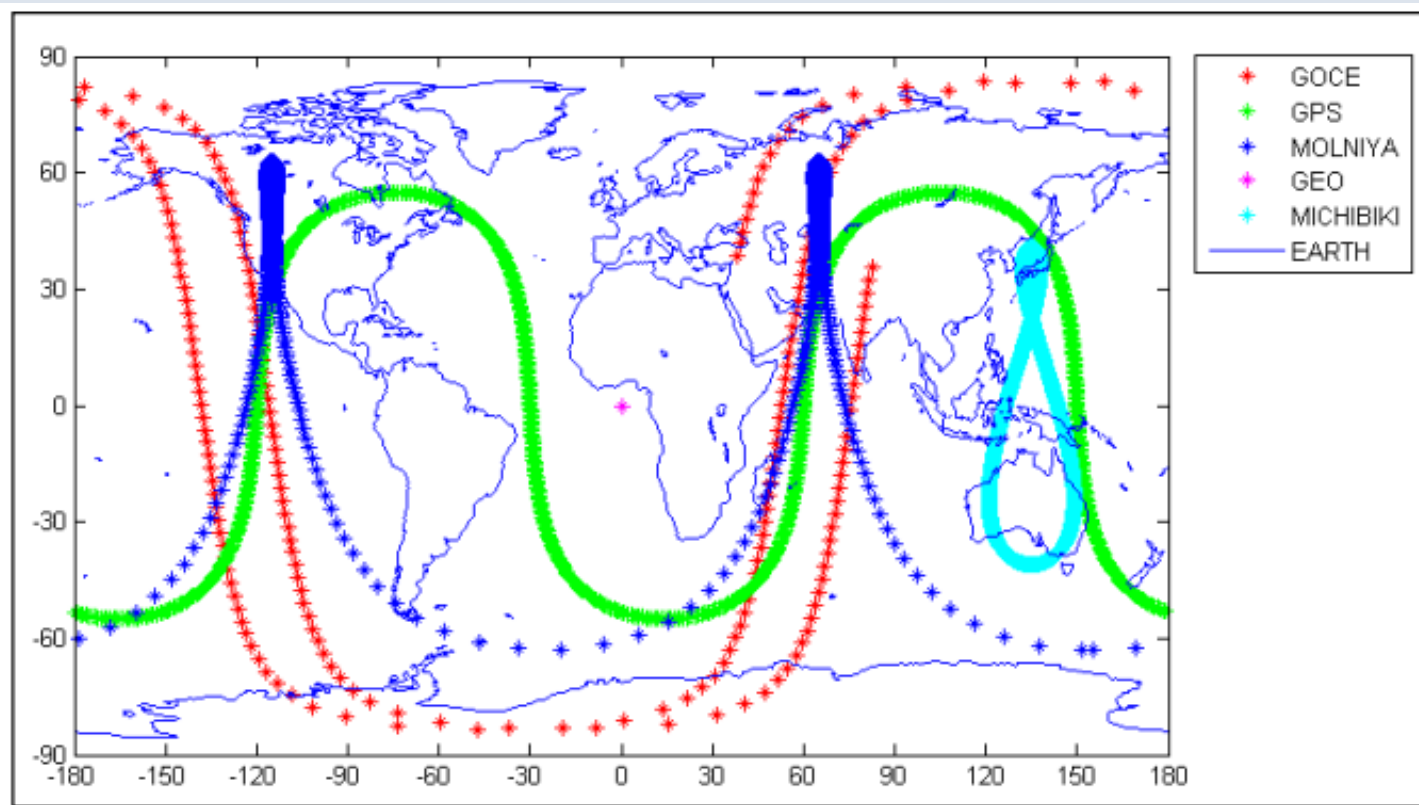


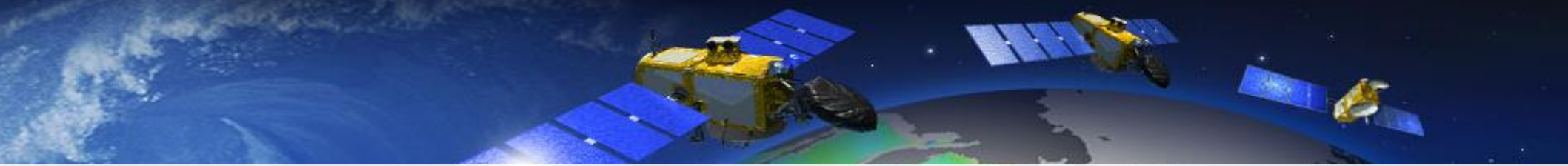


## Task III

Space-fixed system  $\Rightarrow$  Earth-fixed system

Ground-tracks on the Earth surface





## Task IV

Earth-fixed system  $\Rightarrow$  Topocentric system

