

# Waves II: Interference and Diffraction

Interference and diffraction are covered in chapters 41, 42, and 43 of Halliday and Resnick, and chapter 14 of Wang and Ricardo, volume 1. For more mathematical detail, see chapter 9 of Crawford's *Waves*. For lighter reading with neat examples, see chapters I-29 and I-30 of the Feynman lectures. For much more, see chapters 9 and 10 of *Optics* by Hecht, a well-written but somewhat long-winded university level text. There is a total of **66** points.

## 1 Double Slit Interference

### Remark: Coherence

We introduced interference in **W1**. However, in many cases interference for optical light can't be observed at all, because the relative phase between the two waves will oscillate rapidly, making the interference term cancel out. We say waves are "coherent" if their phase relation is stable enough for interference effects to be seen, and this is difficult with natural sources.

If you superpose light of wavelength 600 nm and 600.0001 nm, their frequency difference will be  $\sim 10^9$  Hz. Interference effects will oscillate at this frequency, and thus be impossible to see with the eye. Now, some lamps generate light with transitions between atomic energy levels, so that they ideally only have one output frequency. But you still wouldn't see interference if you used two identically constructed lamps, because tiny effects like thermal fluctuations will make the frequency wobble by enough to wash out interference effects.

So how did anybody ever observe light interference in the 19th century? It helps a lot to use a *single* lamp, since the light it produces is presumably in phase with itself. But that isn't quite enough, because the light received from the lamp at different points ultimately came from different parts of the lamp. The solution was to put a tiny (sub-millimeter) pinhole between the light source and the actual experiment; if the pinhole is small enough, then all the light going through the pinhole at a given moment is in phase with *itself*. After passing through the pinhole, the light spreads out again and can be directed, e.g. to the two slits of a double slit experiment. The pinhole is so effective that the source can even be the Sun.

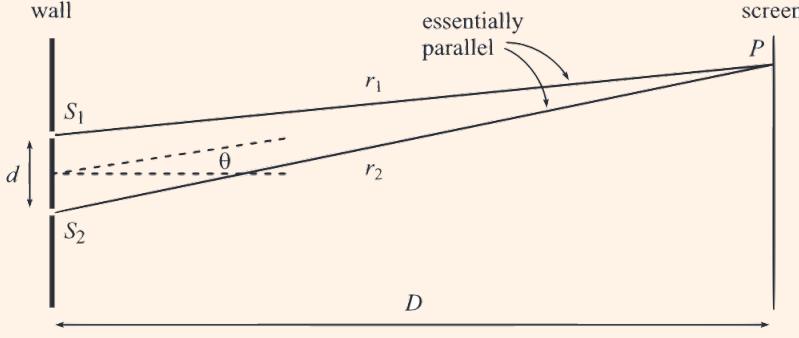
Today, carrying out interference experiments is much easier, because lasers produce extremely coherent light. One can simply fire an ordinary laser pointer at a double slit. If you work even harder, it is possible, though challenging, to observe interference between two *different* lasers of the same type. Anyway, for the problems below we'll generally assume coherence is perfect, but one should remember it's always an important consideration in real setups.

### Example 1

Derive the far-field intensity pattern for the double slit experiment.

**Solution**

We suppose a plane wave with wavenumber  $k$  is incident on two small slits. Using Huygens' principle, the screen absorbs all wavelets except the ones at the slits, so it's as if we have two spherical waves coming from each slit. They then travel to a screen a distance  $D$  away.



The amplitude of the light on the screen at some height  $y$  is

$$A(t) \propto e^{i(kr_1 - \omega t)} + e^{i(kr_2 - \omega t)}$$

where we've switch to complex notation, and the quantities  $r_1$  and  $r_2$  depend on  $y$ . In principle the two terms don't need to have the same magnitude, e.g. if the two slits are not equal in size. Additionally, as you saw in **W1**, for spherical waves the amplitude actually falls as  $1/r$ . However, in the far-field limit we have  $r_1 \approx r_2$ , so we neglect this effect.

At each point on the screen, the amplitude is some constant times the time-varying phase,

$$A(t) \propto (e^{ikr_1} + e^{ikr_2})e^{-i\omega t} = A_0 e^{-i\omega t}.$$

The intensity of the light we see at that point is proportional to  $|A_0|^2$ , using the results about wave energy from **W1**. Since only the coefficient squared matters, we can factor out a common phase to get

$$A_0 \sim e^{-ik\Delta r/2} + e^{ik\Delta r/2}, \quad \Delta r = r_2 - r_1$$

and the intensity is

$$I \propto |A_0|^2 \propto |e^{-ik\Delta r/2} + e^{ik\Delta r/2}|^2 = |2 \cos(k\Delta r/2)|^2 \propto \cos^2(k\Delta r/2).$$

Next, we find the path length difference  $\Delta r$ . For far-field diffraction, using  $D \gg d$ ,

$$\Delta r = r_2 - r_1 \approx d \sin \theta$$

where  $\theta$  is the angle from the slits to the point on the screen we're looking at. Therefore,

$$I(\theta) \propto \cos^2(kd(\sin \theta)/2).$$

This yields a series of bright and dark bands on the screen. Most of the time, we'll only be concerned with the few most central minima and maxima, which for typical parameter values lie at  $\theta \ll 1$ . Then we may use the small angle approximation  $\sin \theta \approx \tan \theta = y/D$ , giving

$$I(y) \propto \cos^2 \frac{kdy}{2D}.$$

We have a periodic pattern of dark and light fringes. The separation between the minima and between the maxima on the screen is  $D\lambda/d$ .

The set of approximations made above gives us the theory of far field (“Fraunhofer”) diffraction. For example, when we calculated the path length difference, which was of order  $d$ , we neglected a higher order term of order  $d(d/D)$ . This is acceptable as long as this is small compared to the wavelength  $\lambda$ . Thus, defining the Fresnel number  $F = d^2/D\lambda$ , we have far field diffraction when  $F \ll 1$ . For  $F \sim 1$  we must use the more complicated near field (Fresnel) diffraction. For  $F \gg 1$ , the bands get spaced so close together that they end up being washed out, and interference effects stop being visible. Here we just use geometrical optics, as covered in **W3**.

