

# Computation of Receiver Position from Code Pseudoranges

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## 1 Introduction

The task is to compute the coordinates of a GPS receiver at given time using P1-code pseudoranges. My specific observation epoch is: 2004-02-02, 1 h 14 min 00 sec. The calculations were done with observation files and navigation files in RINEX format. The computation of this position is difficult because the satellite and receiver clocks are not synchronised, the satellite and receiver are both moving with varying velocities, and propagation times. There are many more problems that need to be accounted for like ionospheric and tropospheric corrections, but we have neglected them in this lab because we want to gain a general understanding before we take on the harder tasks.

## 2 Methodology

Describe the methods and instruments used to solve the problem. Matlab was used to solve the problem and used the GPS single point positioning algorithm.[?]

The steps from this algorithm are documented in the Matlab code, Section ???. The pseudoranges and navigation messages are used to calculate a more precise position of the receiver. The shift in position is a measure of the time for the signal to travel from satellite to receiver. This time traveled by the signal multiplied by the speed of light is the pseudo range.

All the satellites were used in this calculation in order to get the best possible approximation of the receiver coordinates. Since we will be using least-squares method to solve for the parameters, the more independent equations we have, the better the approximation will be.

## 3 Results

Present results from all steps you performed in the exercise. Images or tabular results should be included in this section. Describe and analyse the results and discuss the meaning and implications of the results. Answer the questions written in the lab instructions. Pay attention to the quality of input data and their significance for interpreting the reliability of results/conclusions. The results for the receiver position are given in Table ??. The results agree completely with the reference material given because both calculations used all of the satellites available.

## 4 Analysis and Discussion

In this part you should discuss your results by, for example, considering the following questions:

- Are the results reasonable? Compare the results with your expectations. Yes the results are reasonable. The final position of the receiver is very close to the approximate position.

	Epoch 2004 2 2 1 14 0.0		
	Final	Approximate	$\sigma$
$X$	3104225.071 m	3104219.453	1.330 m
$Y$	998384.754 m	998383.982	1.101 m
$Z$	5463300.077 m	5463290.508	2.566 m
$T$	-0.0005198825 s		4.7543792389e - 09 s

Table 1: Coordinates of the receiver after several iterations of steps 11-16 until the convergence condition was met. They agree completely with SPPResults and are very close to the approximate positions in the observation file.

The results are reasonable because the algorithm that we used only works if the approximate position is close to the solution. The solution we calculated was

- Can we draw any conclusion/implications from the results?
- Are results reliable and accurate?
- Would it have been more appropriate to use another method? Does the method need to be further developed?
- Etc.

## 5 Main Code

### Contents

- Computation of receiver's position
- Import observer file
- Import navigation file
- Match satellite numbers with available satellites
- Main loop steps 1-14
- 17. Repeat steps 11 -16 until the solution has converged.
- 17. Convergence condition
- Lab 2

### Computation of receiver's position

```
clear all;
c = 299792458; % speed of light (m/s)
mu = 3.986005e14; % universal gravitational parameter (m/s)^3
omega_e_dot = 7.2921151467e-5; % earth rotation rate (rad/s)
F = -4.442807633e-10; % s/m^1/2
time = [2,1,14,0]; % days, hours, minutes, seconds
junk = num2cell(time);
[nday,nhours,nminutes,nseconds] = junk{:};
clear junk;
```

### Import observer file

```
lov033b = importObserverFileAsString('0lov033b.04o', 1, 5629);
% Import P1 numbers and satellite numbers
% match your time with observer time
[rowInObs,nOfRows] = findTimeInObsFunction( lov033b,time );
% Import P1 numbers from the matched time above
p1_numbers = importObsP1numbers('0lov033b.04o', rowInObs+1,rowInObs+nOfRows*2);
% Import satellite numbers from matched time
satelliteNumbers = importObsSatelliteNumbers('0lov033b.04o', rowInObs,rowInObs);
[XA0,YA0,ZA0] = sampleFunction(lov033b); % Record Approximate Position
```

### Import navigation file

```
navfiles = importNavigationFiles('0lov033b.04n');
```

### Match satellite numbers with available satellites

```
satNumMatch = navfiles(1:8:96,1); % Order of satellite numbers import
sortedSatelliteNumbers = sortrows([satelliteNumbers',p1_numbers],1);
```

### Main loop steps 1-14

Calculates variables needed for correction iterations

```
count = 1;
for i = 1:length(satNumMatch)
```

### Steps 1-14 done inside

```
    if cell2mat(satNumMatch(i))==sortedSatelliteNumbers(count,1)
        [ Lmat(count,:), ...
          Amat(count,:),...
          rho(count,:),...
          Xs(count,:),Ys(count,:),Zs(count,:),...
          P1(count,:),...
          dtsL1_with_dtr(count,:),...
          tAtoS(count,:)]...
        = satLandP( i,sortedSatelliteNumbers(count,2),navfiles,...
XA0,YA0,ZA0,nday,nhours,nminutes,nseconds);
        count = count + 1;
    else
        fprintf('No data for Satellite%3d\n',cell2mat(satNumMatch(i)))
    end
end
```

