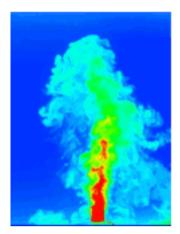
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Detrended structure-function in fully developed turbulence

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Detrended structure-function in fully developed turbulence

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The classical structure-function (SF) method in fully developed turbulence or for scaling processes in general is influenced by large-scale energetic structures, known as the infrared effect. Therefore, the extracted scaling exponents $\zeta(n)$ might be biased due to this effect. In this paper, a detrended structure-function (DSF) method is proposed to extract scaling exponents by constraining the influence of large-scale structures. This is accomplished by removing a first-order polynomial fitting within a window size ℓ before calculating the velocity increment. By doing so, the scales larger than ℓ , i.e. $r \ge \ell$, are expected to be removed or constrained. The detrending process is equivalent to a highpass filter in a physical domain. Meanwhile, the intermittency nature is retained. We first validate the DSF method by using a synthesised fractional Brownian motion for monofractal processes and a lognormal process for multifractal random walk processes. The numerical results show comparable scaling exponents $\zeta(n)$ and singularity spectra D(h)for the original SFs and DSFs. When applying the DSF to a turbulent velocity obtained from a high Reynolds number wind tunnel experiment with $Re_{\lambda} \simeq 720$, the thirdorder DSF demonstrates a clear inertial range with $\mathcal{B}_3(\ell) \simeq 4/5\epsilon \ell$ on the range 10 < $\ell/\eta < 1000$, corresponding to a wavenumber range $0.001 < k\eta < 0.1$. This inertial range is consistent with the one predicted by the Fourier power spectrum. The directly measured scaling exponents $\zeta(n)$ (respectively, singularity spectrum D(h)) agree very well with a lognormal model with an intermittent parameter $\mu = 0.33$. Due to large-scale effects, the results provided by the SFs are biased. The method proposed here is general and can be applied to different dynamics systems in which the concepts of multiscaling and multifractal are relevant.

Keywords: fully developed turbulence; intermittency; detrended structure-function

1. Introduction

Multiscale dynamics is present in many phenomena, e.g. turbulence [1], finance [2,3], geosciences [4,5], etc., to quote a few. It has been found in many multiscale dynamics systems that the self-similarity is broken, in which the concept of multiscaling or multifractal is relevant [1]. This is characterised conventionally by using the structure-functions (SFs), i.e. $S_n(\ell) = \langle \Delta u_\ell(x)^n \rangle \sim \ell^{\zeta(n)}$, in which $\Delta u_\ell(x) = u(x + \ell) - u(x)$ is an increment with separation scale ℓ . Note that for the self-similarity process, e.g. fractional Brownian motion (fBm), the measured $\zeta(n)$ is linear with n. While for the multifractal process, e.g. turbulent velocity, it is usually convex with n. Other methods are available to extract the scaling exponent. For example, wavelet-based methodologies (e.g., wavelet leaders, wavelet transform modulus maxima [5–7]), the Hilbert-based method [8,9] or the scaling analysis

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of probability density function of velocity increments [10], to name a few. Each method has its own advantages and shortcomings. For example, the classical SFs are found to mix information of the large-scale (respectively, known as the infrared effect) and small-scale (respectively, known as the ultraviolet effect) structures [9,11–14]. The corresponding scaling exponent $\zeta(n)$ is thus biased when a large energetic structure is present [9].

Previously, the influence of the large-scale structure has been considered extensively by several authors [12,13,15–18]. For example, Praskvosky et al. [15] found strong correlations between the large scales and the velocity SFs at all length scales. Sreenivasan and Stolovitzky [16] observed that the inertial range of the SFs conditioned on the large-scale velocity shows a strong dependence. Huang et al. [12] showed analytically that the influence of the large-scale structure could be as large as two decades down to the small scales. Blum et al. [13] studied experimentally the non-universal large-scale structure by considering both conditional Eulerian and Lagrangian SFs. They found that both SFs depend on the strength of large-scale structures at all scales. In their study, the large-scale structure velocity is defined as a two-point average, i.e. $\sum u_z(\ell) = [u_z(x) + u_z(x + \ell)]/2$, in which u_z is the vertical velocity in their experiment apparatus. Note that they conditioned SFs on different intensities of $\sum u_z(\ell)$. Later, Blum et al. [18] investigated systematically the large-scale structure conditioned SFs for various turbulent flows. They confirmed that in different turbulent flows the conditioned SFs depend strongly on large-scale structures at all scales.

