# Einstein Consistency — PPN & Binary Pulsar (ASCII-safe)

Low-energy GR limit and classical tests:

• For rho << lambda, the rho^2/(2lambda) term is negligible; dark-radiation C/a^4 redshifts away; Lambda4 small.

• PPN parameters reduce to GR values (gamma≈beta≈1) up to corrections O(rho/lambda, |C|/a^4).

• Binary pulsars: effective 4D dynamics match GR within timing bounds when rho/lambda << 1.

• Solar-System: Shapiro delay and perihelion precession consistent within existing constraints for same limit.

Conclusion: the model preserves Einstein-gravity phenomenology at late times.