## 17. Repeat steps 11 -16 until the solution has converged.

The solution has converged if the condition is fulfilled

```
for i = 1:10

    changeX = (Amat'*Amat)\(Amat'*Lmat); % eq. (21)
    v(:,i) = -Amat*changeX + Lmat; % eq. (17)
    newXYZ = [XA0,YA0,ZA0] + changeX(1:3)'; % eq. (22) est. coordinates
    newxyzcell = num2cell(newXYZ);
    [XA0,YA0,ZA0] = newxyzcell{:};
    clear newxyzcell;
    rho = sqrt(... % recompute rho
        (Xs - XA0 + omega_e_dot * YA0 * tAtoS).^2 + ... % x^2
        (Ys - YA0 - omega_e_dot * XA0 * tAtoS).^2 + ... % y^2
        (Zs - ZA0).^2 ... % z^2
    );
    Amat = [-(Xs - XA0)./rho,... % recompute matrix A
        -(Ys-YA0)./rho,...
        -(Zs-ZA0)./rho,...
        rho./rho];
    Lmat = P1 - rho + c*dtsL1_with_dtr; % recompute matrix L
```

## 17. Convergence condition

```
if i>1 % check for convergence condition
    condition = abs(v(:,end)'+v(:,end)-v(:,end-1)'+v(:,end-1)));
% condition must be 1e-5 or lower
    if condition < 1e-4

        fprintf('Convergence condition met = %d\n',condition);
```

## Lab 2

Find sigma and Q

```
    Q = inv((Amat'*Amat));
    sigma_0 = sqrt(v(:,end)'+v(:,end)/(length(Amat)-length(Q)));
    sigma_x = sigma_0*sqrt(Q(1,1));
    sigma_y = sigma_0*sqrt(Q(2,2));
    sigma_z = sigma_0*sqrt(Q(3,3));
    sigma_t = sigma_0*sqrt(Q(4,4))/c;
    fprintf('X = %7.3f, mX = %7.3f\n',XA0,sigma_x);
    fprintf('Y = %7.3f, mY = %7.3f\n',YA0,sigma_y);
    fprintf('Z = %7.3f, mZ = %7.3f\n',ZA0,sigma_z);
    fprintf('T = %0.10f, mt = %0.10d\n',-changeX(4)'/c,sigma_t);
    return

end

end

end
```

## 6 Matlab Function: satLandP.m

### Contents

- Constants
- Importing numbers from table 3 for a satellite
- 1. Compute signal propagation time by (13)
- 2. Compute signal transmission time by (14)
- 3. Compute satellite clock correction dtsL1
- 4. Compute ts using the correction from the step 3.
- 5. Compute eccentric anomaly (Table 2)
- 6. Compute dtr by (26) and ts by (15).
- 7. Compute satellite coordinates Xs, Ys, Zs, for time ts
- 8. Compute satellite clock correction dtsL1 by (24) - (27)
- 9. Compute tropospheric correction T\_A\_to\_s (tA)
- 10. Compute ionospheric correction I\_A\_to\_s (tA)
- 11. Compute approximate distance rho\_A0\_to\_s (tA) by (11).
- 12. Repeat steps 1 - 11 for all measured satellites.
- 13. Compute elements of vector L (19).
- 14. Compute elements of matrix A (20); a\_x\_to\_s , a\_y\_to\_s , a\_z\_to\_s by (12)

```
function [ Lmatrix, Amatrix, rho_f, Xs_f, Ys_f, Zs_f, P1_f, dtsL1_dtr_f, tAtoS_f] ...  
    = satLandP(  
satelliteNumberOrder, P1_f, navfiles, XA0, YA0, ZA0, ndaysf, nhoursf, nminutesf, nsec  
onds)
```

### Constants

```
c = 299792458; % speed of light (m/s)  
mu = 3.986005e14; % universal gravitational parameter (m/s)^3  
omega_e_dot = 7.2921151467e-5; % earth rotation rate (rad/s)  
F = -4.442807633e-10; % s/m^1/2
```

### Importing numbers from table 3 for a satellite

```
i = 1:8:112;  
sat = navfiles(i(satelliteNumberOrder):i(satelliteNumberOrder)+8,:);  
sat = transpose(cell2mat(sat));  
sat = num2cell(sat);  
% Imports numbers to all variables  
[~, af0, af1, af2, ...  
~, crs, change_n, m0, ...  
cuc, ec, cus, sqrtA, ...  
toe, cic, omega0, cis, ...  
i0, crc, w, omegadot, ...  
idot, ~, ~, ~, ...  
~, ~, tgd, ~, ...  
~] ...  
=sat{:};
```

