Homework # 1 - Observing Stars

Due: Tuesday, January 31 at 2:00pm

Reading: Lecture Notes, Pols chap 1, BOB chap 7 & 8.

Instructions: Homework must be highly legible, on white paper, and with multiple pages stapled. List at the top:

- Your name.
- Collaborators (if applicable).
- Estimated time spent to complete.

Homeworks are due at the beginning of class. Late homeworks will be marked down by 50% until the assignment is graded and returned, and will receive no credit after that. Clearly outline the process for solving the problem: partial credit will be given for presenting the correct steps to solve problem, even if you do not achieve the correct numerical answer.

Problems (3 questions, 20 total points):

- 1. Parallax is critically important for understanding stellar properties. But despite its simplicity (just geometry!), most stars in the sky have been inaccessible to reliable parallax measurements until very recently. (8 pts total)
 - (a) Can you reliably measure parallax to any stars with your eye? (2 pt)
 - (b) Galileo's telescope had a 3 cm aperture. How many stars could he resolve parallax for? (2 pt)
 - (c) The *Hipparcos* satellite was a dedicated astrometry and parallax experiment that operated from 1989-1993. It had an angular resolution of 1 mas (0.001"). How many stars over the entire sky could have parallax measured by *Hipparcos* with at least 10% precision? (*Clearly state your assumption about the density and distribution of stars around us.*) (2 pts)
 - (d) Gaia launched in 2013 and achieves an astrometric precision of 7 μ as $(7 \times 10^{-6}")$. It is still operating, and had a major data release in April 2018. How many stars over the entire sky can Gaia measure parallax with at least 10% precision? (Again, clearly state your assumptions.) (2 pts)
- 2. Let's use an H-R diagram to age-date two star clusters. Download the files cluster1.dat and cluster2.dat from the course HuskyCT page. These are the (B-V) colors and apparent V magnitudes for stars in two open clusters in the Milky Way, compiled from the WEBDA database (http://webda.physics.muni.cz). (8 pts total)
 - (a) Plot both in an H-R diagram of apparent V magnitude versus (B-V) color, labeling your axes and orienting the y-axis to be brightest at the top and faintest at the bottom. You can make two separate H-R diagrams, or color-code each cluster and plot both in the same diagram. (4 pts)
 - (b) What is the approximate age of each cluster? (2 pts)
 - (c) The two clusters are offset on the y-axis. Why? (1 pt)

- (d) One of the clusters has a rather broad main sequence in the H-R diagram, almost like parallel tracks. What's going on here? (1 pt)
- 3. An eclipsing binary star system has a period of 5 days. Two absorption-line components are observed with maximum radial velocities of 50 km/s and 100 km/s. (4 pts total)
 - (a) What is the mass of each star? (2 pts)
 - (b) The fast star is smaller, and takes 2 hr to go from first contact to minimum in the lightcurve. Totality of the eclipse lasts for 10 hr. What is the size of each star? Which one is on the main sequence? (2 pts)

Extra Question for 4710 Honors and 6710 (1 question, 5 total points):

4. Make an H-R diagram of the 314 brightest stars in the sky (limited by apparent magnitude $m_V < 3.55$), drawn from the file hr_neareststars.dat. (You'll note that the columns represent B-V and M_V .) Is this set representative of where most stars live on the H-R diagram? In other words, are any kinds of stars over- or under-represented in a sample limited by apparent brightness? (5 pts)