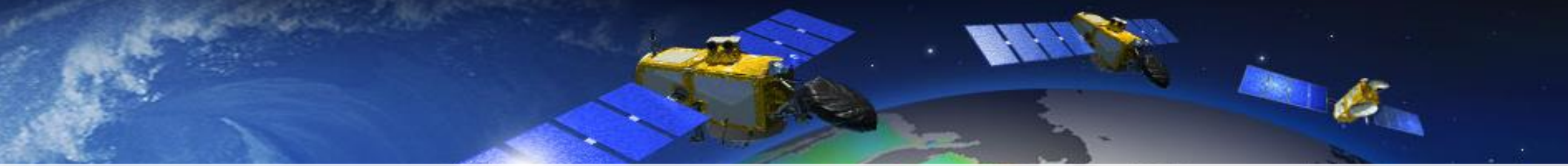


Wettzell

$(4075.53022, 931.78130, 4801.61819)_T$

Key point:

You should move the center of the coordinate system to the Wettzell



## Task IV

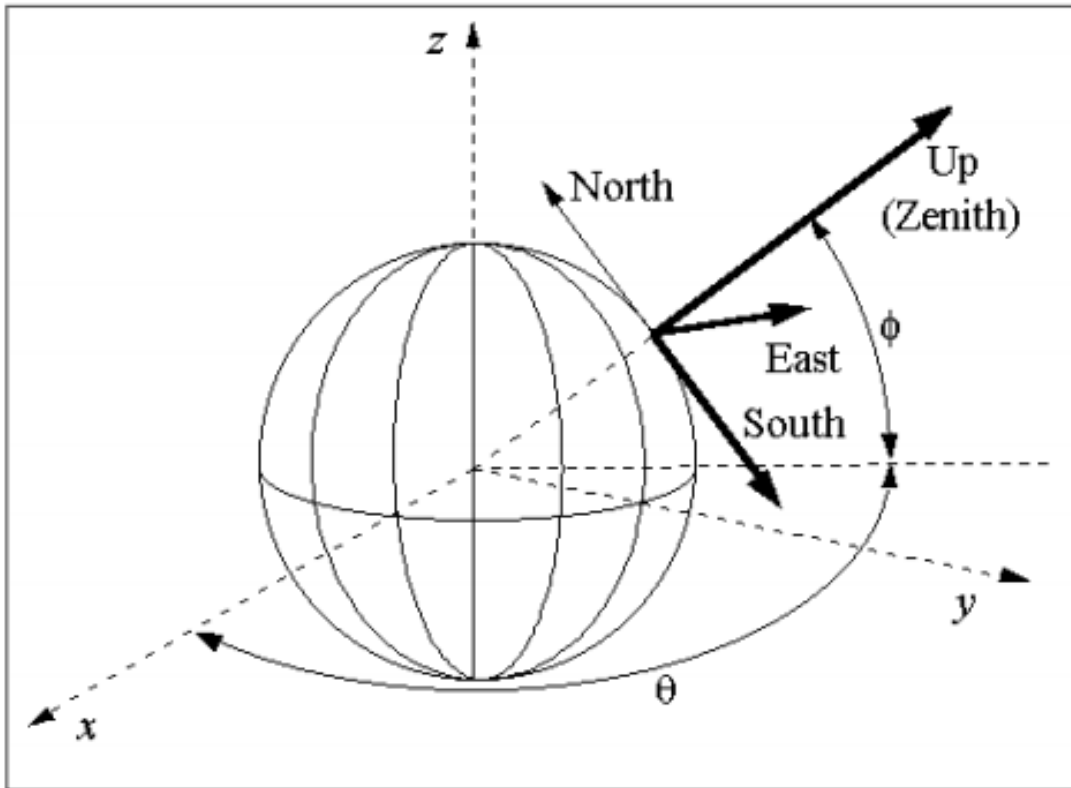
Earth-fixed system  $\Rightarrow$  Topocentric system