In this paper, a detrended structure-function (DSF) method is proposed to extract scaling exponents $\zeta(n)$. This is accomplished by removing a first-order polynomial within a window size ℓ before calculating the velocity increment. This procedure is designated as detrending analysis (DA). By doing so, scales larger than ℓ , i.e. $r \geq \ell$, are expected to be removed or constrained. Hence, the DA acts as a high-pass filter in a physical domain. Meanwhile, the intermittency is still retained. A velocity increment $\Delta u_{i,\ell}(x)$ is then defined within the window size ℓ . An nth-order moment of $\Delta u_{i,\ell}(x)$ is introduced as an nth-order DSF. The DSF is first validated by using a synthesised fBm and a lognormal process with an intermittent parameter $\mu = 0.15$, respectively, for mono-fractal and multifractal processes. It is found that DSFs provide comparable scaling exponents $\zeta(n)$ and singularity spectra D(h) with the ones provided by the original SFs. When applying to a turbulent velocity with a Reynolds number $Re_{\lambda} = 720$, the third-order DSF shows a clear inertial range $10 < \ell/\eta <$ 1000, which is consistent with the one predicted by the Fourier power spectrum $E_u(k)$, e.g. $0.001 < k\eta < 0.1$. Moreover, a compensated height of the third-order DSF is $0.80 \pm$ 0.05. This value is consistent with the famous Kolmogorov four-fifth law. The directly measured scaling exponents $\zeta(n)$ (respectively, singularity spectrum D(h)) agree very well with the lognormal model with an intermittent parameter $\mu = 0.33$. Due to the large-scale effect, known as the infrared effect, the SFs are biased. Note that the scaling exponents are extracted directly without resorting to the Extended-Self-Similarity (ESS) technique. The method is general and could be applied to different types of data, in which the multiscale and multifractal concepts are relevant.

2. Detrending analysis and detrended structure-function

2.1. Detrending analysis

We start here with a scaling process u(x), which has a power-law Fourier spectrum, that is,

$$E(k) = Ck^{-\beta} \tag{1}$$

in which β is the scaling exponent of E(k). Parseval's theorem states the following relation:

$$\langle u(x)^2 \rangle_x = \int_0^{+\infty} E(k) \, \mathrm{d}k \tag{2}$$

in which $\langle \rangle$ is an ensemble average and E(k) is the Fourier power spectrum of u(x) [19]. We first divide the given u(x) into m segments with length ℓ each. A qth-order detrending of the ith segment is defined as

$$u_{i,\ell}(x) = u_i(x) - P_{i,\ell}^q(x), (i-1)\ell \le x \le i\ell,$$
 (3)

in which $P_{i,\ell}^q(x)$ is a *q*th-order polynomial fitting of $u_i(x)$. We consider below only the first-order detrending, i.e. q=1. To obtain a detrended signal, i.e. $u_\ell(x)=[u_{1,\ell}(x),u_{2,\ell}(x)\cdots u_{m,\ell}(x)]$, a linear trend is removed within a window size ℓ . Ideally, scales larger than ℓ , i.e. $r>\ell$, are removed or constrained from the original data u(x). This implies that the DA procedure is a high-pass filter in the physical domain. The kinetic energy of $u_\ell(x)$ is related directly to its Fourier power spectrum, that is,

$$\mathcal{D}_2(\ell) = \langle u_\ell(x)^2 \rangle_x = \int_0^{+\infty} E_\ell(k) \, \mathrm{d}k \simeq \int_{k_\ell}^{+\infty} E(k) \, \mathrm{d}k \tag{4}$$

in which $k_\ell = 1/\ell$ and $E_\ell(k)$ is the Fourier power spectrum of $u_\ell(x)$. This illustrates again that the DA procedure acts as a high-pass filter, in which the lower Fourier modes $k < k_\ell$ (respectively, $r > \ell$) are expected to be removed or constrained. For a scaling process, i.e. $E(k) \sim k^{-\beta}$, it leads to a power-law behaviour, that is,

$$\mathcal{D}_2(\ell) \sim k_\ell^{1-\beta} \sim \ell^{\beta-1}.\tag{5}$$

The physical meaning of $\mathcal{D}_2(\ell)$ is quite clear. It represents a cumulative energy over the Fourier wavenumber band $[k_\ell, +\infty]$ (respectively, scale range $[0, \ell]$). We emphasise here again that the DA acts as a high-pass filter in a physical domain and the intermittency nature of u(x) is still retained.

