

Radio-Interferometric Measurement Equation

Introductory Radio Interferometry Course

Radio Astronomy Techniques and Technologies Group
(RATT)

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Radio-Interferometric Measurement Equation (RIME)

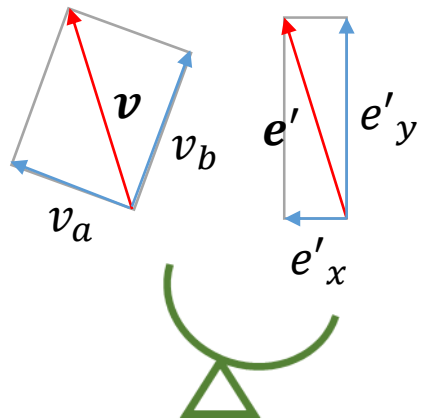
- Compact, intuitive, **matrix**-based way of representing **propagation effects** in radio interferometry.
- Useful for **calibration** (solving for and correcting these propagation effects).

Introduction

e'_x, e'_y : Components of electric field vector
in reference frame of sky, at the observer

v_a, v_b : Voltages measured by antenna feed
(linearly or circularly polarized)

Propagation effects



Antenna

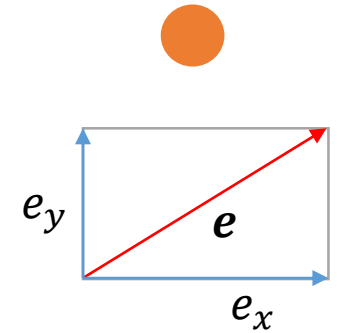
Can be represented
as vectors:

$$\mathbf{e} = \begin{pmatrix} e_x \\ e_y \end{pmatrix}$$

$$\mathbf{e}' = \begin{pmatrix} e'_x \\ e'_y \end{pmatrix}$$

$$\mathbf{v} = \begin{pmatrix} v_a \\ v_b \end{pmatrix}$$

Source



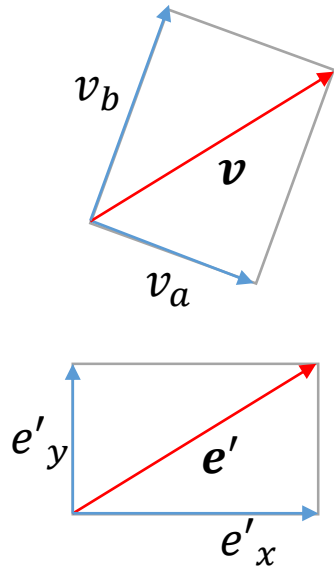
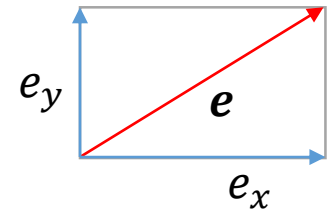
e_x, e_y : Components of electric field vector
in reference frame of sky, at the source

Propagation effects absent

Amplitude and direction of electric vector
remain unchanged during propagation

No propagation effects

Source

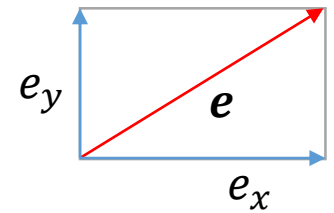


Antenna

Propagation effects present

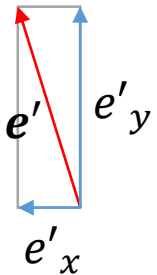
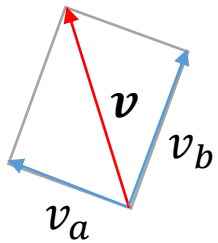
Amplitude and direction of electric vector
change during propagation

Source



Propagation effects

Linear transformation matrix, \mathbf{J}



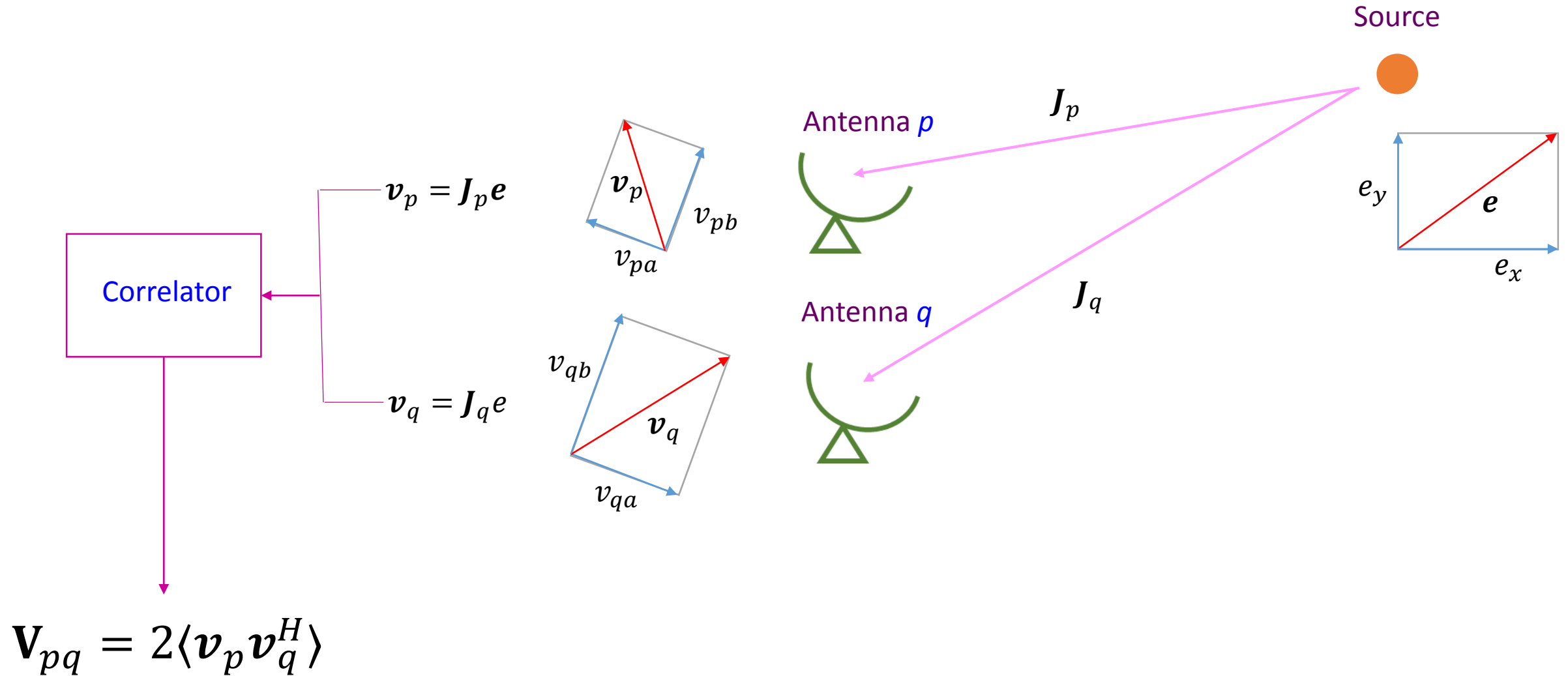
Antenna

Jones matrix

Voltage vector $\leftarrow \mathbf{v} = \mathbf{J}\mathbf{e} \rightarrow$ Electric field vector

$$\begin{pmatrix} v_a \\ v_b \end{pmatrix} = \begin{pmatrix} j_{11} & j_{12} \\ j_{21} & j_{22} \end{pmatrix} \begin{pmatrix} e_x \\ e_y \end{pmatrix}$$

Correlation



Visibility

- The correlator computes the **visibility**, \mathbf{V}_{pq} , on the baseline pq :

$$\begin{aligned}
 \mathbf{V}_{pq} &= 2 \underbrace{\langle \mathbf{v}_p \mathbf{v}_q^H \rangle}_{\text{Average}} \\
 &= 2 \left\langle \begin{pmatrix} v_{pa} \\ v_{pb} \end{pmatrix} \begin{pmatrix} v_{qa}^* & v_{qb}^* \end{pmatrix} \right\rangle \\
 &= 2 \begin{pmatrix} \langle v_{pa} v_{qa}^* \rangle & \langle v_{pa} v_{qb}^* \rangle \\ \langle v_{pb} v_{qa}^* \rangle & \langle v_{pb} v_{qb}^* \rangle \end{pmatrix}
 \end{aligned}$$

