

Classical Mechanics : : Homework 10

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Problem 01

The equations of torque are

$$\begin{aligned}\tau_1 &= I_1\dot{\omega}_1 + (I_3 - I_2)\omega_3\omega_2 \\ \tau_2 &= I_2\dot{\omega}_2 + (I_1 - I_3)\omega_1\omega_3 \\ \tau_3 &= I_3\dot{\omega}_3 + (I_2 - I_1)\omega_2\omega_1\end{aligned}$$

Without losing generality let us assume $I_1 < I_2 < I_3$. So Dzhanibekov Effect, or Tennis Racket Theorem, says that I_2 moment of inertia is the one that is along the unstable principle axis.

Unstability along I_2

Let's look at an initial Ω rotating only along $(0, 1, 0)$ (the unstable axis). Note that we are working in the basis where the unit vectors are the principle axis.

Let's have a slight perturbation along for angular velocity $(1, 0, 0)$ that is δ which says

$$(0, \Omega, 0) \rightarrow (\delta, \Omega, 0)$$

The external torques are zero. Then using the above perturbation to find the rate of change of ω for the initial moment

$$\begin{aligned}\tau_1 &= 0 \implies \dot{\omega}_1 = 0 \\ \tau_2 &= 0 \implies \dot{\omega}_2 = 0 \\ \tau_3 &= 0 = I_3\dot{\omega}_3 + (I_2 - I_1)\Omega\delta \implies \dot{\omega}_3 = -\frac{I_2 - I_1}{I_3}\Omega\delta\end{aligned}$$

Hence after some infinitesimal time Δt we get ω_3 to be

$$\omega_3 \approx -\frac{I_2 - I_1}{I_3}\Omega\delta\Delta t$$

Now, with the new ω_3 iteration, the Euler's equations are (external torque is still 0)

$$\begin{aligned}\dot{\omega}_1 &= -\frac{I_3 - I_2}{I_1} \left(-\frac{I_2 - I_1}{I_3}\Omega\delta\Delta t \right) \Omega = \frac{(I_3 - I_2)(I_2 - I_1)}{I_1 I_3} \Omega^2 \delta \Delta t > 0 \\ \dot{\omega}_2 &= -\frac{I_1 - I_3}{I_2} \delta \left(-\frac{I_2 - I_1}{I_3}\Omega\delta\Delta t \right) = -\frac{(I_3 - I_1)(I_2 - I_1)}{I_3 I_2} \Omega \delta^2 \Delta t\end{aligned}$$

Being a little bit more sloppy with the mathematics we take a derivative

$$\ddot{\omega}_1 = \frac{(I_3 - I_2)(I_2 - I_1)}{I_1 I_3} \Omega^2 \delta$$

$$\ddot{\omega}_2 = -\frac{(I_3 - I_1)(I_2 - I_1)}{I_3 I_2} \Omega \delta^2$$

We can see above that ω_1 's second derivative grows rapidly for a small perturbation δ . This is an exponentially growing motion, instead of an oscillatory one. So the overall direction of rotation does not stay long $(0, \Omega, 0)$ for the whole motion, implying $(0, 1, 0)$ direction of rotation to be unstable.

We picked this to be the direction of I_2 when $I_1 < I_2 < I_3$. So this principal axis is unstable.

This exact same calculation for $(0, \Omega, \delta)$ perturbation will give the exact same exponential growth along perturbation.

Stability along I_1 (or I_3)

Let's start from $\vec{\omega} = (\Omega, 0, 0)$ and then introduce a general perturbation $(\Omega, \epsilon_2, \epsilon_3)$

$$0 = \dot{\omega}_1 + \frac{I_3 - I_2}{I_1} \epsilon_2 \epsilon_3$$

$$0 = \dot{\epsilon}_2 + \frac{I_1 - I_3}{I_2} \Omega \epsilon_3$$

$$0 = \dot{\epsilon}_3 + \frac{I_2 - I_1}{I_3} \Omega \epsilon_2$$

Fix the signs if necessary

$$0 = \dot{\omega}_1 + \frac{I_3 - I_2}{I_1} \epsilon_2 \epsilon_3$$

$$0 = \dot{\epsilon}_2 - \frac{I_3 - I_1}{I_2} \Omega \epsilon_3$$

$$0 = \dot{\epsilon}_3 + \frac{I_2 - I_1}{I_3} \Omega \epsilon_2$$

\implies Contract the notations

$$0 = \dot{\omega}_1 + A \epsilon_2 \epsilon_3$$

$$0 = \dot{\epsilon}_2 - B \Omega \epsilon_3$$

$$0 = \dot{\epsilon}_3 + C \Omega \epsilon_2$$

Let's look at the perturbation's differential equations

$$0 = \dot{\epsilon}_2 - B \Omega \epsilon_3$$

$$0 = \dot{\epsilon}_3 + C \Omega \epsilon_2$$

\implies Take on derivative

$$0 = \ddot{\epsilon}_2 - B \Omega \dot{\epsilon}_3 = \ddot{\epsilon}_2 + BC \Omega \epsilon_2 = 0$$

$$0 = \ddot{\epsilon}_3 + C \Omega \dot{\epsilon}_2 = \ddot{\epsilon}_3 + BC \Omega \epsilon_3 = 0$$

We get simple harmonic equations for the perturbations. So they don't make any noticeable difference other than just small precession like motion.

Given the amplitude of the perturbation being small, we can simply see that $\dot{\omega}_1 \approx 0$. For this we can say that I_1 is a stable principle axis direction.

Because of the symmetry of the equations, just flipping the notations will immediately prove the exact same reasoning for I_3 too. So we proved that Mr. Dzhanibekov was right (or Tennis Racket Theorem is true).

Veritasium in youtube has made a nice video on this matter.

Problem 02

(a)

For the hollow sphere, the inertia tensor is not going to have any off diagonal elements because of the symmetry (symmetry being any principle axis has the same moment of inertia given the origin is in the center of the ball). Also it's going to be a multiple of the Identity Matrix because of the spherical symmetry.

Through any axis crossing the center, the moment of inertia can be easily computed.

$$I_{ij} = \int_V \rho (r^2 \delta_{ij} - x_i x_j) dV$$

For our case I_{11}, I_{22}, I_{33} are the same thing. And any off diagonals are zero.

$$\rho = \frac{M}{\frac{4}{3}\pi(a^3 - b^3)}$$

$$I = \int_V \rho r^2 \sin^2 \theta dV$$

For the sphere

$$dV = r^2 \sin \theta dr d\theta d\phi$$

Take the integral

$$I = \rho \int_b^a \int_0^\pi \int_0^{2\pi} r^4 \sin^3 \theta d\phi d\theta dr$$

ϕ symmetry

$$\int_0^{2\pi} 1 d\phi = 2\pi$$

Latitude integral

$$\int_0^\pi \sin^3 \theta d\theta = \int_0^\pi (1 - \cos^2 \theta) \sin \theta d\theta = \left[-\frac{\cos \theta}{3} + \frac{\cos^3 \theta}{3} \right]_0^\pi = \frac{4}{3}$$

Radius from b to a

$$\int_b^a r^4 dr = \frac{1}{5} [r^5]_b^a = \frac{1}{5}(a^5 - b^5)$$

Put everything together

$$I = \rho \cdot 2\pi \cdot \frac{4}{3} \cdot \frac{1}{5}(a^5 - b^5) = \frac{8\pi}{15} \rho (a^5 - b^5)$$

$$I = \frac{8\pi}{15} \cdot \frac{M}{\frac{4}{3}\pi(a^3 - b^3)} \cdot (a^5 - b^5)$$

$$I = \frac{2M}{5} \cdot \frac{a^5 - b^5}{a^3 - b^3}$$

So the tensor is

$$\hat{I} = \frac{2M}{5} \cdot \frac{a^5 - b^5}{a^3 - b^3} \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

(b)

Angular impulse applied by cue (r distance from center)

$$\tau\Delta t = rF\Delta t = r\Delta p$$

Angular impulse received by ball

$$\Delta L = L - 0 = I\omega$$

They should be equal

$$r\Delta p = I\omega \quad (1)$$

Linear momentum gained is Δp . Invoke no slip condition then

$$\omega = \frac{v}{a} = \frac{\Delta p}{ma}$$

Use this with (1) then

$$r\Delta p = I \frac{\Delta p}{ma} \implies r = \frac{I}{ma}$$

What we get for H is

$$H = r + a = \frac{I}{ma} + a$$

In this case it will be

$$H = \frac{2}{5a} \frac{a^5 - b^5}{a^3 - b^3} + a$$

(c)

