APPLIED PHYSICS LETTERS VOLUME 80, NUMBER 14 8 APRIL 2002

## Interdependence of absorber composition and recombination mechanism in $Cu(In,Ga)(Se,S)_2$ heterojunction solar cells

M. Turcu, O. Pakma, a) and U. Rau<sup>b)</sup>
Institute of Physical Electronics (IPE), University of Stuttgart, Pfaffenwaldring 47, 70569 Stuttgart, Germany

(Received 13 December 2001; accepted for publication 13 February 2002)

Temperature-dependent current-voltage measurements are used to determine the dominant recombination path in thin-film heterojunction solar cells based on a variety of Cu(In,Ga)(Se,S)<sub>2</sub> alloys. The activation energy of recombination follows the band gap energy of the respective Cu(In,Ga)(Se,S)<sub>2</sub> alloy as long as the films are grown with a Cu-poor final composition. Thus, electronic loss in these devices is dominated by bulk recombination. In contrast, all devices based on absorber alloys with a Cu-rich composition prior to heterojunction formation are dominated by recombination at the heterointerface, with activation energies smaller than the band gap energy of the absorber material. These activation energies are independent from the S/Se ratio but increase with increasing Ga/In ratio. © 2002 American Institute of Physics. [DOI: 10.1063/1.1467621]

Cu-chalcopyrite semiconductors CuGaSe2, CuInS2 and their alloys provide absorber materials for the to date most efficient thin-film solar cell technology. Record chalcopyrite solar cells are made from Cu(In<sub>1-x</sub>Ga<sub>x</sub>)Se<sub>2</sub> thin-film absorbers with  $x \approx 0.15-0.35$ , and band gap energy  $E_g \approx 1.1-1.2 \text{ eV.}^1$  The usage of  $Cu(In_{1-x}Ga_x)(Se_{1-y}S_y)_2$  alloys with higher Ga and/or S content and thus with a wider band gap is desirable when aiming towards high open circuit voltage solar cells. However, solar cells based on those wide-gap Cu-chalcopyrites generally deliver a significantly lower device performance compared to the record devices. The main drawback of those wide-gap chalcopyrite devices is the fact that the open circuit voltage  $V_{\rm OC}$  does not increase at the same rate as  $E_g$ . Thus, it appears that the recombination mechanism changes either qualitatively or quantitatively upon alloying CuInSe2 with high contents of Ga and/or S. It is, however, unclear whether this trend results from lower-grade bulk properties of the wide-gap materials<sup>3,4</sup> or from the less favorable band offset between the CdS buffer layer and the wide-gap chalcopyrite.2

The band diagram in Fig. 1(a) illustrates the basic situation. The discontinuity  $\Delta E_C = E_C^{\rm ch} - E_C^{\rm cdS}$  between the conduction band energy  $E_C^{\rm ch}$  of the chalcopyrite and  $E_C^{\rm cdS}$  of the CdS is close to zero in the case of the standard low-gap  ${\rm Cu}({\rm In}_{1-x}{\rm Ga}_x){\rm Se}_2$  (Ref. 5) but can become large in the case of wide-gap chalcopyrites. This is because alloying  ${\rm CuInSe}_2$  with Ga increases  $E_g$  almost exclusively by rising  $E_C^{\rm ch}$  on the energy scale and alloying with S rises  $E_C^{\rm ch}$  by about half the amount of the change in  $E_g$ . The situation in Fig. 1(a) is unfavorable because the interface recombination barrier  $\Phi_b^p$  between the Fermi energy and the valence band edge of the absorber at the CdS/absorber interface is relatively small and electrons from the CdS can (cross)-recombine with holes

from the absorber. Therefore, a high interface recombination current is considered as an important drawback of wide-gap chalcopyrite solar cells.<sup>2</sup>

