19.8 The Runge-Lenz Vector

We studied the hydrogen atom in chapter 11. The Hamiltonian \hat{H} is very simple; it contains a kinetic energy term and a Coulomb potential term $V(r) = -e^2/r$. We calculated the spectrum of this Hamiltonian and found degenerate multiplets. While the energy eigenvalues depend only on a principal quantum number n, for each n there are degenerate multiplets of angular momentum with $\ell = 0, 1, ..., n - 1$.

The large amount of degeneracy in this spectrum asks for an explanation. The hydrogen Hamiltonian has in fact some hidden symmetry: there is a conserved quantum Runge-Lenz vector operator. In the following we discuss the *classical* Runge-Lenz vector and its conservation. In the end-of-chapter problems, you will learn about the quantum Runge-Lenz operator. In chapter 20 this knowledge will be used to give a fully algebraic derivation of the hydrogen atom spectrum.

Since the following analysis is classical, the vectors are not operators and carry no hats. Consider the energy function for a particle of

momentum **p** moving in a central potential V(r):

$$E = \frac{\mathbf{p}^2}{2m} + V(r). \tag{19.8.1}$$

The force **F** on the particle is given by the negative gradient of the potential:

$$\mathbf{F} = -\nabla V = -V'(r)\frac{\mathbf{r}}{r},\tag{19.8.2}$$

Here primes denote derivatives with respect to the argument. Newton's equation sets the rate of change of the momentum equal to the force:

$$\frac{d\mathbf{p}}{dt} = -V'(r)\frac{\mathbf{r}}{r}.\tag{19.8.3}$$

Here, $\mathbf{p} = m\dot{\mathbf{r}}$. You should confirm that for motion in a central potential the angular momentum $\mathbf{L} = \mathbf{r} \times \mathbf{p}$ is conserved:

$$\frac{d\mathbf{L}}{dt} = 0. ag{19.8.4}$$

Let us now calculate the time derivative of $\mathbf{p} \times \mathbf{L}$:

$$\frac{d}{dt}(\mathbf{p} \times \mathbf{L}) = \frac{d\mathbf{p}}{dt} \times \mathbf{L} = -\frac{V'(r)}{r} \mathbf{r} \times (\mathbf{r} \times \mathbf{p})$$

$$= -\frac{mV'(r)}{r} \mathbf{r} \times (\mathbf{r} \times \dot{\mathbf{r}}) = -\frac{mV'(r)}{r} [\mathbf{r}(\mathbf{r} \cdot \dot{\mathbf{r}}) - \dot{\mathbf{r}} r^{2}].$$
(19.8.5)

We now note that

$$\mathbf{r} \cdot \dot{\mathbf{r}} = \frac{1}{2} \frac{d}{dt} (\mathbf{r} \cdot \mathbf{r}) = \frac{1}{2} \frac{d}{dt} r^2 = r\dot{r}.$$
 (19.8.6)

Using this result, the derivative of $\mathbf{p} \times \mathbf{L}$ becomes

$$\frac{d}{dt}(\mathbf{p} \times \mathbf{L}) = -\frac{mV'(r)}{r} \left[\mathbf{r} r \dot{r} - \dot{\mathbf{r}} r^2 \right] = mV'(r)r^2 \left[\frac{\dot{\mathbf{r}}}{r} - \frac{\mathbf{r} \dot{r}}{r^2} \right]
= mV'(r)r^2 \frac{d}{dt} \left(\frac{\mathbf{r}}{r} \right).$$
(19.8.7)

Because of the factor $V'(r)r^2$, the right-hand side fails to be the time derivative of some quantity. But if we focus on potentials for which this

factor is a constant, the right-hand side is a time derivative, and we get a conserved quantity. So assume that for some constant γ we have

$$V'(r)r^2 = \gamma. {19.8.8}$$

It then follows that

$$\frac{d}{dt}(\mathbf{p} \times \mathbf{L}) = m\gamma \frac{d}{dt} \left(\frac{\mathbf{r}}{r}\right) \quad \Rightarrow \quad \frac{d}{dt} \left(\mathbf{p} \times \mathbf{L} - m\gamma \frac{\mathbf{r}}{r}\right) = 0. \tag{19.8.9}$$

