

Modern I: Semiclassical Mechanics

The basics of quantum mechanics can be found in chapters 46 and 47 of Halliday and Resnick, and are covered more thoroughly in chapters 3 through 6 in Krane. Chapter 10 of Krane covers quantum statistical mechanics as used in the final section. For a complete, but advanced treatment of the WKB approximation, see chapter 9 of Griffiths' *Introduction to Quantum Mechanics* (3rd edition). For some nice conceptual discussion, see chapters I-37 and I-38 of the Feynman lectures, or if you're ambitious, essentially all of volume III. There is a total of **82** points.

1 The WKB Approximation

A proper introduction to quantum mechanics would take a whole book. Luckily, there is a “semiclassical” regime of quantum mechanics which can be handled with much less machinery. Historically, this regime was discovered first, by scientists like Bohr, and it suffices to explain many quantum effects. To introduce the WKB approximation, we'll start by considering classical standing waves.

Idea 1

The variation of the phase ϕ of a wave is described by its wavenumber and angular frequency,

$$k = \frac{d\phi}{dx}, \quad \omega = \frac{d\phi}{dt}$$

As covered in **W1**, the group velocity is

$$v = \frac{d\omega}{dk}.$$

A standing wave can form if the wave's phase lines back up with itself after one round trip,

$$\oint k dx = 2\pi n, \quad n \in \mathbb{Z}.$$

A simple case is a string of length L with fixed ends, where we have

$$2kL = 2\pi n$$

which gives the wavenumbers $k_n = \pi n/L$ and hence the standing wave angular frequencies $\omega_n = \pi v n/L$, as we saw in **W1**.

- [1] **Problem 1.** A slightly more subtle case is the case of a string of length L with one fixed and one free end. Show that the standing wave angular frequencies are

$$\omega_n = \frac{\pi v}{L}(n + 1/2).$$

The reason our principle above doesn't give the right answer is that a wave picks up an extra phase shift π when it reflects off a fixed end, so we really should have written

$$\oint k dx = 2\pi(n + 1/2)$$

in this case. We didn't run into any problems for two fixed ends, because in that case we get two phase shifts of π , which have no overall effect.

- [2] **Problem 2.** Suppose a string of length L is hung from the ceiling. The string has mass density μ , and the bottom of the string is held fixed and pulled down with a force $F \gg gL\mu$. If the string were weightless, then the standing wave angular frequencies would simply be $\pi v n/L$, where $v = \sqrt{F/\mu}$. However, the weight causes the tension and hence the wave speed to vary throughout the rope.

- (a) Explain why the wave's angular frequency ω is uniform, i.e. why standing wave solutions are proportional to $\cos(\omega t)$.
- (b) Find the angular frequencies of standing waves, including corrections up to first order in $gL\mu/F$.

This is a more quantitative version of a problem we encountered in **W1**.

In quantum mechanics, the state of a particle is described by a wavefunction $\psi(x, t)$ which obeys the Schrodinger equation. When a particle is confined in a finite volume, there are standing wave solutions analogous to those of classical wave mechanics, which have discrete frequencies.

Idea 2: WKB Approximation

The momentum and energy of a quantum particle obey the de Broglie relations

$$p = \hbar k, \quad E = \hbar \omega.$$

As usual, for nonrelativistic particles, the energy E and momentum p are related by

$$E = \frac{p^2}{2m} + V(x).$$

For a particle with reasonably well-defined momentum, the wavefunction is a wavepacket which travels at the group velocity

$$v_g = \frac{d\omega}{dk} = \frac{dE}{dp} = \frac{p}{m}.$$

In the classical limit, this is simply the ordinary velocity of the particle.

The de Broglie relations also apply for relativistic particles, if E is the relativistic kinetic plus potential energy, and p is the relativistic momentum. (If you want, you can also add mc^2 to E to get the total relativistic energy, but it makes no difference since a constant shift in energy doesn't do anything.) In this case, the group velocity obeys $p = \gamma m v_g$, as expected.

Just as for a classical standing wave, ω is the same everywhere for quantum standing waves. Since energy is related to frequency, these standing waves are also states of definite energy. In the semiclassical limit, the standing waves must satisfy

$$\oint p dx = (2\pi n + \alpha)\hbar = \left(n + \frac{\alpha}{2\pi}\right) h.$$

The extra phase α depends on how the particle gets reflected at the endpoints of its motion.

Remark

The left-hand side of the quantization condition above is precisely the adiabatic invariant from **M4**, which stays the same if we change the system parameters slowly. This ensures the quantization condition is preserved over time, as it must be for self-consistency. If you instead change the system parameters quickly, the integral is not preserved, but that's because the change causes transitions from one energy level to another (i.e. to waves with different n).

- [2] **Problem 3.** Consider a one-dimensional box of length L , with hard walls. We can think of these hard walls as a potential $V(x)$ that is zero inside the box and infinite outside the box. It can be shown that each of these “hard” boundaries contributes π to α , so in this problem we have $\alpha = 2\pi$, which is equivalent to just taking $\alpha = 0$.

- (a) Find the energy levels of a particle of mass m .
- (b) Now suppose the particle is replaced with a photon, with $E = pc$. Find the allowed energies.

The frequencies you found in part (b) correspond to the standing wave frequencies for electromagnetic waves in a box with reflecting (i.e. perfectly conducting) walls.

- [2] **Problem 4.** Now consider a particle of mass m in the potential $V(x) = kx^2/2$. In this case, the particle turns around at a point where the potential energy gradually increases from below the particle's energy E to above it. It can be shown that each of these “soft” boundaries contributes $\pi/2$ to α , so that for this potential we can take $\alpha = \pi$.