- [1] **Problem 1.** Sketch the intensity on the screen as a function of  $\theta$  if  $kd = 8$ . What are the bounds on  $\theta$ ? How many completely dark points exist?
- [3] **Problem 2.** Some basic but important questions about the double slit result.
  - (a) Explain intuitively how the pattern on the screen changes if the slits are brought closer together. Do the same if the wavelength of the light is increased.
  - (b) Suppose the light going through the top slit has a small additional phase shift of  $\theta$ , e.g. because it passes through a piece of glass in the slit. Describe how the interference pattern changes.
  - (c) Now suppose the top slit is twice as wide. Describe the interference pattern.
  - (d) Suppose that coherent white light is passed through the slits. Describe the interference pattern.
  - (e) Usually, the double slit experiment is done by shining a laser on both slits. Now suppose a different laser was aimed on each slit. Assume, somewhat unrealistically, that the lasers have perfectly stable but slightly different frequencies, say  $\Delta f = 1$  Hz. Describe what the interference pattern would look like.

- [3] **Problem 3.**  USAPhO 1999, problem A3. A triple slit experiment.

- [2] **Problem 4.** A pair of slits is separated by a distance  $d_1$ , and two of these pairs are separated by a larger distance  $d_2$ , so that  $d_2 \gg d_1 \gg \lambda$ . Sketch the intensity pattern on the screen for this four-slit apparatus. (Hint: to avoid a complicated computation, factor the expression for the amplitude.)

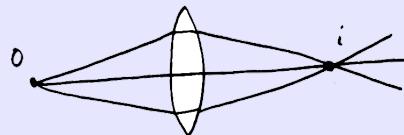
### Idea 1: Image Sources

Some interference problems have complex arrangements of mirrors and lenses. In these cases, actually computing the path length differences can be a nightmare. For instance, you’d have to account for the detailed shape of every lens. Also, you won’t just have to compute the path length, but rather than [optical path length](#), which is the ordinary path length weighted by the index of refraction. This is because the index of refraction affects the wavelength and hence the phase difference.

However, there’s a trick which makes everything much simpler: any point image can be treated like its own light source. That means you can compute path length differences by

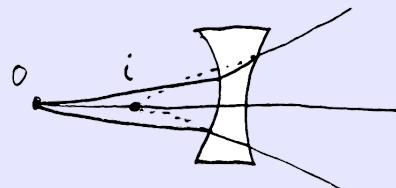
starting from the images, rather than having to go all the way back to the original objects.

For real images, there's a very simple way to see why this works. For instance, consider the setup below, where an object  $o$  is focused with a lens to an image  $i$ .



Fermat's principle of least time tells us that all of the paths shown take the same time, and since phases are directly related to time by  $\Delta\phi = \omega\Delta t$ , it means that all of the rays arrive at the image with the same phase. That means they leave the image with the same phase, so the image can be treated just like a coherent source. (That is, the phase of the light coming from the image doesn't depend on the direction it comes out.) To find the phase of the waves at  $i$ , you can pick any of the paths, most conveniently the one on the symmetry axis.

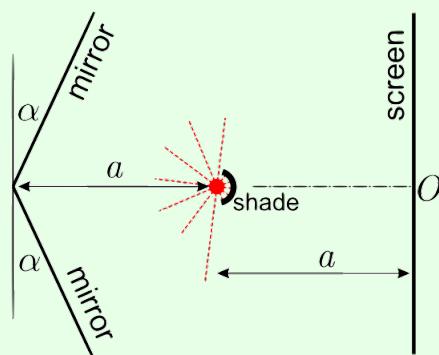
We can also consider virtual images, as shown below.



Here, the Fermat's principle argument doesn't work because the rays never actually meet at  $i$ , but we can use Huygens' principle. The key ideas are that (1) light locally propagates perpendicularly to wavefronts, and (2) the phase on a wavefront is always uniform, by definition. The first point implies the outgoing wavefronts are spheres centered on  $i$ . The second point implies that the phase only depends on the distance from  $i$ , so it can again be treated just like a source. In this way, seemingly impossible questions can be solved instantly.

### Example 2: Kalda 17

Consider the optical setup shown below.



Many light and dark bands appear on the screen, with dark bands separated by distance  $d$ . Assuming that  $\alpha \ll 1$ , find the wavelength  $\lambda$  of the light.

### Solution

This is actually just a double slit interference problem! Each mirror produces a (virtual) image source reflected behind it, and the pattern on the screen results from the interference between the two image sources, just as if there were two slits at those points.

Specifically, let the light source have coordinates  $(0, 0)$ , with the screen at  $x = a$ . Then the image sources are located at  $(-2a, \pm 2\alpha a)$ , so we have a double slit setup with sources  $4\alpha a$  apart from each other, a distance  $3a$  from the screen. Using our existing results,

$$\lambda = (4\alpha a) \frac{d}{3a} = \frac{4\alpha d}{3}.$$

Note that reflection from a mirror changes the phase by  $\pi$ , but that didn't matter in this problem, because both image sources pick up the same phase.

- [3] **Problem 5.** (1) USAPhO 2020, problem B2. A problem on interference with images.

## 2 Thin Film Interference

### Idea 2

In general, the phase of a wave is unaffected by reflection from a rarer medium, and flipped by  $180^\circ$  when reflected from a denser medium; here a “denser” medium is defined as one where the wave speed is lower, e.g. one with a higher index of refraction for light. This is analogous to the result for wave reflection in a string derived in **W1**, and is derived starting from Maxwell's equations in **E8**.

### Example 3

Why does a very thin soap film on a wire loop looks dark when viewed from above?