Hermitian conjugate

Outer product

Average

These 4 quantities are the outputs from the correlator

Correlation

$$\mathbf{v}_p = \mathbf{J}_p \mathbf{e} \quad , \quad \mathbf{v}_q = \mathbf{J}_q \mathbf{e}$$

$$\mathbf{V}_{pq} = 2 \langle \mathbf{v}_p \mathbf{v}_q^H \rangle$$

$$= 2 \langle (\mathbf{J}_p \mathbf{e})(\mathbf{J}_q \mathbf{e})^H \rangle$$

$$= 2 \langle \mathbf{J}_p (\mathbf{e} \mathbf{e}^H) \mathbf{J}_q^H \rangle$$

$$= \langle \mathbf{J}_p (2 \mathbf{e} \mathbf{e}^H) \mathbf{J}_q^H \rangle$$

Coherency, or Brightness

$$\mathbf{V}_{pq} = \langle \mathbf{J}_p (2\mathbf{e}\mathbf{e}^H) \mathbf{J}_q^H \rangle$$

By definition, the coherency, or brightness, \mathbf{B} , is given by:

$$\mathbf{B} = \langle 2\mathbf{e}\mathbf{e}^H \rangle = \begin{pmatrix} I + Q & U + iV \\ U - iV & I - Q \end{pmatrix}$$

$\langle \mathbf{e}\mathbf{e}^H \rangle$ is the coherence of the electromagnetic field with itself,
and is described by the Stokes parameters I, Q, U, V

$$\mathbf{V}_{pq} = \mathbf{J}_p \mathbf{B} \mathbf{J}_q^H$$

Radio-Interferometric Measurement Equation (RIME)

Visibility Brightness

$$\mathbf{V}_{pq} = \mathbf{J}_p \mathbf{B} \mathbf{J}_q^H$$

Jones matrices

$$\begin{pmatrix} v_{aa} & v_{ab} \\ v_{ba} & v_{bb} \end{pmatrix} = \begin{pmatrix} j_{11a} & j_{12a} \\ j_{21a} & j_{22a} \end{pmatrix} \begin{pmatrix} I + Q & U + iV \\ U - iV & I - Q \end{pmatrix} \begin{pmatrix} j_{11b} & j_{12b} \\ j_{21b} & j_{22b} \end{pmatrix}^H$$

Component Jones matrices

The **Jones matrix** for an antenna is a product of several component Jones matrices, corresponding to different corrupting effects along the signal path

Example:

$$J = B G D E P T$$

Diagram illustrating the decomposition of the Jones matrix J into component matrices B , G , D , E , P , and T , each corresponding to a specific physical effect:

- B : Bandpass gain
- G : Instrumental gain
- D : Polarization leakage
- E : Primary beam
- P : Parallax angle feed rotation
- T : Ionospheric Faraday rotation

Component Jones matrices

Jones chain:

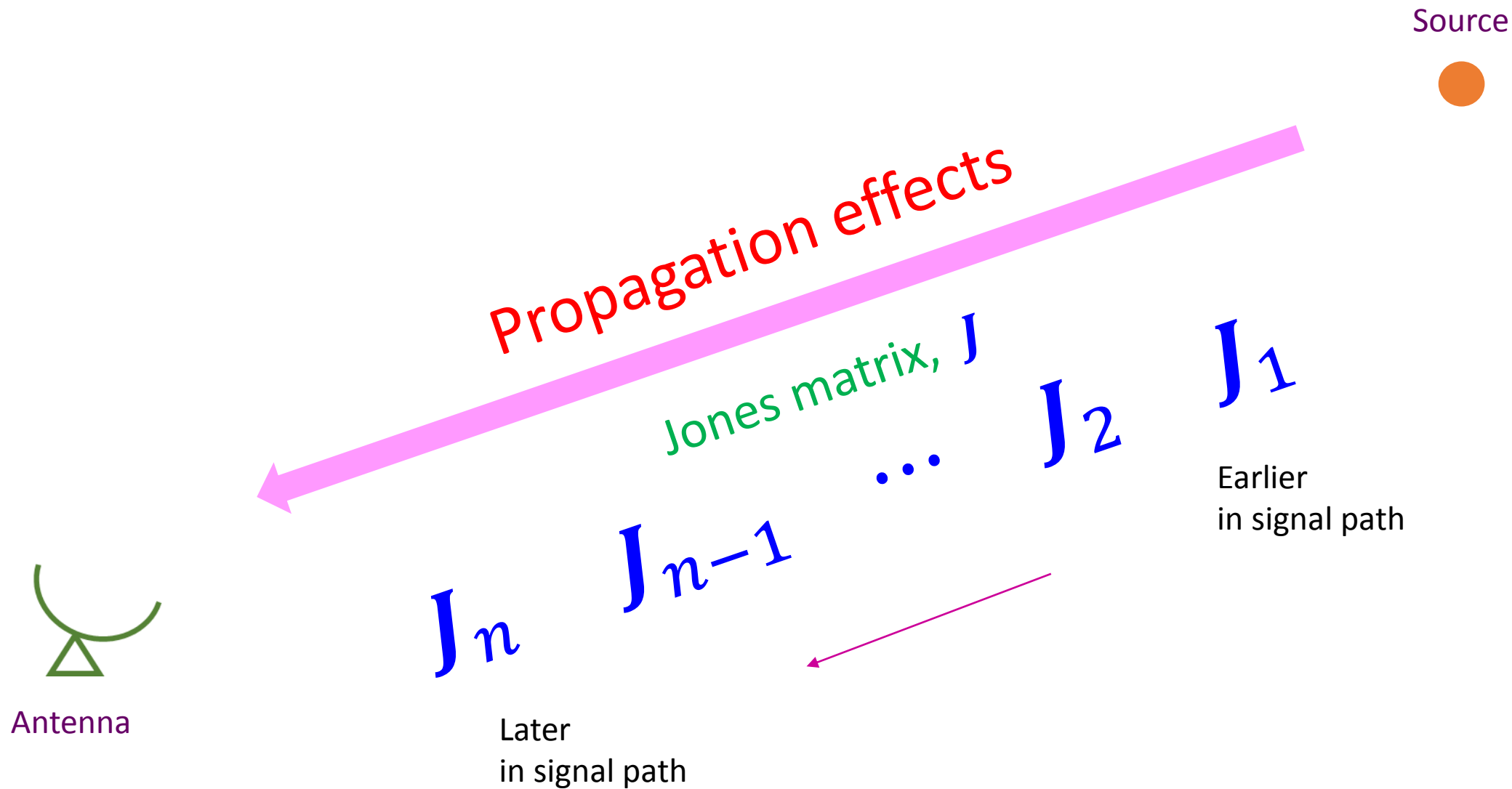
$$J = J_n J_{n-1} \cdots J_2 J_1$$

Later
in signal path



Earlier
in signal path

Component Jones matrices



Component Jones matrices

Antenna p : $J_p = J_{pn} J_{p(n-1)} \cdots J_{p2} J_{p1}$

Antenna q : $J_q = J_{qn} J_{q(n-1)} \cdots J_{q2} J_{q1}$

Visibility Brightness

$$\mathbf{V}_{pq} = J_p \mathbf{B} J_q^H$$

Jones matrices

$$\mathbf{V}_{pq} = J_{pn} J_{p(n-1)} \cdots J_{p2} J_{p1} \mathbf{B} J_{q1}^H J_{q2}^H \cdots J_{q(n-1)}^H J_{qn}^H$$

$$\mathbf{V}_{pq} = J_{pn} \left(J_{p(n-1)} \left(\cdots \left(J_{p2} \left(J_{p1} \mathbf{B} J_{q1}^H \right) J_{q2}^H \right) \cdots \right) J_{q(n-1)}^H \right) J_{qn}^H$$

Calibration

Calibration: Determining and correcting for **propagation effects** in order to compute the **brightness**.

i.e., solve for Jones matrices **J** to compute **B** :

Diagram illustrating the relationship between Visibility, Brightness, and Jones matrices:

$$\mathbf{V}_{pq} = \mathbf{J}_p \mathbf{B} \mathbf{J}_q^H$$

The diagram shows three green labels: 'Visibility' at the top left, 'Brightness' at the top right, and 'Jones matrices' at the bottom center. Three magenta arrows point from these labels to the corresponding terms in the equation: from 'Visibility' to \mathbf{V}_{pq} , from 'Brightness' to \mathbf{B} , and from 'Jones matrices' to \mathbf{J}_p and \mathbf{J}_q^H .