Show that the energy levels are

$$E_n = \hbar\omega_0 \left(n + \frac{1}{2} \right), \quad \omega_0 = \sqrt{\frac{k}{m}}.$$

This system is called the quantum harmonic oscillator, and remarkably, this is the exact answer, even though we used an approximation to get it. This result will be used in several problems below.

- [3] **Problem 5.** ⌚ USAPhO 2015, problem A1. (The use of the WKB approximation in this problem is technically incorrect: the problem takes $\alpha = \pi$ when it actually should take $\alpha = 3\pi/2$. Often, people will apply the WKB approximation in a sloppy way because they're only after rough estimates at small n , or the limiting behavior at $n \gg 1$. In the previous problems, we treated α properly because doing so will give the *exact* correct answer in those cases.)
- [5] **Problem 6.** ⌚ IPhO 2006, problem 1. This is a neat problem which illustrates the effect of a gravitational field on quantum particles, as well as the basics of interferometry, a subject developed further in **W2**. Give this a try even if it looks tough; only the ideas introduced above are needed!

Idea 3: Bohr Quantization

In general, $p dx$ may be replaced by any generalized momentum/position pair. For example,

$$\oint L d\theta = nh.$$

When angular momentum is conserved, the left-hand side is simply $2\pi L$, immediately giving

$$L = n\hbar$$

which is Bohr's quantization condition. In a system of particles rotating together, L stands for the total angular momentum of the system.

Compared to back-and-forth linear motion, covered in idea 2, rotation is different because it's inherently periodic. For rotation, the integer n can be positive or negative, representing a particle going clockwise or counterclockwise. Also, there is no analogue of the α phase factor because the particle just rotates all the way around; it never gets reflected.

Example 1

Find the energy levels and orbit radii of the electron in the hydrogen atom using Bohr quantization.

Solution

We postulate a circular orbit, and quantize the angular momentum. We have

$$\frac{mv^2}{r} = \frac{e^2}{4\pi\epsilon_0 r^2}, \quad L = mvr = n\hbar.$$

Solving the second equation for v and plugging into the first gives

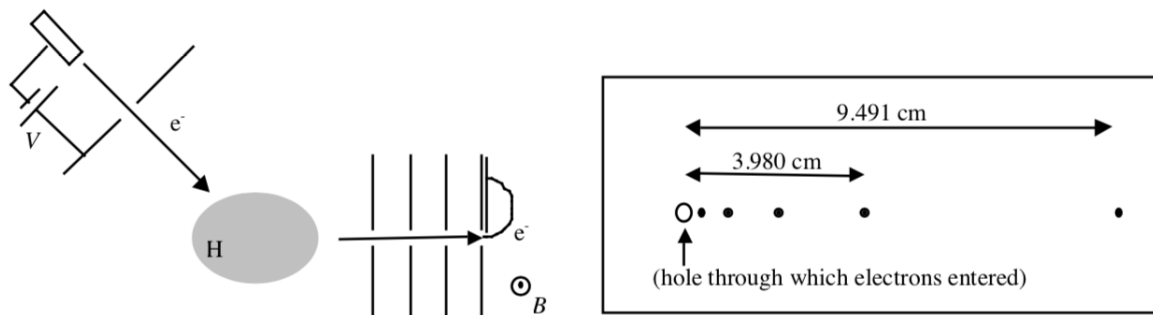
$$r = \frac{4\pi\epsilon_0 \hbar^2}{me^2} n^2 = a_0 n^2$$

where $a_0 = 5.3 \times 10^{-11}$ m is called the Bohr radius; these are the allowed orbit radii. To get the energies, we use the standard result for circular motion with an inverse square force that the total energy is half the potential energy, so

$$E = -\frac{e^2}{8\pi\epsilon_0 r} = -\frac{me^4}{2(4\pi\epsilon_0)^2 \hbar^2} \frac{1}{n^2}.$$

Evidently, they get more and more closely spaced together as n increases. The big constant in front is called the Rydberg, and is equal to 13.6 eV.

- [1] **Problem 7.** Find the energy levels of positronium, a bound state of a positron and electron.
- [2] **Problem 8** (USAPhO 2004). Electrons are accelerated from rest through a potential V into a cloud of cold atomic hydrogen. A series of plates with aligned holes select a beam of scattered electrons moving perpendicular to the plates. Immediately beyond the final plate, the electrons enter a uniform magnetic field B perpendicular to the beam; they curve and strike a piece of film mounted on the final plate.




When the film is developed, a series of spots is observed. The distances between the hole and the two most distant spots are measured. You may assume that the film is large enough to have intercepted all of the electrons, i.e. that there are no spots farther from the hole than those shown. The number of spots shown is not necessarily accurate.

Make the approximation that the mass of the hydrogen atom is much larger than the mass of the electron. Assume that each electron scatters off only one atom, which is initially in the ground state (lowest energy state) and has negligible thermal velocity. Determine B , V , and the total number of spots on the film.

- [2] **Problem 9.** A rotor consists of two particles of mass m connected by a rigid rod of length L .
- Find the energy levels if the particles are not identical.
 - Find the energy levels if the particles are identical. (Hint: recall that the closed loop integrals in the previous ideas are over paths that take the system back to its original state.)

You might find it disturbing that the result is so different if the particles are or aren't *completely* identical, but it's a well-verified fact about molecular rotational energy levels. For example, if the two atoms in the molecule had nuclei of different isotopes, we measure the result of part (a), and if they were the same isotope, we measure the result of part (b).

