

PHYS114 Reference

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1 Forward Model and Photometer

1.1 Forward Model

A mathematical or quantitative description of the problem you're solving. One can use synthetic data and example numbers to test.

1.2 Photometer

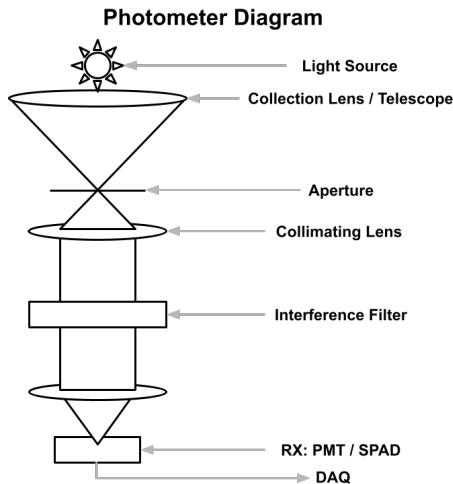


Figure 1: A diagram of a photometer.

- Measures the number of photons from a source.
- IF stands for interference filter.
- RX stands for receiver.

- TX stands for transmitter.
- The receiver can either be a **photomultiplier tube (PMT)** or a **single-photon avalanche diode (SPAD)**.
- **Charge-coupled devices (CCDs)** can be found in smartphones, but scientific-grade CCDs are much more sensitive and expensive.
- **Photon count formula:**

$$N = N_{\text{signal}} + N_{\text{noise}}$$

where

- N is the total number of photons.
- N_{signal} is the number of photons from the desired source.
- N_{noise} is the number of photons from other (undesirable) sources.

- **Noise photon count formula:**

$$N_{\text{noise}} = N_C \Delta t$$

where

- N_C [counts/s] is the noise count rate.
- Δt [s] is the **integration time**, the time interval at which the detector collects photons.

- **Signal photon count formula:**

$$N_{\text{signal}} = BA\Delta t \Omega T_A \nu_{RX}$$

$$\Omega = \frac{A}{r^2} = 2\pi(1 - \cos \alpha)$$

where

- B [photons/(s · m² · str)] is a constant.
- A [m²] for N_{signal} is the area of the telescope.
- Δt [s] is the integration time.
- Ω [str] is the solid angle.
- T_A is the atmospheric transmission.
- ν_{RX} is the receiver efficiency.
- A for N_{noise} is the area of the detector.
- r is the focal length of the telescope.
- α is the half-angle of the cone of light collected by the telescope, also known as the field of view.

Note: $A\Omega$ is a conversed quality called the **etendue**, and it will be explored later.

2 Distributions and Uncertainty

- For **nonstationary** PDFs, the moments are in terms of time.
- For **wide-sense stationary (WSS)** PDFs, only the **first two moments** are stationary, which represents most real-life phenomena.
- For **stationary** PDFs, all the moments are constant and *not* in terms of time.
- **Moment generating function (continuous):**

$$n^{\text{th}} \text{ moment} = \int_{-\infty}^{\infty} x^n \text{ PDF}(x) \partial x$$

- **Moment generating function (discrete):**

$$n^{\text{th}} \text{ moment} = \sum_{i=1}^{\infty} x_i^n \text{ PDF}(x_i)$$

- This class only focuses on the Gaussian, Binomial, and Poisson distributions. The rest are extra information.
- Remember that probabilities multiply. The probability of reproducing an entire dataset is

$$\prod_{i=1}^N \text{PDF}(x_i) = p_1 p_2 \dots p_N$$

- General process of data collection and processing:
 1. Determine PDF.
 2. Create estimators for moments from data.
 3. Determine if estimators are biased or unbiased using **expectation values**.
 4. Determine the chi-squared goodness of fit.
- From strongest to weakest: independent, uncorrelated, and no covariance.

2.1 Gaussian Distribution

Also known as the normal distribution. Represents humans and many natural phenomena.

$$f(x; \mu, \sigma) = \frac{1}{\sigma\sqrt{2\pi}} \exp \left[-\frac{1}{2} \left(\frac{x - \mu}{\sigma} \right)^2 \right]$$

- μ and σ are parameters.

- 1st moment (mean): μ
- 2nd moment (variance): σ^2
- Estimator for 1st moment:

$$\bar{x} = \frac{1}{N} \sum_{i=1}^N x_i$$

- Estimator for 2nd moment:

$$s^2 = \frac{1}{N-1} \sum_{i=1}^N (x_i - \bar{x})^2$$

- 3rd moment (skewness)
- 4th moment (kurtosis)
- Uncertainty for 1st moment:

$$\sigma_{\bar{x}} = \frac{\sigma}{\sqrt{N}}$$

This signals diminishing returns for increasing N .

- However, the above is assuming WSS. In a case without WSS:

$$\bar{x} = \frac{\sum_{i=1}^N \left(\frac{x_i}{\sigma_i^2} \right)}{\sum_{i=1}^N \left(\frac{1}{\sigma_i^2} \right)}$$

$$\sigma_{\bar{x}}^2 = \frac{1}{\sum_{i=1}^N \frac{1}{\sigma_i^2}}$$

2.2 Binomial Distribution

$$\text{PDF}(x; n, p) = \binom{n}{x} p^x (1-p)^{n-x}$$

where

$$\binom{n}{x} = \frac{n!}{x!(n-x)!}$$

- Trials must be **independent**.
- For p large and n large, similar to the **Gaussian distribution**.
- For p low and n large, similar to the **Poisson distribution**.