At this point, it is important to mention, that there are also two different ways to grow a chalcopyrite thin film: High-efficiency devices based on low-gap Cu(In,Ga)Se<sub>2</sub> absorbers<sup>1</sup> use an overall *Cu-poor* composition of the final film. In the very surface region of those films, a layer with a Cu-poorer composition and a larger band gap with respect to the bulk material develops naturally<sup>8,9</sup> as sketched in Fig. 1(b). High-gap materials, especially CuInS<sub>2</sub>, <sup>10</sup> are grown with a final *Cu-rich* composition where the excess CuS has to be removed by a KCN etch from the film before further device processing. <sup>11</sup> After this etching step, the remaining film has a rather stoichiometric composition and there is no Cu-depleted surface layer. For CuGaSe<sub>2</sub>, both the Cu-poor and the Cu-rich growth methods are used. <sup>12</sup> In all Cu-poor

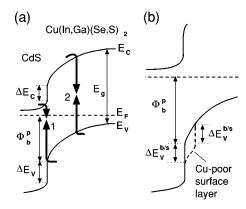


FIG. 1. (a) Band alignments at the interface between CdS and the high-gap  $\operatorname{Cu}(\operatorname{In},\operatorname{Ga})(\operatorname{Se},\operatorname{S})_2$  chalcopyrite and possible recombination paths at the heterointerface and in the space charge region of the absorber. The quantities  $E_C$ ,  $E_V$ ,  $E_g$ , and  $E_F$  denote the conduction band, the valence band, the band gap and Fermi energy, respectively, and  $\Phi_b^p$  indicates the potential barrier for interface recombination. (b) The surface of as-grown Cu-poor chalcopyrite thin films naturally display a Cu-poorer surface layer which widens the band gap of the absorber towards the heterojunction interface by lowering  $E_V$ . As a result the barrier  $\Phi_b^p$  for interface recombination is enlarged.

a)Permanent address: Faculty of Arts and Science, University of Mugla, Turkey.

b) Author to whom all correspondence should be addressed; electronic mail: uwe.rau@ipe.uni-stuttgart.de

 $Cu(In_{1-x}Ga_x)Se_2$ -based solar cells, the limiting factor for the open circuit voltage  $V_{OC}$  is the recombination in the bulk of the absorber material while  $Cu(In,Ga)S_2$  (Refs. 14 and 15) and those  $CuGaSe_2$  devices where the films are prepared under Cu-rich conditions are dominated by interface recombination. Therefore, it seems well possible that the borderline between bulk and interface recombination is a matter of Cu-poor and Cu-rich preparation rather than a result of the difference between small-gap and wide-gap materials as argued earlier.

This letter systematically investigates the composidependence of recombination in ZnO/CdS/ Cu(In,Ga)(Se,S)<sub>2</sub> thin-film solar cells. From temperaturedependent current-voltage measurements (TIV), we find that the activation energy  $E_a$  of the recombination follows the band gap energy  $E_g$  of the absorber as long as the films are prepared with a final Cu-poor composition. Thus, electronic loss in those devices is dominated by bulk recombination. In contrast, all devices based on absorber alloys with a Cu-rich composition prior to heterojunction formation are dominated by recombination at the heterointerface with an activation energy  $E_a < E_g$ . Moreover,  $E_a$  is independent from the S/(S+Se) ratio y but increases as long as  $E_g$  is increased by Ga alloying.

 $Cu(In_{1-x}Ga_x)(Se_{1-y}S_y)_2$  absorber layers are prepared by coevaporation of the elements onto Mo-coated soda lime glass. We use a single-step evaporation procedure with the evaporation rates of all elements kept constant until an approximately 2  $\mu$ m thick film is deposited. Due to the geometrical arrangement of the evaporation sources, we obtain samples with overall Cu-poor and Cu-rich composition from the same preparation run. The absorber material is etched in KCN solution prior to heterojunction formation to remove the excess Cu(Se,S) secondary phases<sup>11</sup> from the Cu-rich part of the film. Device finalization is accomplished by chemical bath deposition of a 50 nm thick CdS buffer layer and by sputter deposition of a 300 nm thick ZnO window layer followed by the evaporation of an Al metal grid as a front contact. The absorber composition is determined by energy dispersive x-ray spectroscopy. TIV measurements are made using a liquid-N<sub>2</sub> cooled cryostate and a Keithley 2400 source meter. A halogen lamp serves for illuminating the sample and a set of neutral density filters for adjusting the light intensity. For the determination of the absorber band gap energy, we use spectral quantum efficiency measurements.

