



THE UNIVERSITY *of* EDINBURGH
School of Physics
and Astronomy

Senior Honours Project Template for Writing a Report

Jan Kocka
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Abstract

In this project we continue the work on data from $e4\nu$. We further inspect the data generated with GENIE, mainly looking at the Delta resonances and compare then to data from the CLAS experiment.

Declaration

I declare that this project and report is my own work.

Signature:

Date: 31/10/2021

Supervisor: Dr. Cheryl Patrick

10 Weeks

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1 Background and Motivation

This project concerns neutrinos and their properties and so a brief description of the relevant properties and their scientific significance is presented. Neutrinos are 3 uncharged leptons of different flavours (each corresponding to the other 3, charged, leptons – electron, μ and τ of ascending masses). Neutrinos interact only via the weak force, the possible Feynman diagrams of their interactions with the other fundamental particles through the W and Z bosons are shown in fig. 1. Their existence was originally proposed by Pauli in the 1930s to explain the beta decay, and their mass was long thought to be 0, most notably the standard model assumes this. However at the end of the last century neutrinos have been observed to undergo oscillations which implies non-zero mass. Since then they have been at the forefront of research, nowadays they are accepted to have mass and there are questions about them being their own antiparticles (Majorana) and possibly violating charge-parity symmetry. The answers to these questions could significantly further our understanding of the universe and explain the dominance of matter over antimatter.

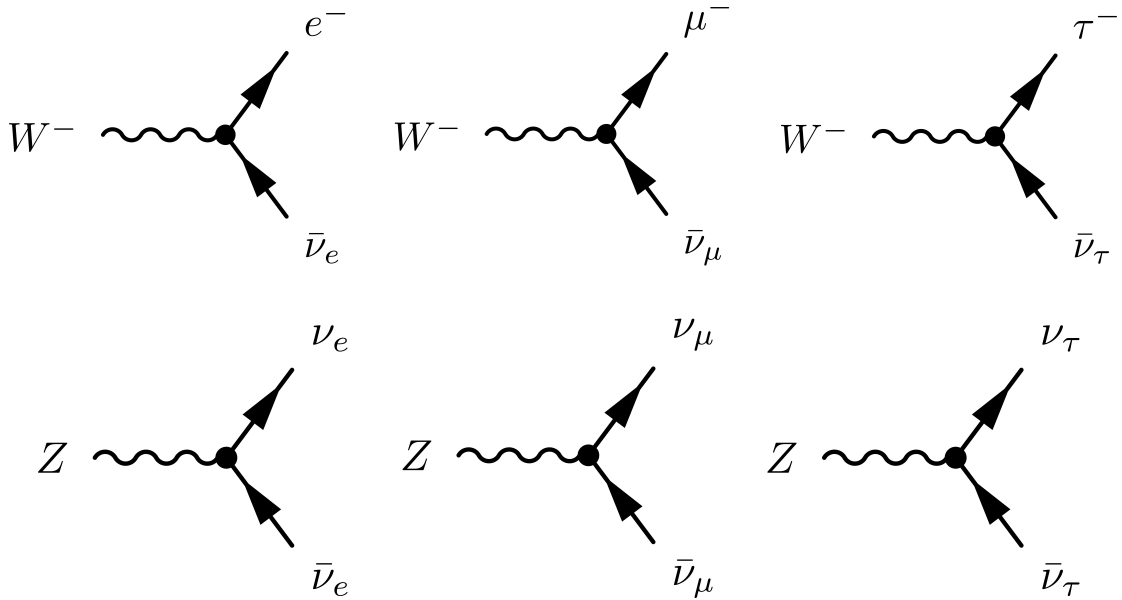


Figure 1: Summary of the necessary fundamental interactions of neutrinos, all other such interactions can be achieved by suitably "rotating" these vertices. This may involve switching some particles for their antiparticles (including the W^- becoming a W^+), say the first could be changed to represent a e^- being converted to a ν_e by an incoming W^+ . Figures taken from [1].

1.1 Neutrino oscillations

In particle physics particle oscillations is a phenomenon that takes place when the particle mass eigenstates which govern its evolution in time are not the same as the state in which it is observed. This phenomenon is not exclusive to neutrinos, the neutral Kaon is known to oscillate to its own antiparticle and vice-versa [2].

In the case of neutrinos, we say there are 3 flavours and the way we differentiate them is by the weak interaction, so say an electron capture takes place when an electron interacts with a proton to form a neutron and a neutrino. This is a weak interaction process and so far all interactions we have observed conserve the 3 separate lepton numbers (electron, μ and τ again) thus we observe that neutrino to be an electron neutrino, so we say it is collapsed into the electron neutrino eigenstate.

Further, that electron eigenstate can be expressed in terms of the mass eigenstates, and if the 2 eigenstate sets aren't identical it will have multiple non-zero components. Then as the mass eigenstates travel differently (the neutrino has a definite E and so through $E^2 = m^2 + p^2$ we get that the mass eigenstates have different momenta and so speeds), the relative components of the state at different positions as it travels will change periodically – oscillate. Thus the probabilities of measuring the neutrino in any of the 3 flavour states also oscillates.

A very important property resulting from the two eigenstate sets being different is the transformation matrix between them. For neutrinos this is the PMNS¹ matrix, it is referred to as U and is such that

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle, \text{ and } |\nu_i\rangle = \sum_\alpha U_{i\alpha}^\dagger |\nu_\alpha\rangle \quad (1)$$

where $|\nu_\alpha\rangle$ are the 3 flavour eigenstates and $|\nu_i\rangle$ the mass eigenstates. The exact form of this matrix then determines many of the neutrino properties, notably if all the matrix elements are real than they maintain Charge-Parity (CP) symmetry.

Finally to complete the picture, the probability of a neutrino measured to be of flavour α to be measured again in the flavour β is given by

$$P(\alpha \rightarrow \beta) = \sum_i |U_{\alpha i} U_{i\beta}^\dagger|^2 + 2 \operatorname{Re} \sum_{j>i} U_{\alpha i} U_{i\beta}^\dagger (U_{\alpha j} U_{j\beta}^\dagger)^* \exp\left(-i \frac{\Delta m_{ij}^2}{2}\right) \frac{L}{E} \quad (2)$$

where E is it's energy, L the distance between our measurements and Δm_{ij}^2 the mass difference of the mass eigenstates i, j squared (for more detail see [3]).

Finally, these probabilities are something we can measure, there are many active experiments working on this from Super-Kamiokande to MINOS, DUNE, Hyper-Kamiokande and many more. Which is where $e4\nu$ comes in, in all of these experiments the neutrinos are either from the sun, microwave cosmic radiation or secondary sources from accelerators[3]. None of these

¹Stands for Pontecorvo–Maki–Nakagawa–Sakata matrix, it is the analogue of the CKM matrix for the mixing of quarks.

2 Results & Discussion

3 Conclusion

References

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- ³K. Zuber, *Neutrino physics*, Third edition., Series in High Energy Physics, Cosmology & Gravitation (Taylor & Francis, Boca Raton, FL, 2020).