RADIATIVE ENERGY TRANSPORT

of interaction = Modx effective cross section for alsorption and ocathering photon covers path distance 'dx'

number density

NOW,

for # of interactions = 1 and dx=1

: [L= 1] - This distance 'l' is called "meon free posh"

and there will be variety of absorbers and scatteress, each with its own density no and como section of

$$L = \frac{1}{\sum n_i \sigma_i} = \frac{1}{f \kappa}$$
 $\begin{cases} f \Rightarrow \text{ mass density} \end{cases}$

Thomson Scattering of photons on free electrons

The Thomson cross section is

$$E_{\perp} = \frac{3}{3} \left(\frac{e^2}{m_e c^2} \right)^2 = 6.7 \times 10^{-25} \text{ cm}^2$$

Independent of Tand ho.

and in interior of sun, we can assume all hydrogen are lonised, or, ne & S/mH (as there is lelectron per atom of mass mH)

So, Mean free path for electron scattering will be,
$$l_{es} = \frac{1}{1.4} \approx \frac{m_H}{1.4} \approx \frac{1.7 \times 10^{-24} \text{ g}}{1.4 \text{ g cm}^{-3} \times 6.7 \times 10^{-25} \text{ cm}^2} \approx 2 \text{ cm}$$
Solar rean density

Now, this was calculated mean free path but in reality, density of sun is higher in places where electron scattering is dominant.

And owed to this reason, the mean free path of typical photon is less,

.: L & I mm.

Thus, photons can travel only a tiny distance inside the Sun before being scattered or absorbed and re-emitted in a new direction. Since the new direction is random, the emergence of photons from the Sun is necessarily a random walk process.

D - charge in position of a photon after N steps li - leeh step described by li having length l.

$$D = 1, + 1_{2} + 1_{3} + 1_{4} + \cdots + 1_{N}$$
and the square of linear diplena cowerd,
$$D^{2} = |1_{1}|^{2} + |1_{2}|^{2} + |1_{3}|^{2} + \cdots + |1_{N}|^{2} + 2(|1_{1}|^{2} + |1_{3}|^{2} + \cdots + |1_{N}|^{2} + |1_{N}|^$$

and its depending is $\langle D^2 \rangle = NI^2$

will be zero as it is sum over many vector dost products

so, total time taken for photon to energy from centre to surjace,

$$T_{rw} \approx \frac{L}{C} \frac{r_0^2}{\ell^2} = \frac{r_0^2}{Lc} = \frac{(7 \times 10^{10} \text{ cm})^2}{10^{-1} \text{ cm } \times 3 \times 10^{10} \text{ cm s}^{-1}}$$

Each slep requires
$$T = \frac{1.6 \times 10^{12} \text{ Sec}}{2000 \text{ years}}$$

$$N = \frac{1.6 \times 10^{12} \text{ Sec}}{2000 \text{ years}}$$