The evaluation of the TIV measurements uses the short circuit current densities  $j_{\rm SC}$  and the open circuit voltages  $V_{\rm OC}$  at the various temperatures and illumination intensities. The interdependence of  $j_{\rm SC}$  and  $V_{\rm OC}$  is given by  $^{16}$ 

$$j_{\text{SC}} = j_0 \exp\left(\frac{qV_{\text{OC}}}{AkT}\right) = j_{00} \exp\left(\frac{qV_{\text{OC}}}{AkT}\right) \exp\left(\frac{-E_a}{AkT}\right),$$
 (1)

where A and  $j_0$  are the ideality factor and the saturation current density of the diode, kT/q is the thermal voltage,  $j_{00}$  is a weakly temperature-dependent prefactor, and  $E_a$  is the activation energy of recombination. Resolving Eq. (1) for  $V_{\rm OC}$  yields

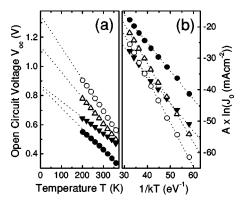


FIG. 2. (a) Extrapolation of the open circuit voltage  $V_{\rm OC}$  toward 0 K for Cu-rich (full symbols) and Cu-poor (open symbols) ZnO/CdS/Cu(In<sub>1-x</sub>Ga<sub>x</sub>)(Se<sub>1-y</sub>S<sub>y</sub>)<sub>2</sub> heterojunction solar cells. (b) The corrected saturation current density A  $\ln(j_0)$  versus inverse thermal energy 1/kT for the corresponding heterojunctions in (a). The slopes provide the activation energies  $E_a$  of the recombination process; closed circles: x=0,y=0.41,  $E_g=1.15$  eV, and  $E_a=0.90$  eV; closed triangles: x=0,y=0.87,  $E_g=1.43$  eV, and  $E_a=0.89$  eV; open triangles: x=0.25, y=0.27,  $E_g=1.22$  eV, and  $E_a=1.25$  eV; open circles: x=0.31, y=0.54,  $E_g=1.49$  eV, and  $E_a=1.40$  eV.

$$V_{\rm OC} = \frac{E_a}{q} - \frac{AkT}{q} \ln \left( \frac{j_{00}}{j_{SC}} \right). \tag{2}$$

If A,  $j_{SC}$ , and  $j_{00}$  are independent of temperature T, a plot of  $V_{OC}$  versus T should yield a straight line and the extrapolation of this line to T = 0 K gives the activation energy. Reorganizing Eq. (1) yields the relationship

$$A \ln(j_0) = \frac{-E_a}{kT} + A \ln(j_{00}). \tag{3}$$

Thus, a plot of the corrected saturation current density  $A \ln(j_0)$  versus the inverse thermal energy 1/kT should yield a straight line where the slope provides  $E_a$ . This activation energy is either the band gap energy  $E_g$  in case of bulk recombination [path 2 in Fig. 1(a)], or the barrier  $\Phi_b^p$  in case of interface recombination (path 1).

The extrapolation of  $V_{\rm OC}$  to T=0 K is illustrated in Fig. 2(a) for several devices made from Cu-rich or Cu-poor Cu(In,Ga)(Se,S)<sub>2</sub> absorbers. The extrapolated values correspond to the band gap energy of the absorbers for the Cu-poor devices, whereas the extrapolation for the Cu-rich devices yields  $V_{\rm OC}(0)$  which is considerably smaller than  $E_g/q$ . The modified Arrhenius plots of  $A \ln(j_0)$  versus 1/kT for the same samples as in Fig. 2(a) are shown in Fig. 2(b), where the saturation currents are obtained from the analysis of the illuminated current–voltage characteristics by plotting  $j_{\rm SC}$  versus  $V_{\rm OC}$  for different illumination intensities. The slopes of the Arrhenius plots yield activation energies which are similar to the extrapolated values  $V_{\rm OC}$  (T=0) from Fig. 2(a). Thus both methods yield consistent results.

