

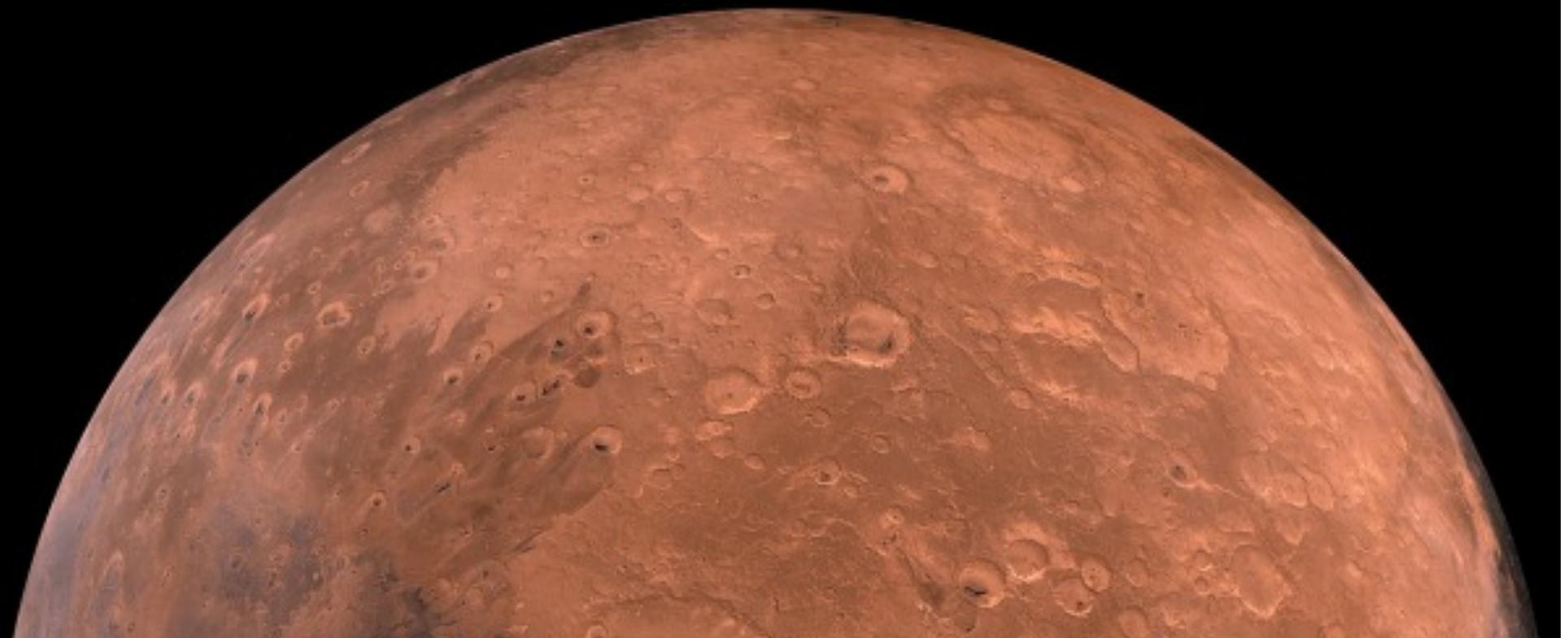
Investigation of Martian Dust Lifting

Laura Kulowski

Department of Mathematics,
Brown University, Providence, Rhode
Island

Helen Wang

Atomic and Molecular Physics
Division, Harvard-Smithsonian
Center for Astrophysics,
Cambridge, Massachusetts



Why Study Martian Dust Lifting?

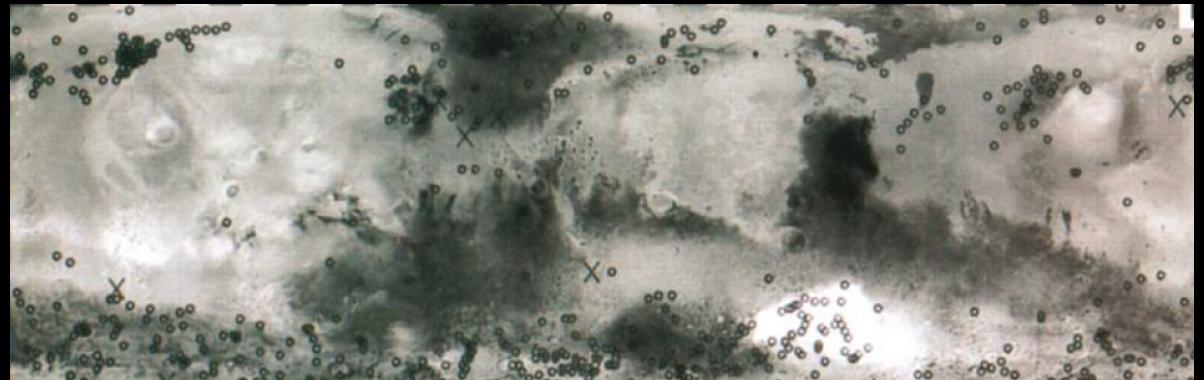
- First analysis of this kind

- Ground our understanding in actual events

- Temporal and spatial distribution can be used to put constraints on the Martian dust cycle

- Frequency of lifting helps us understand surface-atmosphere interaction

- Improve General Circulation Models (GCM)



Spatial distribution of dust storms between March of 1999 and December of 1999 [Cantor et al., 2001]