$b = 0$ Solid Sphere

$$H = \frac{2}{5a} a^2 + a = \frac{2}{5} a + a = \frac{7}{5} a$$

$b \rightarrow a$ Thin Sphere

$$\begin{aligned} H &= \frac{2}{5a} \frac{a^5 - b^5}{a^3 - b^3} + a \\ &= \frac{2}{5a} \frac{a^5 - (a - \varepsilon)^5}{a^3 - (a - \varepsilon)^3} + a \\ &= \frac{2}{5a} \frac{a^5 - a^5(1 - \frac{\varepsilon}{a})^5}{a^3 - a^3(1 - \frac{\varepsilon}{a})^3} + a \\ &\approx \frac{2}{5a} \frac{a^5 - a^5(1 - 5\frac{\varepsilon}{a})}{a^3 - a^3(1 - 3\frac{\varepsilon}{a})} + a \\ &= \frac{2}{5a} \frac{a^5(1 - 1 + 5\frac{\varepsilon}{a})}{a^3(1 - 1 + 3\frac{\varepsilon}{a})} + a \\ &= \frac{2}{3a} a^2 + a \\ &= \frac{2}{3} a + a = \frac{5}{3} a \end{aligned}$$

Problem 03

Call $I_1 = I_2 = I_x = I_y$. Then

$$\vec{\omega} = (\omega_1, 0, \omega_3)$$

Let the angular velocity of the space station be $\vec{\omega} = (\omega_1, 0, \omega_3)$ in the body frame before the torque is applied. The torque $\vec{\Gamma} = (0, 0, \Gamma)$ acts along the z -axis, which is a principal axis. The Euler equations for a rigid body in the body frame are

$$\begin{aligned} I_1\dot{\omega}_1 - (I_2 - I_3)\omega_2\omega_3 &= 0 \\ I_2\dot{\omega}_2 - (I_3 - I_1)\omega_3\omega_1 &= 0 \\ I_3\dot{\omega}_3 - (I_1 - I_2)\omega_1\omega_2 &= \Gamma \end{aligned}$$

Initially when $\vec{\omega} = (\omega_1, 0, \omega_3)$ and $I_1 = I_2$

$$\begin{aligned} I_1\dot{\omega}_1 &= 0 \\ I_2\dot{\omega}_2 - (I_3 - I_1)\omega_3\omega_1 &= 0 \\ I_3\dot{\omega}_3 &= \Gamma \end{aligned}$$

Let's solve this problem for a general $\vec{\omega} = (\omega_1, \omega_2, \omega_3)$ and the initial condition will be invoked as $\vec{\omega}_0 = (\omega_1^0, 0, \omega_3^0)$

$$\begin{aligned} \dot{\omega}_1 - \frac{I_2 - I_3}{I_1}\omega_2\omega_3 &= 0 \\ \dot{\omega}_2 - \frac{I_3 - I_1}{I_2}\omega_3\omega_1 &= 0 \\ \dot{\omega}_3 - \frac{I_1 - I_2}{I_3}\omega_1\omega_2 &= \frac{\Gamma}{I_3} \end{aligned}$$

$$I_1 = I_2 = I$$

$$\begin{aligned} \dot{\omega}_1 - \frac{I - I_3}{I}\omega_2\omega_3 &= 0 \\ \dot{\omega}_2 - \frac{I_3 - I}{I}\omega_3\omega_1 &= 0 \\ \dot{\omega}_3 &= \frac{\Gamma}{I_3} \end{aligned}$$

Define

$$\frac{I - I_3}{I} = \Lambda$$

Apply

$$\begin{aligned} \dot{\omega}_1 - \Lambda\omega_2\omega_3 &= 0 \\ \dot{\omega}_2 + \Lambda\omega_3\omega_1 &= 0 \\ \dot{\omega}_3 = \frac{\Gamma}{I_3} \implies \omega_3 &= \frac{\Gamma}{I_3}t + \omega_3^0 \end{aligned}$$

$$\begin{aligned} \ddot{\omega}_1 - \Lambda\omega_3\dot{\omega}_2 &= 0 \implies \ddot{\omega}_1 - \Lambda\omega_3(-\Lambda\omega_3\omega_1) = 0 \implies \ddot{\omega}_1 + \Lambda^2\omega_3^2\omega_1 = 0 \\ \ddot{\omega}_2 + \Lambda\omega_3\dot{\omega}_1 &= 0 \implies \ddot{\omega}_2 + \Lambda\omega_3(\Lambda\omega_2\omega_3) = 0 \implies \ddot{\omega}_2 + \Lambda^2\omega_3^2\omega_2 = 0 \end{aligned}$$

Equations of motion we are left with

$$\begin{aligned} \ddot{\omega}_1 + \Lambda^2\omega_3^2\omega_1 &= 0 \\ \ddot{\omega}_2 + \Lambda^2\omega_3^2\omega_2 &= 0 \\ \omega_3 &= \frac{\Gamma}{I_3}t + \omega_3^0 \end{aligned}$$

$$\vec{\omega} = \begin{bmatrix} \omega_1^0 \cos\left(\frac{I-I_3}{I} \left[\frac{\Gamma}{I_3} t + \omega_3^0\right] t\right) \\ \omega_1^0 \sin\left(\frac{I-I_3}{I} \left[\frac{\Gamma}{I_3} t + \omega_3^0\right] t\right) \\ \frac{\Gamma}{I_3} t + \omega_3^0 \end{bmatrix}$$

Description of motion Rotation along axis of symmetry of the object increases because of the constant torque being provided.

The precession frequency is given by function of time

$$\Omega = \sqrt{\Lambda^2 \omega_3^2} = \frac{I - I_3}{I} \left(\frac{\Gamma}{I_3} t + \omega_3^0 \right)$$

The precession increases overtime and gets faster arbitrarily.

Problem 04

I have attached my handwritten scratch work for this problem.

(a)

Solving for acceleration vector

Top to bottom particles be named 1, 2, 3. Let the positive force direction be picked so that it points along the direction of *decreasing* θ_i . So in the diagram given in the problem set, for given θ_1 , the force points out of the screen towards the reader.

The force on 1 from 2 that points along the rim of the circle

$$F_{12} = \frac{kQ(Q\eta)}{b^2 + (a\theta_1 - a\theta_2)^2} \sin \alpha = \frac{\eta k Q^2 (a(\theta_1 - \theta_2))}{(b^2 + (a\theta_1 - a\theta_2)^2)^{3/2}} \approx \frac{\eta k Q^2 a(\theta_1 - \theta_2)}{b^3}$$

Using the above equation we can solve for the force from third particle 3

$$F_{13} \approx -\frac{kQ^2 a(\theta_1 - \theta_3)}{(2b)^3}$$

The force on 2

$$F_{21} \approx -\frac{\eta k Q^2 a(\theta_1 - \theta_2)}{b^3}$$

$$F_{23} \approx \frac{\eta k Q^2 a(\theta_2 - \theta_3)}{b^3}$$

The force on 3

$$F_{31} \approx \frac{kQ^2 a(\theta_1 - \theta_3)}{(2b)^3}$$

$$F_{32} \approx -\frac{\eta k Q^2 a(\theta_2 - \theta_3)}{b^3}$$

Now $F_i = ma\ddot{\theta}_i$ hence defining two constants

$$A = \frac{\eta k Q^2 a}{mb^3(a)} \quad B = \frac{kQ^2 a}{m(2b)^3 a}$$

The system of forces can be written as¹.

$$\begin{aligned} F_{12} &= A(\theta_1 - \theta_2) & F_{13} &= -B(\theta_1 - \theta_3) \\ F_{21} &= -A(\theta_1 - \theta_2) & F_{23} &= A(\theta_2 - \theta_3) \\ F_{31} &= B(\theta_1 - \theta_3) & F_{32} &= -A(\theta_2 - \theta_3) \end{aligned}$$

Total force on F_1 being $F_1 = F_{12} + F_{13}$ we can write the matrix form of the

$$\begin{bmatrix} \ddot{\theta}_1 \\ \ddot{\theta}_2 \\ \ddot{\theta}_3 \end{bmatrix} = \begin{bmatrix} (A - B) & -A & B \\ -A & 2A & -A \\ B & -A & (A - B) \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix}$$

We have gotten the required form

$$\begin{bmatrix} (A - B) & -A & B \\ -A & 2A & -A \\ B & -A & (A - B) \end{bmatrix} = \begin{bmatrix} \alpha & \gamma & \delta \\ \gamma & \beta & \gamma \\ \delta & \gamma & \alpha \end{bmatrix}$$

¹Please don't be bothered about the ma factor in A, B , consider it 1

(a+b) Solving for the Eigenpairs (Eigenvalues and Eigenvectors)

For the matrix we found above, I am just going to put that into Wolfram Alpha to reduce the computational burden from myself.

The associated eigenvalues and eigenvectors I get are

$$\begin{aligned}\lambda_1 &= 0 & \vec{\lambda}_1 &= \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix} \\ \lambda_2 &= 3A & \vec{\lambda}_2 &= \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix} \\ \lambda_3 &= A - 2B & \vec{\lambda}_3 &= \begin{bmatrix} -1 \\ 0 \\ 1 \end{bmatrix}\end{aligned}$$

The eigenfrequencies are

$$\begin{aligned}\omega_2^2 &= \frac{3\eta k Q^2}{mb^3} \\ \omega_3^2 &= \frac{\eta k Q^2}{mb^3} - \frac{k Q^2}{4mb^3} = \frac{k Q^2}{mb^3} \left(\eta - \frac{1}{4} \right)\end{aligned}$$

The first one doesn't correspond to any oscillatory motion but just simple translation of the three masses in parallel.