### Solution

For the soap film, we consider interference between two paths for the light: bouncing off the top surface, or transmitting through and bouncing off the bottom surface. These have almost the same phase from their path length, since the soap film is thin, but the former has an extra  $180^\circ$  phase shift. So the two destructively interfere, making the soap film look dark.

### Remark

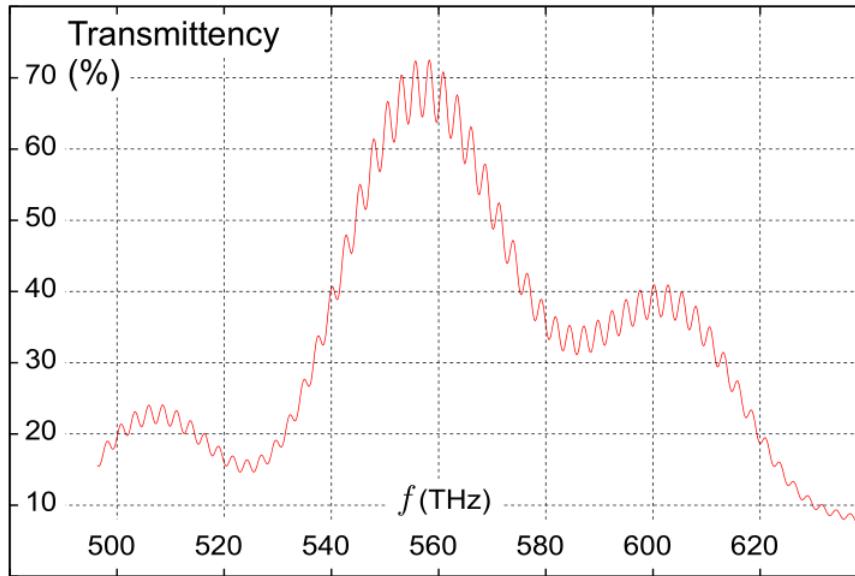
This is the usual high school textbook analysis, but the real situation is a bit more subtle. First off, there are actually infinitely many possible paths for the light to take, and sometimes many of these paths are important, as you saw in **W1** for the Fabry–Perot interferometer

for  $r \approx 1$ . However, the two paths we considered were indeed the most important by far.

Second off, the amplitudes upon reflection and transmission must be computed using the results you found in **W1**, and generally won't have the same magnitude. This means that generically we don't get complete destructive interference, just a lowered intensity. Finally, these reflection and transmission coefficients will vary significantly with angle according to Fresnel's equations, as shown in **E8**. We typically ignore this by focusing on normal incidence.

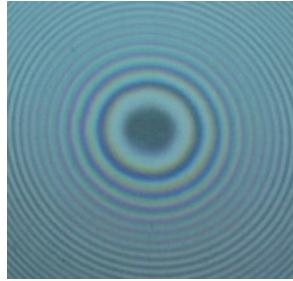
By the way, you might be wondering why we're specializing to thin films; why isn't there thick film interference? Technically there could be interference fringes, but they would be too close to see even if everything was perfect. And in reality, they would then get blurred together due to imperfections in the surfaces, and the spread of frequencies and incidence directions in the incoming light.

- [2] **Problem 6.** The anti-reflection coating on your glasses consists of a thin layer of material whose index of refraction is between that of air ( $n = 1$ ) and glass ( $n = 1.5$ ). The coating is designed to eliminate the reflection of green light,  $\lambda = 550$  nm.
- Accounting for only the two most significant paths for the light, find the minimum possible thickness of the coating, and its index of refraction. (You'll need results for reflection and transmission coefficients from **W1** or **E8**, and a computer to numerically solve an equation.)
  - Roughly how much does the next most significant path contribute to the reflected intensity?
- [2] **Problem 7** (EFPhO 2004). A thick glass plate is coated by a thin transparent film. The emission spectrum of the system at normal incidence is as shown.



The refractive index of the film is  $n = 1.3$ . Find the thickness of the film.

- [2] **Problem 8.** Newton's rings are an interference pattern formed when a lens is placed on a flat glass surface and illuminated from above by light of wavelength  $\lambda$ . For concreteness, suppose the side of the lens touching the surface is spherical, with radius of curvature  $R$ . When viewed from above, one sees an interference pattern with circular fringes.



The most important reflection paths are (1) reflection off the top surface of the lens, (2) reflection off the bottom surface of the lens, and (3) reflection off the flat surface. However, the first reflection path has a very different path length from the other two, which means it won't give rise to visible interference fringes, as explained in the remark above. Instead, the first reflection path just adds some background intensity everywhere, preventing the dark fringes from being perfectly dark. Thus, in this problem we'll only consider the second and third paths.

- (a) Explain why the center of the pattern is dark.
  - (b) Find the radii of the bright and dark fringes, i.e. the values of  $r$  where there is a local minimum or maximum of the intensity. For simplicity, assume  $r \ll R$ .
- [3] **Problem 9** (Kalda). A hall of a contemporary art installation has white walls and a white ceiling, lit with a monochromatic green light of wavelength  $\lambda = 550 \text{ nm}$ . The floor of the hall is made of flat transparent glass plates. The lower surfaces of the glass plates are matte and painted black; the upper surfaces are polished and covered with thin transparent film. A visitor standing in the room will see circular concentric bright and dark strips on the floor, centered around himself. A curious visitor observes that the stripe pattern depends on their height, and upon lowering themselves, sees a maximum of 20 stripes. The film's index of refraction is 1.4 and the glass's is 1.6. Determine the thickness of the film.

### 3 Diffraction

Next, we turn to diffraction.

#### Example 4

Find the interference pattern of a diffraction grating, a set of  $N$  identical slits in a row, each separated by a distance  $d$ .