2.2. Detrended structure-function

The above-mentioned DA can remove/constrain the large-scale influence, known as the infrared effect. This could be utilised to redefine the SF to remove/constrain the large-scale structure effect as follows. After the DA procedure, the velocity increment can be defined within a window size ℓ as

$$\Delta u_{i,\ell}(x) = u_{i,\ell}(x + \ell/2) - u_{i,\ell}(x) \tag{6}$$

in which *i* represents the *i*th segment. We will show in the next subsection why we define an increment with a half width of the window size. An *n*th-order DSF is then defined as

$$\mathcal{B}_n(\ell) = \langle \Delta u_{i,\ell}(x)^n \rangle_x \,. \tag{7}$$

For a scaling process, we expect a power-law behaviour, that is,

$$\mathcal{B}_n(\ell) \sim \ell^{\zeta(n)}$$
 (8)

in which the scaling exponent $\zeta(n)$ is comparable with the one provided by the original SFs. To access negative orders of n (respectively, the right part of the singularity spectrum D(h), see definition below), the DSFs can be redefined as

$$\mathcal{B}_n(\ell) = \langle X_\ell(i)^n \rangle \tag{9}$$

in which $X_{\ell}(i) = \langle |\Delta u_{i, \ell}(x)| \rangle_{(i-1)\ell \leq x \leq i\ell}$ is a local average for the *i*th segment. A power-law behaviour is expected, i.e. $\mathcal{B}_n(\ell) \sim \ell^{\zeta(n)}$. It is found experimentally that when q > 0, Equations (7) and (9) provide the same scaling exponents $\zeta(n)$. In the following, we do not discriminate these two definitions for DSFs.

2.3. An interpretation in the time-wavenumber analysis frame

To understand better the filter property of the detrending procedure and DSFs, we introduce here a weight function $W(\ell, k)$, that is,

$$M_2(\ell) = \int_0^{+\infty} \mathcal{W}(\ell, k) E(k) \, \mathrm{d}k \tag{10}$$

in which E(k) is the Fourier power spectrum of u(x) and $M_2(\ell)$ is a second-order moment, which could be one of $\mathcal{D}_2(\ell)$ or $\mathcal{B}_2(\ell)$ or $\mathcal{S}_2(\ell)$. The weight function $\mathcal{W}(\ell,k)$ characterises the contribution of the Fourier component to the corresponding second-order moment. Note that an integral constant is neglected in Equation (10). For the second-order SFs, one has the following weight function [1,12]:

$$W_{SF}(\ell, k) = 1 - \cos(2\pi k\ell). \tag{11}$$

For a scaling process, one usually has a fast-decaying Fourier spectrum, i.e. $E(k) \sim k^{-\beta}$ with $\beta > 0$. Hence, the contribution from the small-scale part (respectively high-wavenumber Fourier mode) is decreasing. The SFs might be more influenced by the large-scale part for large values of β [12,14,20]. For the detrended data, the corresponding weight function should ideally be as follows:

$$\mathcal{W}_{\mathrm{DA}}(\ell, k) = \begin{cases} 0, & \text{when } k \le k_{\ell} \\ 1, & \text{when } k > k_{\ell} \end{cases}$$
 (12)

The DSFs (the combination of the DA and SF) have a weight function

$$W_{DSF}(\ell, k) = \begin{cases} 0, & \text{when } k \le k_{\ell} \\ 1 - \cos(\pi k \ell), & \text{when } k > k_{\ell} \end{cases}$$
 (13)