$$\mathbf{B} = \mathbf{J}_p^{-1} \mathbf{V}_{pq} (\mathbf{J}_q^H)^{-1}$$

Direction-independent and direction-dependent effects

Propagation effects can be of two kinds:

- Direction-independent effects
- Direction-dependent effects

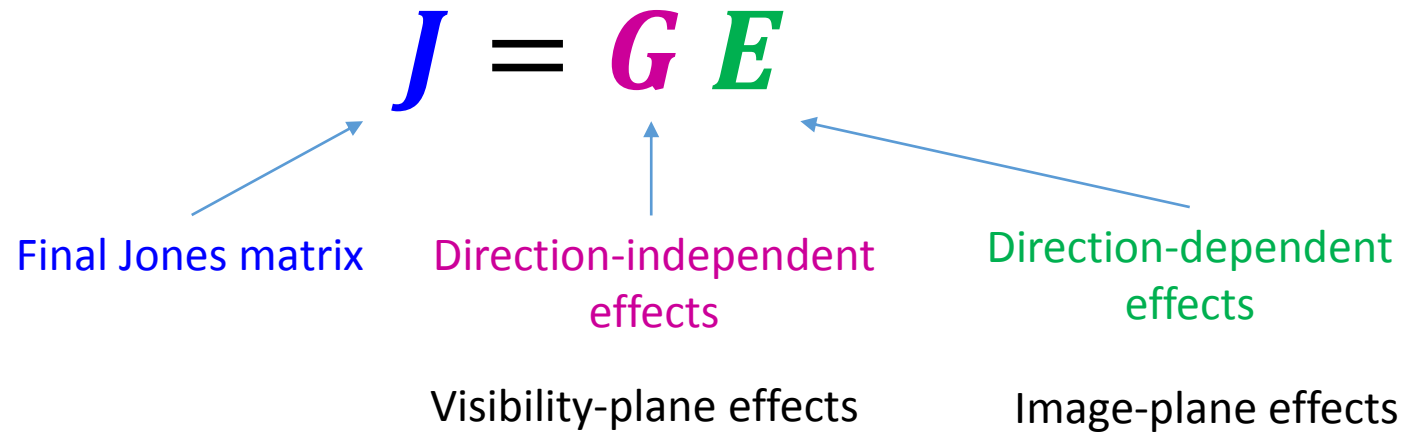
These effects can be represented by different Jones matrices:

The diagram illustrates the equation $J = G E$ where J is blue, G is magenta, and E is green. Three blue arrows point from labels below to the matrices: from 'Final Jones matrix' to J , from 'Direction-independent effects' to G , and from 'Direction-dependent effects' to E .

$$J = G E$$

Final Jones matrix Direction-independent effects Direction-dependent effects

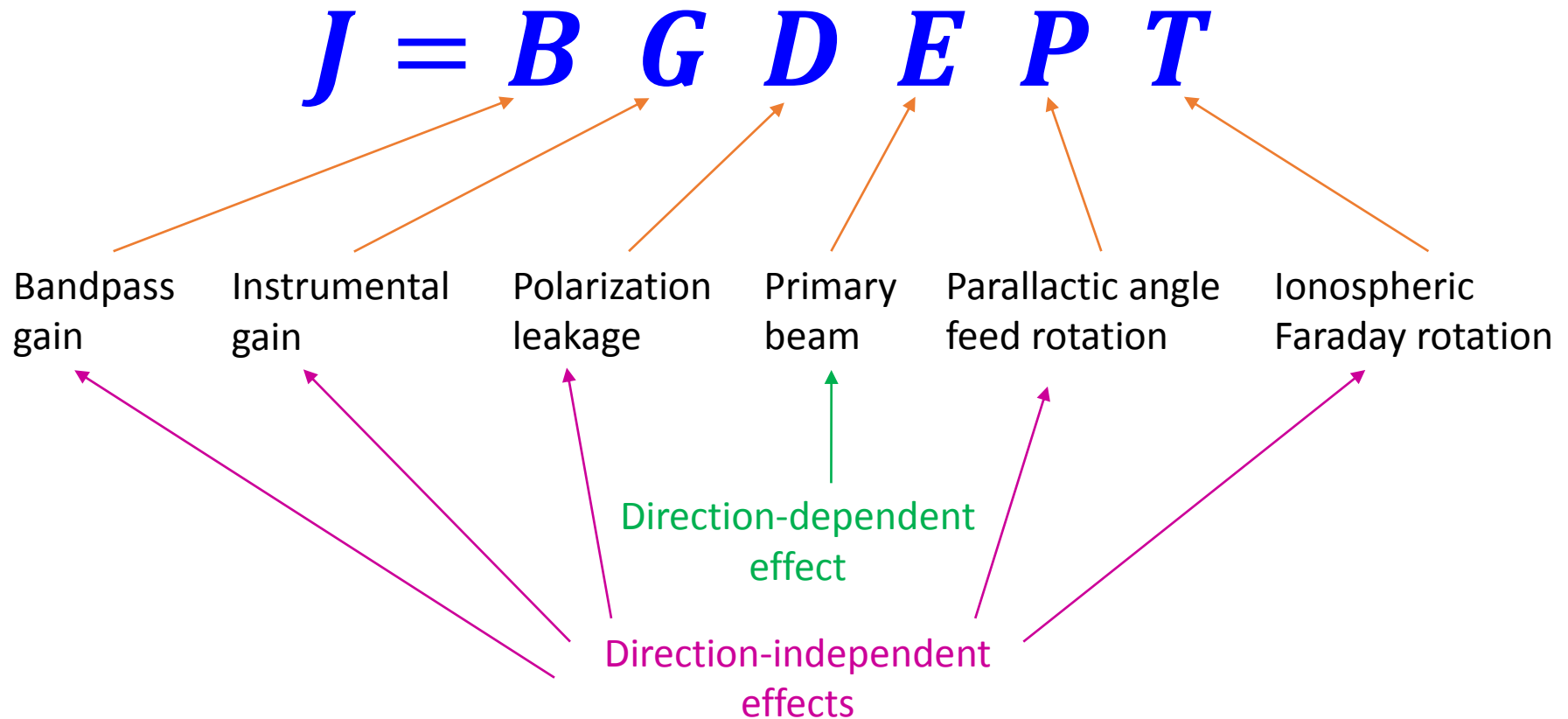
Direction-independent and direction-dependent effects



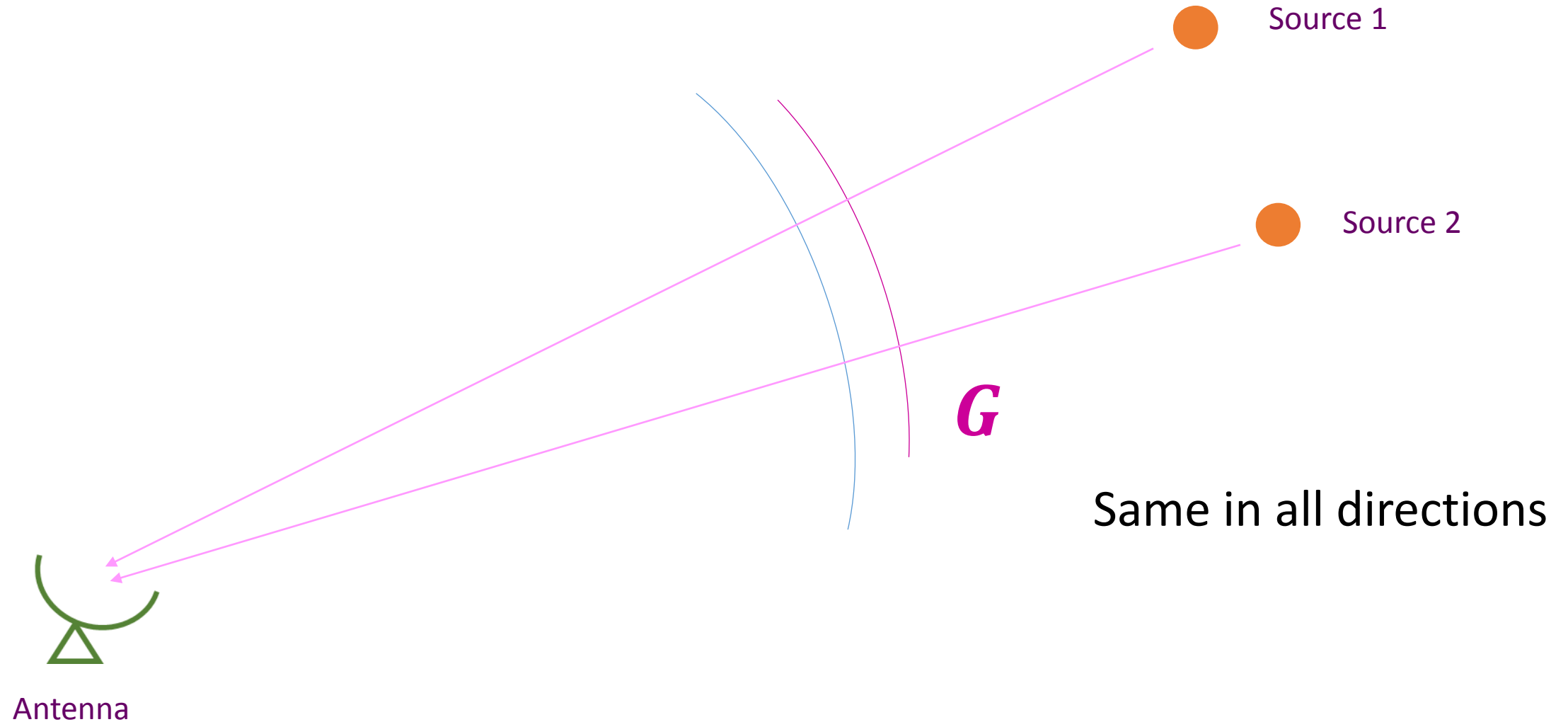
$$\mathbf{V}_{pq} = J_p \mathbf{B} J_q^H$$