#### Solution

Defining  $\Delta r = d \sin \theta$  as before, the amplitude is

$$A \sim 1 + e^{ik\Delta r} + e^{2ik\Delta r} + \dots + e^{(N-1)ik\Delta r} = \frac{e^{ikN\Delta r} - 1}{e^{ik\Delta r} - 1}.$$

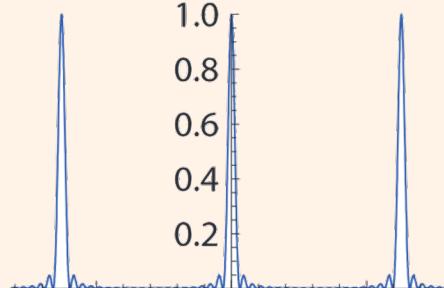
Factoring out a common phase, we have

$$A \sim \frac{e^{ikN\Delta r/2} - e^{-ikN\Delta r/2}}{e^{ik\Delta r/2} - e^{-ik\Delta r/2}} = \frac{\sin(Nk\Delta r/2)}{\sin(k\Delta r/2)}$$

so the intensity is

$$I \propto \frac{\sin^2(Nk\Delta r/2)}{\sin^2(k\Delta r/2)}.$$

The normalized intensity is plotted below as a function of  $\theta$ .



The numerator yields rapid oscillations which aren't very visible; their envelope is given by the slow oscillations in the denominator. These slow oscillations are the ones we care about; they are the diffraction peaks and occur when

$$\frac{k\Delta r}{2} = n\pi, \quad d \sin \theta = \frac{2\pi n}{k} = n\lambda, \quad n \in \mathbb{Z}.$$

This is intuitive, because at the maxima, the contributions from each slit are in phase, as the path length difference between adjacent slits is a multiple of  $\lambda$ .

- [3] **Problem 10.** For practical applications of diffraction gratings, we usually focus on the intensity maxima. However, around each maximum there are also secondary maxima.

- (a) Argue that the first minimum occurs when there is a phase difference of  $2\pi/N$  between adjacent slits, then compute its angle. This is easiest to see using phasors, i.e. by drawing the individual terms in the amplitude  $A$  as vectors in the complex plane.
- (b) Show that each secondary maximum is *half* as wide as the central maximum.
- (c) Let the intensity at the central maximum be  $I_0$ . Assuming  $N \gg 1$ , use phasors to show that the intensity at the  $k^{\text{th}}$  adjacent secondary maximum is roughly  $I_0/((k + 1/2)^2\pi^2)$ .

This final result shows that in general, the secondary maxima are much dimmer, and can be neglected; we will ignore them for almost all problems below.

### Example 5

Use the previous example to get the interference pattern for a single wide slit of width  $a$ .

### Solution

We can think of a single slit as the limit of a diffraction grating with  $Nd = a$ , where  $N \rightarrow \infty$  and  $d \rightarrow 0$ . Taking these two limits simultaneously is a bit delicate. Starting with our previous result

$$A \sim \frac{\sin(Nkd \sin \theta/2)}{\sin(kd \sin \theta/2)}$$

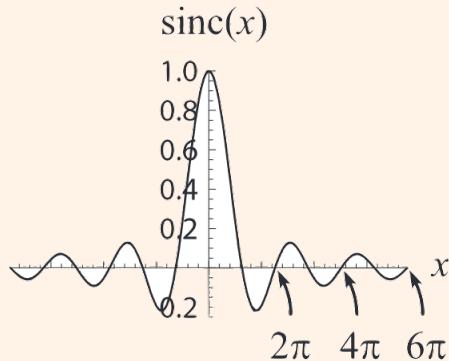
we may substitute  $Nd = a$  in the numerator. Only  $d$  remains in the denominator, so the  $d \rightarrow 0$  limit allows us to use the small angle approximation. We thus have

$$A \sim \frac{\sin(ka \sin \theta/2)}{kd \sin \theta/2}.$$

But as  $d \rightarrow 0$  the expression blows up, because we're taking the number of slits to infinity while keeping the amplitude from each slit constant. To get a consistent limit, we normalize by dividing the amplitude by  $N$ , giving  $Nd = a$  in the denominator for

$$A \sim \frac{\sin \beta}{\beta} \quad \beta = \frac{ka \sin \theta}{2}$$

The amplitude is proportional to the sinc function, shown below.



What we're really doing here is zooming in on the central maximum of the diffraction grating; the other maxima have been removed by sending the slit spacing to zero.

### Remark: Uncertainty Principle

There's a neat way to rephrase our results. In the far field limit and small angle approximation, an opening at height  $z$  gives a wave with amplitude  $e^{i(k/D)yz}$  at height  $y$  on the screen. If we think of a "slit function"  $f(z)$  which is equal to one at holes and zero elsewhere, then the amplitude at the screen is simply the Fourier transform of  $y$ ! Phasors are just a visual way to compute the Fourier transform.

We won't use this language explicitly below, but it can add some intuition if you know it. For example, we know from **W1** that the products of the widths of any function and its Fourier transform are bounded. For example, a wavepacket of width  $\Delta x$  with Fourier components of width  $\Delta k$  obeys  $\Delta x \Delta k \gtrsim 1$ .

In this case, the Fourier pair is screen height  $y$  and the scaled slit height  $(k/D)z$ . (Don't get confused with the notation here; now  $k$  is fixed while  $z$  varies.) Hence the uncertainty principle says

$$\Delta y \Delta z \gtrsim \frac{D}{k} \sim D\lambda$$

which you can check holds for all the examples we've seen so far. The uncertainty principle provides a simple explanation for why making the slits narrower makes the pattern wider.

In fact, this is equivalent to the Heisenberg uncertainty relation  $\Delta y \Delta p_y \gtrsim \hbar$  for photons passing through the slit, as you can verify. This makes sense, as we should be able to calculate the diffraction pattern in terms of either the whole light wave, or in terms of what happens to each of the photons in the light wave.

### Remark

If we take the limit  $a \rightarrow \infty$  for the single slit, the central maximum becomes an infinitely sharp, bright point. But in reality, a light will just uniformly illuminate the screen.

The problem is that when  $a$  gets too high, the approximations of Fraunhofer diffraction break down, and we must switch to Fresnel diffraction. Fresnel diffraction augments Fraunhofer diffraction with two additional effects.