- [3] **Problem 10.**  [INPhO 2020, problem 3.](#)
- [3] **Problem 11.** [INPhO 2016, problem 6.](#) Unfortunately, this question automatically comes with the solutions, but it's still useful to work through.

Remark

In popular science, people sometimes speak of “quantizing” a system as similar to making everything discrete. But as you’ve seen above, it’s more complicated than that. For instance, position never becomes discrete; instead, we integrate over it.

The general rule in quantum mechanics is that confinement to a finite “size” causes the conjugate variable to become discrete. For example, above you looked at several examples of particles bound to potentials. These are confined in space, and hence have discrete orbits in phase space by idea 2, and thus discrete energies. But a free particle not bound to a potential can have any energy, because $E = p^2/2m$ and there is no condition at all on p . On the other hand, angles are always confined to the finite range $[0, 2\pi]$, which is why the angular momentum of any system is quantized.

- [2] **Problem 12.** Neutrinos are quantum particles with extremely low masses. Because their masses are so low, they are almost always produced moving near the speed of light. In the Sun, neutrinos are produced in a superposition of two types, one with mass m_1 and another with mass m_2 .

- (a) Suppose the neutrino is produced with definite energy E . The part of the neutrino wave with mass m_1 has wavenumber k_1 , and the part with mass m_2 has wavenumber k_2 . Find an approximate expression for $k_1 - k_2$ in the ultrarelativistic limit $E \gg m_1, m_2$.
- (b) Find the distance L that the neutrino travels so that the relative phase between these waves changes by 2π . This is the characteristic wavelength of “neutrino oscillations”. Evaluate it numerically assuming the typical values $E = 10 \text{ keV}$, $m_1 c^2 = 0.010 \text{ eV}$, and $m_2 c^2 = 0.013 \text{ eV}$.

2 Higher Dimensions

Idea 4

For a system with more than one degree of freedom, the WKB quantization condition holds for each individually,

$$\oint p_i dx_i = \left(n_i + \frac{\alpha_i}{2\pi}\right) h.$$

In this case, there can be multiple quantum states with a given energy, in which case we say that energy level is degenerate; the number of states with that energy is called the degeneracy.

- [2] **Problem 13.** Consider a particle of mass m in a two-dimensional box of width and length L , with hard walls. This is the two-dimensional analogue of problem 3, and again $\alpha_1 = \alpha_2 = 0$.

- (a) Write down the energy of the state with quantum numbers n_1 and n_2 .
- (b) What is the lowest energy level with a degeneracy of greater than 2?

- [4] **Problem 14.** Consider a particle of mass m in the potential $V(x, y) = kr^2/2$. This is the two-dimensional analogue of problem 4. It is also the potential experienced by an electron in the obsolete “plum pudding” model of the atom, where they are embedded in a ball of uniform charge density.

- (a) By working in Cartesian coordinates, find all of the energy levels, as well as the number of states within each energy level, called the degeneracy.
- (b) [A] Now repeat the exercise in polar coordinates. In this case the integrals

$$\oint p_r dr, \quad \oint L d\theta$$

are quantized. Find the energy levels and their degeneracies. Note that for the radial motion, you will have to use the effective potential, as covered in **M6**. You will run into a difficult integral, so you may use the fact that

$$\int_{C-\sqrt{C^2-D^2}}^{C+\sqrt{C^2-D^2}} dx \sqrt{\frac{2C}{x} - \frac{D^2}{x^2} - 1} = (C - |D|)\pi$$

valid for $|D| \leq C$. How does your answer compare to that of part (a)?

Remark

Sommerfeld applied an analysis like that of part (b) of problem 14 to the Bohr model, yielding the semiclassical orbits which are ellipses with the nucleus at the focus. (In fact, if you're so inclined, you can do this too, using the same provided integral.) This accounted for the quantum numbers n and ℓ in hydrogen. The quantum number m comes from additionally quantizing L_z , which implies that the elliptical orbits can only occur in certain planes, an idea known as “space quantization”. Sommerfeld even managed to compute relativistic corrections to the energy levels.

With all this included, the Bohr theory provides a complete description of the energy levels of hydrogen, except that (1) the $\ell = 0$ orbitals are missing, since they would have to go straight through the nucleus, (2) space quantization seems artificial and breaks rotational symmetry, and (3) the number of states isn't quite right, a deficiency that would later be fixed by including spin. Many complicated attempts were made to patch these problems, or to extend the theory to multi-electron atoms, but they were forgotten after the modern theory of quantum mechanics (in terms of the Schrodinger equation) appeared.

However, what you've learned above is not completely irrelevant today. The correspondence principle is the idea that quantum results should yield smoothly transition to classical ones in the limit $\hbar \rightarrow 0$, which in this context means sending the quantum numbers to infinity. And that's exactly what happens. For high quantum numbers, you can superpose atomic orbitals of nearby energy to create a sharply peaked wavefunction, just like how we could create wavepackets from plane waves in **W1**. These peaks act like localized classical particles, following the Bohr model's orbits. Thus, the Bohr model is still useful for studying [Rydberg atoms](#), which are hydrogen-like atoms excited to very high energy levels. For some more discussion of Sommerfeld's results, see [this paper](#).

[3] **Problem 15** (Cahn). A crude model of an electron bound to an atom is a particle of mass m attached to a one-dimensional spring, with spring constant k and hence angular frequency $\omega = \sqrt{k/m}$. Consider two such atoms.