$$\mathbf{V}_{pq} = G_p (E_p \mathbf{B} E_q^H) G_q^H$$

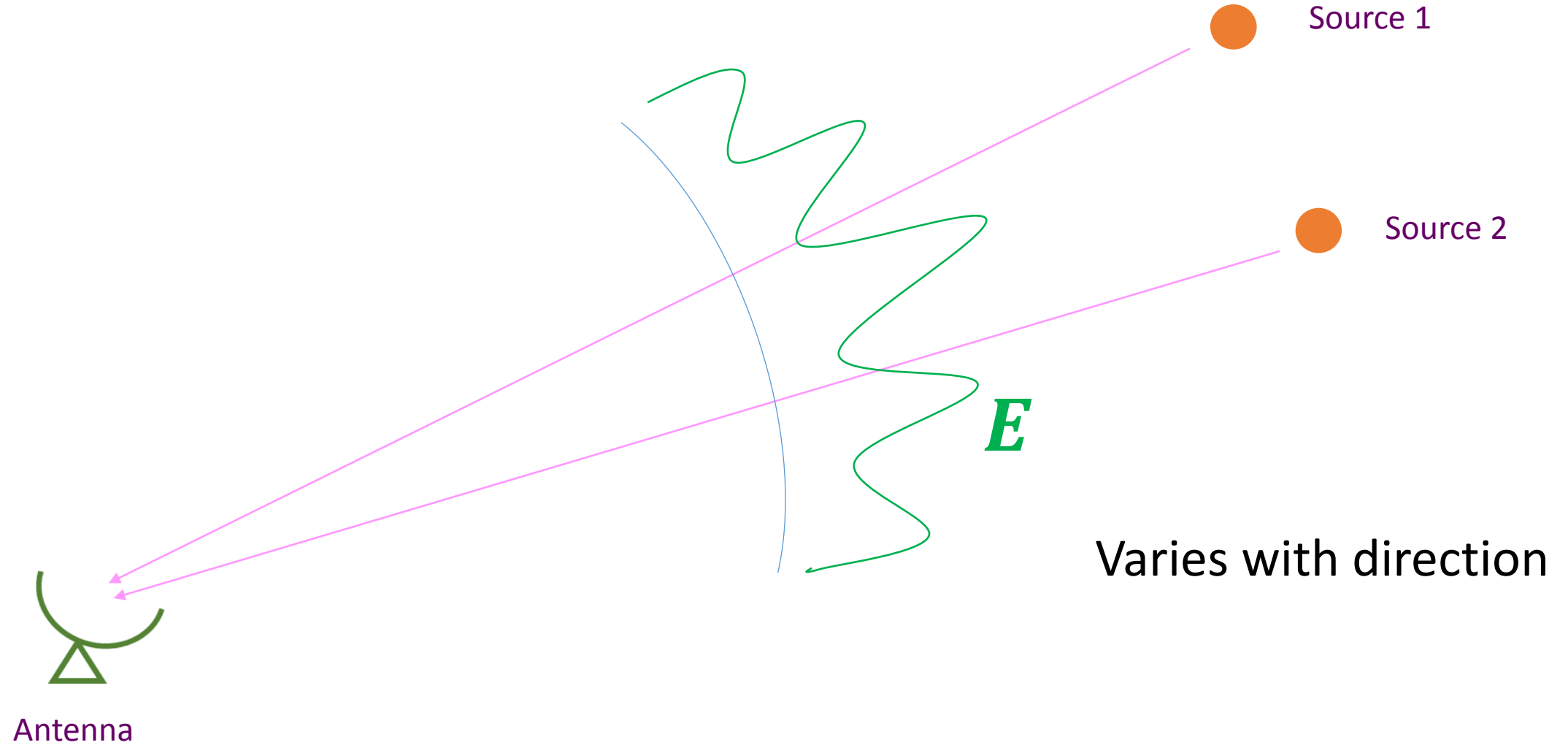
Direction-independent and direction-dependent effects



Direction-independent effects



Direction-dependent effects



Structure of Jones matrices

Most Jones matrices have a simple form:

- $J_{\text{rotation}} = \begin{bmatrix} e^{j\theta} & 0 \\ 0 & e^{-j\theta} \end{bmatrix}$

- $J_{\text{leakage}} = \begin{bmatrix} 1 & D^R \\ D^L & 1 \end{bmatrix}$

- $J_{\text{gain}} = \begin{bmatrix} G^R & 0 \\ 0 & G^L \end{bmatrix}$

(in a circularly polarized basis)

Rotation matrices

- $J_{\text{rotation}} = \begin{bmatrix} e^{j\theta} & 0 \\ 0 & e^{-j\theta} \end{bmatrix}$

Examples:

Parallactic angle feed rotation:

Parallactic angle
↓

$$\mathbf{P} = \begin{bmatrix} e^{-j\alpha} & 0 \\ 0 & e^{j\alpha} \end{bmatrix}$$

Ionospheric Faraday rotation:

$$\mathbf{T} = \begin{bmatrix} e^{j\chi} & 0 \\ 0 & e^{-j\chi} \end{bmatrix}$$

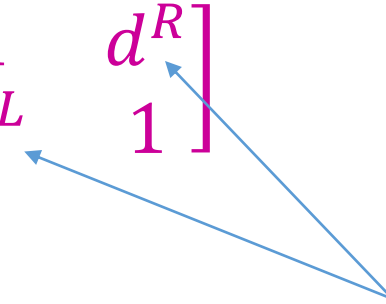
↑
Faraday rotation angle

Leakage matrices

- $J_{\text{leakage}} = \begin{bmatrix} 1 & D^R \\ D^L & 1 \end{bmatrix}$

Examples:

Polarization leakage: $\mathbf{D} = \begin{bmatrix} 1 & d^R \\ d^L & 1 \end{bmatrix}$



Polarization leakage terms

Gain matrices

- $J_{\text{gain}} = \begin{bmatrix} G^R & 0 \\ 0 & G^L \end{bmatrix}$

Examples:

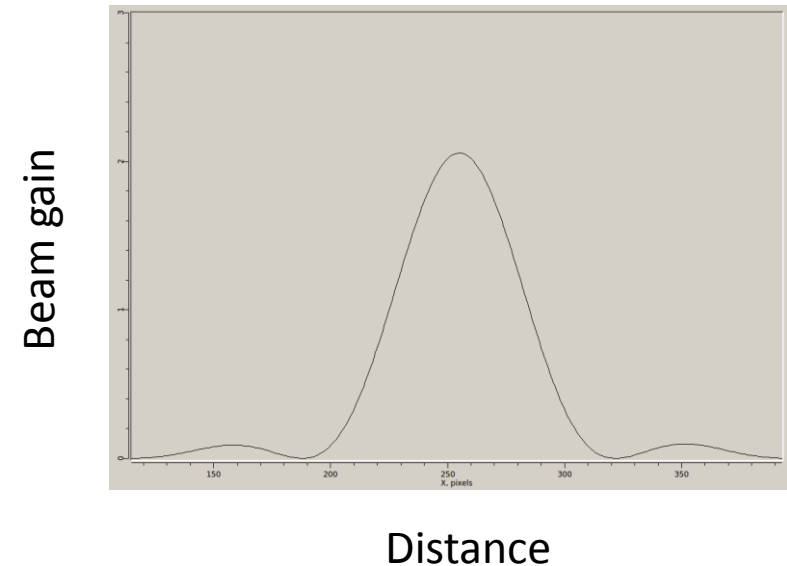
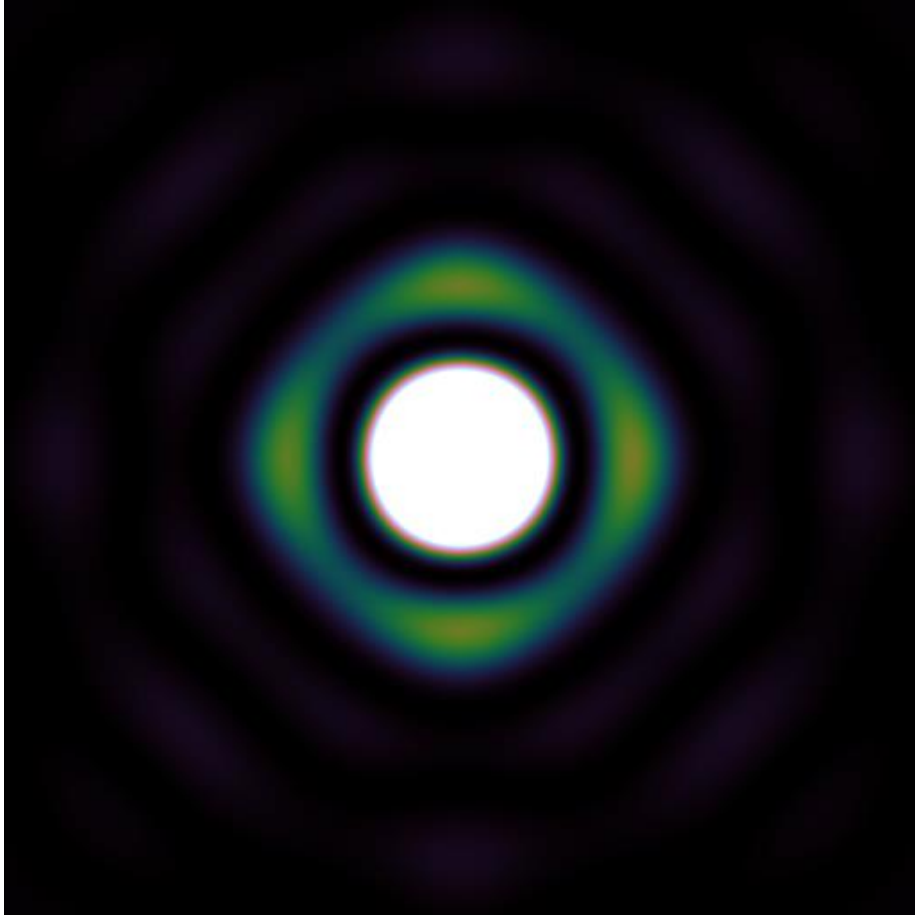
Instrumental gain: $\mathbf{G} = \begin{bmatrix} g^R & 0 \\ 0 & g^L \end{bmatrix} = \begin{bmatrix} a^R e^{j\phi^R} & 0 \\ 0 & a^L e^{j\phi^L} \end{bmatrix}$

Bandpass gain: $\mathbf{B} = \begin{bmatrix} B^R & 0 \\ 0 & B^L \end{bmatrix} = \begin{bmatrix} b^R(\nu) e^{j\psi^R(\nu)} & 0 \\ 0 & b^L(\nu) e^{j\psi^L(\nu)} \end{bmatrix}$

Direction dependent effects: Primary beam

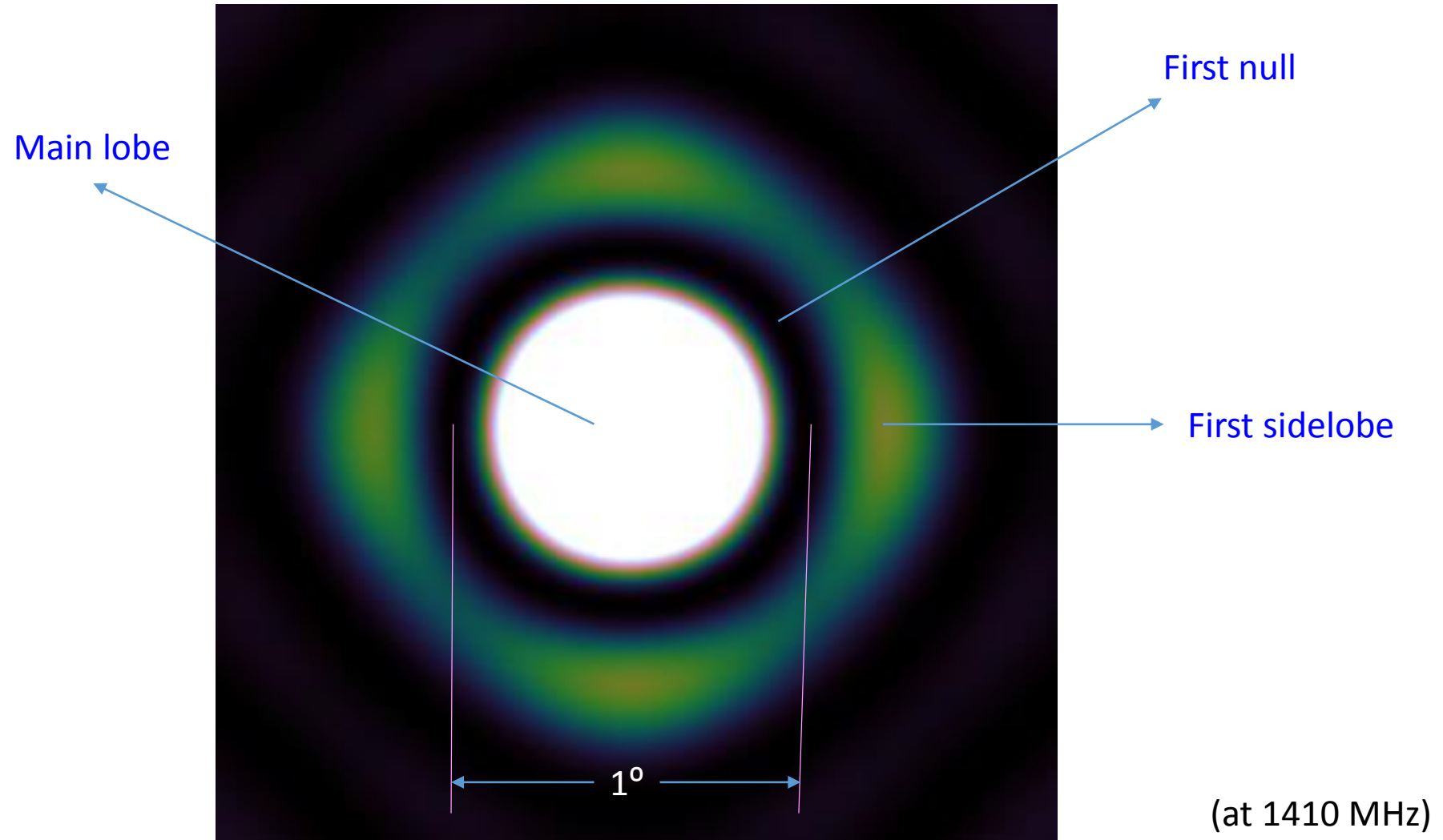
- The primary beam of the antenna is the most important direction-dependent effect.
- Becomes important in wide-field, wide-band observations.
- The primary beam pattern has a multiplicative effect in the image plane, convolutional effect in the visibility plane.
- We will consider the example of a JVLA (Jansky Very Large Array) antenna here.