The complicated vector inside the parentheses is constant in time. The condition (19.8.8) on the potential implies that

$$\frac{dV}{dr} = \frac{\gamma}{r^2} \quad \Rightarrow \quad V(r) = -\frac{\gamma}{r} + c_0. \tag{19.8.10}$$

This is the most general potential for which we get a conserved vector. For $c_0 = 0$ and $\gamma = e^2$, we have the hydrogen atom potential $V(r) = -e^2/r$. For this case the conservation equation reads

$$\frac{d}{dt}\left(\mathbf{p} \times \mathbf{L} - me^2 \frac{\mathbf{r}}{r}\right) = 0. \tag{19.8.11}$$

Factoring a constant we obtain the unit-free conserved **Runge-Lenz** vector **R** associated with the hydrogen atom classical dynamics:

$$\mathbf{R} = \frac{1}{me^2} \mathbf{p} \times \mathbf{L} - \frac{\mathbf{r}}{r}, \qquad \frac{d\mathbf{R}}{dt} = 0.$$
 (19.8.12)

The conservation of the Runge-Lenz vector is a property of inverse-squared central forces. The second term in \mathbf{R} is the inward-directed unit radial vector.

To familiarize ourselves with the Runge-Lenz vector, we first examine its value for a circular orbit, as shown in figure 19.8. With counterclockwise motion, the vector \mathbf{L} points out of the page, and $\mathbf{p} \times \mathbf{L}$ points radially outward. The vector \mathbf{R} is thus a competition between the outward-pointing first term along $\mathbf{p} \times \mathbf{L}$ and the inward-pointing second term along $-\hat{\mathbf{r}}$. If these two terms did not cancel, the result would be a radial vector, outward or inward, but in any case not conserved as it rotates

with the particle. This cannot happen, therefore the two terms must cancel. Indeed, for a circular orbit

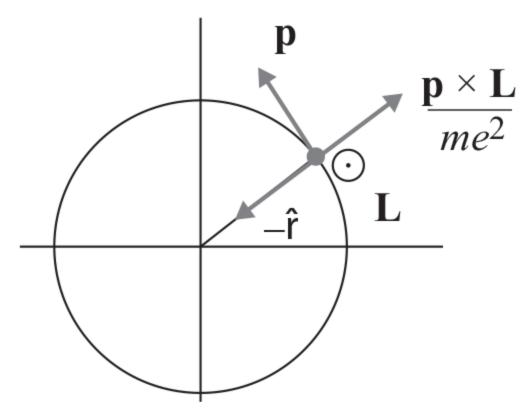


Figure 19.8The Runge-Lenz vector vanishes for a circular orbit.

$$m\frac{v^2}{r} = \frac{e^2}{r^2} \quad \Rightarrow \quad \frac{mv^2r}{e^2} = 1 \quad \Rightarrow \quad \frac{(mv)(mvr)}{me^2} = 1 \quad \Rightarrow \quad \frac{pL}{me^2} = 1, \tag{19.8.13}$$

which states that in a circular orbit the first term in **R** is a unit vector. Since it points outward, it cancels with the second term, and the Runge-Lenz vector vanishes for a circular orbit.

We now argue that for an elliptic orbit the Runge-Lenz vector is not zero. Consider figure 19.9, showing a particle in counterclockwise motion around an elliptic orbit. One focus of the ellipse is at the origin $\mathbf{r} = 0$. At all times the conserved \mathbf{L} points off the page. At the aphelion A, the point farthest away from the focal center, the first term in \mathbf{R} points outward, and the second term point inward. Thus, if \mathbf{R} does not vanish it must be a vector along the axis joining the focus and the aphelion, a horizontal vector on the figure. Now consider point B, shown directly above the focus

of the orbit. Here, \mathbf{p} is no longer perpendicular to the radial vector, and therefore $\mathbf{p} \times \mathbf{L}$ is no longer radial. As you can see, it points slightly to the left of the vertical. Since \mathbf{R} is conserved, and we know it is horizontal, its pointing to the left allows us to conclude that in an elliptic orbit \mathbf{R} is a nonzero vector pointing *from* the focus *to* the perihelion C, the point of closest approach in the orbit.