The second and third one correspond to the following motion.

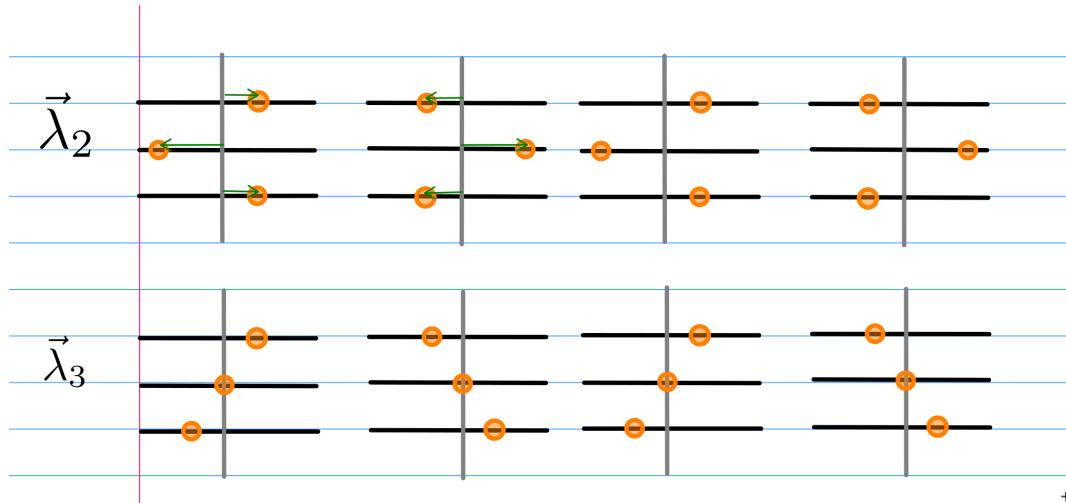


Figure 1: ./ss/11/1.png

For $\vec{\lambda}_2$ is an alternating motion where $\vec{\lambda}_3$ keeps the collinear between 3 particles.

(c)

Discussing unstable states, the first possible way how that could occur is if the η is too small, so the effects of the middle charge becomes obsolete, then the mutual forces between 1 and 3 would be too strong and that can break

the equilibrium. It's apparent if we see the eigenfrequency ω_3 where we required $\eta > \frac{1}{4}$ so that $\omega_3^2 > 0$ and not imaginary.

η can be indefinitely big (compared to 1) and still possibly keep the equilibrium.

(d)

Inspired by Morin Chapter 3 on Oscillations we can write the solution in the following way

$$\begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix} = A \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix} \cos(\omega_2 t + \phi_2) + B \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix}$$

Here for $t = 0$ also the speed

$$\frac{d}{dt} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix} = A \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix} (-\omega_2) \sin(\omega_2 t + \phi_2) = \vec{0} \implies \phi_2 = 0$$

So the cases are for $t = 0$

$$\begin{aligned} A \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix} + B \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix} &= \begin{bmatrix} 0 \\ \phi \\ 0 \end{bmatrix} \\ \begin{bmatrix} A + B \\ -2A + B \\ A + B \end{bmatrix} &= \begin{bmatrix} 0 \\ \phi \\ 0 \end{bmatrix} \\ \implies A &= -B \\ A &= -\frac{\phi}{3} \end{aligned}$$

Finally

$$\begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix} = -\frac{\phi}{3} \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix} \cos(\omega_2 t) + \frac{\phi}{3} \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix}$$

Where $\omega_2 = \sqrt{3\eta k Q^2 / mb^3}$

Problem 05

Force on any mass, using pure logic of springs

$$\begin{aligned} F_i &= -F_{i-1} + F_{i+1} = -k(\theta_i - \theta_{i-1}) + k(\theta_{i+1} - \theta_i) \\ &= -k\theta_i + k\theta_{i-1} + k\theta_{i+1} - k\theta_i \\ &= -2k\theta_i + k\theta_{i-1} + k\theta_{i+1} \end{aligned}$$

(a)

$$\ddot{\theta}_i = \frac{k}{m} (-2\theta_i + \theta_{i-1} + \theta_{i+1})$$

(b)

The chain of equation are

$$\begin{aligned} F_1 &= -2k\theta_1 + k\theta_N + k\theta_2 \\ F_2 &= -2k\theta_2 + k\theta_1 + k\theta_3 \\ F_3 &= -2k\theta_3 + k\theta_2 + k\theta_4 \\ &\vdots \\ F_{N-1} &= -2k\theta_{N-1} + k\theta_{N-2} + k\theta_N \\ F_N &= -2k\theta_N + k\theta_{N-1} + k\theta_1 \end{aligned}$$

Finding the matrix representation

$$\begin{bmatrix} F_1 \\ F_2 \\ F_3 \\ \vdots \\ F_N \end{bmatrix} = \begin{bmatrix} -2k & k & 0 & 0 & \cdots & 0 & k \\ k & -2k & k & 0 & \cdots & 0 & 0 \\ 0 & k & -2k & k & \cdots & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\ k & 0 & 0 & 0 & \cdots & k & -2k \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \\ \vdots \\ \theta_N \end{bmatrix}$$

This can be written as

$$\begin{bmatrix} F_1 \\ F_2 \\ F_3 \\ \vdots \\ F_N \end{bmatrix} = -k \begin{bmatrix} 2 & -1 & 0 & 0 & \cdots & 0 & -1 \\ -1 & 2 & -1 & 0 & \cdots & 0 & 0 \\ 0 & -1 & 2 & -1 & \cdots & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\ -1 & 0 & 0 & 0 & \cdots & -1 & 2 \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \\ \vdots \\ \theta_N \end{bmatrix}$$

Focusing on the matrix

$$\begin{bmatrix} 2 & -1 & 0 & 0 & \cdots & 0 & -1 \\ -1 & 2 & -1 & 0 & \cdots & 0 & 0 \\ 0 & -1 & 2 & -1 & \cdots & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\ -1 & 0 & 0 & 0 & \cdots & -1 & 2 \end{bmatrix} = 3I - \begin{bmatrix} 1 & 1 & 0 & 0 & \cdots & 1 \\ 1 & 1 & 1 & 0 & \cdots & 0 \\ 0 & 1 & 1 & 1 & \cdots & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots \\ 1 & 0 & 0 & 0 & \cdots & 1 \end{bmatrix} = 3I - \mathcal{M}$$

We are interested on this \mathcal{M} .

- The cyclic nature of \mathcal{M} can give us some motivation to guess the eigenvectors. We will attack this problem with Polynomials.²
- I searched up for some papers that deals with this problem, here's one for alternating signs and general case [Paper 1 Link]
- Inspired, we will look at the N -th root of 1 which is η . I take assistance from MATH 382 Computation Complex Analysis courework where the N -th root η n -th term can be given by $\eta_n = \exp(2\pi i n/N)$ and obviously $0 \leq n < N$.
- Given the condition of η ,

$$\begin{bmatrix} 1 \\ \eta \\ \eta^2 \\ \vdots \\ \eta^{N-1} \end{bmatrix}$$

is an eigenvector with eigenvalue $\eta^{-1} + 1 + \eta$ in this case.

- The eigenvalues of the entire matrix $3I - \mathcal{M}$ hence are given by $3 - (\eta^{-1} + 1 + \eta) = 2 - \eta^{-1} - \eta$.
- Because of N different roots of 1 as given by η_n , the eigenvalues are for $0 \leq n < N$

$$\lambda_n = 2 - \eta_n^{-1} - \eta_n = 2 - (e^{-2\pi i n/N} + e^{2\pi i n/N})$$

Turning to cosines

$$\lambda_n = 2 - 2 \cos(2\pi n/N) \implies \lambda_n = 4 \sin^2(\pi n/N)$$

And the corresponding eigenvector is

$$|4 \sin^2(\pi n/N)\rangle = \begin{bmatrix} 1 \\ \eta_n \\ \eta_n^2 \\ \vdots \\ \eta_n^{N-1} \end{bmatrix}$$

- Since the numbers n and $N - n$ yield the same value for λ_n in $4 \sin^2(\pi n/N)$, the eigenvalues tend to come in pairs (except for when $n = 0$ and $n = N/2$ if N is even). Because of this we can form real linear combinations of the two corresponding complex eigenvectors given in $[1 \ \eta_n \ \eta_n^2 \ \dots \ \eta_n^{N-1}]$.
- The two vectors

$$V_n^+ = \frac{1}{2} (V_n + V_{N-n}) = \begin{bmatrix} 1 \\ \cos(2\pi n/N) \\ \cos(4\pi n/N) \\ \vdots \\ \cos(2(N-1)\pi n/N) \end{bmatrix}$$

and

$$V_n^- = \frac{1}{2i} (V_n - V_{N-n}) = \begin{bmatrix} 0 \\ \sin(2\pi n/N) \\ \sin(4\pi n/N) \\ \vdots \\ \sin(2(N-1)\pi n/N) \end{bmatrix}$$

both have eigenvalue $\lambda_n = \lambda_{N-n}$ as discussed in the last itemized point. Note that we see **degeneracy** here.