Figure 3 plots the activation energy  $E_a$  obtained from corrected Arrhenius plots of various  $\operatorname{Cu}(\operatorname{In}_{1-x}\operatorname{Ga}_x)(\operatorname{Se}_{1-y}\operatorname{S}_y)_2$  devices as a function the absorber band gap energy  $E_g$ . Let us first discuss the absorbers grown with a Cu-rich final composition. Here, the activation energy of all  $\operatorname{CuIn}(\operatorname{Se}_{1-y}\operatorname{S}_y)_2$  devices is approximately 0.9 V independently from  $E_g$  and the Cu-rich  $\operatorname{Cu}(\operatorname{In}_{1-x}\operatorname{Ga}_x)$   $(\operatorname{Se}_{1-y}\operatorname{S}_y)_2$  devices  $(x\approx 0.25)$  exhibit  $E_a$  of around 1.1 V again independent from  $E_g$ . Thus,  $E_a < E_g$  holds in all Cu-

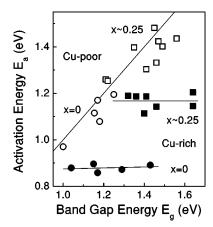


FIG. 3. Activation energies  $E_a$  of the recombination process as a function of the band gap energy  $E_g$  of Cu-poor (open symbols) and Cu-rich (full symbols)  $\operatorname{Cu}(\operatorname{In}_{1-x}\operatorname{Ga}_x)(\operatorname{Se}_{1-y}\operatorname{S}_y)_2$  absorbers with various y and with x=0 (circles) and  $x\approx 0.25$  (squares). The activation energy for the Cu-poor devices follows the band gap energy of the absorber material. In the case of Cu-rich devices, the activation energy is independent from the band gap energy  $E_g$  as long as the variation of  $E_g$  is achieved by varying the S/(S + Se) ratio y at a constant  $\operatorname{Ga}/(\operatorname{Ga}+\operatorname{In})$  ratio x.

rich devices which are therefore dominated by interface recombination and the measured  $E_a$  corresponds to the potential barrier  $\Phi^p_b$  [see Fig. 1(a)]. Interestingly, this barrier is not affected by the S/Se ratio, whereas an increase of the Ga/(Ga+In) ratio x increases  $\Phi^p_b$  by about the same amount as  $E_g$  is increased by the Ga admixture. Apparently, the decrease of  $E_V$  expected under S/Se alloying  $^{6,7}$  is compensated by a decreasing Fermi energy position at the heterointerface, such that the barrier  $\Phi^p_b$  remains constant. In turn, the increase of  $\Phi^p_b$  with increasing Ga/(Ga+In) ratio x must result from an upward shift of the Fermi level at the interface because  $E_V$  remains constant when alloying CuInSe $_2$  with  $Ga.^{6,7}$ 

The activation energies derived from the Cu-poor devices in Fig. 3 follow the band gap energy  $E_g$  regardless whether the band gap variation is achieved by alloying with S or Ga. Thus, *all* Cu(In,Ga)(Se,S)<sub>2</sub> devices grown with a final Cu-poor composition are dominated by bulk recombination. The fact is, however, hard to explain for wide-gap chalcopyrites, as  $\Phi_b^p = E_g - \Delta E_C$  is much smaller than  $E_g$  for large  $\Delta E_C$  and the interface recombination should prevail under these circumstances. The fundamentally different recombination behavior of Cu-poor chalcopyrite devices com-

pared to their Cu-rich grown counterparts can only be explained when taking into account the surface layer with a higher band gap energy than that of the bulk material. This band gap widening towards the surface is accommodated by a bulk/surface valence band offset  $\Delta E_V^{b/s}$ , which is around 0.4 eV in case of CuInSe2. 5,8 Figure 1(b) shows that such a band offset directly increases the barrier  $\Phi_b^p$  to a value  $\Phi_b^p$  which is now sufficiently large to eliminate interface recombination. 14,17 Thus, the present results suggest that the enlargement of the band gap energy towards the film surface is an important and operative element in *all* Cu-poor Cu(In,Ga)(Se,S)2 films.

