Astronomy

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1 Identification of Celestial Bodies

Our night sky is on a cloudless night is one littered with light. These lights are celestial bodies extremely far away that could be distant stars, or even galaxies. Though despite the vast number of celestial bodies that make them seem endless and identical, some bodies still have some distinct qualities to them that help us identify them when viewed from a different location/time.

One of these formations of stars or asterism that is part of the the constellation, Ursa Major is the Big Dipper. The Big Dipper is made up of bright, identifiable, stars that when viewed, looks like a dipper, spoon, ladle. We only know these stars because they were positioned in a shape that made them identifiable when observed in relation with every other star of the asterism. That is the basic idea behind constellations as well. These groups of stars only received wide recognition due to the fact that they created recognizable shapes when observed as a group.

This is what we are trying to achieve without the help of an observable visual pattern. By taking a patch of night sky that includes all the celestial bodies you would see in that patch, we will try to devise an algorithm that when given a smaller patch of our patch, we can identify the celestial bodies regardless if the patch has been transformed, rotated, or dilated in anyway.

We would start off with obtaining our patch of sky. We try to get a mostly uniform patch of sky and we would do this by first deciding on what part of the sky we would want. When describing locations on Earth, we use longitude and latitude while when we are describing locations in the sky, we would use right ascension and declination. In our case, we pick our right ascension and declination to range from 180 degrees to 181.62 degrees and 43.85 degrees to 45 degrees respectively.

We then would use these position parameters and make a query to the Sloan Digital Sky Survery database in order to get information on the celestial bodies inside our patch of sky.

```
[93]: #Code provided by Dr. John Ringland
    #import the libraries that we need
    import requests
    from io import StringIO
    import pandas as pd

#Our query string
    q = '''SELECT TOP 100000
        p.objid,p.ra,p.dec,p.u,p.g,p.r,p.i,p.z,
        s.specobjid, s.class, s.z as redshift
FROM PhotoObj AS p
        JOIN SpecObj AS s ON s.bestobjid = p.objid
```

```
WHERE

p.ra BETWEEN 180 AND 181.62 and
p.dec BETWEEN 43.85 AND 45

"""

#Sloan Digital Sky Survey website

url = 'http://skyserver.sdss.org/dr14/SkyServerWS/SearchTools/SqlSearch?

→cmd={}&format=csv'.format(q.replace(" ","%20"))

r = requests.get(url)

#If we successfully get a response from the server, we take the returned CSV

→file and read it into a dataframe

if r.status_code==200:

csv = StringIO(r.text)

df = pd.read_csv(csv,skiprows=1)

display(df.head())

else:

print("unsuccessful")
```

```
objid
                                       dec
                             ra
                                                           g
                                                                     r
 1237661872939073866 180.474860 44.912173 22.65754 21.29497
                                                              20.37337
1 1237661850394166115 180.105016 44.454457 24.89294 23.83855
                                                              23.36518
2 1237661872402203235 180.638896 44.533391 22.40390 20.91432
                                                              19.37721
3 1237661850931036788 180.195379 44.859022 24.79819 21.98236
                                                              20.36770
4 1237661850931167390 180.538095 44.877138 26.30246 21.03177 19.24720
         i
                              specobjid class redshift
0 20.26740 20.14058 7477226183001612288
                                          STAR -0.000734
1 19.88595 18.42799 1541401554513324032
                                          STAR -0.000313
2 18.59760 18.08331 1541401004757510144 GALAXY 0.293244
3 19.55420 19.23905 7478483749643984896 GALAXY 0.488021
 18.53533 18.21984 7478495019638169600 GALAXY 0.408716
```

Now that we have a database with the coordinates of the celestial bodies in our patch, we just need to give each body a nickname to distinguish them and to project them onto the tangent plane so we can have a x and y value for each celestial body.

```
dfb['y'] = dfb['dec'] #creating the y coordinate value
```

This code creates a copy of our initial dataframe and modifies it by adding x and y values to each celestial body.

Now using this code, we added nicknames to the dataframe to distinguish each celestial body more easily.

We can see that we now have a sorted by brightness dataframe with x,y,coordinates,and a nickname.

We now need a method in order to classify each of our celestial bodies. Just like the stars in a constellation, they are meaningless and indistinct by themselves but when viewed in context with other closer stars, they start to become distinguishable. We will then need to classify each star relating to two other stars.