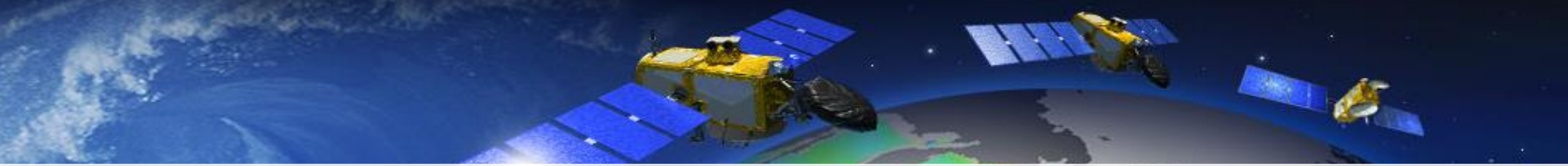
Transformation

$$\mathbf{d} = \mathbf{R} - \mathbf{r}$$

Rotation

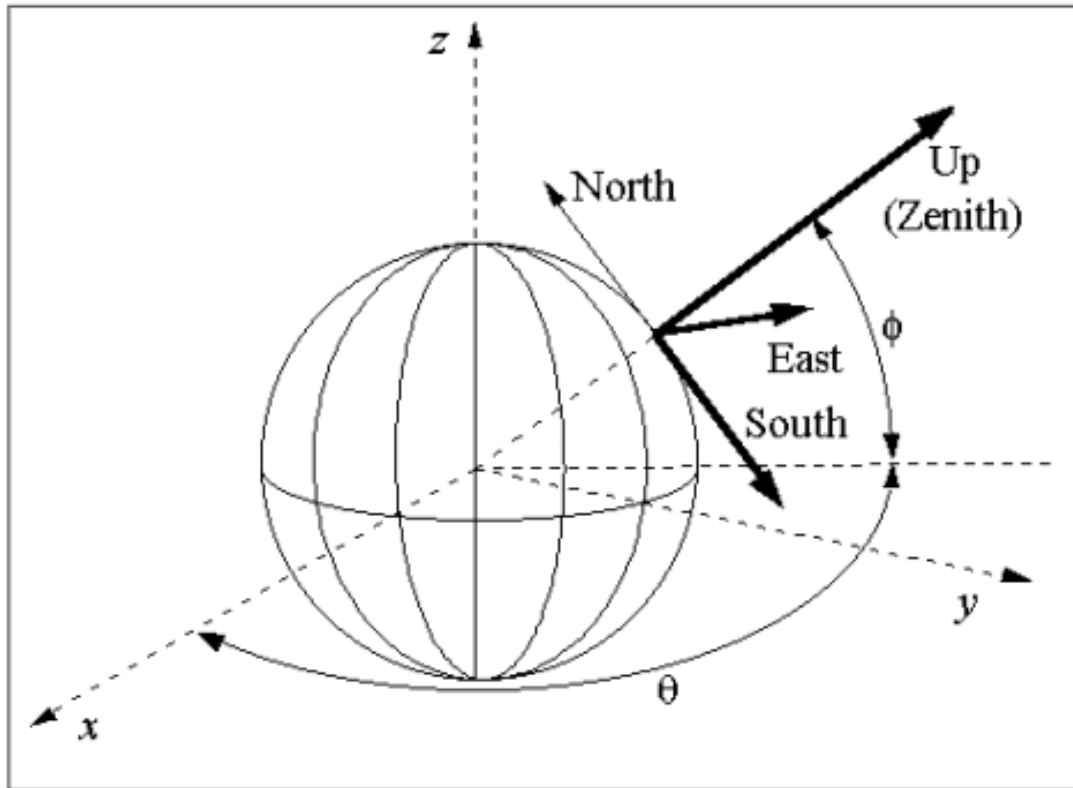
$$\mathbf{d}_{\text{hori}} = Q_1 R_2(90 - \varphi) R_3(\lambda) \mathbf{d}$$