Comparing with the original SFs, the DSFs defined here can remove/constrain the large-scale effect. Figure 1 shows the corresponding $W(\ell, k)$ for the SF, DA and DSF, respectively. The detrended scale ℓ is illustrated by a vertical line, i.e. $k_{\ell} = 1/\ell$. We note here that with the definition of Equation (6), $\mathcal{B}_2(\ell)$ provides a better compatible interpretation with the

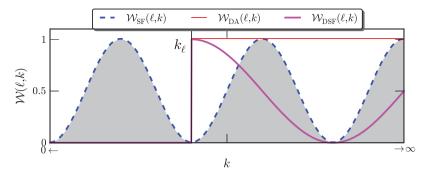


Figure 1. (Colour online) An illustration of the weight function $W(\ell, k)$ for different methods: SF $W_{SF} = 1 - \cos(2\pi k\ell)$ (dashed line), first-order DA W_{DA} (thin solid line) and DSF W_{DSF} (thick solid line). The detrended scale ℓ is demonstrated by a vertical solid line with $k_{\ell} = 1/\ell$. Ideally, scales larger than ℓ , i.e. $r < \ell$ (respectively, $k < k_{\ell}$), are expected to be removed after the detrending process.

Fourier power spectrum E(k), since we have $\mathcal{W}_{DSF}(\ell, k_{\ell}) = 1$. This is the main reason why we define the velocity increment with the half size of the window width ℓ .

We provide some comments on Equation (10). The above argument is exactly valid for linear and stationary processes. In reality, the data are always non-linear and non-stationary for some reasons (for more details, see Ref. [21]). Therefore, Equation (10) holds approximately for real data. Another comment has to be emphasised here for the detrending procedure. Several approaches might be applied to remove the trend [22,23]. However, the trend might be linear or non-linear. Therefore, different detrending approaches might provide different performances. In the present study, we only consider the first-order polynomial detrending procedure, which is efficient for many types of data.

3. Numerical validation

3.1. Fractional Brownian motion

We first consider here the fBm as a typical mono-scaling process. Fractional Brownian motion is a Gaussian self-similar process with a normal distribution increment, which is characterised by H, Hurst number, 0 < H < 1 [24–27]. A Wood–Chan algorithm is used to synthesise the fBm with a Hurst number H = 1/3. We perform 100 realisations with a data length 10^5 points each. A power-law behaviour is observed on a large range of scales for $-4 \le n \le 4$. The corresponding singularity spectrum is

$$h = \zeta'(n), \ D(h) = \min_{n} \{ hn - \zeta(n) + 1 \}.$$
 (14)

Ideally, one should have a single point of singularity spectrum with h=1/3 and D(1/3)=1. However, in practice, the measured singularity spectrum D(h) is always lying in a narrow band. Figure 2 shows the measured singularity spectrum D(h) for SFs (\square) and DSFs (\circ) for $-4 \le n \le 4$, in which the inset shows the singularity spectra D(h) estimated on the range $0 \le n \le 4$. Visually, both estimators provide the same D(h) and the same statistical error, which is defined as the standard deviation from different realisations.

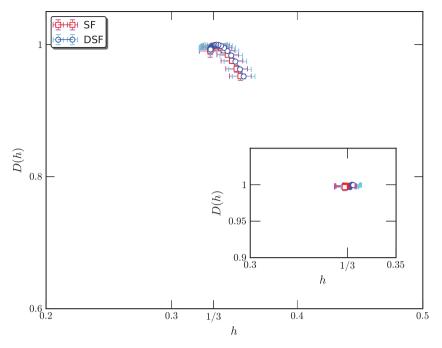


Figure 2. (Colour online) Measured singularity spectrum D(h) for an fBm with H = 1/3 on the range $-4 \le n \le 4$. The inset shows the singularity spectrum D(n) on the range $0 \le n \le 4$. The error bar is the standard deviation estimated from 100 realisations. Ideally, one should have h = 1/3 and D(1/3) = 1. Both methods provide the same h and D(h) and statistical error.