- (a) Write down the energy levels of the system, assuming the atoms are completely independent. How many states correspond to each energy?
- (b) Let the electrons have positions x_i relative to their respective equilibrium positions. Now suppose the atoms are brought close together, causing the electrons to repel. For simplicity, we represent this in terms of an extra potential energy term $k'x_1x_2$, where k' is small. Find the new energy levels of the system exactly. (Hint: this can be done with a clever change of variables. However, you have to be careful because changing to new coordinates x'_i also requires changing the momenta; after all, if we didn't, then the quantization condition of idea 2 would change, leading to different energy levels! If K is the kinetic energy, and you are using momentum variables x_i , then the momenta should be defined as $p_i = \partial K / \partial \dot{x}_i$.)
- (c) Your answer should not make sense for large k' . Physically, what is going on?

Part (b) gives a simple example of how energy levels “split” in the presence of interactions.

Example 2

A nonrelativistic particle of mass m is in a cubical box with side length L and hard walls. Find the approximate number of quantum states with energy at most E_0 , where E_0 is large.

Solution

Using the same reasoning as in previous problems, we apply “hard wall” boundary conditions, requiring the wavefunction to go to zero at the boundary. Thus, the wavefunction is

$$\psi \propto \sin(k_x x) \sin(k_y y) \sin(k_z z)$$

where

$$k_i = \frac{\pi}{L} n_i, \quad n_i \text{ positive integer}$$

and the energy is

$$E = \frac{p^2}{2m} = \frac{\hbar^2 k^2}{2m}.$$

The simplest way to proceed is to think in terms of “momentum space”, an abstract space whose axes are p_x , p_y , and p_z . The allowed states form a grid in the first octant of momentum space, with a volume of $(\pi\hbar/L)^3$ for each state. The surface $E = E_0$ corresponds to a sphere of radius $\sqrt{2mE_0}$. Therefore the number of states with energy at most E_0 is approximately

$$N = \frac{1}{8} \left(\frac{4}{3} \pi (2mE_0)^{3/2} \right) \left(\frac{\pi\hbar}{L} \right)^{-3}.$$

Now let’s solve the problem a slightly different way: suppose the box has *periodic* boundary conditions, so that the right side is identified with the left side, and so on. In this case, the wavefunctions can all be written in the form

$$\psi \propto \exp(i(k_x x + k_y y + k_z z))$$

but now the allowed values of the wavenumbers are

$$k_i = \frac{2\pi}{L} n_i, \quad n_i \text{ integer}.$$

The allowed states form a grid in all of momentum space, not just the first octant, with a volume of $(2\pi\hbar/L)^3$ for each state. That is, while the volume around each state is eight times as large, the states now occupy eight octants instead of one. Then the overall density of states is still the same, and the number of states with energy at most E_0 is approximately

$$N = \left(\frac{4}{3} \pi (2mE_0)^{3/2} \right) \left(\frac{2\pi\hbar}{L} \right)^{-3}$$

which matches the result for hard walls. The point of this computation is to show that when we care about the statistical properties of many states, the boundary conditions won’t matter. In practice, you’ll see both kinds of boundary conditions quite often.

If you find the differences between the two boundary conditions confusing, you’re not alone. In his original derivation of blackbody radiation, Lord Rayleigh used “hard wall” boundary conditions but also allowed negative n_i , leading to a factor of 8 error. Jeans corrected it, which is why the result is now called the Rayleigh–Jeans law.

- [4] **Problem 16.** ⌚ Do the following [JPhO problem](#). This pedagogical problem introduces the WKB approximation and phase space, reviewing everything covered above, and applies it to “clusters” of atoms. You can skip sections I and III, which are covered elsewhere on this problem set.

Remark

Because matter has wave properties, particles such as electrons can exhibit interference effects, like those covered in **W2**. For example, you can run a double-slit experiment firing individual electrons at a time and seeing where they hit the screen, which will gradually build up an [interference pattern](#). You’ll see an example of this in the next problem.

- [4] **Problem 17.** ⌚ IPhO 1993, problem 3. (Hint: to do the final part of this problem, you should *not* try to compare the total path lengths traversed by the electrons. That would be very hard, and worse, it won’t give the right answer, because the potential from the wire also affects the electrons’ phases. Instead, you should use the facts about wavefronts mentioned in **W2**. That is, waves always propagate perpendicular to wavefronts, and all points on a wavefront have the same phase.)

3 The Uncertainty Principle

Idea 5: Heisenberg Uncertainty

So far we have treated a quantum particle as having a well-defined position and momentum, but in reality the uncertainties in the position and momentum obey

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

where, as in **P2**, the uncertainties may be interpreted as standard deviations. The “semi-classical limit” used in the rest of this problem set simply corresponds to the case where the required uncertainty is relatively small, which is reached for energy levels $n \gg 1$. Occasionally, Olympiad questions will ask you to use the Heisenberg uncertainty principle to make a very rough estimate. In these cases, the constant factors will not matter.

Idea 6: Energy-Time Uncertainty

There are two commonly used versions of the energy-time uncertainty principle. If the energy of a system is only measured for a finite time Δt , it must have a finite uncertainty ΔE in its energy. In addition, if a system significantly changes its state in time Δt , then its energy must have been uncertain by a finite amount ΔE . In both cases, we have

$$\Delta E \Delta t \geq \frac{\hbar}{2}.$$

A third common statement of the energy-time uncertainty principle is “for a short time Δt , a system can violate energy conservation by an amount ΔE ”. This is wrong, because quantum systems always conserve energy; systems that naively seem to violate energy conservation simply didn’t have a well-defined energy in their initial state to begin with. However, thinking this way will usually get you the right answers, essentially because of dimensional analysis.

