## Function GET\_MW - single-thread version

built-in abundance tables

Calling syntax:

Function parameters:

- 0. Lparms 5-element long integer array of dimensions and global (for all voxels) integer parameters (see below).
- 1. Rparms 3-element double array of global (for all voxels) real parameters (see below).
- 2. Parms array of LOS parameters,  $15 \times Nz$  elements, double. Parms[\*, i] represents the parameters for *i*th voxel (see below).
- 3. T\_arr array of temperatures where DEM/DDM are specified, NT elements, double, in K. The temperature grid is assumed to be the same in all voxels, and the same for both DEM and DDM.
- 4. DEM\_arr array of DEMs, NT × Nz, double, in cm<sup>-6</sup> K<sup>-1</sup>. DEM\_arr[\*, i] represents the DEM for ith voxel.
- 5. DDM\_arr array of DDMs, NT × Nz, double, in cm<sup>-3</sup> K<sup>-1</sup>. DDM\_arr[\*, i] represents the DDM for ith voxel.
- 6. RL input/output array,  $7 \times Nf$ , double. RL[\*, i] corresponds to *i*th frequency (see below).

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NT, DEM\_key, DDM\_key]

- 0. Nz number of voxels along LOS;
- 1. Nf number of frequencies in the spectrum;
- 2. NT number of temperatures in the T\_arr array; must be  $\geq 2$  otherwise DEM/DEM are ignored;
  - 3. DEM\_key global DEM on/off key.
    - a. 0: DEM is enabled: it can be used in all or some voxels, depending on the local DEM on/off keys (see below).
    - b.  $\neq$  0: DEM is disabled for all voxels, regardless of the local DEM on/off keys.
  - 4. DDM\_key global DDM on/off key: same as above, but for DDM.

Array of global real parameters Rparms:

Rparms =  $[S, f_0, \Delta f]$ 

- 0. S visible source area, in cm<sup>2</sup>.
- 1.  $f_0$  starting frequency of the spectrum, in Hz:
  - a. is used, only if  $f_0 > 0$ ;
  - b. if  $f_0 \le 0$ , the frequencies are taken from the RL[0, \*] array.
- 2.  $\Delta f$  logarithmic frequency step (is used only if  $f_0 > 0$ ).

Array of parameters Parms (for a single voxel, 15 parameters):

- 0. Parms[0] =  $\Delta z$  voxel length, in cm.
- 1. Parms[1] =  $T_0$  plasma temperature, in K (is used if DEM or DDM are not specified).
- 2. Parms[2] =  $n_0$  either electron concentration or total atomic concentration (depending on other parameters), in cm<sup>-3</sup> (is used if DEM or DDM are not specified).
  - 3. Parms[3] = B magnetic field strength, in G.
  - 4. Parms[4] =  $\theta$  viewing angle, in degrees.
  - 5. Parms[5] =  $\varphi$  magnetic field azimuthal angle, in degrees.
  - 6. Parms[6] emission mechanism flag (rounded to the nearest integer):
    - a. 0: all emission mechanisms (gyroresonance + free-free + contribution of neutrals) are included;
    - b. 1: gyroresonance is off;
    - c. 2: free-free is off;
    - d. 4: contribution of neutrals is off.

Several flags can be combined by usual or bitwise summation: e.g.,  $mechanism\ flag = 2 + 4\ turns\ off\ both\ free-free\ and\ contribution\ of\ neutrals,\ etc.$ 

- 7. Parms[7] =  $s_{\text{max}}$  maximum cyclotron harmonic number.
- 8. Parms[8] =  $n_p$  proton concentration, in cm<sup>-3</sup> (is used if DEM or DDM are not specified, and the temperature is low).
  - 9. Parms[9] =  $n_{\rm HI}$  neutral hydrogen concentration, in cm<sup>-3</sup>.
  - 10. Parms[10] =  $n_{\text{HeI}}$  neutral helium concentration, in cm<sup>-3</sup>.
  - 11. Parms[11] local DEM on/off key:
    - a. 0: DEM is used (provided that  $NT \ge 2$  and DEM is enabled by the global key);
    - b.  $\neq$  0: DEM in this voxel is ignored even if it is specified;  $T_0$  and  $n_0$  are used instead.
  - 12. Parms[12] local DDM on/off key: same as above, but for DDM.
  - 13. Parms[13] element abundance model:
    - a. 0: coronal (default);
    - b. 1: photospheric (Caffau);
    - c. 2: photospheric (Scott).
- 14. Parms[14] = Vox\_ID voxel type (coronal / chromospheric / etc.), currently ignored.

Input/output array RL:

0. First row (RL[0, \*]) – emission frequencies, in GHz. On input, this array is used if  $f_0$  = Rparms[1]  $\leq$  0; otherwise, the frequencies are computed using the  $f_0$  and  $\Delta f$  parameters:  $f_1 = f_0 10^{\Delta f}$ ,  $f_2 = f_1 10^{\Delta f}$ , etc. On output, this array contains the computed or pre-defined emission frequencies.