- Eigenfrequency corresponding to the two above normal modes are

$$\lambda_n = 4 \sin^2(\pi n/N) \implies \omega_n = \sqrt{\frac{k}{m}} (2 \sin(\pi n/N))$$

²This problem was part of my self-study before going to International Physics Olympiad so there's a lot of nostalgia attached to this funny matrix.

(c)

$$\omega_n = 2\sqrt{\frac{k}{m}} \sin(\pi n/4)$$

For $N = 4$, if $n = 0$ we find $\lambda_0 = \omega_0 = 0$ and $V_0 = (1, 1, 1)$.

For $n = 1$ we find $\lambda_1 = \sqrt{k/m}2 \sin(\pi/4) = \sqrt{2}\sqrt{k/m}$. Modes

$$V_4^+ = \begin{bmatrix} 1 \\ \cos(2\pi n/4) \\ \cos(4\pi n/4) \\ \cos(6\pi n/4) \end{bmatrix}$$

$$V_4^- = \begin{bmatrix} 0 \\ \sin(2\pi n/4) \\ \sin(4\pi n/4) \\ \sin(6\pi n/4) \end{bmatrix}$$

$$V_4^+ = \begin{bmatrix} 1 \\ \cos(2\pi/4) \\ \cos(4\pi/4) \\ \cos(6\pi/4) \end{bmatrix} = \begin{bmatrix} 1 \\ 0 \\ -1 \\ 0 \end{bmatrix}$$

$$V_4^- = \begin{bmatrix} 0 \\ \sin(2\pi/4) \\ \sin(4\pi/4) \\ \sin(6\pi/4) \end{bmatrix} = \begin{bmatrix} 0 \\ 1 \\ 0 \\ -1 \end{bmatrix}$$

For $n = 2$ we find we find $\omega_2 = 2\sqrt{k/m}$

$$V_4^+ = \begin{bmatrix} 1 \\ \cos(2\pi/2) \\ \cos(4\pi/2) \\ \cos(6\pi/2) \end{bmatrix} = \begin{bmatrix} 1 \\ -1 \\ 1 \\ -1 \end{bmatrix}$$

$$V_4^- = \begin{bmatrix} 0 \\ \sin(2\pi/2) \\ \sin(4\pi/2) \\ \sin(6\pi/2) \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

For $n = 3$ we find $\omega_3 = \sqrt{2}\sqrt{k/m}$. This will give the same eigenvectors as we had seen that n and $N - n$ are the same. So $n = 1$ and $N - 1 = 3$ will be the same case. Same for $n = 4$ where $\omega_4 = 2\sqrt{k/m}$.

So the eigenmodes are (where degeneracy is present with two different eigenmodes for one eigenfrequency)

$$\text{for } \sqrt{2k/m} \implies \begin{bmatrix} 0 \\ 1 \\ 0 \\ -1 \end{bmatrix} \text{ and } \begin{bmatrix} 1 \\ 0 \\ -1 \\ 0 \end{bmatrix}$$

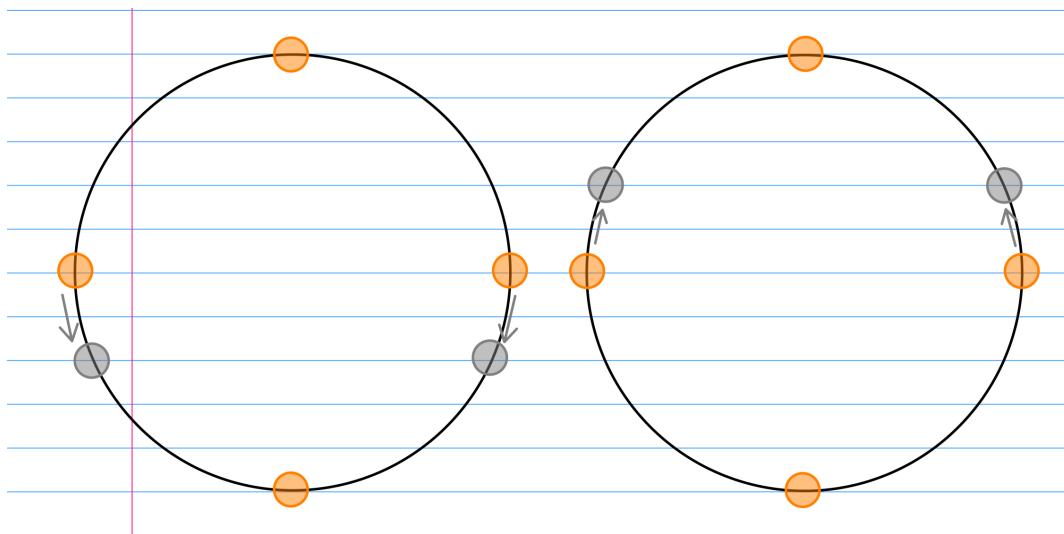


Figure 2: The mode $(0, 1, 0, -1)$

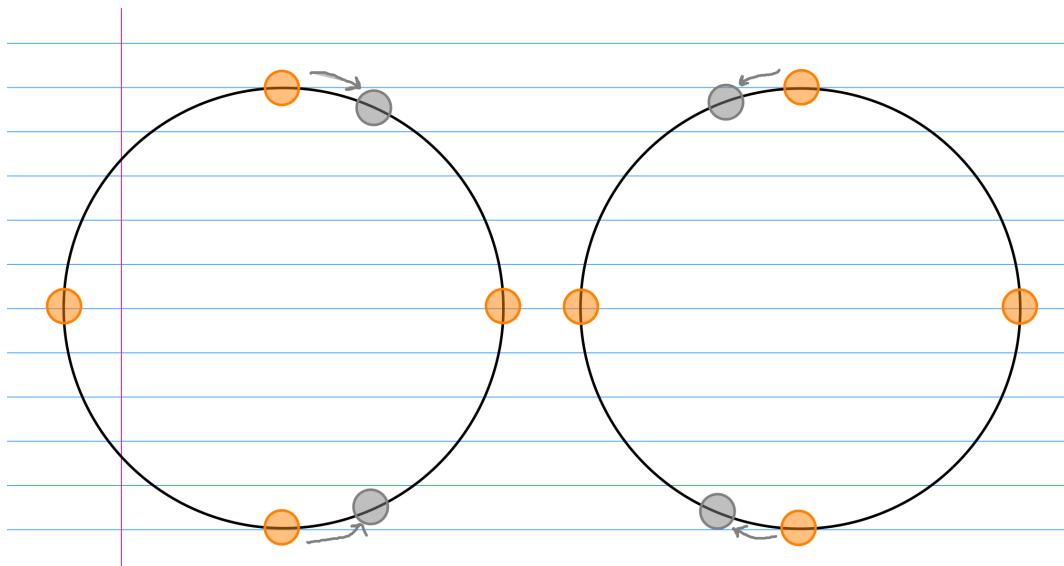


Figure 3: The mode $(1, 0, -1, 0)$

$$\text{for } 2\sqrt{k/m} \Rightarrow \begin{bmatrix} 1 \\ -1 \\ 1 \\ -1 \end{bmatrix} \text{ and } \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

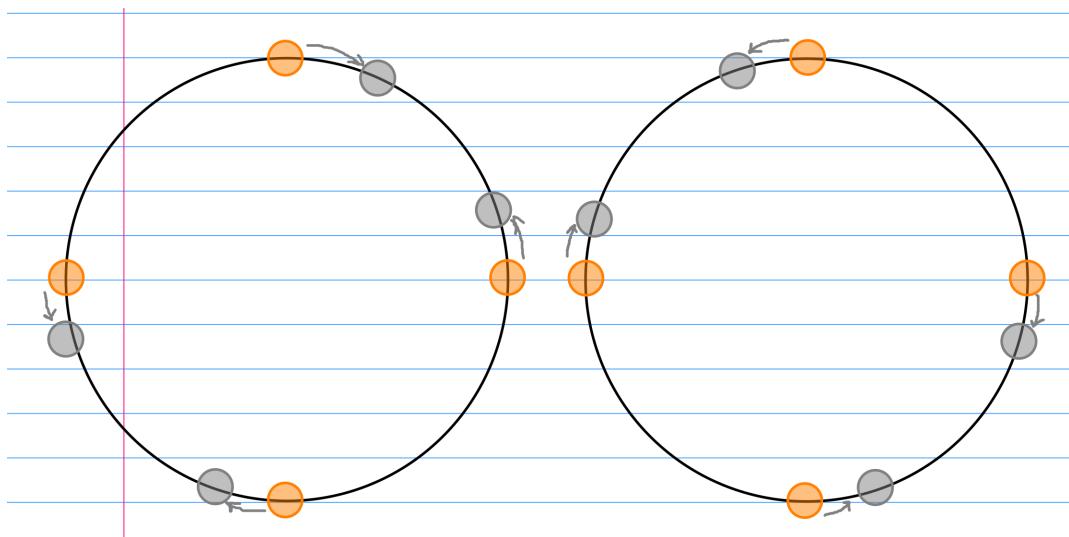


Figure 4: The mode for $(1, -1, 1, -1)$

Problem 5

(a)