Example 3

Consider once again a particle of mass m attached to a one-dimensional spring, with natural angular frequency ω . Use the uncertainty principle to estimate the minimum possible energy of the particle, and compare it with the result of problem 4.

Solution

Suppose the uncertainties in position and momentum are Δx and Δp . Then the potential energy is of order $k(\Delta x)^2/2$ and the kinetic energy is of order $(\Delta p)^2/2m$. Dropping constants,

$$E \sim k(\Delta x)^2 + \frac{(\Delta p)^2}{m} \gtrsim k(\Delta x)^2 + \frac{\hbar^2}{(\Delta x)^2 m}$$

where we applied the uncertainty principle. The ground state minimizes the energy, which is achieved when $(\Delta x)^2 \sim \hbar/\sqrt{km}$. In this case, the energy is of order $k\hbar/\sqrt{km} \sim \hbar\sqrt{k/m} \sim \hbar\omega$, which is just what we found earlier. (A similar derivation can be used to derive the energy of the ground state of hydrogen, along with the Bohr radius; try it!)

Remark

We can also “solve” the above problem with the energy-time uncertainty principle incorrectly. The only timescale in the problem is $1/\omega$, so

$$\Delta E \gtrsim \frac{\hbar}{\Delta t} \sim \hbar\omega$$

so $E \gtrsim \hbar\omega$. However, in reality the ground state has *no* energy uncertainty; its energy is simply the ground state energy. Another way of saying this is that a particle can hang out in the ground state forever, so Δt is infinite and hence ΔE is zero. This incorrect derivation gives the right answer just because it’s the only possible answer by dimensional analysis. Thus, a sloppy problem might ask you to do it.

Example 4

Consider a single slit diffraction experiment, where photons of wavelength λ pass through a slit of width a . If the screen is a large distance D away, roughly how wide is the resulting diffraction pattern on the screen?

Solution

The photon has a momentum $p_x = \hbar k = h/\lambda$, and passing through the slit necessarily gives it a transverse momentum uncertainty of

$$\Delta p_y \sim \frac{h}{a}$$

where we dropped order one constants, which means an angle uncertainty of

$$\Delta\theta \sim \frac{\Delta p_y}{p_x} \sim \frac{\lambda}{a}.$$

Therefore, using basic geometry, the size of the pattern on the screen is

$$\Delta y \sim D\Delta\theta \sim \frac{D\lambda}{a}.$$

This is the approximate width of the central maximum for single slit diffraction, as we found in **W2**. The reason the result is the same is that light acts like a wave both classically and quantum mechanically; the quantum version of the derivation is just the same as the classical version, but with “everything multiplied by h ”. What’s new about this derivation is that it also applies for matter particles, which have $\lambda = h/p$.

Example 5

The Higgs boson has a mass of 125 GeV and a lifetime of about $\tau = 1.6 \times 10^{-22}$ s. About what percentage uncertainty must a measurement of a Higgs boson’s mass have?

Solution

Decay is a significant change in the particle’s state, and this change happens over a time τ , which means the energy uncertainty is

$$\Delta E \sim \frac{\hbar}{\tau} = 7 \times 10^{-13} \text{ J} = 0.004 \text{ GeV}.$$

When we measure the Higgs boson’s mass, we really measure the $E = mc^2$ energy released when it decays, so the unavoidable uncertainty of the mass is $\Delta E/E \sim 0.003\%$. (But the actual measured uncertainties are much higher, due to a variety of other effects.)

- [1] **Problem 18** (Krane 4.39). An apparatus is used to prepare an atomic beam by heating a collection of atoms to a temperature T and allowing the beam to emerge through a hole of diameter d in one side of the oven. Show that the uncertainty principle causes the diameter of the beam, after traveling a length L , to be larger than d by an amount of order $L\hbar/d\sqrt{mk_BT}$, where m is the mass of an atom.
- [2] **Problem 19** (Insight 8.26). When helium is cooled to extremely cold temperatures, it becomes a superfluid, an exotic type of liquid that can flow with zero dissipation. These strange properties occur because quantum mechanical effects are large, making the quantum uncertainty in the position of each helium atom on the same order as the separation between atoms.
- (a) Given that superfluid helium has density ρ and a helium atom has mass m , estimate the temperature T at which helium becomes a superfluid. This is closely related to, but not quite the same thing as Bose–Einstein condensation, a phase transition that bosons undergo at low temperatures.
- (b) Numerically evaluate T , given that $\rho \sim 100 \text{ kg/m}^3$ and $m \sim 7 \times 10^{-27} \text{ kg}$.
- [4] **Problem 20**. A neutron is inside a small cubical box of side length d . Ignore gravity.
- (a) Estimate the minimum possible pressure on the walls using the uncertainty principle, dropping all numeric factors. In the next two parts, we’ll calculate the pressure more carefully.

- (b) Calculate the average pressure on the walls by treating the neutron as a classical particle bouncing back and forth, with the same momentum as expected for the ground state in the WKB approximation.
- (c) Calculate the average pressure on the walls by finding the energy E of the ground state using the WKB approximation, and the definition of pressure, $P = -\partial E/\partial V$. (This actually gives the exact answer. Of course, by dimensional analysis, taking $P \sim E/V$ would also produce the right answer, up to a constant factor.)
- (d) Now suppose that $N \gg 1$ neutrons are inside the box. Neutrons are fermions, as explained in the next section, and hence no two can share the same quantum state. Neglecting any interactions between the neutrons, estimate the minimum possible pressure on the walls, using either the method of part (b) or (c). How does it scale with the number density $n = N/V$?