Other rows – emission intensities, as observed from the Earth, in sfu:

- 1. RL[1, \*] left polarization, weak mode coupling;
- 2. RL[2, \*] right polarization, weak mode coupling;
- 3. RL[3, \*] left polarization, strong mode coupling;
- 4. RL[4, \*] right polarization, strong mode coupling;
- 5. RL[5, \*] left polarization, exact mode coupling.
- 6. RL[6, \*] right polarization, exact mode coupling.

On input, these arrays specify the emission intensities at the start of the line-of-sight; on output, they contain the emission intensities at the end of the line-of-sight.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

## Function GET\_MW - single-thread version

user-defined abundance tables

Calling syntax:

Function parameters:

- 0. Lparms 8-element long integer array of dimensions and global (for all voxels) integer parameters (see below).
- 1-5. Rparms, Parms, T\_arr, DEM\_arr, DDM\_arr same as in the version with built-in abundance tables; for Parms, see a note below.
- 6. fzeta\_arr array of frequencies where the  $\zeta$ -function is specified, Nf\_zeta elements, double, in Hz. The frequency grid is assumed to be the same in all voxels and for all supplied abundance sets.
- 7. Tzeta\_arr array of temperatures where the  $\zeta$ -function is specified, NT\_zeta elements, double, in K. The temperature grid is assumed to be the same in all voxels and for all supplied abundance sets.
- 8. zeta\_arr array of  $\zeta$ -function values, Nf\_zeta × NT\_zeta × N\_zeta elements, double. This array is the same for all voxels. zeta\_arr[\*, \*, m] represents the 2D  $\zeta$ -function table for mth abundance set.
  - 9. RL same as in the version with built-in abundance tables.

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NT, DEM\_key, DDM\_key, Nf\_zeta, NT\_zeta, N\_zeta]

- 0-4. Nz, Nf, NT, DEM\_key, DDM\_key same as in the version with built-in abundance tables.
  - 5. Nf\_zeta number of frequencies where the  $\zeta$ -function is specified.
  - 6. NT\_zeta number of temperatures where the  $\zeta$ -function is specified.
  - 7. N\_zeta number of supplied 2D  $\zeta$ -function tables (abundance sets).

Array of parameters Parms (for a single voxel):

Most of parameters are the same as in the version with built-in abundance tables, except Parms[13] that specifies the element abundance model index m, so that the 2D  $\zeta$ -function table given by zeta\_arr[\*, \*, m] is used in this voxel.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

## Function GET\_MW\_SLICE - multi-thread version

*built-in abundance tables* 

Calling syntax:

res = call\_external(libname, 'GET\_MW\_SLICE', Lparms\_M, Rparms\_M, \$
Parms\_M, T\_arr, DEM\_arr\_M, DDM\_arrM, RL\_M)

Function parameters:

- 0. Lparms\_M 6-element long integer array of dimensions and global (for all voxels and all LOSs) integer parameters (see below).
- 1. Rparms\_M array of real parameters common for all voxels within each LOS, 3 × Npix, double (see below).
- 2. Parms\_M array of voxel parameters,  $15 \times Nz \times Npix$  elements, double (see below).
- 3. T\_arr array of temperatures where DEM/DDM are specified, NT elements, double, in K. This parameter is the same as in the GET\_MW function: the temperature grid is assumed to be the same in all voxels and all LOSs, and the same for both DEM and DDM.
  - 4. DEM\_arr\_M array of DEMs, NT × Nz × Npix, double, in cm<sup>-6</sup> K<sup>-1</sup> (see below).
- 5. DDM\_arr\_M array of DDMs, NT × Nz × Npix, double, in cm<sup>-3</sup>  $K^{-1}$  (see below).
  - 6. RL\_M input/output array,  $7 \times Nf \times Npix$ , double (see below).

Array of dimensions and global integer parameters Lparms\_M: Lparms\_M = [Npix, Nz, Nf, NT, DEM\_key, DDM\_key]

0. Npix – number of LOSs.

Other elements ( $1^{st}$  to  $5^{th}$ ) are the same as the  $0^{th}$  to  $4^{th}$  elements of the Lparms array in the GET MW function. In particular:

- a. all LOSs have the same number of voxels Nz;
- b. the number of frequencies Nf is the same for all LOSs;
- c. the global DEM and DDM on/off keys are related to all voxels within all LOSs.

Other parameters: sub-arrays Rparms\_M[\*, i], Parms\_M[\*, \*, i], DEM\_arr\_M[\*, \*, i], DDM\_arr\_M[\*, \*, i] and RL\_M[\*, \*, i] correspond respectively to the parameters Rparms, Parms, DEM\_arr, DDM\_arr and RL of the single-thread GET\_MW function, for ith LOS.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

Function GET\_MW\_SLICE – multi-thread version for user-defined abundance tables: not implemented yet.