3.2. Multifractal random walk with lognormal statistics

We now consider a multifractal random walk with lognormal statistics [28–30]. A multiplicative discrete cascade process with lognormal statistics is performed to simulate a multifractal measure $\epsilon(x)$. The larger scale corresponds to a unique cell of size $L = \ell_0 \lambda_1^N$, where ℓ_0 is the largest scale considered and $\lambda_1 > 1$ is a dimensional scale ratio. In practice, for a discrete model, this ratio is often taken as $\lambda_1 = 2$ [9,30]. The next scale involved corresponds to λ_1 cells, each of size $L/\lambda_1 = \ell_0 \lambda_1^{N-1}$. This is iterated and at step p ($1 \le p \le N$) λ_1^p cells are retrieved. Finally, at each point the multifractal measure $\epsilon(x)$ is the product of n cascade random variables, that is,

$$\epsilon(x) = \prod_{m=1}^{N} W_m(x),\tag{15}$$

where $W_m(x)$ is the random variable corresponding to position x and level m in the cascade [30]. Following the multifractal random walk idea [28,29], a non-stationary multifractal time series can be synthesised as

$$u(x) = \int_0^x \epsilon(x')^{1/2} dB(x'), \tag{16}$$

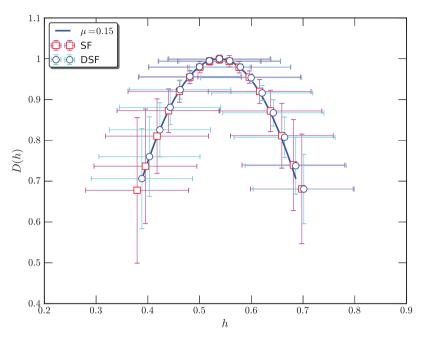


Figure 3. (Colour online) Measured singularity spectrum D(h) for the lognormal process with an intermittent parameter $\mu = 0.15$. The errorbar is the standard deviation from the 100 realisations. The theoretical singularity curve is illustrated by a solid line. Both estimators provide the same singularity spectra D(h) and statistical error.

where B(x) is Brownian motion. Taking a lognormal statistic for ϵ , the scaling exponent $\zeta(n)$ for the SFs, i.e. $\langle \Delta u_{\ell}(x)^n \rangle \sim \ell^{\zeta(n)}$, is written as

$$\zeta(n) = \frac{n}{2} - \frac{\mu}{2} \left(\frac{n^2}{4} - \frac{n}{2} \right),\tag{17}$$

where μ is the intermittency parameter ($0 \le \mu \le 1$) characterising the lognormal multifractal cascade [30].

Synthetic multifractal time series are generated following Equation (16). An intermittent parameter $\mu=0.15$ is chosen for m=17 levels each, corresponding to a data length of 131, 072 points each. A total of 100 realisations are performed. The statistical error is then measured as the standard deviation from these realisations. Figure 3 shows the corresponding measured singularity spectra D(h), in which the theoretical value is illustrated by a solid line. Graphically, the theoretical singularity spectra D(h) are recovered by both estimators. The statistical errors are again found to be the same for both estimators.

We would like to provide some comments on the performance of these two estimators. For the synthesised processes, they have the same performance since there is no intrinsic structure in these synthesised data. However, for the real data, as mentioned above, they possess non-stationary and non-linear structures [21]. Therefore, as shown below, they might have different performances.

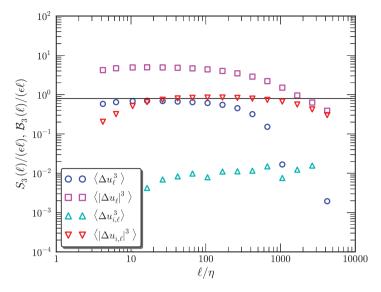


Figure 4. (Colour online) Measured compensated third-order moments $S_3(\ell)/(\epsilon\ell)$ and $\mathcal{B}_3(\ell)/(\epsilon\ell)$ from experimental homogeneous and nearly isotropic turbulent flows. They are, respectively, third-order SFs with (\square) and without (\circ) absolute value, and third-order DSFs with (∇) and without (\triangle) absolute value. The horizontal solid line indicates Kolmogorov's four-fifth law. An observed plateau for $\mathcal{B}_3(\ell)/\epsilon\ell$ indicates an inertial range on the range $10 < \ell/\eta < 1000$, corresponding to a wavenumber range $0.001 < k\eta < 0.1$. Roughly speaking, a plateau for $S_3(\ell)/\epsilon\ell$ indicates an inertial range on the range $10 < \ell/\eta < 100$. The heights of the inertial range are, respectively, 0.67 ± 0.02 (\circ), 4.84 ± 0.14 (\square), 0.0098 ± 0.0024 (\triangle) and 0.80 ± 0.05 (∇), in which the statistical error is the standard deviation obtained from the inertial range. Note that the inertial range is $10 < \ell/\eta < 100$ for the SFs and $10 < \ell/\eta < 1000$ for the DSFs. The corresponding scaling exponents ζ (3) are 0.95 ± 0.02 , 0.84 ± 0.03 , 1.15 ± 0.07 and 0.99 ± 0.03 , respectively. The statistical error is the 95% fitting confidence on the inertial range.