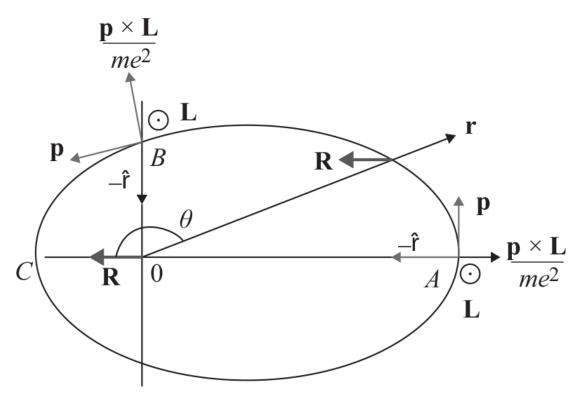


Figure 19.9 In an elliptic orbit, the Runge-Lenz vector is a vector along the major axis of the ellipse and points from the focus to the perihelion *C*.

Since \mathbf{R} vanishes for circular orbits, the length R of \mathbf{R} must measure the deviation of the orbit from circular. In fact, the magnitude R of the Runge-Lenz vector is precisely the eccentricity of the orbit. To see this we form the dot product of \mathbf{R} with the radial vector \mathbf{r} :

$$\mathbf{r} \cdot \mathbf{R} = \frac{1}{me^2} \mathbf{r} \cdot (\mathbf{p} \times \mathbf{L}) - r. \tag{19.8.14}$$

Referring to figure 19.9, let θ be the angle measured at the origin, increasing clockwise and defined with $\theta = 0$ the direction to the perihelion.

The angle between \mathbf{r} and \mathbf{R} is then θ and we get

$$rR\cos\theta = \frac{1}{me^2} L \cdot (\mathbf{r} \times \mathbf{p}) - r = \frac{1}{me^2} L^2 - r.$$
 (19.8.15)

Collecting terms proportional to r,

$$r(1+R\cos\theta) = \frac{L^2}{me^2} \Rightarrow \boxed{\frac{1}{r} = \frac{me^2}{L^2}(1+R\cos\theta)}.$$
 (19.8.16)

This is one of the standard presentations of an elliptic orbit, and R appears at the place one conventionally has the eccentricity e, thus e = R. If R = 0, the orbit is circular because r does not depend on θ . The identification of R with e follows from the definition

$$e = \frac{r_{\text{max}} - r_{\text{min}}}{r_{\text{max}} + r_{\text{min}}}.$$
 (19.8.17)

Here r_{\min} and r_{\max} are, respectively, the minimum and maximum distances to the focus located at the center of force.

Exercise 19.13. Use equations (19.8.16) and (19.8.17) to confirm that, indeed, e = R.

This analysis thus far has been classical. Quantum mechanically, some things must be changed; happily, not that much! The definition of \mathbf{R} only has to be changed to guarantee that \hat{R} is a Hermitian operator. Hermitization gives

$$\hat{\mathbf{R}} \equiv \frac{1}{2me^2} \left(\hat{\mathbf{p}} \times \hat{\mathbf{L}} - \hat{\mathbf{L}} \times \hat{\mathbf{p}} \right) - \frac{\hat{\mathbf{r}}}{r}. \tag{19.8.18}$$

Here \hat{r} is the position vector operator (not a unit vector).

Exercise 19.14. Confirm that \hat{R} defined above is Hermitian and reduces to R when vector operators become classical vectors.

The quantum mechanical conservation of $\hat{\mathbf{R}}$ is the statement that it commutes with the hydrogen Hamiltonian:

$$[\hat{\mathbf{R}}, \hat{H}] = 0. \tag{19.8.19}$$

The required calculation (problem 19.10) is the quantum analogue of the above classical calculation that showed that the time derivative of \mathbf{R} is zero. Moreover, the length squared of the operator $\hat{\mathbf{R}}$ is also of interest. The result (problem 19.11) is

$$\hat{\mathbf{R}}^2 = 1 + \frac{2}{me^4} \hat{H} (\hat{\mathbf{L}}^2 + \hbar^2). \tag{19.8.20}$$

These facts above will be used in section 20.8 to show that the symmetries generated by \hat{R} and the angular momentum \hat{L} determine completely the spectrum of the hydrogen atom.