1. The amplitude of each wavelet falls off as  $1/r$ .
2. The amplitude of each wavelet is proportional to the “obliquity factor”  $(1 + \cos \theta)/2$ , where  $\theta$  is the angle from its original forward direction of propagation. (Strictly speaking, this factor appears in Fraunhofer diffraction too, but in that case it’s not too important, because all the wavelets that reach a given point of the screen have about the same  $\theta$ .)

Both effects matter, but it suffices to consider the first to fix the problem. This amplitude falloff implies that in the case  $a \gg D$ , the illumination at each point on the screen mostly comes from points on the slit within a distance  $D$ , not from the entire slit. Since every point on the screen can see such a range of points, the screen is uniformly illuminated.

For points on the screen near the edge of the slit, there is a gradual shadow, along with some interference bands from “edge diffraction”. In the limit  $D \gg \lambda$  these residual diffraction peaks get very close and blur together, leaving only a smooth shadow. This is just as expected, as in this case we have  $F \gg 1$  and geometrical optics should apply.

For a derivation of Fresnel diffraction starting from the wave equation, see section 10.4 of Hecht. Incidentally, Fresnel diffraction came first historically, since reaching the simpler Fraunhofer regime  $F \ll 1$  requires manufacturing tiny optical instruments. This is yet another example of how the textbook treatment we enjoy today is easier. We can start with the simple case, but the pioneers had to get it all right at once.

### Remark: Interference vs. Diffraction

Interference is the name for the fact that when waves superpose, their energy doesn’t just add; it can become larger than the sum (constructive interference) or smaller (destructive interference). Diffraction is the name for the fact that waves do not need to keep going in a straight line when they hit an obstacle.

What’s confusing is that double slit “interference”, single slit “diffraction”, and a many-slit “diffraction” grating all involve *both* interference and diffraction. Why is it called diffraction

when there are one or many holes, but not when there's two? I was very confused about this in high school, but I'm pretty sure there is no difference; it's just historical convention. (However, one pattern is that things that have maxima at larger angles tend to be called "diffraction", because it's more apparent that the direction of the light has been changed.)

**[2] Problem 11.** Comparing the single and double slit.

- (a) Show that the minima for the single slit occur when

$$a \sin \theta = n\lambda, \quad n \neq 0.$$

- (b) Note that this looks almost identical to the result for the *maxima* of a double slit with separation  $a$ . Explain the difference using phasors.  
(c) In our analysis of the double slit, we didn't account for the small but nonzero width of each slit. Using the idea of problem 4, sketch the diffraction pattern accounting for this.

For tips on accurately measuring diffraction patterns, see the handout on experimental methods.

**[3] Problem 12 (MPPP 123).** Some imperfect diffraction gratings. For this problem, you can ignore secondary maxima. Use the small angle approximation throughout, and neglect any diffraction effects from the finite widths of the slits.

- (a) In an imperfect diffraction grating, the slits have equal widths, but the distances between the slits are alternately  $d$  and  $3d$ . Sketch the resulting diffraction pattern, indicating the relative heights of the maxima.  
(b) In another imperfect diffraction grating, the slits are evenly spaced, but their widths are alternatively  $a$  and  $b$ , where  $a \approx b$ . Sketch the resulting diffraction pattern, indicating the relative heights of the maxima.

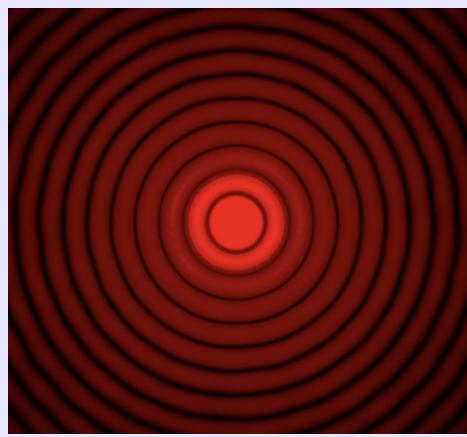
It may be useful to refer to problem 4.

## 4 Higher Dimensions

In these problems we tackle interference and diffraction effects in more than one dimension, which can be used to infer the structure of molecules and crystals.

### Idea 3

We've shown that in a single narrow, tall slit, the central maximum is a tall band, bounded by minima at  $\theta = \pm\lambda/d$ . If we instead had a circular slit, the central maximum is a circle, bounded by minima at  $\theta \approx 1.22\lambda/d$ . The radii of the higher-order minima then get closer and closer spaced as one moves outward. The resulting pattern is called an Airy disc.



You can straightforwardly write down an integral that gives the intensity  $I(r)$ , but the integral can only be performed in terms of special functions, called Bessel functions.

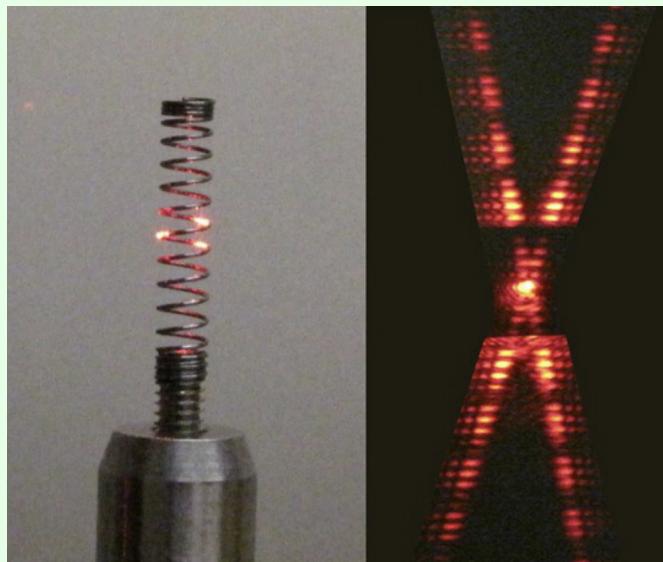
#### Idea 4: Babinet's Principle

Consider all of the rays  $R$  that strike a point  $P$  on the screen. If the intensity at  $P$  is zero, then the rays must completely destructively interfere. That means that if we split  $R$  into two sets of rays  $R_1$  and  $R_2$  in any way, then the amplitudes due to the rays  $R_1$  and the rays  $R_2$  must be equal and opposite, which means either set of rays alone would produce the same intensity at  $P$ . This is Babinet's principle.