## Task IV

Earth-fixed system  $\Rightarrow$  Topocentric system



Transformation

$$\mathbf{d} = \mathbf{R} - \mathbf{r}$$

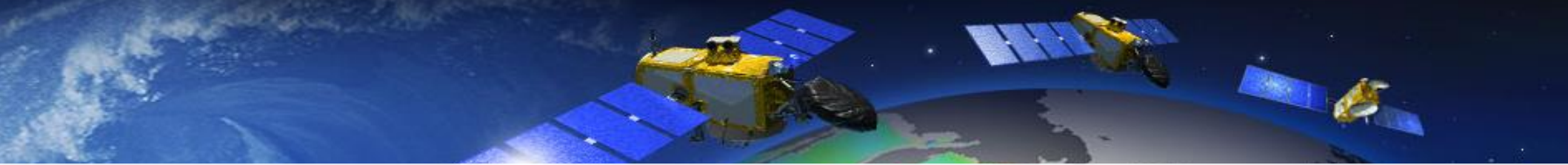
Rotation

$$\mathbf{d}_{\text{hori}} = \mathbf{Q}_1 \mathbf{R}_2(90 - \varphi) \mathbf{R}_3(\lambda) \mathbf{d}$$

Notice:

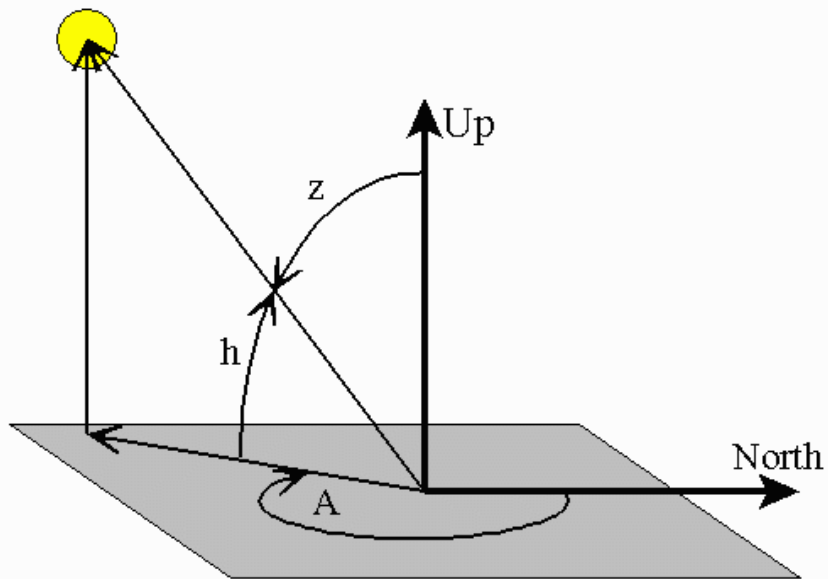
$$\mathbf{Q}_1 = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Left hand  $\rightarrow$  right hand



## Task IV

Earth-fixed system  $\Rightarrow$  Topocentric system



$h$  = elevation  
angle, measured  
up from horizon

$z$  = zenith angle,  
measured from  
vertical

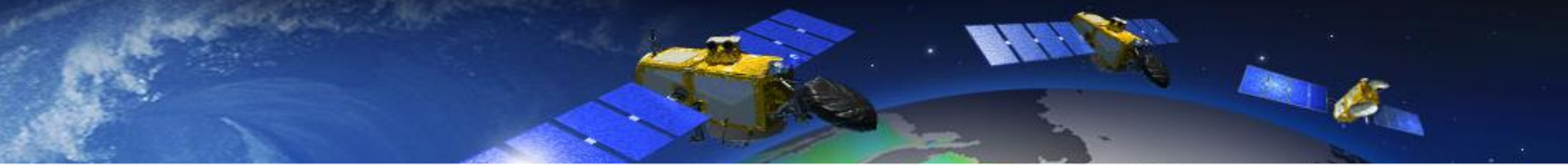
$A$  = Azimuth angle,  
measured clockwise  
from North