- n and p are parameters.
- 1st moment: np
- 2nd moment: $np(1 - p)$
- 3rd moment: $\frac{(1-p)-p}{\sqrt{np(1-p)}}$
- Estimator for 1st moment:

$$\bar{x} = \frac{1}{N} \sum_{i=1}^N x_i$$

- Estimator for 2nd moment:

$$s^2 = \frac{1}{N-1} \sum_{i=1}^N (x_i - \bar{x})^2$$

2.3 Poisson Distribution

Represents scenarios with counting and waiting for events.

$$\text{PDF}(x; \lambda) = \frac{\lambda^x e^{-\lambda}}{x!}$$

- λ is the parameter.
- 1st moment: λ
- 2nd moment: λ
- 3rd moment: $\frac{1}{\sqrt{\lambda}}$
- Estimator for 1st moment:

$$\bar{x} = \sum_{i=1}^N x_i$$

- Estimator for 2nd moment:

$$s^2 = \sum_{i=1}^N x_i$$

- Percent error:

$$\frac{1}{\sqrt{\bar{x}}} \times 100\%$$

2.4 Laplace Distribution

$$f(x; \mu, b) = \frac{1}{2b} \exp\left[-\frac{|x - \mu|}{b}\right]$$

- μ and b are parameters.
- μ is a location parameter.
- $b > 0$ is called "diversity."
- 1st moment: μ
- 2nd moment: $2b^2$
- 3rd moment: \emptyset

2.5 Gamma Distribution

$$f(x; \alpha, \beta) = \frac{\beta^\alpha}{\Gamma(\alpha)} x^{\alpha-1} e^{-\beta x}$$

where Γ is the gamma function.

- α and β are parameters.
- α is the shape parameter.
- $\frac{1}{\beta}$ is the scale parameter.
- 1st moment: $\frac{\alpha}{\beta}$
- 2nd moment: $\frac{\alpha}{\beta^2}$
- 3rd moment: $\frac{2}{\sqrt{\alpha}}$

2.6 Assumptions

We often assume that

1. Stationary or WSS PDF: $\sigma_i \approx \sigma$
2. Each measurement is independent and identically distributed (IID).
3. **Method of Maximum Likelihood (MML)**
 - Also known as Maximum Likelihood Method (MLM).
 - Maximum probability of reproducing the data.
 - An alternate method is **Maximum Entropy Method (MEM)** that seeks to add as much disorder as possible to the datapoints, but this class only focuses on MML.
4. An implicit assumption is the zero-mean process. We assume or subtract off the mean to make calculations much easier.

2.7 Expectation Values

$$\langle x_i \rangle = E[x_i] = \mu \quad (\text{typically})$$

Examples:

- We expect the \bar{x} estimator to yield μ : $\langle \bar{x} \rangle = \mu$
- And similarly for s^2 : $\langle s^2 \rangle = \sigma^2$

A rule for expectation values is that you ignore constants:

$$\langle c \rangle = c$$

You can also ignore summations:

$$\left\langle \sum_{i=1}^N x_i \right\rangle = \sum_{i=1}^N \langle x_i \rangle$$

2.8 Uncertainty Propagation

- Uncertainty is represented by σ . For example, the uncertainty in \bar{x} is denoted as $\sigma_{\bar{x}}$.
- Uncertainty can either propagate in a **worst case** or **even-steven** (informally named by our professor) scenario.
- Basics for worst case propagation: Suppose there are height measurements h_1 and h_2 with uncertainties σ_{h_1} and σ_{h_2} . Adding the two heights together also adds their uncertainties: $(h_1 \pm \sigma_{h_1}) + (h_2 \pm \sigma_{h_2}) = (h_1 + h_2) \pm (\sigma_{h_1} + \sigma_{h_2})$. This is the simplest example, but this can get much more complex, especially for operations like multiplication.
- **Even-steven formula** (Taylor series approximation):

$$\begin{aligned} \sigma_x^2 &\approx \sigma_u^2 \left(\frac{\partial x}{\partial u} \right)^2 + \sigma_v^2 \left(\frac{\partial x}{\partial v} \right)^2 + \dots + 2\sigma_{uv}^2 \left(\frac{\partial x}{\partial u} \right) \left(\frac{\partial x}{\partial v} \right) + \dots \\ \sigma_{uv}^2 &= \frac{1}{N-1} \sum_{i=1}^N (u_i - \bar{u})(v_i - \bar{v}) \end{aligned}$$

– σ_{uv}^2 is the covariance, which is 0 when both u and v are **independent**.

2.9 Chi-Squared Goodness of Fit

Given a histogram:

$$\chi^2 \equiv \sum_{j=1}^n \frac{(h(x_j) - NP(x_j))^2}{\sigma_j(h)^2} = \sum_{i=1}^N \frac{(x_i - \bar{x})^2}{\sigma_i^2}$$

where

- n is the number of bins.
 - $h(x_j)$ is the value of the j -th bin.
 - N is the total number of measurements.
 - $P(x_j)$ is the probability of the j -th bin.
 - $\sigma_j(h)^2$ is the variance in the j -th bin.
 - χ^2 is a measure of the goodness of fit of a model to the data.
 - A smaller χ^2 indicates a better fit. Ideal value is 0, but in practice it is around n because the numerator and denominator of the summation are around the same order of magnitude.
 - **reduced chi-squared:**
- $$\chi_{\nu}^2 = \frac{\chi^2}{\nu}$$
- $$\nu = n - n_c$$
- n_c is the number of constraints. Usually 1.
 - ν is the number of degrees of freedom. Usually $n - 1$ with no outliers.