As a concrete example, consider shining a laser pointer at a wall. There will be a bright spot on the wall at the exact location the laser hits, and darkness everywhere else. Consider some dark point  $P$ . If we had instead passed the laser through two slits, we would only get the rays  $R_1$  going through the slits, and we would generally get some nonzero intensity at  $P$ , due to the double slit interference pattern. Babinet's principle tells us we would get the exact same intensity at  $P$  if we put two slit-shaped obstacles in the way, because then we would get precisely the rays  $R_2$  which don't hit the slits. In other words, the diffraction pattern from an obstacle is precisely the same as the diffraction pattern from a complementary slit.

#### Example 6: BPhO 2016.5

When a laser pointer hits a spring, the following pattern is produced on a screen behind it.



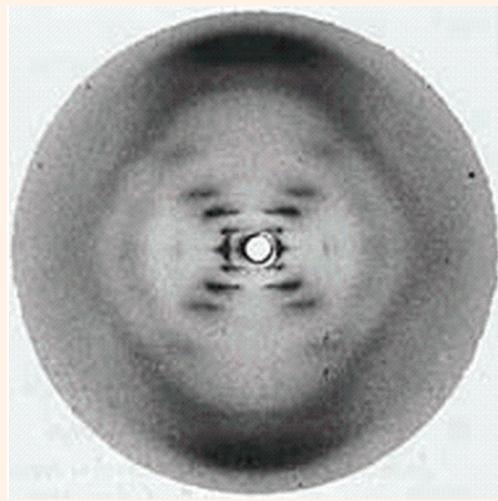
Explain why this happens, and what we can learn about the spring.

### Solution

If we look at the spring along the direction the laser pointer is going, it's essentially two sets of obstructions, one going up and to the right (the front of the spring, in the picture), and one going up and to the left (the back of the spring). By Babinet's principle, the resulting diffraction pattern should be the same as if we had two sets of *openings* instead. Thus, we expect to see two independent diffraction patterns, one due to each of these obstructions. The angle between these two patterns is twice the angle that the spiral path of the spring makes with the horizontal.

Now focus on one interference pattern. By Babinet's principle, it's basically a single slit pattern, which is indeed what we see. However, from the reflection of the laser in the picture, we see that the laser beam is wide enough to hit two separate turns of the spring. The result is a double slit pattern multiplied by a single slit pattern, where the latter yields a minimum at approximately every 5 double slit minima. The spacing between the single slit minima tells us the thickness of the wire, and the spacing between the double slit minima tells us the spacing between the turns in the spring. Combining this with what we know about the angle of the spring tells us about the radius of the spring. Thus, we can figure out essentially everything about its 3D shape.

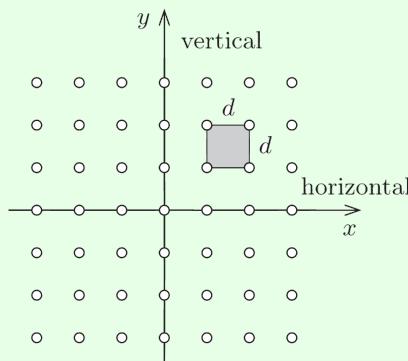
The coolest thing about this problem is that all the logic here applies to diffraction off any helical structure. For example, below is the X-ray diffraction pattern of DNA, taken by Rosalind Franklin in 1952.



As discussed in detail [here](#), such data was used to discover the structure of DNA.

### Example 7: MPPP 125

An opaque sheet is perforated by many small holes arranged in a square grid of side length  $d$ . It is illuminated by light of wavelength  $\lambda$ , and a screen lies a distance  $D$  behind it.



Assuming  $D \gg d \gg \lambda$ , find the locations of the primary diffraction maxima on the screen.

### Solution

Let  $(x, y)$  denote coordinates on the sheet, and  $(x', y')$  denote coordinates on the screen, with the same center. When we considered one-dimensional diffraction gratings, we found that light which originates from  $y$  and hits point  $y'$  on the screen has a path length difference  $yy'/D$  relative to light that came from point  $y = 0$ . A similar argument shows light which comes from  $(x', y')$  gets a path length difference

$$\Delta\ell = \frac{xx' + yy'}{D}$$

relative to light coming from  $x = y = 0$ .

For a square grid,  $(x, y) = (nd, md)$  for integers  $n$  and  $m$ , giving a path length difference

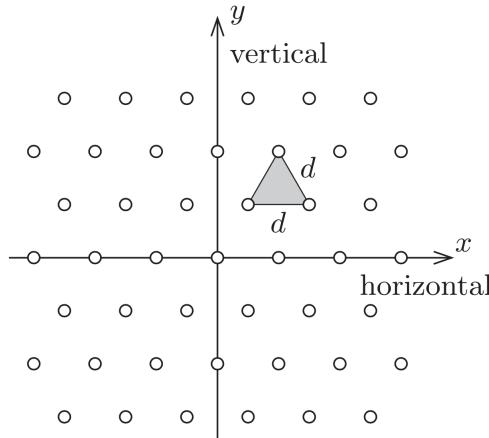
$$\Delta\ell_{n,m} = \frac{d}{D}(nx' + my').$$

We get a diffraction maximum at  $(x', y')$  when the light from each hole arrives in phase, which means this quantity must be a multiple of  $\lambda$  for all  $n$  and  $m$ . This occurs precisely when

$$x' = \frac{\lambda D}{d}n', \quad y' = \frac{\lambda D}{d}m'$$

for integers  $n'$  and  $m'$ . That is, the diffraction maxima also form a square grid of side length  $\lambda D/d$ . Notice again that the diffraction pattern is “inverse” to the pattern on the sheet. It gets bigger when the sheet gets smaller; for instance, if the sheet is compressed horizontally, the maxima on the screen are stretched horizontally.