$$\tan A = \frac{y}{x}$$

$$\cos Z = \frac{z}{\sqrt{x^2 + y^2 + z^2}}$$

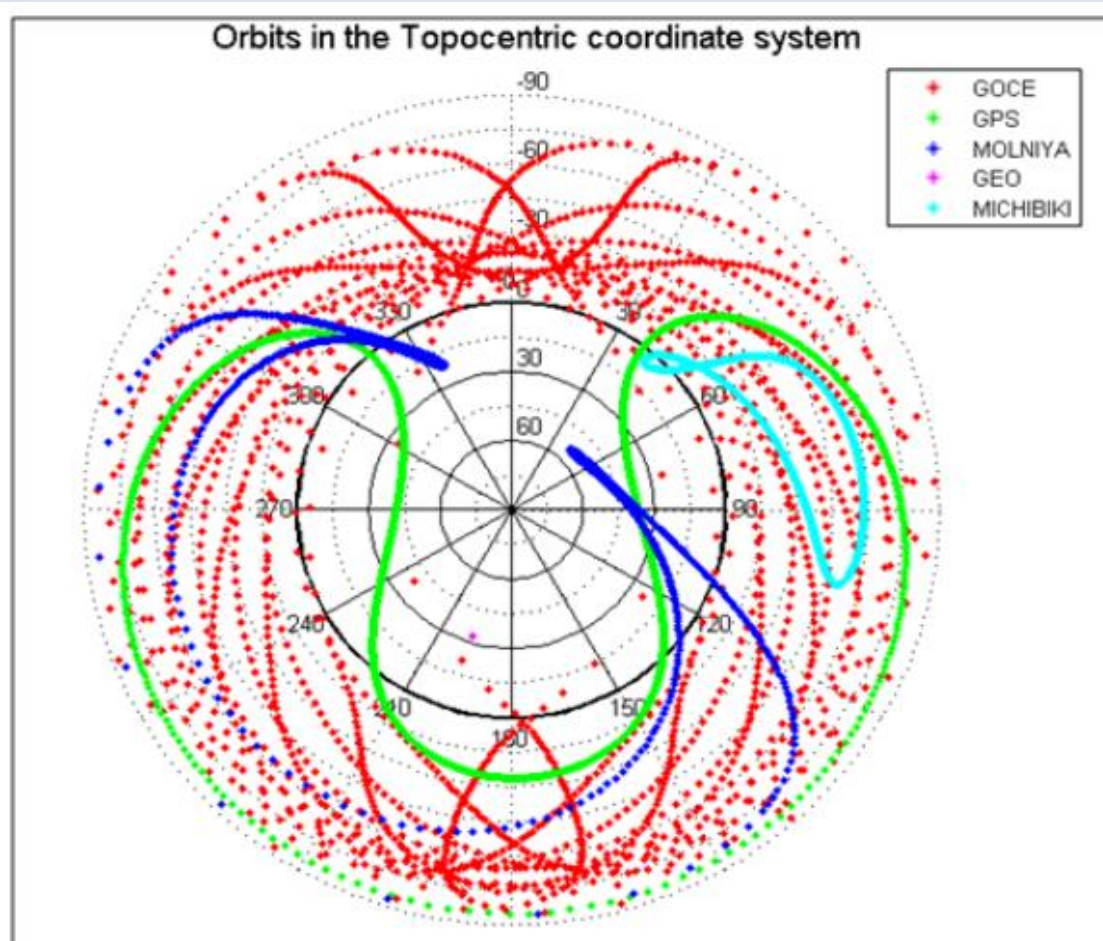
$$E = \frac{\pi}{2} - z$$



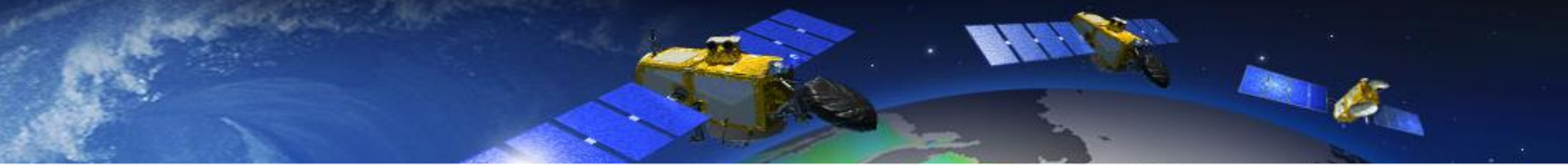


