

THE EQUATION OF STATE

Assuming, a gas composed of three different kind of particles each with its own m_i and density n_i

$$\bar{m} = \frac{n_1 m_1 + n_2 m_2 + n_3 m_3}{n_1 + n_2 + n_3} = \frac{\rho}{n}$$

And gas pressure,

$$\left. \begin{array}{l} P_g = nKT \\ \downarrow \\ P_g = \frac{\rho}{\bar{m}} KT \end{array} \right\} \text{ and } n = \frac{\rho}{\bar{m}}$$

The mean mass will depend on the chemical abundance and ionization state of the gas. And for completely ionised pure hydrogen,

$$\bar{m} = \frac{m_H}{2}$$

$$n = \frac{2n_H}{\text{complete ionization of Hydrogen results in this case (electron + proton)}} + 3n_{He} + \sum \frac{A}{2} n_A = \frac{\rho}{m_H} \left(2x + \frac{3}{4}y + \frac{1}{2}z \right)$$

↓
In this case three particles 2 electrons and a nucleus.

$$\left[\begin{array}{l} n_H = \frac{X\rho}{m_H} \\ n_{He} = \frac{Y\rho}{4m_H} \end{array} \right] \quad \left[\begin{array}{l} n_A = \frac{Z_A \rho}{A m_H} \end{array} \right]$$

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$$= \frac{\rho}{2m_H} \left(3x + \frac{y}{2} + 1 \right) \quad \left\{ \text{where } x + y + z = 1 \right\}$$

$$\Rightarrow \frac{\bar{m}}{m_H} = \frac{\rho}{n m_H} = \frac{2}{1 + 3x + 0.5y}$$

RADIATION PRESSURE

$$P = \frac{F}{A} = \frac{dP/dt}{A} = \frac{2}{c} \int_{\pi/2}^{\pi} B \cos^2 \theta \sin \theta d\theta d\phi$$

$$= \frac{4\pi}{3c} B = \frac{1}{3} u \rightarrow \text{Energy density}$$

or,

$$P_{\text{rad}} = \frac{1}{3} u = \frac{1}{3} a T^4 \quad \left\{ \begin{array}{l} \text{rewriting upon case of} \\ \text{pressure due to thermal} \\ \text{photon inside a star} \end{array} \right.$$

$$P = P_g + P_{\text{rad}}$$

$$\boxed{P = \frac{\rho K T}{\bar{m}} + \frac{1}{3} a T^4}$$