- [3] **Problem 13** (MPPP 126). Continuing on the previous example, suppose the holes are instead arranged in a triangular grid with side length  $d$ .



Find the primary diffraction maxima on the screen. What kind of grid do they form?

- [3] **Problem 14.** [AuPhO 2015, problem 14](#). Print out the relevant part of the [answer booklet](#) as well. This problem contains several instructive examples of higher-dimensional diffraction patterns.

For a much harder multi-dimensional diffraction problem, beyond the scope of the Olympiad, see [Physics Cup 2019, problem 5](#).

## 5 Technological Applications

### Example 8

Gratings split light into its components. If a grating can just resolve the two wavelengths  $\lambda$  and  $\lambda + \Delta\lambda$ , its resolving power is  $R = \lambda/\Delta\lambda$ . Compute the resolving power of a diffraction grating with  $N$  slits at the  $n^{\text{th}}$  order maximum.

**Solution**

Conventionally, we say that two diffraction peaks are distinguishable if the maximum of one falls outside the first minimum of the other. We know the  $n^{\text{th}}$  order maximum for wavelength  $\lambda$  occurs when  $d \sin \theta = n\lambda$ , as here the path length difference between adjacent slits is  $n\lambda$ . Furthermore, the first minimum around this maximum occurs when there is an extra net path length difference of  $\lambda$  across the entire diffraction grating, i.e. when

$$Nd \sin \theta = Nn\lambda + \lambda.$$

Setting the value of  $\sin \theta$  equal to that for wavelength  $\lambda + \Delta\lambda$ , we see that we can just resolve these two wavelengths if

$$\frac{n(\lambda + \Delta\lambda)}{d} = \frac{(Nn + 1)\lambda}{Nd}, \quad R = \frac{\lambda}{\Delta\lambda} = Nn.$$

Note that the resolving power is also the number of wavelengths by which the longest and shortest possible paths to the diffraction maximum differ (i.e. the paths through the very top and very bottom slits). The fact that a larger distance may be used to resolve smaller wavelength differences is another manifestation of the uncertainty principle.

- [2] **Problem 15.** Many common diffraction gratings reflect light rather than transmitting it.

- (a) We may crudely model a reflective diffraction grating as a mirror with  $N$  small notches, spaced a distance  $d$  apart. The notches do not reflect light, but the rest of the mirror serves as a source of Huygens wavelets when light is incident on the grating. Show that, unlike the transmission gratings we considered above, the zeroth order maximum of a reflective grating is much brighter than the others.
- (b) This feature is undesirable because the zeroth order maximum is useless for distinguishing different wavelengths. Instead, most modern diffraction gratings are blazed, as shown.



For concreteness, suppose that light is incident straight downward. How should the blaze angle  $\gamma$  be chosen so that the  $n^{\text{th}}$  order maximum is the brightest?

Reflective diffraction gratings are more flexible and more common than transmission gratings. (CDs, DVDs, and “holographic” trading cards and stickers all use reflective diffraction gratings, and you can even make them on [chocolate](#).) Textbooks focus on transmission gratings largely because they make the diagrams a little cleaner.

- [2] **Problem 16 (PPP 127).** When a particular line spectrum is examined using a diffraction grating with 300 lines/mm with the light at normal incidence, it is found that a line at  $24.46^\circ$  contains both red (640 – 750 nm) and blue/violet (360 – 490 nm) components. Are there any other angles at which the same would be observed?
- [3] **Problem 17.** Diffraction limits the resolution of optical instruments.

- (a) Suppose that light of wavelength  $\lambda$  enters through an aperture of width  $D$ . As a result, the light diffracts, which causes the angle of the light's propagation to pick up an additional spread of order  $\theta$ . Estimate  $\theta$ .
- (b) The diameter of a human pupil is about 3 mm. Estimate the size of the smallest text that a human being could read from 5 m away.
- (c) A typical amateur telescope has an aperture of order 10 cm. The Sun has a radius of  $7 \times 10^8$  m. Estimate the furthest possible distance, in light years, that such a telescope could resolve a Sun-sized star. (Stars further away than this will just show up as blurry points.)

### Remark

The effect described in problem 17 motivates astronomers to build [ever larger telescopes](#). The largest examples are radio telescopes, such as the Arecibo observatory that [collapsed in 2020](#), because diffraction effects are more significant the larger the wavelength is. However, the telescope doesn't have to be one big piece. An array of telescopes can effectively be combined into a telescope whose radius, for the purposes of the diffraction limit, is the total width of the array, as in the [Very Large Array](#). The Event Horizon Telescope was able to resolve a black hole  $5 \times 10^7$  ly away because it combined telescopes spaced around the entire Earth.

### Example 9

How close does a Sun-like star have to be in order to see it with the naked eye in daylight?

### Solution

Let the distance to the Sun be  $d$ , and the distance to the star be  $D$ . Then the ratio of intensities of the two is naively

$$\frac{I_{\text{star}}}{I_{\text{Sun}}} = \left(\frac{d}{D}\right)^2.$$

This suggests the star is hard to see if  $D > d$ , which is always true. But this is too pessimistic, because the light from the Sun comes from all directions in the sky, while the light from the star comes from only a single direction. The actual ratio we want to calculate is

$$\frac{I_{\text{star}}/\Omega_{\text{star}}}{I_{\text{Sun}}/2\pi} = \left(\frac{d}{D}\right)^2 \frac{2\pi}{\Omega_{\text{star}}}$$

where  $\Omega_{\text{star}}$  is the apparent solid angle of the star in the sky.