## Task IV

Earth-fixed system  $\Rightarrow$  Topocentric system

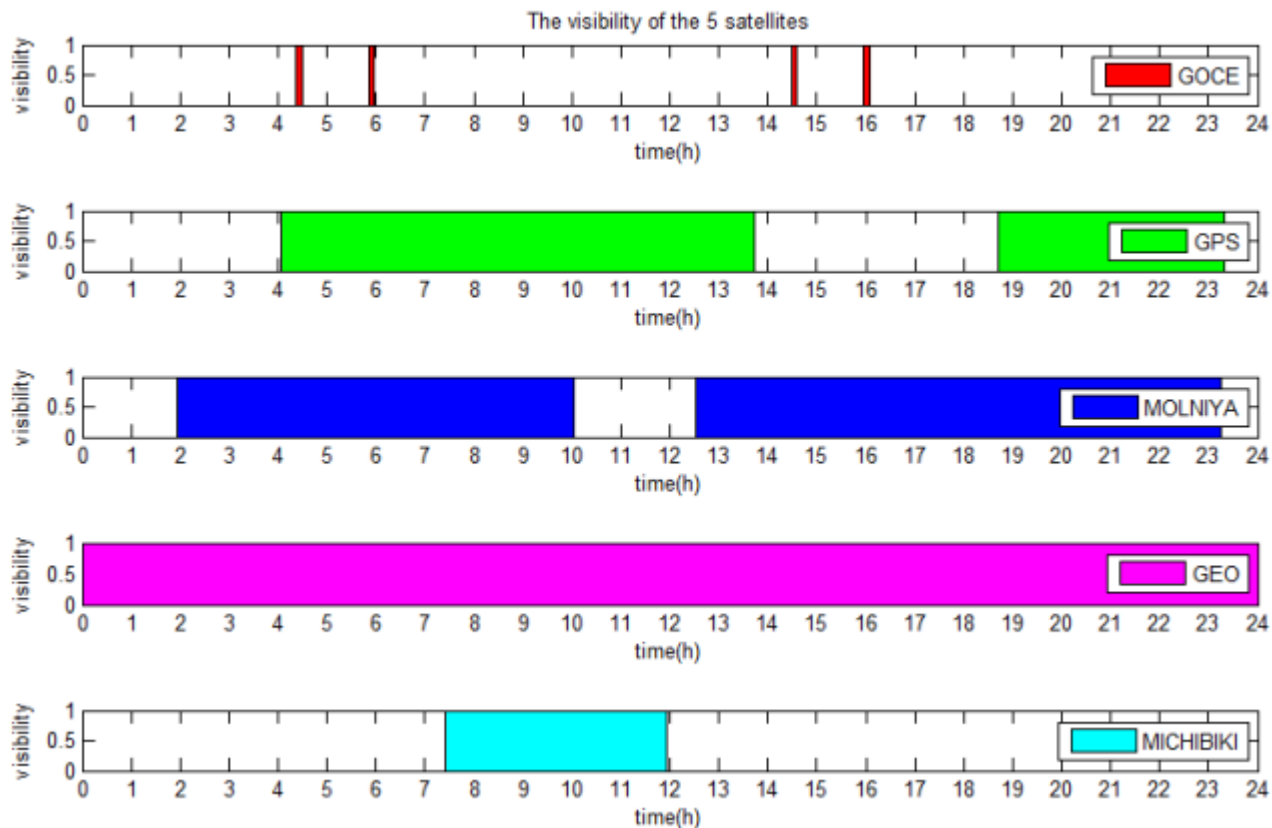






## Task IV

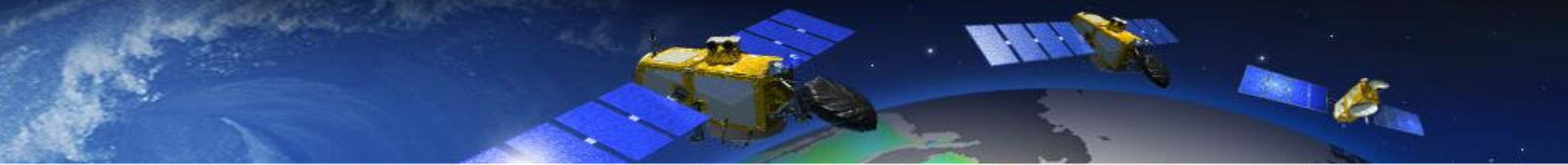
Earth-fixed system  $\Rightarrow$  Topocentric system



