

Problems: Expansion history of the universe and differential equations

ISSYP 2016

Throughout these problems, consider a cosmological expansion history with the following parameters:

$$H_0 = 0.069 \text{Gyr}^{-1} \quad \Omega_m = 0.31 \quad \Omega_{\text{rad}} = 4.15 \times 10^{-5} \quad \Omega_\Lambda = 1 - \Omega_m - \Omega_{\text{rad}} \quad (1)$$

Assume that dark energy is a cosmological constant.

I. EXPANDING UNIVERSE

1. Using Friedmann's equation and the definition of Ω_i , calculate numerical values for the density ρ_{tot} of the universe today, the matter density ρ_m today, and the radiation density ρ_{rad} today.
2. Find expressions for the matter density $\rho_m(a)$, the radiation density $\rho_{\text{rad}}(a)$, the total density $\rho_{\text{tot}}(a)$, and the Hubble parameter $H(a)$, as functions of a . Note that we're using a as the time coordinate here, i.e. all quantities are evaluated at a time when the scale factor of the universe is a .
3. Galaxies emit light in a narrow spectral line with wavelength 1216 Å (the "Lyman alpha" line). Consider a distant galaxy which is redshifted so that the line is observed at 2800 Å. How much has the universe expanded between the time when the light was emitted and now? How much has the total density of the universe ρ_{tot} decreased since then?
4. What was the scale factor of the universe when the densities of matter and dark energy were equal? When the densities of matter and radiation were equal?

II. SETTING UP DIFFERENTIAL EQUATIONS

5. *Radioactive decay, part 1.* Consider a radioactive sample which decays at the rate $r = 0.0001 \text{ yr}^{-1}$, meaning that in one year, 0.01% of the sample decays. Write a differential equation for the mass of the sample as a function of time.
6. *Radioactive decay, part 2.* Now consider a "decay chain". Suppose that radioactive element A decays to element B with rate $r_1 = 0.0001 \text{ yr}^{-1}$, and element B decays in turn to element C with rate $r_2 = 0.0003 \text{ yr}^{-1}$. Write differential equations for the masses of samples A , B , and C as functions of time.
7. *Mixing.* Consider a tank which initially contains 100 liters of pure water. At time $t = 0$, an input valve is turned on which feeds salt water into the tank. Assume that the salt water instantaneously mixes with the water in the tank, so that the concentration of salt is always uniform throughout the tank. To keep the total volume of water in the tank constant at 100 liters, an output valve is also turned on (at $t = 0$) which drains water from the tank at the same rate. Suppose that the rate of water in (and out) of the tank is 0.1 liters per second, and the concentration of salt in the input stream is 100 grams per liter. Write a differential equation for the total mass of salt M in the tank as a function of time.
8. *Free-fall with air resistance.* A brick of mass $m = 2 \text{ kg}$ is dropped out of an airplane at altitude 10000 m. Recall that the acceleration of the brick due to gravity is 9.8 m/s^2 . Assume that the effect of air resistance is to supply a drag force proportional to velocity: $F = kv$, where $k = 0.1 \text{ kg/s}$. Find differential equations which can be solved to find the location of the brick at time t .
9. *Roadrunner and coyote.* Consider a "roadrunner" who starts at the origin $(0,0)$ of the (x,y) plane, and runs in the x -direction at constant velocity 10 m/s . The roadrunner is being chased by a "coyote", who starts at the point $(0,100)$ and runs with velocity 12 m/s . Assume that the coyote always runs directly toward the roadrunner's instantaneous position. Find differential equations which could be solved to determine how long it takes the roadrunner to catch the coyote.

III. USING PYTHON

10. Make a length-50 list whose n -th element (for $0 \leq n < 50$) is equal to $(1.1)^n + 5n^2$.
11. The Fibonacci numbers $F_0 = 1$, $F_1 = 1$, $F_2 = 2$, $F_3 = 3$, $F_4 = 5$, $F_5 = 8$, \dots are defined by the recursion relation $F_n = F_{n-1} + F_{n-2}$. Write a python program to calculate F_{50} .
12. Make a list of 100 numbers t_i uniformly spaced between 0 and (2π) . Make another length-100 list containing the values of $\cos(t_i)$. By calling `matplotlib.pyplot.plot()` with these lists as arguments, make a plot of t vs $\cos(t)$.
13. Choose one of the differential equations you derived in the last section, and write a short python program to integrate it and plot the result.

IV. INTEGRATING FRIEDMANN'S EQUATION WITH PYTHON

Warning: everything in this section is supposed to be a challenge! Don't worry if you find it rough going at first. Friedmann's equation is the differential equation:

$$\frac{da}{dt} = aH(a) \quad (2)$$

where $H(a)$ has a somewhat messy algebraic form that you derived in question 2 above.

For numerically integrating this equation I suggest making time be the *dependent variable* and using $\log a$ (not a !) as the independent variable. In other words, your numerical integration would take equal steps in $\log a$, and evolve the value of t (the elapsed time since the big bang).

Something which isn't supposed to be obvious, but which you can show if you've taken some calculus, is that the form of Friedmann's equations in these variables is:

$$\frac{dt}{d \log a} = \frac{1}{H(a)} \quad (3)$$

or equivalently, defining $u = \log(a)$ for notational clarity

$$\frac{dt}{du} = \frac{1}{H(e^u)} \quad (4)$$

Write a python program to integrate this differential equation and tabulate t as a function of u . Based on your tabulated values, answer the following questions:

14. What is the age of the universe?
15. Consider the redshifted galaxy from question 3 above. How long ago was the light from the galaxy emitted?
16. How different would the age of the universe be if there were no dark energy, but the matter and radiation densities ρ_m and ρ_{rad} were the same?
17. Determine whether the expansion of the universe is accelerating or decelerating today, by determining whether da/dt is increasing or decreasing with the expansion. Would the answer be different if there were no dark energy?
18. Warning: this one is particularly difficult, I'll give some hints tomorrow! What is the comoving size of the universe? I.e. if we consider the furthest point we can look back to, how far away is that point today?