Primary beam amplitude variation with distance from center



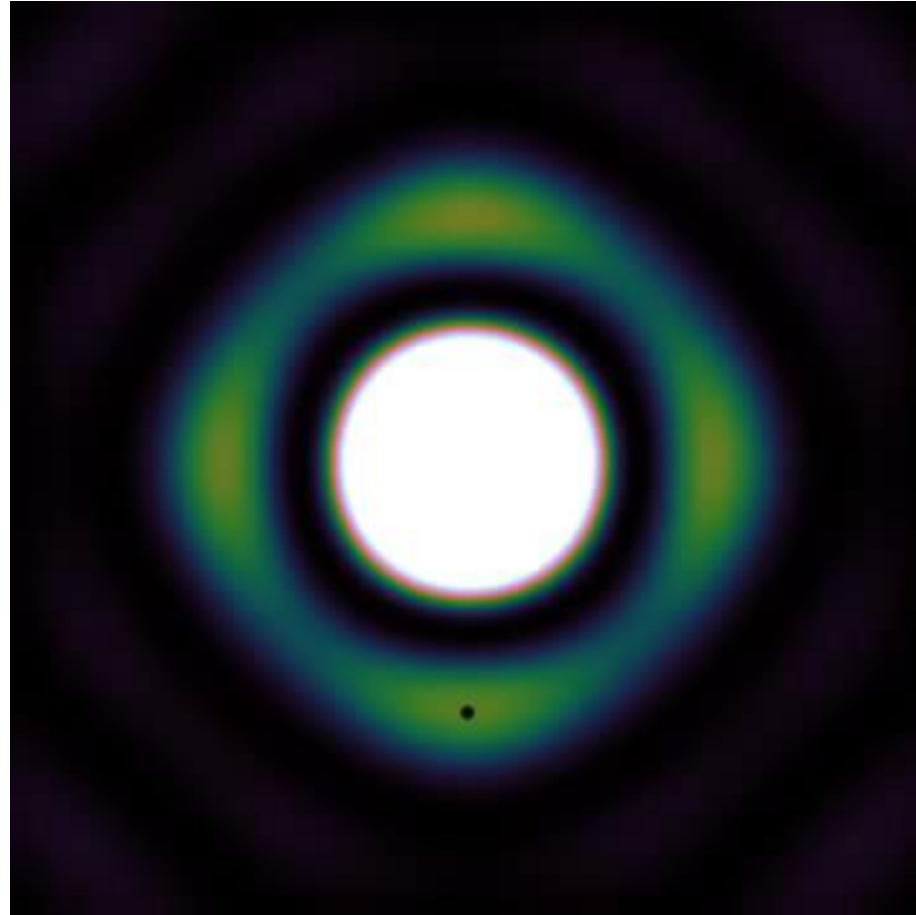
Horizontal cross-section of the beam
through the center

JVLA primary beam

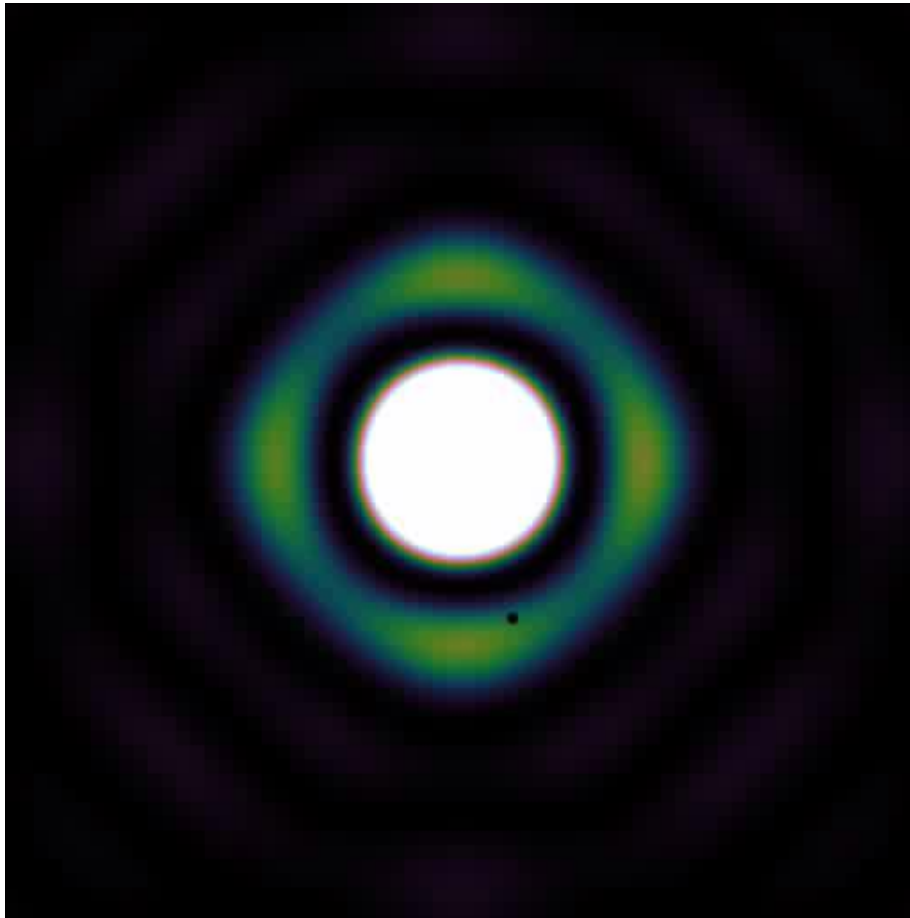


Primary beam rotation

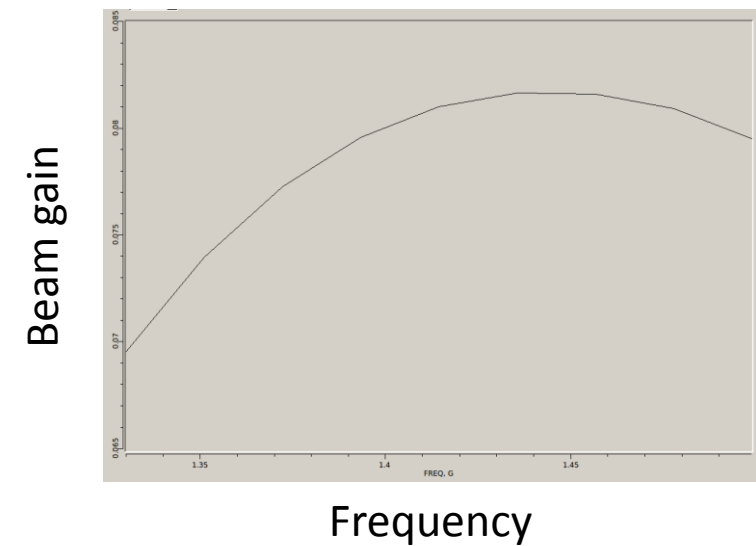
An EVLA antenna has an alt-azimuth mount;
the primary beam rotates during the course of an observation



Variation of primary beam with frequency



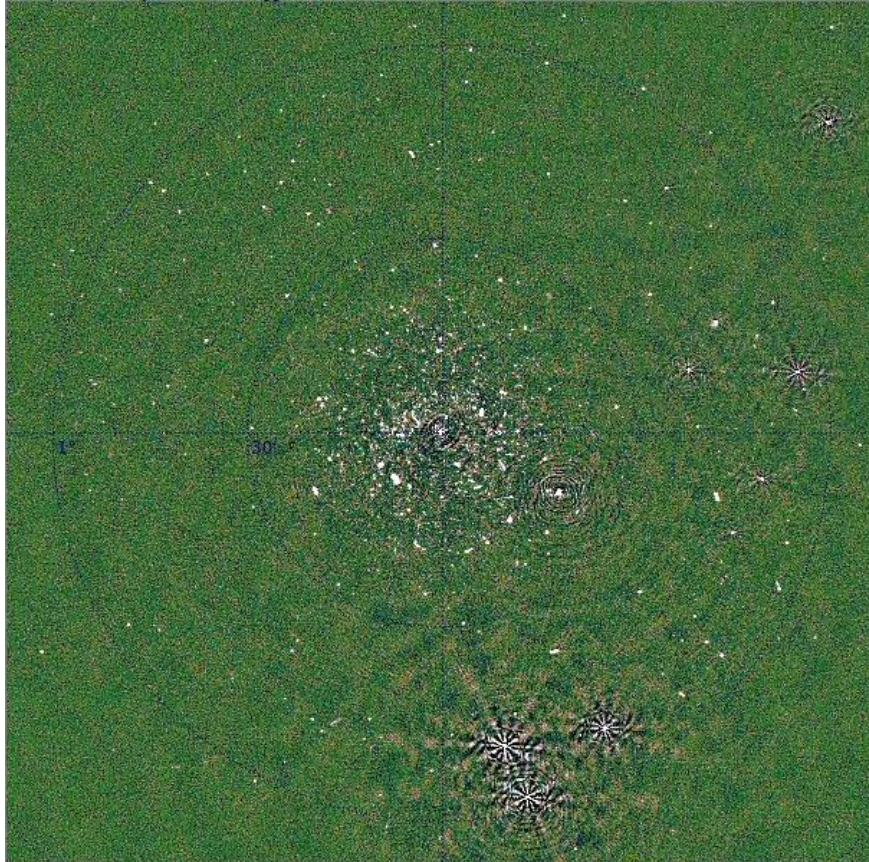
The beam pattern becomes more compact with increasing frequency



