

Supplement: Derivations, Junctions, and Term Map

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A. Action and Junctions

Start from a bulk action $S_5 = (1/2 \kappa_5^2) \int d^5x \sqrt{-g_5} (R_5 - 2 \Lambda_5) + \int_{\text{brane}} d^4x \sqrt{-g} (-\lambda + L_{\text{matter}})$. Apply Israel junction conditions across the brane: $[K_{\mu\nu}] = -\kappa_5^2 (T_{\mu\nu} - (1/3) g_{\mu\nu} (T - \lambda))$.

B. Quadratic Term $\Pi_{\mu\nu}$

$\Pi_{\mu\nu} = -(1/4) T_{\mu\alpha} T^{\alpha}_{\nu} + (1/12) T T_{\mu\nu} + (1/8) g_{\mu\nu} T_{\alpha\beta} T^{\alpha\beta} - (1/24) g_{\mu\nu} T^2$. This generates the high-energy ρ^2 correction when reduced to FRW.

C. Projected Weyl $E_{\mu\nu}$

$E_{\mu\nu} = C_{ABCD} n^A n^C g_{\mu}^B g_{\nu}^D$, with $E^{\mu}_{\mu} = 0$. For FRW, this yields an effective radiation term $\rho_{\text{dr}} = C/a^4$ ("dark radiation").

D. Friedmann Reduction

For a flat FRW brane with perfect fluid $p = w \rho$, the modified Friedmann equation becomes $H^2 = (8\pi G/3) \rho (1 + \rho/(2\lambda)) + \Lambda_4/3 + C/a^4$. Early-time solution in radiation era: $a(t) \sim t^{1/4}$.

E. Observable Map

Break frequency $f_{\text{br}}(\lambda)$ scales with the crossover energy density $\sim \lambda$; use $f_{\text{br}} \sim k_0 * \lambda^{1/4}$. Relate C/a^4 to ΔN_{eff} via $\rho_{\text{dr}}/\rho_{\text{gamma}} = (7/8) (4/11)^{4/3} * \Delta N_{\text{eff}}$. Provide bounds priors from BBN/CMB.

F. Symbols and Units

λ : brane tension (energy density). C : dark-radiation constant. M_5 : 5D Planck mass. k : warp parameter. H : Hubble parameter. a : scale factor. ρ : energy density. All $c = 1$ conventions unless otherwise stated.