This in turn is given by the diffraction limit: if your pupils have radius  $r$ , then

$$\Omega_{\text{star}} \sim (\Delta\theta)^2 \sim (\lambda/r)^2.$$

The star should be visible above daylight if the ratio above is at least one or so, which means the maximum distance is

$$D \sim \frac{r}{\lambda} d \sim \frac{3 \text{ mm}}{600 \text{ nm}} (1 \text{ AU}) \sim 5 \times 10^3 \text{ AU} \sim 0.1 \text{ ly}.$$

This is still closer than the closest other star, so you would need a telescope to see any.

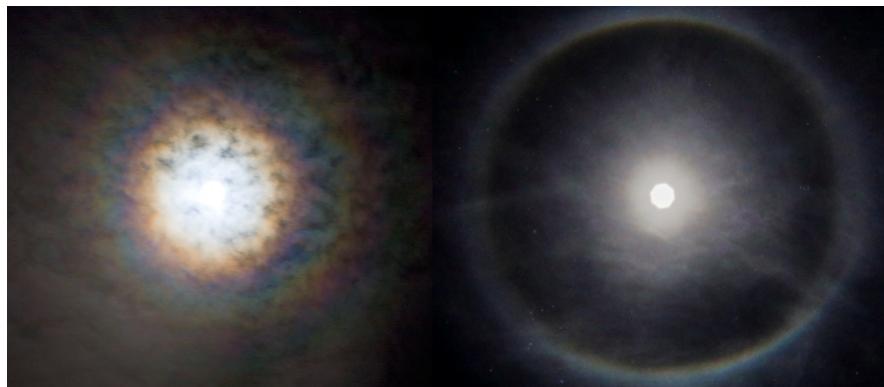
Notice how this differs from a microscope! Microscopes are used to resolve finer details on a small object. But most telescopes can't resolve *any* of the details of any but the nearest stars. Increasing the size of the telescope has two benefits: increasing the amount of light that goes through, and improving the contrast due to decreasing the blurring due to diffraction.

- [3] **Problem 18** (PPP 126). A compact disc contains approximately 650 MB of information. Estimate the size of one bit on a CD using an ordinary ruler. Confirm your estimate using a laser pointer. (If you can't find a CD, a DVD will also work.)
- [4] **Problem 19.** [EPhO 2005, problem 5](#). A subtle interference problem.

## 6 Real World Examples

These questions are not neat and self contained – they illustrate real physical phenomena, for which you'll have to guess an appropriate physical model. Of course, you have the massive advantage of knowing that all of the problems involve interference and diffraction (or do they?).

- [4] **Problem 20.** This problem is about some neat atmospheric phenomena. For some parts, it will be useful to use Babinet's principle.



- (a) On a foggy night, there are many tiny water droplets in the air. On such nights one can see a ring around the moon, called a lunar corona, shown at left above. The ring is usually reddish in color. If one looks very carefully on a good night, one can see a blue ring outside the red ring and a blueish-white region inside the red ring. On other nights, one can only see a white haze around the moon. Explain these observations.
- (b) The size of the corona depends on the atmospheric conditions. Estimate the diameter of the water droplets in the air if the first red ring around the moon appears to have a diameter 4 times that of the moon. The angular diameter of the moon in the sky is  $0.5^\circ$ .
- (c) On a cold night, there are many thin hexagonal ice crystals in the air. On such nights one can see a much larger, sharper ring around the moon, called a  $22^\circ$  halo, shown at right above. The size of the halo does not depend on the size of the crystals. Explain these observations.
- (d) In the photo used in part (c), the moon is shaped like an octagon. Why?

- (e) On a cold and exceptionally calm night, the results will be different.



Instead of a circle, one will see two “moon dogs”, bright spots displaced about  $22^\circ$  from the moon horizontally. In addition, lights on the ground will produce vertical “light pillars”. Explain these observations.

- [2] **Problem 21** (Povey). Consider a reflective metal tube, such as a length of copper pipe, with length  $L$  and radius  $r$ . If you place a diffuse light source at one end of the tube, on the axis of symmetry, and look at it from the other end, with your eye also on the axis of symmetry, then you will see both the light source and bright circular rings around it. Why does this happen? Assuming the light has wavelength  $\lambda$ , calculate the angles of the bright rings.
- [2] **Problem 22.** Take a pair of glasses, exhale on them to fog them up, and put them on and look at a light. You should see something strange; why does it happen?
- [2] **Problem 23.** Sometimes, a cloud will display a colorful pattern, as shown.



What is the explanation for this phenomenon, and why is it relatively rare? While you’re at it, what’s going on with the sun in this photo?

- [3] **Problem 24.** A student noticed an odd pattern when the light from a streetlamp shined through their open window. The window is fitted with a metal mesh screen and a curtain. Photos were taken with the curtain up (left) and down (right).



For scale, the streetlamp was about 30 m away, the distance between the metal wires was 1.4 mm, the diameter of each wire was 0.4 mm, and the curtain was woven from fibers whose width was comparable to that of a human hair.

- (a) Explain everything you can about the pattern on the left.
- (b) Explain everything you can about the pattern on the right.
- (c) What, if anything, can we learn about the streetlamp from either picture?

[4] **Problem 25.** One day, somebody sent me a photo of a weird pattern on their phone.



The phone was on a desk, about a third of a meter away from the camera.

- (a) Does the desk light emit a roughly continuous spectrum (typical for incandescent or good LED lights) or a sharply peaked spectrum (typical for fluorescent or cheap LED lights)?
- (b) Qualitatively explain everything about the pattern seen. In particular, explain the geometrical pattern of the colored dots, the way the colors are distributed, why the colored dots cover the entire phone, and why one dot is large and white.
- (c) How are the pixels on the phone laid out?
- (d) Estimate the pixel spacing on the phone, looking up or estimating any numbers necessary.
- (e) Using a suitable source of light, such as a laser pointer, determine the resolution of your own phone screen as accurately as possible. (You can look up relevant wavelengths of light; also

note that it won't work with all phones; older ones may fare better.) Prepare a lab report with a data table and an uncertainty estimate, as explained in **P2**, and compare your result against the advertised value.