We want to consider the points in terms of a triangle where each celestial body is a vertex of a triangle. We can do this by taking taking triples of three points and creating a unique triple for every combination of points in the patch.

```
[313]: def listOfTriples(df):
    xs = df["x"].tolist() #writes the column of the dataframe to a list
    ys = df["y"].tolist()
    listofXY = []
    for i in range(len(xs)):
        coordinates = (xs[i],ys[i]) #creates a triple of x and y values to use_
    →as coordinates
    listofXY.append(coordinates)
    return listofXY
```

This function made a list of 2 item lists that contains the x and y coordinates but now we need to create a 3 item list for each combination of points.

```
[202]: def triangles():
    listOfCombinations = []
    trips = listOfTriples()
    for i in range(len(trips)):
```

```
for j in range(i+1,len(trips)):
    for k in range(j+1,len(trips)):
        listOfCombinations.append([trips[i],trips[j],trips[k]])#grabs

→unique triples of every coordinate pair in triples

return listOfCombinations
```

Now that we have successfully created our list of unique combinations of coordinates, we now essentially have a list of triangles.

To distinguish each triangle and to be able to pinpoint it back to our known points, we need to establish a value to each triangle to identify it.

We can standardize this triangle by having the longest side always be positioned horizontally such that one of the points of the longest side is at the origin and the other is on the x-axis with the distance between them being the length of the longest side of the triangle but we will later standardize this triangle further by normalizing the sides of the triangle by the length of the longest side.

Through the use of trigonometry, we will pick the point that is not on the x-axis after normalizing the triangle to be the represented point of the triangle.

The first step in doing this is to determine a consistent direction we want to be using for all triangles.

We choose for the relative direction of the triangle to be clockwise. So that the side clockwise from the longest is the side opposite point B, which is one of the points of the longest side of the triangle resting on the x-axis. The function above does that for us by swapping the order of the triples and giving us the correct orientation.

```
[204]: import math

def betsyP(triple):
    working = isCounterClockwise(triple)
    A = working[0] #point A
    B = working[1] #point B
    C = working[2] #point C
```

```
a = math.sqrt((C[0]-B[0])**2+(C[1]-B[1])**2) #line BC
b = math.sqrt((A[0]-C[0])**2+(A[1]-C[1])**2) #line AC
c = math.sqrt((B[0]-A[0])**2+(B[1]-A[1])**2) #line AB
longest = max(a,b,c)
if (a == longest):
    ap = 1
    bp = b/a
    cp = c/a
    x = ((cp**2) + (ap**2) - (bp**2))/2
    y = abs(math.sqrt((cp**2) - (x**2)))
    return (x,y)
if (b == longest):
    ap = a/b
    bp = 1
    cp = c/b
    x = ((ap**2)+(bp**2)-(cp**2))/2
    y = abs(math.sqrt((ap**2)-(x**2)))
    return (x,y)
if (c == longest):
    ap = a/c
    bp = b/c
    cp = 1
    x = ((bp**2)+(cp**2)-(ap**2))/2
    y = abs(math.sqrt((bp**2)-(x**2)))
    return (x,y)
```

Now that we have our algorithm to return distinguished points determined by the coordinates of three points, we can now apply it to our triangle points.

We now have a dictionary with our distinguishing points acting as keys for us to use to access the names of the points of the triangles, one entry for every unique combination of x and y

coordinates.

