

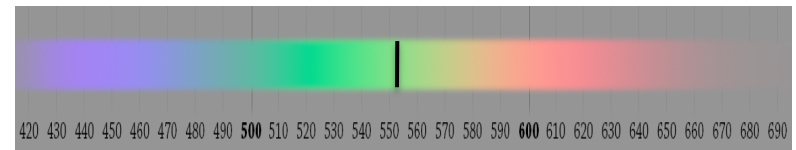
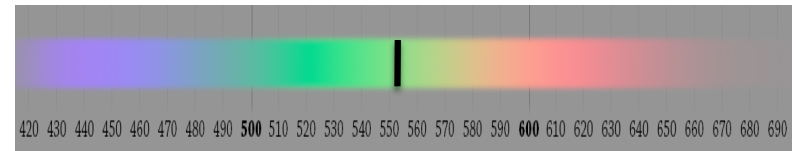
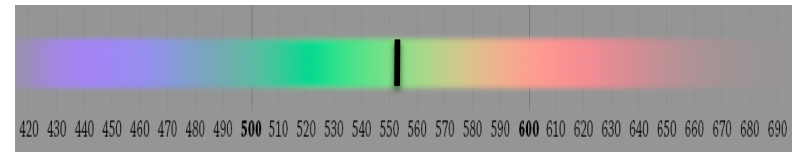
Introductory Astronomy

Week 4: Stars

Clip 12: Spectroscopic Binaries

Spectroscopic Binaries

- If stars too close to **resolve** can distinguish periodic **Doppler** shift:
Spectroscopic Binary
- If we can see both stars:
double-line binary
- Often see only one



Learning from Spectrum

- Measure: v_1, v_2, P $v = c(1 - \lambda/\lambda_0)$

$$R_{1,2} = \frac{v_{1,2}P}{2\pi} \quad M_1 R_1 = M_2 R_2 \quad M_1 v_1 = M_2 v_2$$

$$M = M_1 + M_2 = M_1(1 + v_1/v_2) = M_1(v_2 + v_1)v_2$$

$$R = R_1 + R_2 = \frac{Pv_1}{2\pi} + \frac{Pv_1}{2\pi} = (v_1 + v_2)\frac{P}{2\pi}$$

Finding Masses

$$P^2 = \frac{4\pi^2}{GM} R^3 \quad R = \frac{P}{2\pi} (v_1 + v_2) \quad M = (v_1 + v_2) M_1 / v_2$$

$$M = \frac{4\pi^2 R^3}{GP^2} = \frac{4\pi^2 P^3}{(2\pi)^3 GP^2} (v_1 + v_2)^3 \quad M_1 = \frac{v_2 P}{2\pi G} (v_1 + v_2)^2$$

- We can find **masses!** $M_1 = \frac{P}{1 \text{ y}} \frac{v_2 (v_1 + v_2)^2}{(29.78 \text{ km/s})^3} M_\odot$
- **Caveat:** Orbit may **tilt**, we measure only v_r
- If we only see one star only know v_1

Credits

- Rendered spectrum: Wikimedia/Spigget
[http://en.wikipedia.org/wiki/
File:Rendered_Spectrum.png/](http://en.wikipedia.org/wiki/File:Rendered_Spectrum.png/)