The large pressure you will find in part (d) is known as degeneracy pressure. It supports compact objects such as white dwarfs and neutron stars, as you'll investigate in **X3**.

[3] **Problem 21.**  USAPhO 2018, problem B2.

[3] **Problem 22.** Classically, an electron orbiting a proton with angular frequency ω_o emits radiation with angular frequency $\omega_c = \omega_o$, as covered in **E7**. On the other hand, quantum mechanically the energy levels are discrete, and using the de Broglie relation $\Delta E = \hbar\omega$ indicates the angular frequencies of radiation emitted when the electron drops between energy levels are discrete as well. The classical and quantum models thus seem to be radically different, but in the limit $n \rightarrow \infty$ where quantum effects become negligible, the two should match.

- (a) Suppose that the electron can orbit the proton in circular orbits with discrete radii r_n . For the n^{th} orbit, compute the angular frequency ω_c of the emitted radiation according to classical mechanics.
- (b) Now suppose the electron drops from the n^{th} energy level to the $(n-1)^{\text{th}}$ energy level. Compute the angular frequency ω_q of the emitted radiation according to quantum mechanics, assuming the orbits have radii r_n .
- (c) In the limit $n \rightarrow \infty$, the results of parts (a) and (b) should coincide. Therefore, by equating these results, infer how r_n depends on n , and thus how L depends on n . If all goes well, you should recover the result of Bohr quantization.

The reasoning here is exactly how Bohr came up with Bohr quantization in the first place. (The de Broglie relation we had to use was motivated earlier through Planck's law, as we showed in **T2**.)

[4] **Problem 23.**  IPhO 2005, problem 3. You may skip part 4, since it's quite similar to another problem on this problem set.

4 Bosons and Fermions

So far, we've only consider the quantum states of individual particles. Now we'll consider what happens when we put many of these particles together. We will assume the particles do not interact, which means their quantum states are just the same as the quantum states for individual particles. If the particles are fermions, they obey the Pauli exclusion principle, which means no two can occupy the same quantum state. If they are bosons, there is no such restriction; we'll consider bosons first.

- [3] **Problem 24.** This problem is a modification and clarification of USAPhO 2011 A4. (In fact, the answers differ, so don't compare against the USAPhO solution!) Consider a simplified model of the electromagnetic radiation inside a cubical metal box of side length L at temperature T . In this model, modes of the electric field have spatial dependence

$$E(x, y, z) = E_0 \sin(k_x x) \sin(k_y y) \sin(k_z z)$$

where one corner of the box lies at the origin and the box is aligned with the x , y , and z axes. For simplicity, we will treat the electric field as a *scalar*, even though it actually has direction.

- The electric field must be zero everywhere at the sides of the box. What condition does this impose on the k_i ?
- Show that for a typical box at room temperature, $L \sim 1$ m and $T \sim 300$ K, the quantity $k_B T$ is much greater than $\hbar \omega_0$, where ω_0 is the lowest nonzero mode angular frequency. Therefore, for the rest of the problem, you may approximate $k_B T \gg \hbar \omega_0$.
- Each permitted value of the triple (k_x, k_y, k_z) corresponds to a mode, which can be occupied by any number of photons. Each photon has an energy $E = \hbar \omega$, where $\omega = ck$ is the angular frequency of the mode. How many modes have an energy per photon of at most $k_B T$?
- As a crude approximation, suppose that in thermal equilibrium, each mode with energy per photon at most $k_B T$ contains exactly one photon, while all other modes contain no photons. Compute the total energy of the photons in the box. (Answer: $(k_B T)^4 L^3 / 8\pi^2 \hbar^3 c^3$.)

Note that the procedure here is different from what we did above. Before, we started with particles and quantized $\oint p dx$ to get the allowed quantum states. Here, we're treating a situation with many particles (photons), which are excitations of an underlying field (the electromagnetic field). In this case, we found the (normal) modes of the classical field, then quantized by saying that photons could occupy these modes. This is the methodology of quantum field theory.

- [4] **Problem 25.** The final result of the problem above is correct dimensionally, but has incorrect numerical factors because of the crude approximations made. In this problem we'll do a more careful analysis to get the right result. This question is self-contained, but background from **T1** and **T2** will be helpful.

- Consider a quantum mode that can support photons of energy E . The mode can be occupied by any whole number of photons. Thus, using the Boltzmann distribution, the probability of having n photons is

$$p_n \propto e^{-nE/k_B T}.$$

Show that the expected number of photons in the mode is

$$\langle n \rangle = \frac{1}{e^{E/k_B T} - 1}.$$

This is the Bose–Einstein distribution.

- Sketch $\langle n \rangle$ as a function of E . How does it behave at high and low E , and do those results make physical sense?

- (c) Using the Bose–Einstein distribution, show that the total energy is

$$U = \frac{L^3 \hbar}{\pi^2 c^3} \int_0^\infty d\omega \frac{\omega^3}{e^{\hbar\omega/k_B T} - 1}$$

where ω is the angular frequency. You'll have to multiply by a factor of two, because there are two independent photon polarizations for each mode we found above. (Note that if we open the box, the photons will fly out, and the frequency distribution of the emitted light will be given by the integrand; this yields Planck's law for blackbody radiation.)