Studying the problem

The initial configuration of the string

$$q(x, 0) = f(x) = \begin{cases} \frac{2h}{L}x & 0 \leq x \leq \frac{L}{2} \\ 2h - \frac{2h}{L}x & \frac{L}{2} \leq x \leq L \end{cases}$$

$$\frac{\partial q(x, t)}{\partial t} \Big|_{t=0} = \sum_{n=1}^{\infty} d_n \Omega_n \phi_n(x) = g(x) = 0 \implies \boxed{d_n = 0}$$

The general solution to the string equation (assumed solution is separable between time and position)

$$q(x, t) = \sum_{n=1}^{\infty} [c_n \cos(\Omega_n t) + d_n \sin(\Omega_n t)] \phi_n(x)$$

For $t = 0$ we get,

$$q(x, 0) = f(x) = \sum_{n=1}^{\infty} c_n \phi_n(x)$$

We are interested on finding the general solution of $q(x, t)$ that will hold for the future given this initial condition. The variables of our equation are obviously x, t and what we need to find out is c_n, d_n . The next sub-section will find out a solution for c_n (d_n is trivially zero given zero initial velocity).

Solving for c_n

Let us do the following computation now. Let us multiply both sides of the above equation with $\phi_p(x)$ where p represents the p -th term while we take a summation over the index of n .

$$f(x) \phi_p(x) = \sum_{n=1}^{\infty} c_n \phi_n(x) \phi_p(x)$$

Just so that we can invoke the inner product between orthonormal bases, we can take an integral with the following way

$$\begin{aligned} \int_0^L dx f(x) \phi_p(x) &= \int_0^L dx \left(\sum_{n=1}^{\infty} c_n \phi_n(x) \phi_p(x) \right) \\ &= \sum_{n=1}^{\infty} c_n \int_0^L dx \phi_n(x) \phi_p(x) \\ &= \sum_{n=1}^{\infty} c_n \delta_{np} \frac{L}{2} \\ &= c_p \frac{L}{2} \end{aligned}$$

This above gives us the p -th term

$$c_p = \frac{2}{L} \int_0^L dx f(x) \phi_p(x)$$

Using the explicit equation for the bases and also looking at the piecewise function, we can write,

$$\begin{aligned}
c_p &= \frac{2}{L} \int_0^L dx f(x) \sin\left(\frac{p\pi x}{L}\right) \\
&= \frac{2}{L} \left(\int_0^{\frac{L}{2}} f(x) \sin\left(\frac{p\pi x}{L}\right) + \int_{\frac{L}{2}}^L f(x) \sin\left(\frac{p\pi x}{L}\right) \right) \\
&= \frac{2}{L} \left(\int_0^{\frac{L}{2}} \frac{2h}{L} x \sin\left(\frac{p\pi x}{L}\right) + \int_{\frac{L}{2}}^L \left(2h - \frac{2h}{L}x\right) \sin\left(\frac{p\pi x}{L}\right) \right) \\
&= \frac{2}{L} \left(\frac{hL}{\pi^2 p^2} \left[2 \sin\left(p\frac{\pi}{2}\right) - \pi p \cos\left(p\frac{\pi}{2}\right) \right] - \frac{hL}{\pi^2 p^2} \left[2 \sin\left(\pi p\right) - 2 \sin\left(p\frac{\pi}{2}\right) - \pi p \cos\left(p\frac{\pi}{2}\right) \right] \right) \\
&= \frac{8h}{\pi^2 p^2} \sin\left(p\frac{\pi}{2}\right) \left[1 - \cos\left(p\frac{\pi}{2}\right) \right]
\end{aligned}$$

Hence if I write this huge mess properly

$$c_p = \frac{8h}{\pi^2 p^2} \sin\left(\frac{p\pi}{2}\right) \left[1 - \cos\left(\frac{p\pi}{2}\right) \right]$$

Discussion on odd and even modes

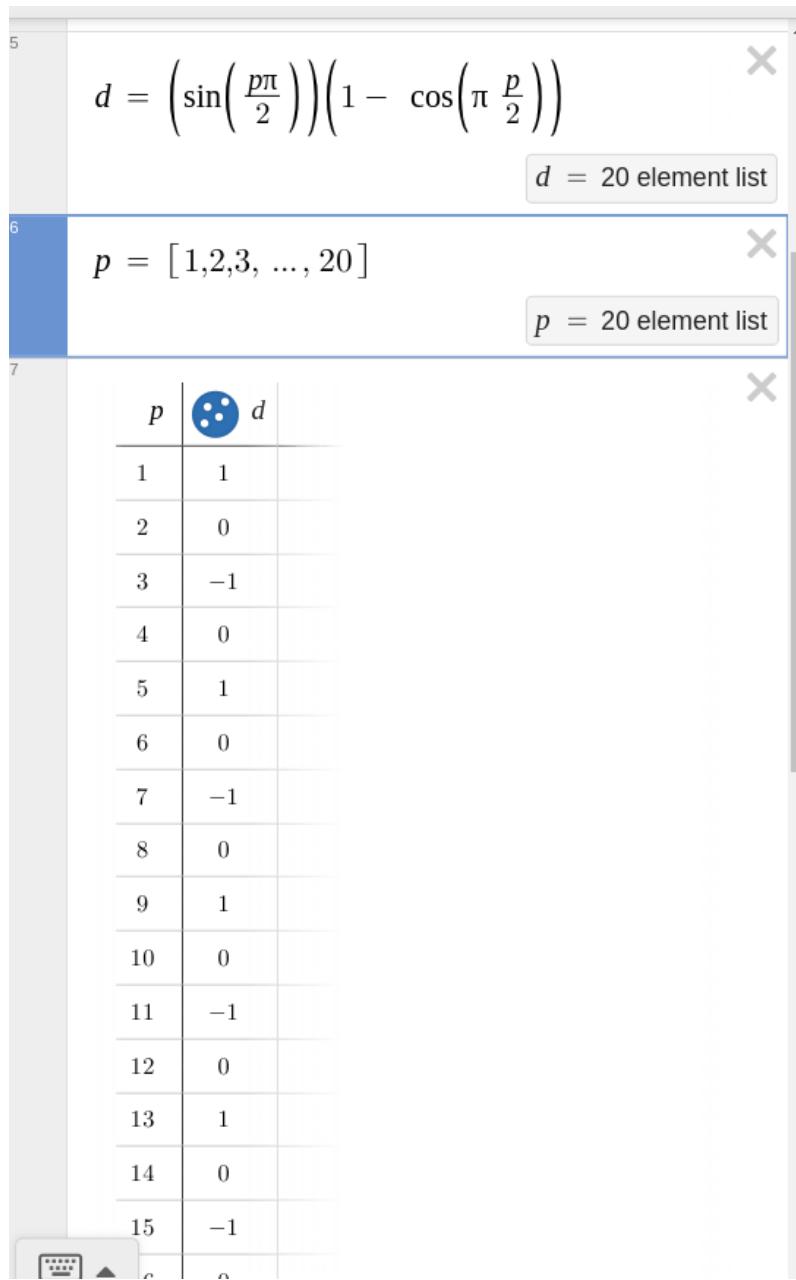


Figure 5: ss/graph-cn.png

We can see that every *even* numbers is giving us a 0 for c_p (d term in graph). The *odd* terms alter between -1 and 1 . The contribution of even and odd modes is obvious from here.

Plotting the displacement with time

Let's set $L = 1$ and we will look at $0 \leq x \leq 1$.

As given $\Omega_n = k_n = \frac{n\pi}{L} = n\pi$.

For h let's pick $h = 0.4$.