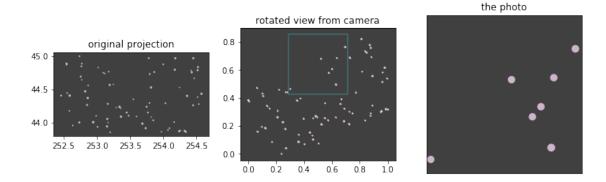
```
[379]: #Code provided by Dr. John Ringland
      def take_photo(starprojection_in):
          def random(lo,hi):
              return lo + (hi-lo)*np.random.rand()
          starprojection = starprojection_in.copy()
          starprojection['brightness'] -= starprojection['brightness'].min()
          starprojection['brightness'] += .1
          starprojection['brightness'] /= starprojection['brightness'].max()
          sp = starprojection[['x','y']].values
          plt.figure(figsize=(12,4))
          plt.subplot(131,aspect=1,facecolor='#404040')
          \#plt.plot(sp[:,0],sp[:,1],'o',color='\#ffddff',markersize=1,alpha=0.7)
          plt.scatter(sp[:,0],sp[:,1],c='#ffddff', s=10*starprojection['brightness'],u
       \rightarrowalpha=0.75, linewidth=0)
          plt.title('original projection')
          sp -= sp.min(axis=0)
          sp /= sp.max()
          # data now lies in unit square
          theta = 2*np.pi*np.random.rand() # choose a rotation angle
          c,s = np.cos(theta),np.sin(theta)
          rotation = np.array([[c,-s],[s,c]])
          rsp = (np.dot(rotation,sp.T)).T
          # scale to unit square again
          rsp -= rsp.min(axis=0)
          rsp /= rsp.max()
          w = random(.2, .6)
          xlo = random(0, 1-w)
          xhi = xlo + w
          h = random(.15,.5)
          ylo = random(0, 1-h)
          yhi = ylo + h
          plt.subplot(132,aspect=1,facecolor='#404040')
          plt.plot(rsp[:,0],rsp[:,1],'o',color='#ffddff',markersize=1,alpha=0.7)
          plt.scatter(rsp[:,0],rsp[:,1],c='#ffddff',u

→s=10*starprojection['brightness'], alpha=0.75, linewidth=0 )
          box = np.array([[xlo,xhi,xhi,xlo,xlo],[ylo,ylo,yhi,yhi,ylo]])
          plt.plot(*box,'#40a0a0',alpha=0.5)
          plt.title('rotated view from camera')
          starprojection['xprime'] = rsp[:,0]
```

```
starprojection['yprime'] = rsp[:,1]
  photo = starprojection[ (starprojection['xprime']>=xlo) & \
                            (starprojection['xprime'] <= xhi) & \</pre>
                            (starprojection['yprime']>=ylo) & \
                            (starprojection['yprime']<=yhi) ]</pre>
   \#ibox = np.dot(np.linalq.inv(rotation),box)
  plt.subplot(133,aspect=1,facecolor='#404040')
   #plt.
\rightarrow plot(photo['xprime'], photo['yprime'], 'o', color='#ffddff', markersize=2, alpha=0.
\rightarrow 7
  plt.scatter(photo['xprime'],photo['yprime'],c='#ffddff',__
⇒s=100*starprojection['brightness'], alpha=0.75, linewidth=0)
  plt.xlim(xlo,xhi)
  plt.ylim(ylo,yhi)
  plt.xticks([])
  plt.yticks([])
  plt.title('the photo')
   #plt.axis('off')
  photo = photo.rename(columns={'nickname':'nickname for checking_
→answers','brightness':'brightness for plotting only'})
  return photo[['xprime','yprime','nickname for checking answers','brightness_
→for plotting only']].reset index(drop=True)#, 'brightness']]
```

With this code, when running it with our data-frame, we can get a transformed set of x and y coordinates for the celestial bodies in a portion of our patch. Though the bodies are transformed, the relational position of the celestial bodies are still held and through our algorithm we should be able to figure out which celestial bodies are in this transformed portion.

```
[381]: example = take_photo(sortedDF)
      example.head()
[381]:
                     yprime nickname for checking answers
           xprime
      0 0.629262 0.497374
                                                   Kameran
      1 0.694801 0.765416
                                                   Andria
      2 0.601289 0.608364
                                                   Adalia
      3 0.578063 0.581012
                                                     Pius
      4 0.521253 0.681715
                                                     Emari
         brightness for plotting only
      0
                             0.818968
                             0.768076
      1
      2
                             0.548289
      3
                             0.539558
      4
                             0.495120
```



This shows a dataframe of the augmented x and y coordinates. We will do what we did with our original dataframe and create unique combinations of triples of coordinates, which we will find the distinguishing coordinate for, then compare back to our dictionary to find the names of the celestial bodies.

```
[389]: def listOfTriplesK():
          xs = example["xprime"].tolist() #writes the column of the dataframe to a_
       \hookrightarrow list
          ys = example["yprime"].tolist()
          listofXY = []
          for i in range(len(xs)):
               coordinates = (xs[i], ys[i]) #creates a triple of x and y values to use
       \rightarrowas coordinates
               listofXY.append(coordinates)
          return listofXY
      fff = listOfTriplesK()
      def trianglesK():
          listOfCombinations = []
          for i in range(len(fff)):
               for j in range(i+1,len(fff)):
                   for k in range(j+1,len(fff)):
                       listOfCombinations.append([fff[i],fff[j],fff[k]])#grabs unique_
       → triples of every coordinate pair in triples
          return listOfCombinations
[391]: | uuu = trianglesK()
```

Because the augmented coordinates could not exactly match the keys of the dictionary, we need a different method to find the approximation or closest distinguishing point that is closest to the distinguishing points of the augmented coordinates.