- (d) **[A]** Using an appropriate substitution, show that U is a dimensionful constant times the dimensionless integral


$$\int_0^\infty dx \frac{x^3}{e^x - 1}.$$

To evaluate this integral, expand the denominator as a power series, integrate each term individually, and use the fact that the Riemann zeta function obeys

$$\zeta(s) = \sum_{n=1}^{\infty} \frac{1}{n^s}, \quad \zeta(4) = \frac{\pi^4}{90}.$$

When the smoke clears, you should find that

$$U = \frac{\pi^2 (k_B T)^4 L^3}{15 (\hbar c)^3}.$$

- [5] Problem 26.**  APhO 2002, problem 1. This useful problem covers the other common example of a quantized bosonic field. In the above problems, we quantized the electromagnetic field to get photons. Here, we quantize a displacement field to get phonons.

Idea 7

In problems 24 and 25, we handled a system of bosons (specifically photons) by considering the modes the photons could occupy, then calculating how many photons were in each mode. This was the easiest route. If we had instead fixed the number of photons, then counted the ways they could be distributed among the mode, the combinatorics would have been a complete nightmare, because multiple photons can occupy the same mode.

Fermions, which obey the Pauli exclusion principle, are simpler, because no two can be in the same state. For instance, if there are n noninteracting fermions in a system, then the lowest energy state of the whole system consists of having one fermion occupy the lowest energy state, the second occupy the second-lowest energy state, and so on. (Accounting for interactions makes the problem much more complicated, because it means the energy of a state depends on whether other states are occupied. However, you can explain a surprising amount while completely neglecting interactions.)

Example 6: Tremaine–Gunn Bound

Suppose all of the dark matter in the galaxy is composed of a single kind of fermionic particle, of mass m . The escape velocity of the galaxy is of order $v_{\text{esc}} \sim 10^{-3}c$, and the dark matter density near Earth is $\rho \sim 0.3 \text{ GeV}/(c^2 \text{ cm}^3)$. What's the minimum possible value of m ?

Solution

The reason there's a minimum possible value of m is that, as m gets smaller, we need more dark matter particles. But the Pauli exclusion principle tells us that if we want to add more particles, they need to have higher and higher energy, and at some point the particles will have so much energy they won't be bound to the galaxy at all.

To get a rough estimate, let's suppose the galaxy has length scale L , so that we need at least $N \sim \rho L^3/m$ dark matter particles. They need to have energy less than $E_0 \sim m_0 v_{\text{esc}}^2$. Plugging this into the final result of example 2 and dropping all numeric factors gives

$$\frac{\rho L^3}{m} \gtrsim \frac{m^3 v_{\text{esc}}^3 L^3}{\hbar^3}$$

which yields the bound

$$m \lesssim \left(\frac{\rho \hbar^3}{v_{\text{esc}}^3} \right)^{1/4} \sim 10 \text{ eV}/c^2.$$

A few decades ago, neutrinos were leading dark matter candidates, since they are light fermionic particles that interact very weakly with ordinary matter. But we now know that the neutrino mass is well below this bound, so that nice idea doesn't work. On the other hand, dark matter could still be composed of bosonic particles of much lighter mass.

- [2] **Problem 27.** Consider a system with many noninteracting fermions, and many quantum states. Each quantum state can be either empty or occupied by a fermion. We want to find the probability that a given quantum state, of energy E , is occupied.

- (a) To put a fermion in this state, we need to remove a fermion from some other state. Suppose the energy released by doing this, suitably averaged, is μ . (This is the chemical potential, and it depends on the temperature, the number of fermions, and the number of states and their energies.) Using the Boltzmann distribution, show that the probability of occupancy is

$$\langle n \rangle = \frac{1}{e^{(E-\mu)/k_B T} + 1}.$$

This is the Fermi–Dirac distribution.

- (b) Sketch $\langle n \rangle$ as a function of E for small but nonzero temperature, as well as the limit attained for zero temperature.

- [3] **Problem 28.** In this problem we'll consider the energy of the conducting electrons in a solid at low temperatures. Model a solid as a cubical box of volume $V = L^3$ with periodic boundary conditions.

- (a) Suppose there are $N \gg 1$ electrons in total. If the temperature is low enough, they will fill all

of the quantum states up to energy $\mu = E_F$, where E_F is called the Fermi energy. Show that

$$E_F = \frac{h^2}{2m} \left(\frac{3N}{8\pi V} \right)^{2/3}.$$

Be careful to account for the fact that an electron's spin can have two possible values.

- (b) A sodium crystal has one conduction electron per atom. (The rest of the electrons are bound to individual atoms, and don't matter for the purposes of this problem.) The density and molar mass are

$$\rho = 0.971 \times 10^3 \text{ kg/m}^3, \quad M = 0.023 \text{ kg/mol}.$$

Find N/V and E_F , and use this to evaluate the typical speed v_F of an electron.

- (c) Find a rough estimate for the temperature range at which the above results apply. Is it satisfied for typical solids at room temperature?

[3] **Problem 29** (MIT). [A] This is an advanced problem that is only placed here because the final result is neat. An integer N can be *partitioned* by writing it as a sum of positive integers, and the partition function $p(N)$ is the number of unique ways this can be done. For example,

$$4 = 1 + 1 + 1 + 1 = 1 + 1 + 2 = 2 + 2 = 1 + 3 = 4$$

which implies $p(4) = 5$. Counting the number of partitions of an integer is a very hard combinatorics problem, but we can get an estimate for large N using string theory.