Using hte c_p we derived above, the displacement function $q(x, t)$ is then,

$$q(x, t)_{(10)} = \sum_{n=1}^{10} c_n \cos(n\pi t) \sin(n\pi x)$$

Expanding c_n ,

$$q(x, t)_{(10)} = \frac{8(0.4)}{\pi^2} \sum_{n=1}^{10} \frac{1}{n^2} \sin\left(\frac{n\pi}{2}\right) \left[1 - \cos\left(\frac{n\pi}{2}\right)\right] \cos(n\pi t) \sin(n\pi x)$$

$$\text{And } \tau = \frac{2\pi}{\Omega_1} = \frac{2\pi}{\pi} = 2$$

The first three non zero terms of the sum are

$$\frac{8h}{\pi^2} \left(\cos(\pi t) \sin(\pi x) + \left(-\frac{1}{9}\right) \cos(3\pi t) \sin(3\pi x) + \left(\frac{1}{25}\right) \cos(5\pi t) \sin(5\pi t) \right)$$

Plots of $q(x, t)$ where $\tau = 2$

$$q(x, t)_{(10)} = \frac{8(0.4)}{\pi^2} \sum_{n=1}^{10} \frac{1}{n^2} \sin\left(\frac{n\pi}{2}\right) \left[1 - \cos\left(\frac{n\pi}{2}\right)\right] \cos(n\pi t) \sin(n\pi x) \quad \text{and } x \in [0, 1]$$

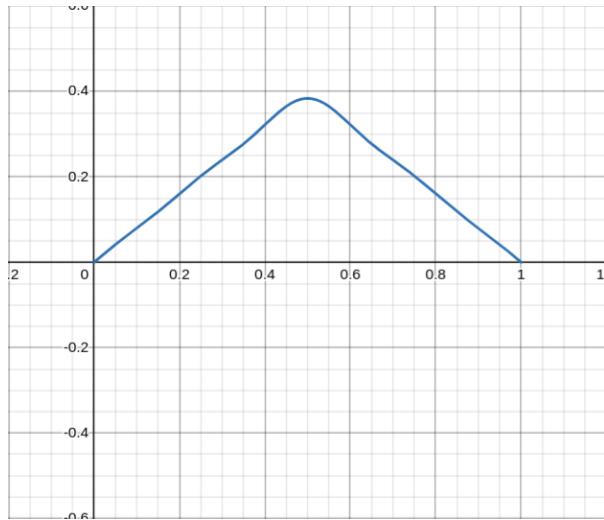


Figure 6: $t = 0$ and also $t = \tau$

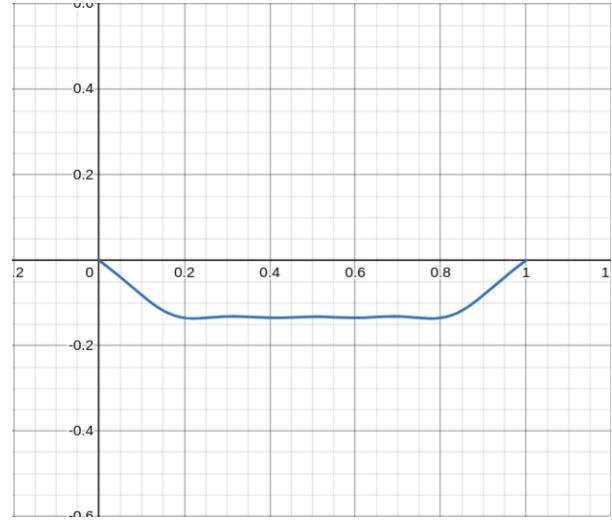


Figure 8: $t = \frac{\tau}{3} = \frac{2}{3}$

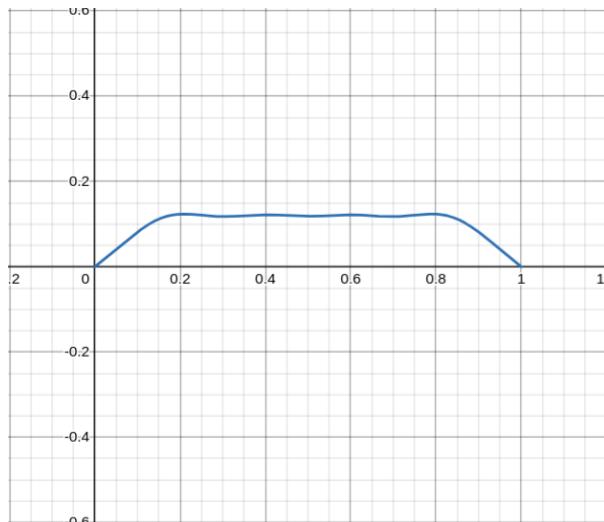


Figure 7: $t = 0.35$

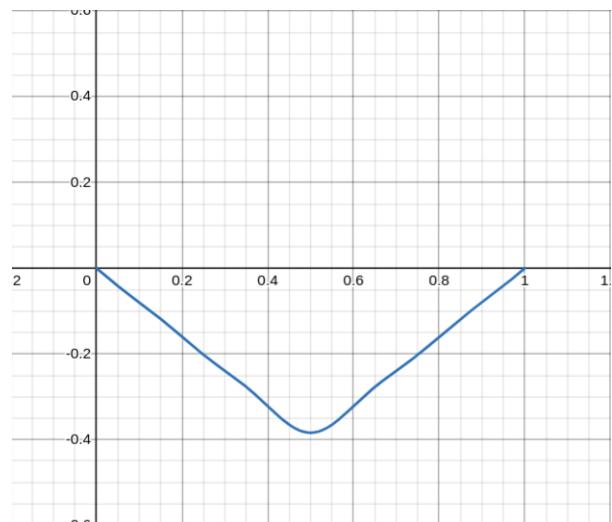


Figure 9: $t = \frac{\tau}{2} = 1$

(b)

Studying the Problem

Initially the string is tight and straight hence

$$q(x, 0) = \sum_{n=1}^{\infty} c_n \phi_n(x) = f(x) = 0 \implies [c_n = 0]$$

$$\frac{\partial q(x, t)}{\partial t} \Big|_{t=0} = g(x) = v_0 \theta \left(a - \left| x - \frac{L}{2} \right| \right)$$

Where $\theta(x)$ is a Heaviside step function (outputs 1 whenever input is 0 or positive). I am not going to waste my and graders time by re-writing everything I wrote above, the procedure we are going to follow is same as above.

Computation of d_n

$$\begin{aligned} g(x) &= \sum_{n=1}^{\infty} d_n \Omega_n \phi_n(x) \\ \int_0^L dx g(x) \phi_p(x) &= \sum_{n=1}^{\infty} \int_0^L dx d_n \Omega_n \phi_n(x) \phi_p(x) \\ \int_{\frac{L}{2}-a}^{\frac{L}{2}+a} v_0 \phi_p(x) &= d_p \Omega_p \frac{L}{2} \\ \int_{\frac{L}{2}-a}^{\frac{L}{2}+a} v_0 \sin(k_p x) &= d_p k_p \frac{L}{2} \\ \frac{v_0}{k_p} [-\cos(k_p x)]_{x=\frac{L}{2}-a}^{x=\frac{L}{2}+a} &= d_p k_p \frac{L}{2} \\ \cos\left(k_p \frac{L}{2} - k_p a\right) - \cos\left(k_p \frac{L}{2} + k_p a\right) &= \frac{d_p k_p^2 L}{2v_0} \\ \cos\left(\frac{n\pi}{2} - \frac{n\pi a}{L}\right) - \cos\left(\frac{n\pi}{2} + \frac{n\pi a}{L}\right) &= \frac{d_p n^2 \pi^2}{2v_0 L} \end{aligned}$$

This gives us

$$d_p = \frac{2v_0 L}{n^2 \pi^2} \left[\cos\left(\frac{n\pi}{2} - \frac{n\pi a}{L}\right) - \cos\left(\frac{n\pi}{2} + \frac{n\pi a}{L}\right) \right] = \frac{4v_0 L}{n^2 \pi^2} \sin\left(\frac{n\pi}{2}\right) \sin\left(\frac{n\pi a}{L}\right)$$