4. Application to turbulent velocity

We consider here a velocity database obtained from a high Reynolds number wind tunnel experiment in the Johns Hopkins University with Reynolds number $Re_{\lambda}=720$. A probe array with four X-type hot wire anemometry is used to record the velocity with a sampling wavenumber of 40 kHz in the streamwise direction x/M=20, in which M is the size of the active grid. These probes are placed in the middle height and along the centre line of the wind tunnel to record the turbulent velocity simultaneously for a duration of 30 seconds. The measurement is then repeated for 30 times. Finally, we have $30 \times 4 \times 30 \times (4 \times 10^4)$ data points (number of measurements \times number of probes \times duration time \times sampling wavenumber). Therefore, there are 120 realisations (number of measurements \times number of probes). The Fourier power spectrum $E_u(k)$ of the longitudinal velocity reveals a nearly two decades inertial range on the wavenumber range $0.001 < k\eta < 0.1$ with a scaling exponent $\beta \simeq 1.65 \pm 0.02$ (see Ref. [31]). This corresponds to time scales $10 < \ell/\eta < 1000$. Here η is the Kolmogorov scale. Note that we convert our results into spatial space by applying Taylor's frozen hypothesis [1]. More details about this database can be found in Ref. [31].

To determine the inertial range in real space, we plot the measured compensated thirdorder moments in Figure 4 for the SFs $(S_3(\ell)/(\epsilon \ell))$ with (\Box) and without (\circ) absolute values) and DSFs $(\mathcal{B}_3(\ell)/(\epsilon \ell))$ with (∇) and without (\triangle) absolute values), respectively. A horizontal solid line indicates Kolmogorov's four-fifth law. A plateau is observed for $\mathcal{B}_3(\ell)/(\epsilon \ell)$ on

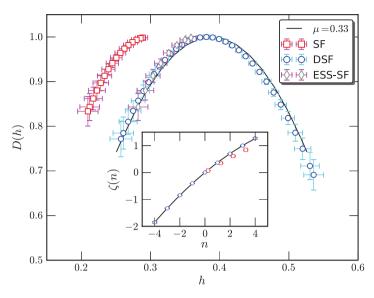


Figure 5. (Colour online) Measured singularity spectrum D(h). The error bar is the standard deviation from 120 realisations. The inset shows the corresponding scaling exponents $\zeta(n)$. For comparison, the lognormal model with an intermittent parameter $\mu = 0.33$ is illustrated by a solid line.

the range $10 < \ell/\eta < 1000$, which agrees very well with the inertial range predicted by $E_{\nu}(k)$, i.e. on the range $0.001 < k\eta < 0.1$. The corresponding height and scaling exponent are 0.80 ± 0.05 with absolute value (0.0098 ± 0.0024 without absolute value) and $\zeta(3) =$ 0.99 ± 0.03 ($\zeta(3) = 1.15 \pm 0.07$), respectively. The statistical error is the standard deviation obtained on the range $10 < \ell/\eta < 1000$. Note that Kolmogorov's four-fifth law indicates a linear relation $\langle \Delta u_{\ell}^{3} \rangle = -4/5\epsilon \ell$. It is interesting to note that, despite the sign, we have $\langle |\Delta u_{i,\ell}|^3 \rangle = 4/5\epsilon \ell$ on nearly two-decade scales. For comparison, the third-order SFs are also shown. Roughly speaking, a plateau is observed on the range $10 < \ell/\eta < 100$. This inertial range is shorter than the one predicted by the Fourier analysis or DSFs, which is now understood as the large-scale influence. The corresponding height and scaling exponent are 0.67 ± 0.02 without absolute value (4.84 \pm 0.14 with absolute value) and 0.95 \pm 0.02 (0.84 ± 0.03) , respectively. Therefore, the DSFs provide a better indicator of the inertial range since they remove/constrain the large-scale influence. We therefore estimate the scaling exponents for $\mathcal{B}_n(\ell)$ on the range $10 < \ell/\eta < 1000$ for $-4 \le n \le 4$ directly without resorting to the ESS technique [32,33]. For the SFs, we calculate the scaling exponents $\zeta(n)$ on the range $10 < \ell/\eta < 100$ for $0 \le n \le 4$ directly.