- (a) Consider an ideal string with hard boundary conditions and fundamental angular frequency ω . Show that the number of distinct quantum states with energy $N\hbar\omega$ is $p(N)$.
- (b) Now suppose the string is at temperature T , where T is chosen so that the expected energy is $N\hbar\omega$. In the thermodynamic limit $N \gg 1$, find a relation between N and T . You may use the result $\zeta(2) = \pi^2/6$.
- (c) By approximating the entropy as $S \approx k_B \log p(N)$, show that

$$\frac{\hbar\omega}{kT} = \frac{d \log p(N)}{dN}.$$

Combine this with your previous result to find an estimate for $p(N)$.

To check your answer, the celebrated Hardy–Ramanujan formula (which is more accurate than the very rough treatment we give above) is

$$p(N) \sim \frac{1}{4\sqrt{3}N} \exp \left(\pi \sqrt{\frac{2N}{3}} \right).$$

While we only considered a simple nonrelativistic string here, calculations of this sort show up in the thermodynamics of string theory. For further discussion, see chapter 22 of Zwiebach.

Example 7: Casimir Forces

Consider an infinite uniform string, on which waves propagate with speed v . The string is held fixed by pins at two points separated by a distance L . When the string is in its ground state, what is the resulting force between the pins?

Solution

Of course, in classical mechanics the ground state would just be $y(x) = 0$, and the force would just be the tension T . But there is an additional quantum mechanical contribution, which arises because each of the standing waves between the pins, with angular frequencies $\omega_n = n\pi v/L$, carries a so-called “zero point” energy $\hbar\omega_n$.

As usual, force can be found by differentiating energy, $F = -dE/dx$. By summing up the zero point energy in all of the standing waves, we naively get

$$E = \sum_n \frac{\hbar\omega_n}{2} = \frac{\pi\hbar v}{2L} \sum_{n=1}^{\infty} n = E_0 \sum_{n=1}^{\infty} n = \infty$$

which is rather unhelpful. This result is off for two reasons. First, even when there aren’t any pins, the string still has standing waves, and these waves *also* have a naively infinite energy density. When we move the pins a bit, we change both the zero point energy outside the pins and inside, by infinite amounts, but the *net* change is finite, giving a finite force.

Quantitatively, the energy in between the pins due to the standing waves is

$$E_{\text{plate}} = E_0 \sum_{n=1}^{\infty} n = \infty$$

and the energy we would have had there if the pins didn’t exist is the “continuous” sum,

$$E_{\text{vac}} = E_0 \int_0^{\infty} x \, dx = \infty.$$

The difference should be finite, but we can’t just subtract infinity with infinity, which brings us to the second problem: none of these quantities are *actually* infinite. Any real string will have a finite maximum oscillation frequency – for instance, the wavelength certainly can’t get smaller than the atomic spacing. Alternatively, even if we had an idealized string where E_{vac} was actually infinite, no real pin can perfectly block waves of all frequencies. For sufficiently high frequencies the waves won’t be affected by the pins, so that the sum in E_{plate} eventually behaves like the integral in E_{vac} , leaving a finite difference between the two.

In other words, the difference between E_{plate} and E_{vac} in reality comes from only low n and x . Therefore, let’s “regulate” the two expressions above so that they’re unchanged in this regime, but match each other at high n and x . The simplest way to do this is to take

$$E_{\text{plate}} = E_0 \sum_{n=1}^{\infty} n e^{-\epsilon n}, \quad E_{\text{vac}} = E_0 \int_0^{\infty} x e^{-\epsilon x} \, dx = \frac{E_0}{\epsilon^2}$$

for small ϵ . To handle the sum, let $\alpha = e^{-\epsilon}$, so that

$$E_{\text{plate}}/E_0 = \alpha + 2\alpha^2 + 3\alpha^3 + \dots$$

Now we use the usual trick for arithmetic-geometric series. Note that

$$\alpha E_{\text{plate}}/E_0 = \alpha^2 + 2\alpha^3 + 3\alpha^4 + \dots$$

Subtracting, we find

$$(1 - \alpha)E_{\text{plate}}/E_0 = \alpha + \alpha^2 + \alpha^3 + \dots = \frac{\alpha}{1 - \alpha}.$$

We thus conclude that

$$E_{\text{plate}} = \frac{e^{-\epsilon}}{(1 - e^{-\epsilon})^2} E_0 = E_{\text{vac}} - \frac{1}{12} E_0 + O(\epsilon)$$

where we used a result from **P1**. Finally, when we take ϵ to zero, the difference is simply

$$E = E_{\text{plate}} - E_{\text{vac}} = -\frac{1}{12} E_0.$$

Differentiating gives the force,

$$F = \frac{\pi \hbar v}{24L^2}$$

which turns out to be attractive. Not only is this finite, it's right! Experiments have measured this “Casimir force” precisely for light between two conductors, where $v = c$, and confirmed the expected results.

You're probably suspicious about this derivation because it depends on the arbitrary choice of an exponential suppression. What if the sums and integrals were regulated at high n and x in a different way? Shouldn't the answer depend on the details of the string and pin? Remarkably, the answer is no: the regulator doesn't matter. If you try others, such as $e^{-\epsilon n^2}$ or $1/n^\epsilon$, you'll get the same result; you can find a general proof in chapter 15 of Schwartz's *Quantum Field Theory and the Standard Model*. The reason is that the effect comes from physics at low frequencies, so it doesn't matter how you regulate the high frequencies.

It is for precisely this reason that you will sometimes see the mysterious equation

$$1 + 2 + 3 + \dots = -\frac{1}{12}.$$

It's not really true. Instead, what it physically means is that the difference between the regulated sum and integral is $-1/12$ for any reasonable regulator.