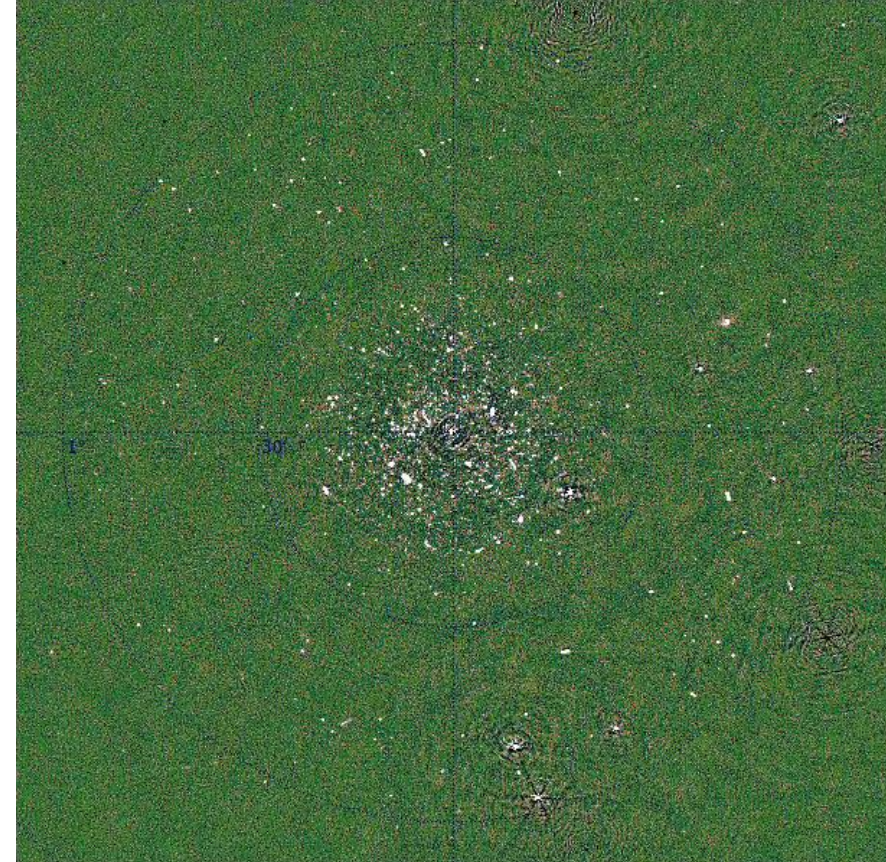
Beam-induced spectral variation
for the source represented by a dot

Incorporating primary beam in calibration

EVLA image of the field around the radio source 3C147



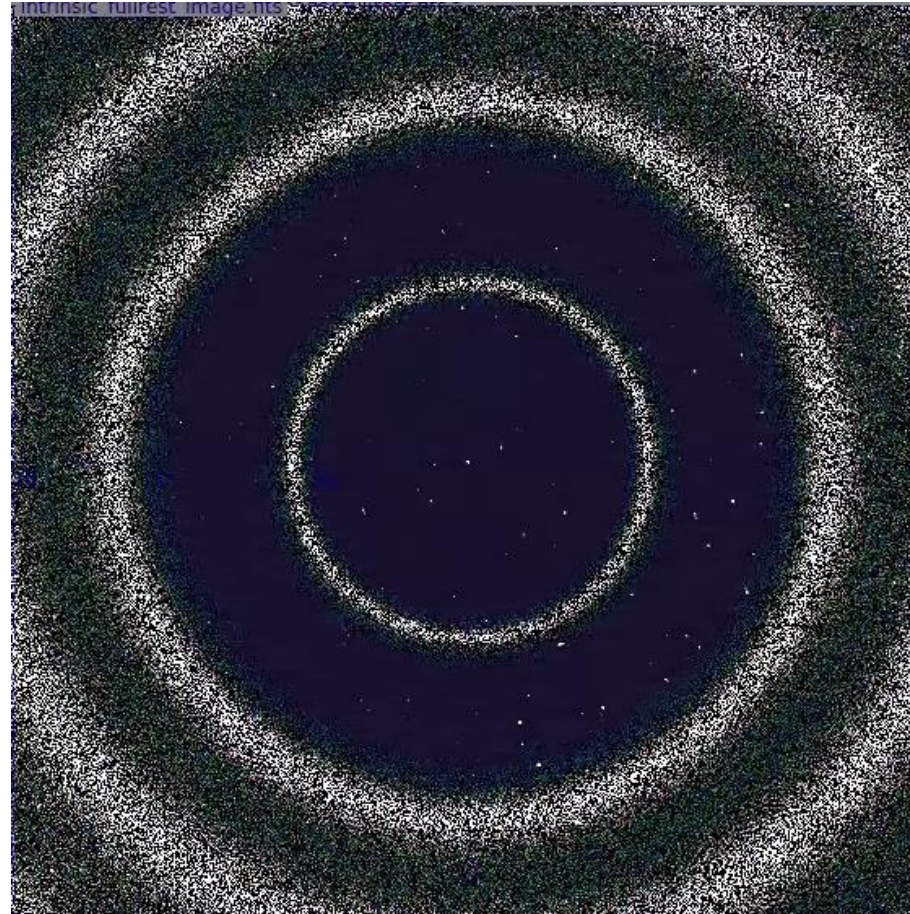
Calibration **without primary beam** included



Calibration **with primary beam** included

Effect of primary beam on noise over the field of view

EVLA image of the field around the radio source 3C147



Calibration procedure

1. Start with visibility data, \mathbf{V}_{pq} , and initial brightness model, \mathbf{B} .
2. Solve $\min_J |\mathbf{V}_{pq} - J_p \mathbf{B} J_q^H|$ for J s.
3. Calculate residual visibility data $\mathbf{V}_{pq}^{\text{residual}} = \mathbf{V}_{pq} - J_p \mathbf{B} J_q^H$.
4. Image $\mathbf{V}_{pq}^{\text{residual}}$ to create a residual image, I .
5. Perform a source-finding procedure to find sources in the residual image, and add these to the initial model \mathbf{B} to form a new, updated model \mathbf{B}^{new} .
6. Set $\mathbf{B} = \mathbf{B}^{\text{new}}$, and repeat steps 2-5 until the residual image I is noise-like.

Differential gains

- Differential gain solutions encompass the **unknown and unmodeled direction-dependent effects** in the signal path.
- The Jones matrix in the direction of source s is then given by:

$$J^{(s)} = G E \Delta E^{(s)}$$

Final Jones matrix
in the direction
of source s

Direction-independent
effects

Known and modeled
direction-dependent
effects

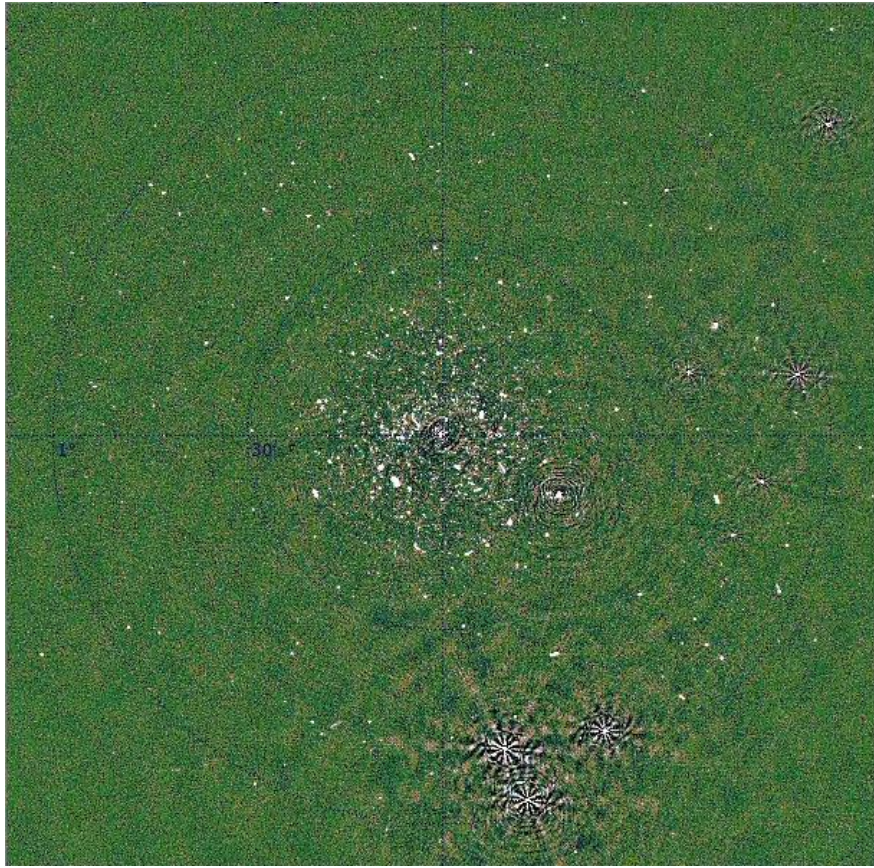
Differential gain:
Unknown/unmodeled
direction-dependent
effects in the direction
of source s