Figure 5 shows the measured singularity spectra D(h) for $-4 \le n \le 4$, in which the error bar is a standard deviation from 120 realisations. The inset shows the corresponding scaling exponents $\zeta(n)$. For comparison, the lognormal model $\zeta(n) = n/3 - \mu/18(n^2 - 3n)$ with an intermittent parameter $\mu = 0.33$ is shown as a solid line. Visually, the DSFs curve fully recovers the lognormal curve not only on the left part $(n \ge 0)$, but also on the right part $(n \le 0)$. Due to the large-scale contamination, the SFs underestimate the scaling exponents $\zeta(n)$ when $n \ge 0$ [11,12]. This leads to an overestimation of the left part of the singularity spectrum D(h) (see \square in Figure 5). However, if one resorts to the ESS algorithm when measuring the SF scaling exponent $\zeta(n)$, the corresponding singularity spectrum D(h) is then horizontally shifted to the theoretical curve. This has been interpreted as that the ESS technique suppresses the finite Reynolds number effect. We show here that

if one removes/constrains the effect of large-scale motions, one can retrieve the scaling exponent $\zeta(n)$ (singularity spectrum D(h)) without resorting to the ESS technique, or, in other words, the finite Reynolds number effect manifests at large-scale motions, which is usually anisotropic too.

5. Conclusion

In this paper, we introduce a DSF analysis to remove/constrain the influence of large-scale motions, known as the infrared effect. In the first step of our proposal, the first-order polynomial trend is removed within a window size ℓ . By doing so, the scales larger than ℓ , i.e. $r \ge \ell$, are expected to be removed/constrained. In the second step, a velocity increment is defined with a half of the window size. The DSF proposal is validated by the synthesised fBm for the mono-fractal process and a lognormal random walk for the multifractal process. The numerical test shows that both SF and DSF estimators provide a comparable performance for synthesised processes without intrinsic structures.

When applying to the turbulent velocity obtained from a high Reynolds number wind tunnel experiment, the third-order DSFs show a clearly inertial range on $10 < \ell/\eta < 1000$ with a linear relation $\mathcal{B}_3(\ell) \simeq 4/5\epsilon \ell$. The inertial range provided by DSFs is consistent with the one predicted by the Fourier power spectrum. Note that, despite the sign, Kolmogorov's four-fifth law is retrieved for the third-order DSFs. The corresponding third-order SFs are biased by the large-scale structures, known as the infrared effect. It shows a shorter inertial range and underestimates the third-order scaling exponent $\zeta(3)$. The scaling exponents $\zeta(n)$ are then estimated directly without resorting to the ESS technique. The corresponding singularity spectrum D(h) provided by the DSFs fully recovers the lognormal model with an intermittent parameter $\mu = 0.33$ on the range $-4 \le n \le 4$. However, the classical SFs overestimate the left-part singularity spectrum D(h) (underestimate the corresponding scaling exponents $\zeta(n)$ on the range 0 < n < 4. This has been interpreted as the finite Reynolds number effect and can be corrected by using the ESS technique. Here, to the best of our knowledge, we show for the first time that if one removes/constrains the influence of the large-scale structures, one can recover the lognormal model without resorting to the ESS technique.

The method we proposed here is general and applicable to other complex dynamical systems as well, in which the multiscale statistics are relevant. It can also be applied systematically to more turbulent velocity databases with different Reynolds numbers to see whether the finite Reynolds number effect manifests on large-scale motions as well as we show for high Reynolds number turbulent flows.

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