Discussion on Even and Odd modes

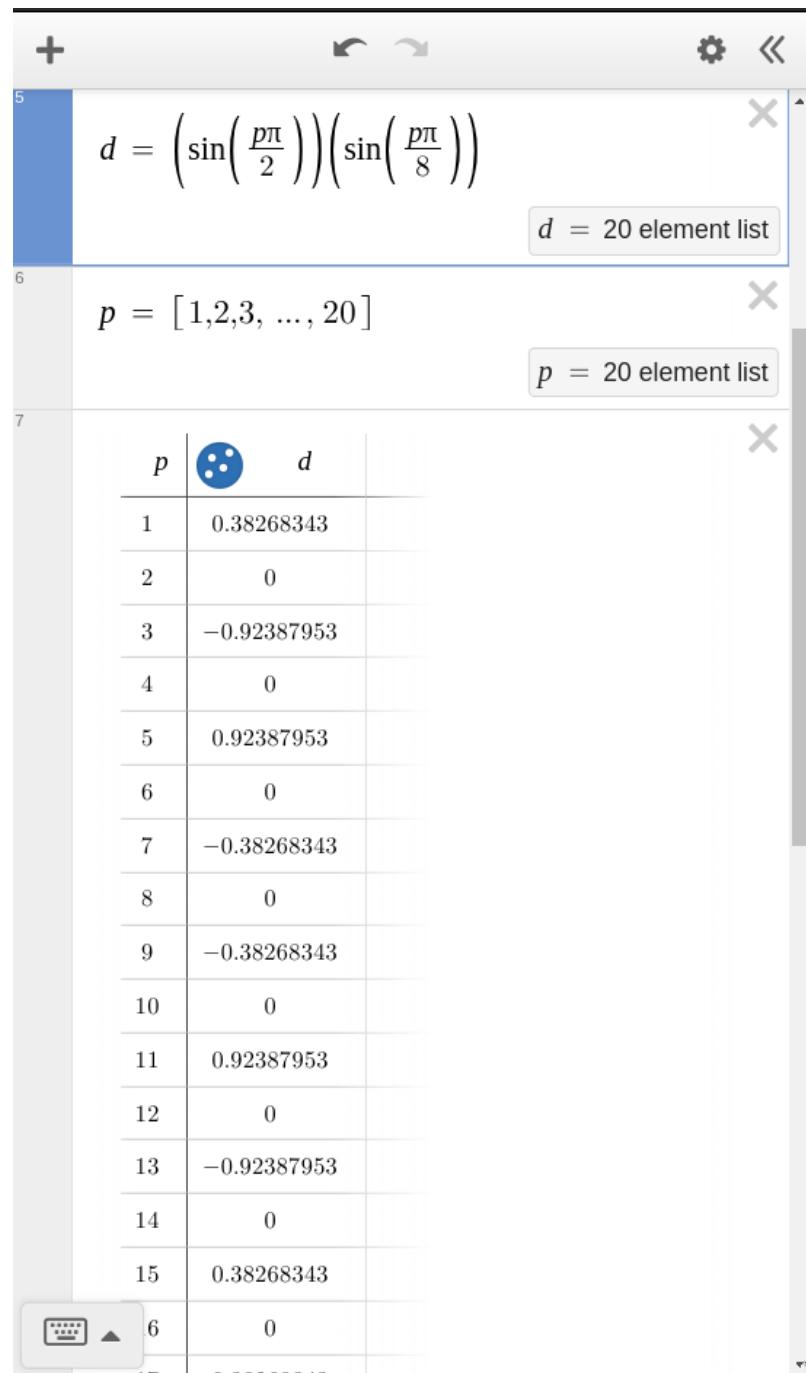


Figure 10: ss/graph-dn

Like before we can see every *even* modes are 0.

The first three terms of the expansion

$$\frac{4v_0L}{\pi^2} ((0.382)\sin(\pi t)\sin(\pi x) + (-0.923)\sin(3\pi t)\sin(3\pi x) + (0.923)\sin(5\pi t)\sin(5\pi x))$$

Plots of $q(x, t)$ where $\tau = 2$

$$q(x, t)_{(10)} = \frac{4(1.5)}{\pi^2} \sum_{n=1}^{10} \frac{1}{n^2} \sin\left(\frac{n\pi}{2}\right) \sin(n\pi(1/8)) \sin(n\pi t) \sin(n\pi x) \quad \text{and } x \in [0, 1]$$