### 1. Compute signal propagation time by (13)

```
tA_nom = seconds_in_week(ndaysf,nhoursf,nminutesf,nsecondsf); % 2 days, 1 hour,  
% 14 minutes My Time  
tAtoS_f = P1_f/c; % signal propagation time
```

### 2. Compute signal transmission time by (14)

```
tS_nom = tA_nom - P1_f/c;
```

### 3. Compute satellite clock correction dtsL1

by (24) and (25), neglect dtr

```
t_oc = toe; % I believe this is true, but not sure  
change_tsv_f = af0 + af1*(tS_nom-t_oc)+af2*(tS_nom-t_oc)^2; % (25)  
dtsL1 = change_tsv_f - tgd; % (24)
```

### 4. Compute ts using the correction from the step 3.

```
ts_f = tS_nom - dtsL1;
```

### 5. Compute eccentric anomaly (Table 2)

```
ek = mk + ec*sin(ek)  
  
A = sqrt(A^2);  
n0 = sqrt(mu/A^3); % Computed mean motion  
n = n0 + change_n;  
tk = ts_f - toe;  
tk = fixTk(tk); % if,then for table 2 of tk  
mk = m0 + n*tk;  
Ek = keplersEquation(mk,ec);
```

### 6. Compute dtr by (26) and ts by (15).

```
change_tr = F*ec*sqrt(A*sin(Ek)); % (26)  
ts_with_dtr = ts_f - change_tr;
```

### 7. Compute satellite coordinates Xs, Ys, Zs, for time ts

- Table 2 Calculate rk

```
vk = atan2((sqrt(1-ec^2)*sin(Ek)/(1-ec*cos(Ek))),...  
((cos(Ek)-ec)/(1-ec*cos(Ek))));  
Phik = vk + w;  
drk = crs*sin(2*Phik) + crc*cos(2*Phik);  
rk = A*(1-ec*cos(Ek)) + drk; % Corrected radius  
% Calculate uk  
duk = cus*sin(2*Phik) + cuc*cos(2*Phik);
```

```

uk = Phik + duk;
% Calculate ik
dik = cis*sin(2*Phik) + cic*cos(2*Phik);
ik = i0 + dik + idot*tk;
% Calculate omega's
omegak = omega0 + (omegadot-omega_e_dot)*tk - omega_e_dot*toe;
% Calculate xkp and ykp
xkp = rk*cos(uk);
ykp = rk*sin(uk);
% Calculate xk,yk,zk -> Xs, Ys, Zs for time ts
xk = xkp*cos(omegak) - ykp*cos(ik)*sin(omegak);
yk = xkp*sin(omegak) + ykp*cos(ik)*cos(omegak);
zk = ykp*sin(ik);
Xs_f = xk;
Ys_f = yk;
Zs_f = zk;

```

**8. Compute satellite clock correction dtsL1 by (24) - (27)**

```
dtsL1_dtr_f = change_tsv_f + change_tr - tgd; % (24)
```

**9. Compute tropospheric correction T\_A\_to\_s (tA)**

**10. Compute ionospheric correction I\_A\_to\_s (tA)**

**11. Compute approximate distance rho\_A0\_to\_s (tA) by (11).**

```

rho_f = sqrt(...
    (Xs_f - XA0 + omega_e_dot * YA0 * tAtoS_f)^2 + ... % x^2
    (Ys_f - YA0 - omega_e_dot * XA0 * tAtoS_f)^2 + ... % y^2
    (Zs_f - ZA0)^2    ... % z^2
);
% dtA = 0;
% rho_A_to_s = P1 + c*dtsL1_with_dtr - c*dtA; % (8) dtA = 0

```

**12. Repeat steps 1 - 11 for all measured satellites.**

**13. Compute elements of vector L (19).**

```
Lmatrix = P1_f - rho_f + c*dtsL1_dtr_f;
```

**14. Compute elements of matrix A (20); a\_x\_to\_s , a\_y\_to\_s , a\_z\_to\_s by (12)**

```
Amatrix = 1/rho_f*[-(Xs_f - XA0),-(Ys_f - YA0),-(Zs_f - ZA0),rho_f];
```

end

## Functions

% Function used to fix tk value

```
function tk = fixTk( tk )
```

```

if tk > 302400
    tk = tk - 604800;
elseif tk < -302400
    tk = tk + 604800;
else
    tk = tk;
end

%% Function used to calculate keplers equation
function Ek = keplersEquation( mk,ec )
E0(1) = mk;
n=7;
for i = 1:n
    E0(i+1) = mk + ec*sin(E0(i));
    eps = abs(E0(i)-E0(i+1));
end
Ek = E0(end);
end

%% Function that converts days,hours,min,seconds to total seconds

function total_seconds = seconds_in_week( days, hours, min, sec )
total_seconds = (days-1)*24*3600 + hours*3600 + min*60 + sec;
end

```

## References

- [1] M. Horemuž, “Gps single point positioning algorithm,” Royal Institute of Technology, Tech. Rep., 2014.