The authors thank our colleagues at the IPE, M. Balboul, A. Jasenek, Q. Nguyen, K. Orgassa, F. Pfisterer, and H. W. Schock for fruitful discussions as well as J. H. Werner for continuous support. This work was supported by the German Federal Ministry for Education and Research (BMBF) and by the German–Israeli Foundation (GIF).

<sup>&</sup>lt;sup>1</sup>M. A. Contreras, B. Egaas, K. Ramanathan, J. Hiltner, A. Swartzlander, F. Hasoon, and R. Noufi, Prog. Photovoltaics 7, 311 (1999).

<sup>&</sup>lt;sup>2</sup>R. Herberholz, V. Nadenau, U. Rühle, C. Koble, H. W. Schock, and B. Dimmler, Sol. Energy Mater. Sol. Cells 49, 227 (1997).

<sup>&</sup>lt;sup>3</sup>G. Hanna, A. Jasenek, U. Rau, and H. W. Schock, Phys. Status Solidi A 179, R7 (2000).

<sup>&</sup>lt;sup>4</sup>J. Reiß, J. Malmström, A. Werner, I. Hengel, R. Klenk, and M. C. Lux-Steiner, Mater. Res. Soc. Symp. Proc. 668, H9.4.1 (2001).

<sup>&</sup>lt;sup>5</sup> M. Morkel, L. Weinhardt, B. Lohmüller, C. Heske, E. Umbach, W. Riedl, S. Zweigart, and F. Karg, Appl. Phys. Lett. **79**, 4482 (2001).

<sup>&</sup>lt;sup>6</sup>S.-H. Wei and A. Zunger, J. Appl. Phys. **78**, 3846 (1995).

<sup>&</sup>lt;sup>7</sup>M. Turcu, I. M. Kötschau, and U. Rau, Appl. Phys. A: Mater. Sci. Process. **73**, 769 (2001).

<sup>&</sup>lt;sup>8</sup>D. Schmid, M. Ruckh, F. Grunwald, and H. W. Schock, J. Appl. Phys. **73**, 2902 (1993).

<sup>&</sup>lt;sup>9</sup> D. Schmid, M. Ruckh, and H. W. Schock, Solar Energy Mater. Sol. Cells 41, 281 (1996).

<sup>&</sup>lt;sup>10</sup> J. Klaer, J. Bruns, R. Henninger, K. Siemer, R. Klenk, K. Ellmer, and D. Bräunig, Semicond. Sci. Technol. 13, 1456 (1998).

<sup>&</sup>lt;sup>11</sup>T. Walter, A. Content, K. O. Velthaus, and H. W. Schock, Solar Energy Mater. Sol. Cells 26, 357 (1992).

<sup>&</sup>lt;sup>12</sup> V. Nadenau, U. Rau, A. Jasenek, and H. W. Schock, J. Appl. Phys. 87, 584 (2000)

<sup>&</sup>lt;sup>13</sup> U. Rau, M. Schmidt, A. Jasenek, G. Hanna, and H. W. Schock, Solar Energy Mater. Sol. Cells 67, 137 (2001).

<sup>&</sup>lt;sup>14</sup>R. Klenk, Thin Solid Films **387**, 135 (2001).

<sup>&</sup>lt;sup>15</sup>I. Hengel, A. Neisser, R. Klenk, and M. C. Lux-Steiner, Thin Solid Films 361, 458 (2000).

<sup>&</sup>lt;sup>16</sup>U. Rau, A. Jasenek, H. W. Schock, F. Engelhardt, and T. Meyer, Thin Solid Films 361, 299 (2000).

<sup>&</sup>lt;sup>17</sup>T. Dullweber, G. Hanna, U. Rau, and H. W. Schock, Solar Energy Mater. Sol. Cells 67, 145 (2001).