## Judge the Visibility

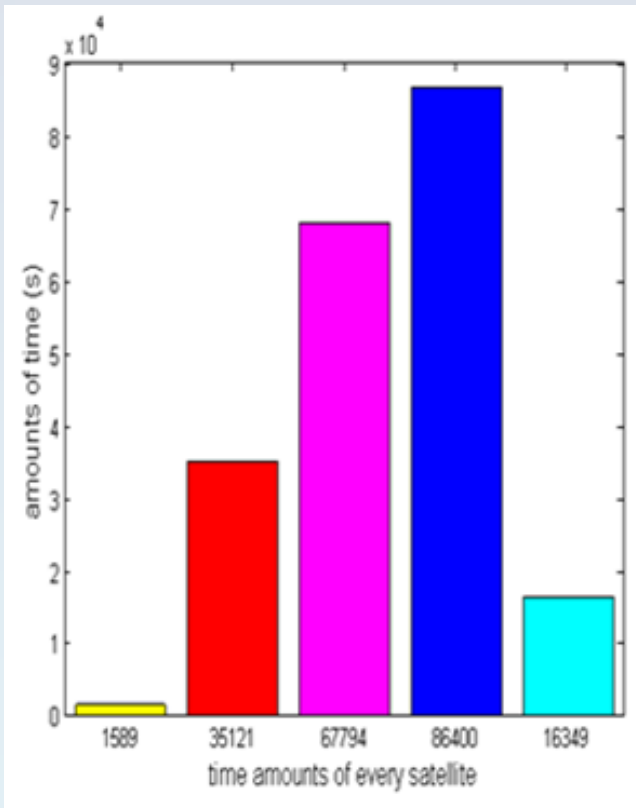
Rule:

Elevation angle  $> 0$  ----  $\rightarrow$  visible



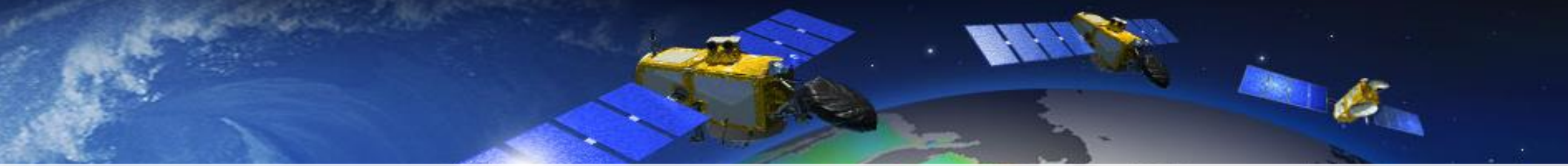
## Task IV

Earth-fixed system  $\implies$  Topocentric system



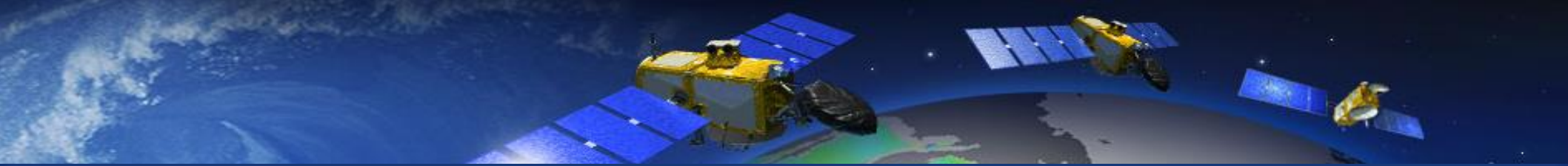
Satellites	GOCE	GPS	MOLNIYA	GEO	MICHIBIKI
Visibility time/s	1589	35121	67794	86400	16349
Percentage/%	1.84	40.65	78.47	100	18.96

Or some other formats you like.....



## Tips in programming

- Avoid using loop
- Avoid putting all the code in one function and trying to design the structure of your task
- Check your input parameters and typing errors
- Try to use the debugging of MATLAB and be patient
- No hesitation to get help from you tutor



**Thank you for your attention !**