Differential gains

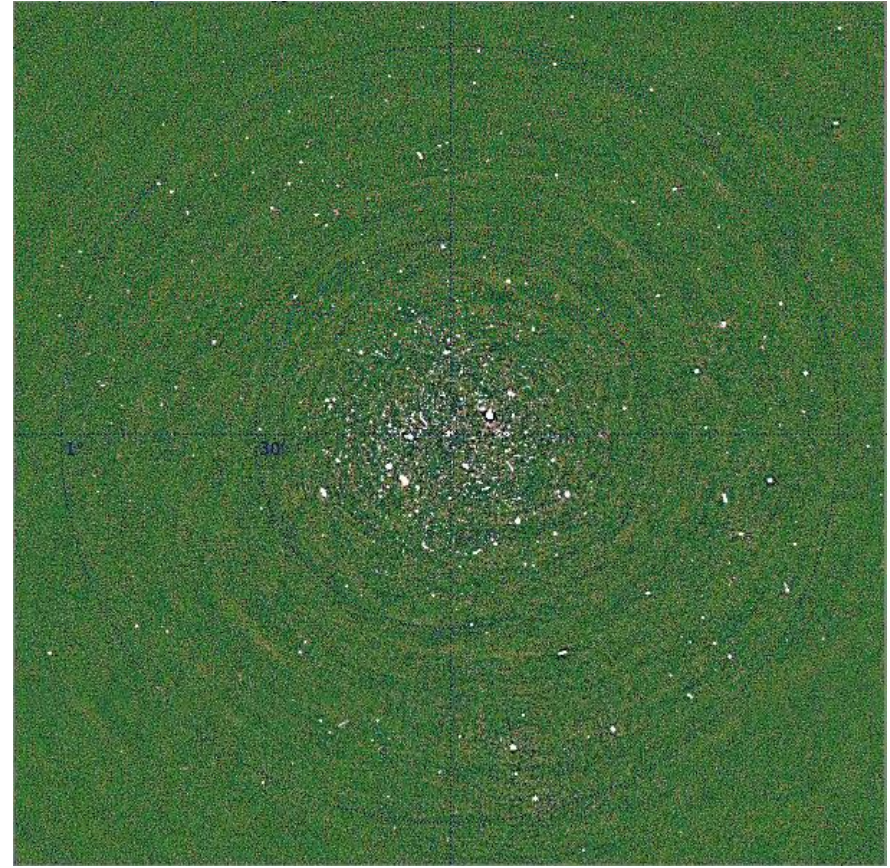
- **Differential gain solutions** are computed (in the direction of a few bright sources) and applied after regular calibration in order to correct for leftover, uncalibrated effects.

Incorporating differential gains in calibration

(Without primary beam incorporated in calibration)



Without differential gain solutions



With differential gain solutions applied

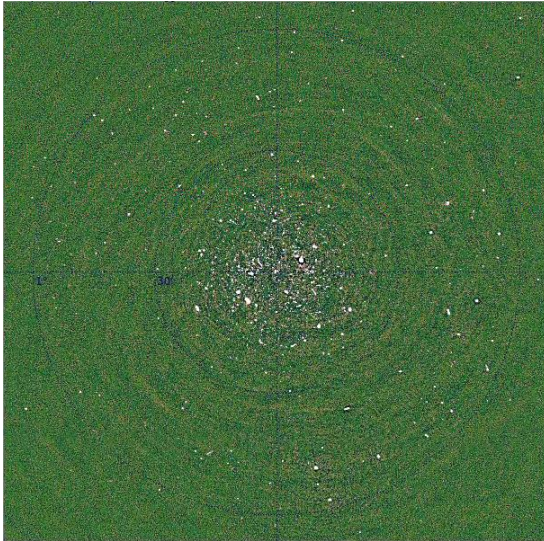
Differential gains

- As more corrupting effects are modeled and accounted for, the calibration becomes more comprehensive, and differential gain solutions approach unity.

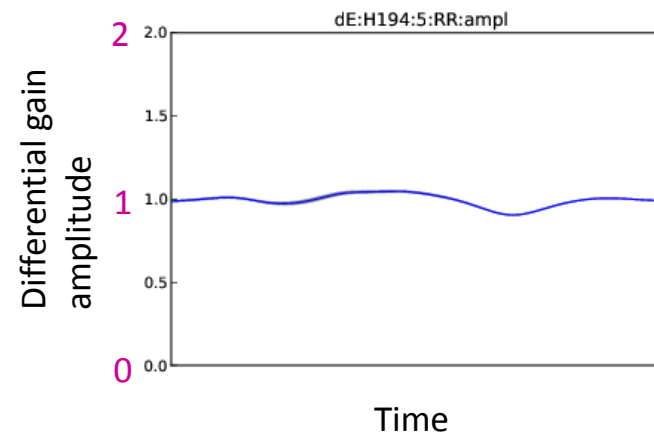
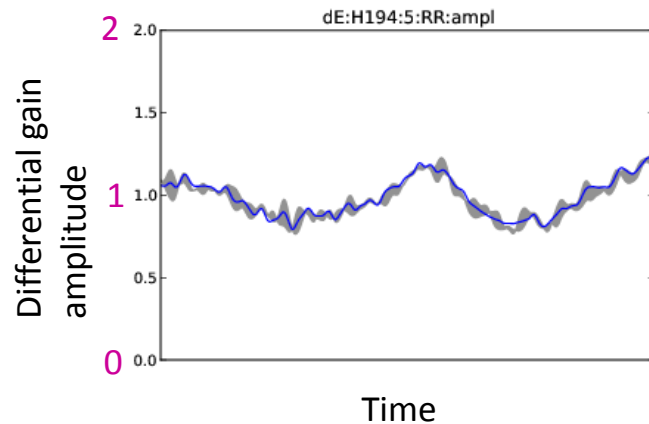
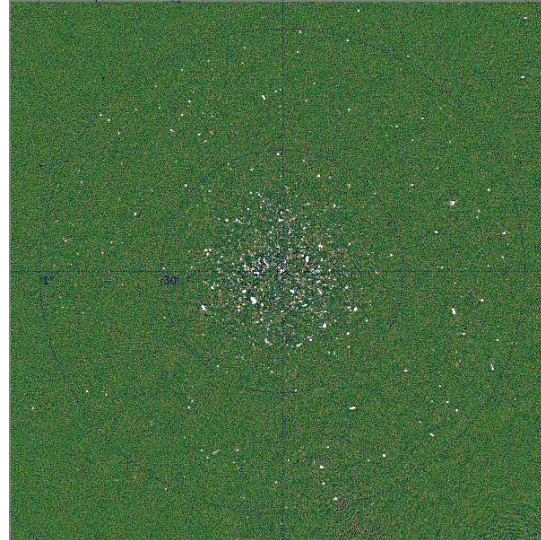
Incorporating primary beam in calibration

Differential gain plots

Without primary beam



With primary beam



- Flattened differential gain curves, ~ 1 over the whole range
- Residual variation due to remaining uncorrected direction-dependent effects (like antenna pointing errors)

References

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- *14th Synthesis Imaging Workshop [lecture slides](#)* (2014), National Radio Astronomy Observatory, Socorro, New Mexico, USA
- Oleg Smirnov's [RIME lecture](#) from *3GC3 Workshop and Interferometry School* (2013), Port Alfred, South Africa

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