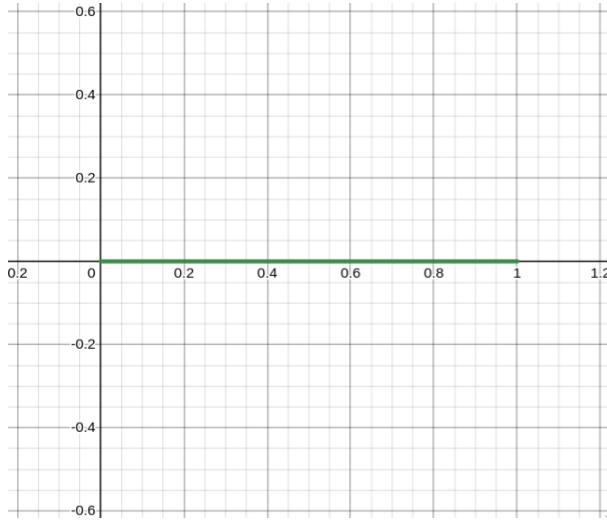


Figure 11: $t = 0$ and also $t = \tau$

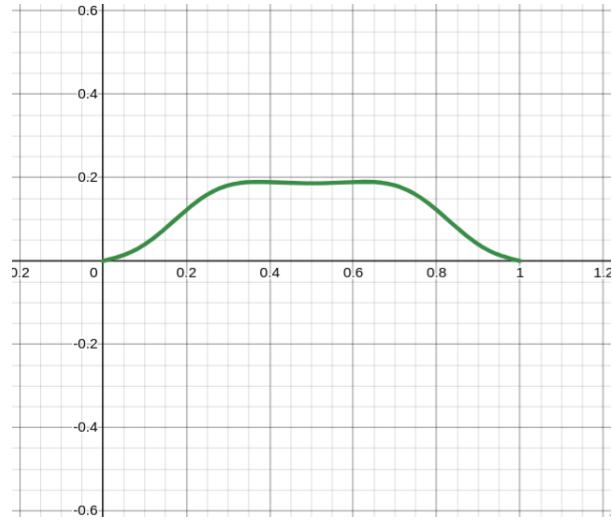


Figure 13: $t = \frac{\tau}{3} = \frac{2}{3}$

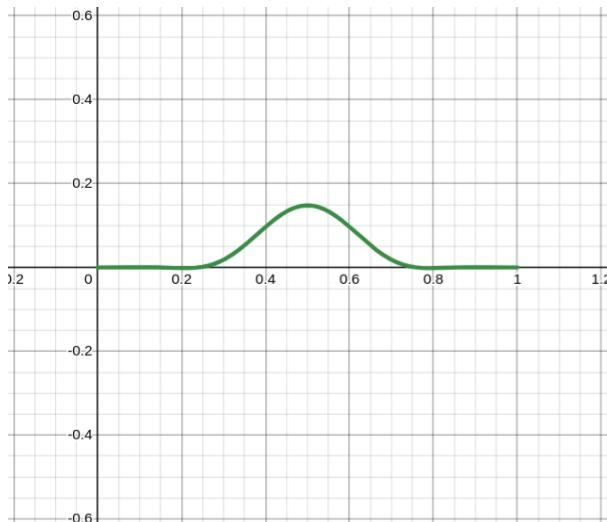


Figure 12: $t = \frac{\tau}{20} = 0.1$

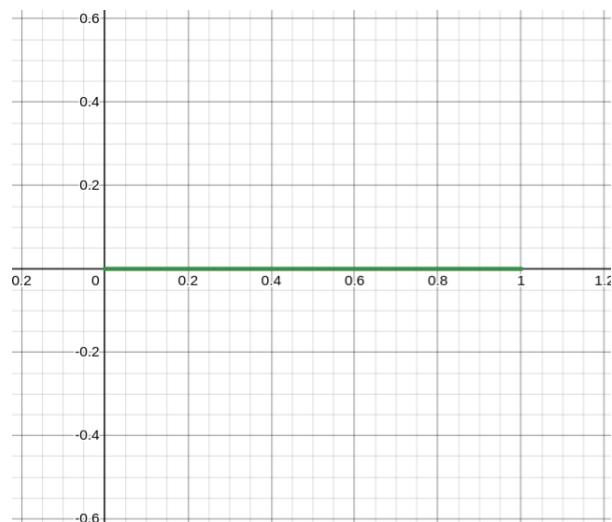


Figure 14: $t = \frac{\tau}{2} = 1$

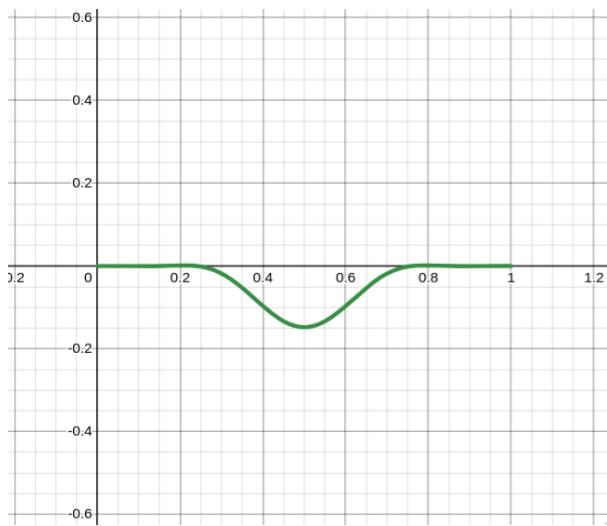


Figure 15: $t = 1.1$

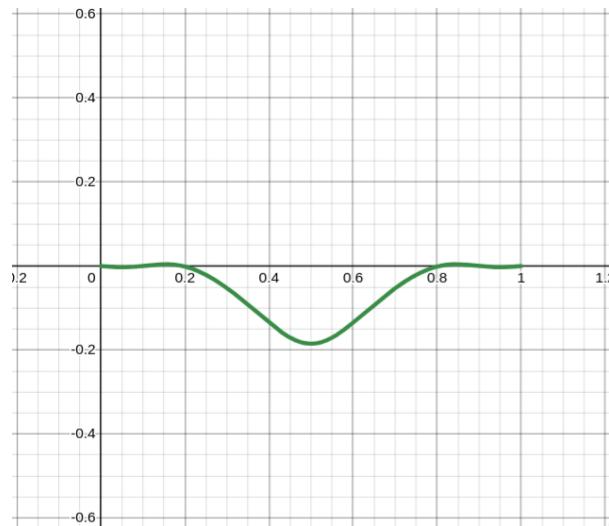


Figure 17: $t = 1.85$

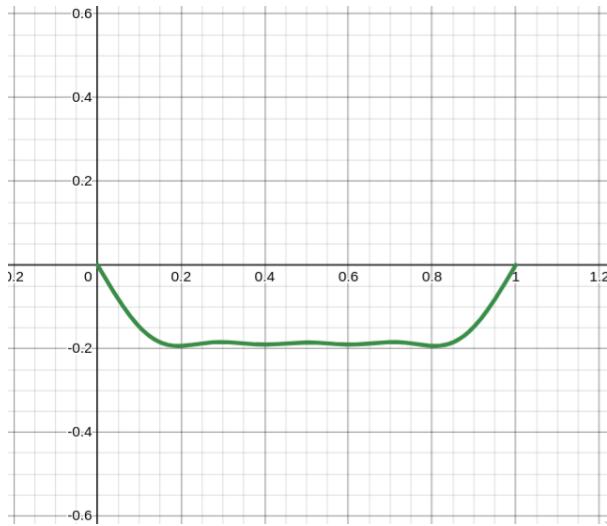


Figure 16: $t = 1.5$

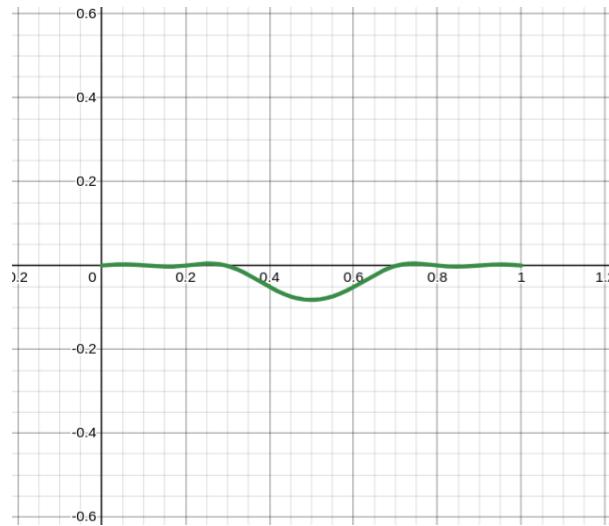
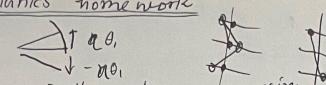


Figure 18: $t = 1.95$

Handwritten scratchwork as promised

classical mechanics homework

$(\theta_1, \theta_2, \theta_3)$ oscillations for the particles
 These should be symmetric

potential energy in the system can make the system oscillate missing total kinetic energy in the system and potential energy / second derivative

second derivative. $F = -\frac{d^2}{dx^2} \rightarrow \omega = \sqrt{\frac{k''(m)}{m}}$

there is constraint, let's look at the force

$$F_{12} = \frac{\eta k Q (\theta_2)}{b^2 + (a\theta_1 - a\theta_2)^2} \sin \alpha$$

$$\sin \alpha = \frac{a(\theta_1 - \theta_2)}{\sqrt{b^2 + (a\theta_1 - a\theta_2)^2}} \rightarrow F_{12} = \frac{\eta k Q^2 [a(\theta_1 - \theta_2)]}{(b^2 + (a\theta_1 - a\theta_2)^2)^{3/2}}$$

It's only from an keeping first order terms,

$$\frac{\eta k Q^2 [a(\theta_1 - \theta_2)]}{b^3 [1 + \frac{a(\theta_1 - \theta_2)}{b^2}]^{3/2}} \approx \frac{\eta k Q^2}{b^3} \cdot \frac{a(\theta_1 - \theta_2)}{1 + \frac{3}{2} \frac{a(\theta_1 - \theta_2)}{b^2}} \approx \frac{\eta k Q^2 a (\theta_1 - \theta_2)}{b^3}$$

$$\text{so, } F_{12} \approx \frac{\eta k Q^2 a (\theta_1 - \theta_2)}{b^3} \quad \& \quad F_{13} = -\frac{\eta k Q^2 a (\theta_1 - \theta_3)}{(2b)^3}$$

first particle force

on the second particle

$$F_{21} \approx \frac{\eta k Q^2 a (\theta_2 - \theta_1)}{b^3} = -\frac{\eta k Q^2 a (\theta_1 - \theta_2)}{b^3}$$

$$F_{23} \approx \frac{\eta k Q^2 a (\theta_2 - \theta_3)}{b^3}$$

on third particle

$$F_{31} \approx \frac{\eta k Q^2 a (\theta_3 - \theta_1)}{(2b)^3} \quad F_{32} \approx -\frac{\eta k Q^2 a (\theta_2 - \theta_3)}{b^3}$$

+ - + -

① $F_{12} = A(\theta_1 - \theta_2) \quad F_{13} = -B(\theta_1 - \theta_3)$

② $F_{21} = -A(\theta_1 - \theta_2) \quad F_{23} = A(\theta_2 - \theta_3)$

③ $F_{31} = B(\theta_1 - \theta_3) \quad F_{32} = -A(\theta_2 - \theta_3)$

$\theta_1 > \theta_2 > \theta_3$

then what

$$F_1 = A\theta_1 - A\theta_2 - B\theta_1 + B\theta_3 = (A-B)\theta_1 + (-A)\theta_2 + (B)\theta_3$$

$$F_2 = -A\theta_1 + A\theta_2 + A\theta_2 - A\theta_3 = (-A)\theta_1 + (2A)\theta_2 + (-A)\theta_3$$

$$F_3 = B\theta_1 - B\theta_3 - A\theta_2 + A\theta_3 = (B)\theta_1 + (-A)\theta_2 + (A-B)\theta_3$$

Now

$$\begin{bmatrix} \ddot{x}_1 \\ \ddot{x}_2 \\ \ddot{x}_3 \end{bmatrix} = \begin{bmatrix} (A-B) & -A & B \\ -A & 2A & -A \\ 0 & -A & (A-B) \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix}$$

$$\ddot{x}_1 = \alpha_1 \theta_1$$

$$\ddot{x}_2 = \alpha_2 \theta_2$$

$$\ddot{x}_3 = \alpha_3 \theta_3$$

$$A = \frac{\eta k Q^2 a}{mb^3 (a)}$$

$$B = \frac{\eta k Q^2 a}{m(2b)^3 (a)}$$

Figure 19: ..//phys311/ss/11/76.jpeg

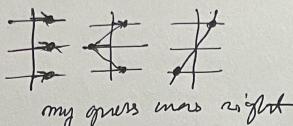
$$\begin{bmatrix} \ddot{\theta}_1 \\ \ddot{\theta}_2 \\ \ddot{\theta}_3 \end{bmatrix} = \begin{bmatrix} A-B & -A & B \\ -A & 2A & -A \\ B & -A & A-B \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix}$$

$A =$

eigenvectors and eigenvalues for this case are

$$\lambda_1 = 0, \lambda_2 = 3\omega, \lambda_3 = \omega - 2\omega$$

$$\hat{x} \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix}, \begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix}, \begin{bmatrix} 1 \\ 0 \\ 1 \end{bmatrix}$$



my guess was right

$$\begin{bmatrix} \ddot{\theta}_1 \\ \ddot{\theta}_2 \\ \ddot{\theta}_3 \end{bmatrix} = \begin{bmatrix} \omega^2 & 0 & 0 \\ 0 & \omega^2 & 0 \\ 0 & 0 & \omega^2 \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \end{bmatrix}$$

problem 05

pick i th mass, displacement of each is $\Delta\theta$

$$\begin{array}{l} \uparrow F_{i+1} \\ \downarrow F_{i-1} \end{array} \quad F_i = -F_{i-1} + F_{i+1} = -K(\theta_i - \theta_{i-1}) + K(\theta_{i+1} - \theta_i) \\ = -K\theta_i + K\theta_{i-1} + K\theta_{i+1} - K\theta_i \\ = -2K\theta_i + K\theta_{i-1} + K\theta_{i+1}$$



$$\begin{aligned} F_1 &= -2K\theta_1 + K\theta_N + K\theta_2 \\ F_2 &= -2K\theta_2 + K\theta_1 + K\theta_3 \\ F_3 &= -2K\theta_3 + K\theta_2 + K\theta_4 \\ &\vdots \\ F_{N-1} &= -2K\theta_{N-1} + K\theta_{N-2} + K\theta_N \\ F_N &= -2K\theta_N + K\theta_{N-1} + K\theta_1 \end{aligned}$$

$$\begin{bmatrix} \ddot{\theta}_1 \\ \ddot{\theta}_2 \\ \ddot{\theta}_3 \\ \vdots \\ \ddot{\theta}_N \end{bmatrix} = \begin{bmatrix} -2K & K & 0 & \dots & 0 & 0 & K \\ K & -2K & K & \dots & 0 & 0 & 0 \\ 0 & K & -2K & K & \dots & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\ 0 & 0 & 0 & 0 & \dots & K & -2K \end{bmatrix} \begin{bmatrix} \theta_1 \\ \theta_2 \\ \theta_3 \\ \vdots \\ \theta_N \end{bmatrix}$$

-2K K crongy ?

Figure 20: ..//phys311/ss/11/76.jpeg