Data Collection



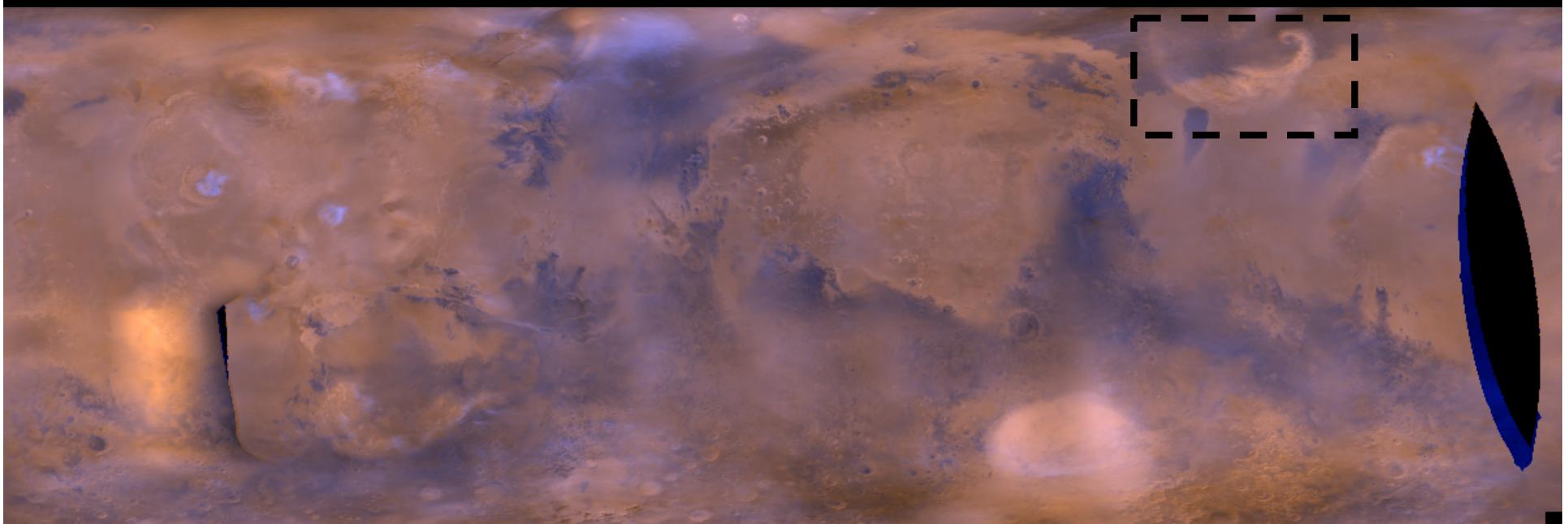
(a)



(b)

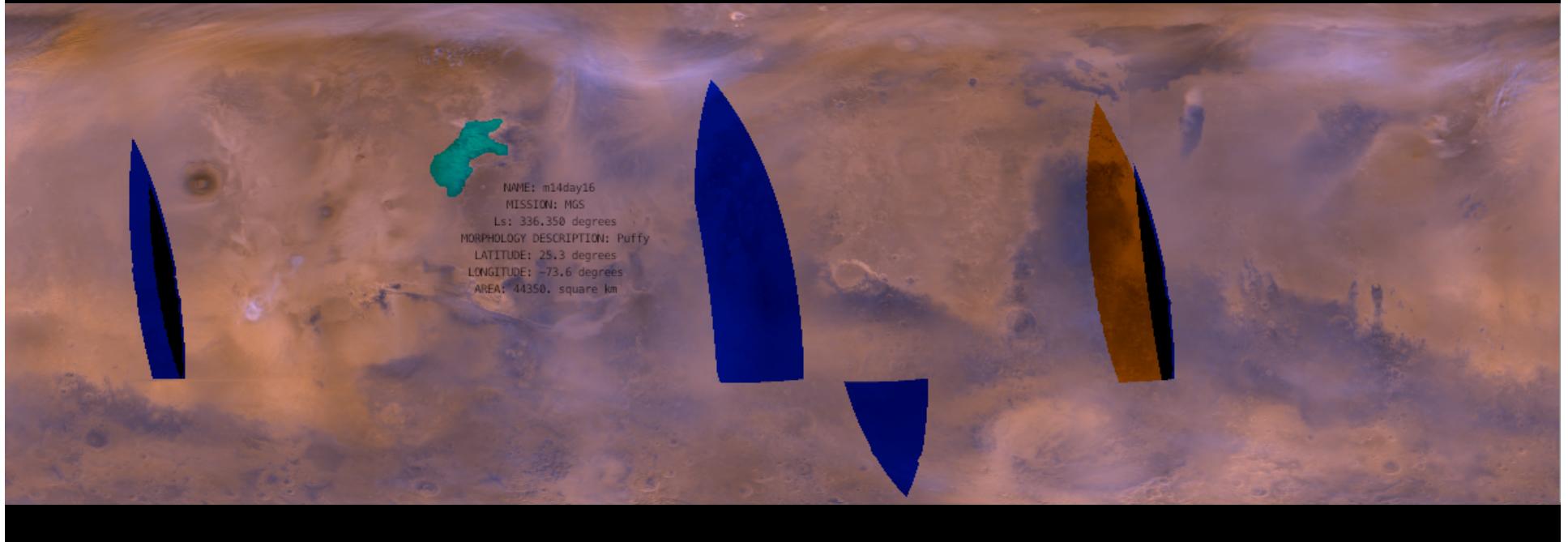
Sample MOC global map swaths: (a) raw red (left) and blue (right) filter images and (b) processed red and blue filter images. The blue image is flipped with respect to the red image [Wang and Ingersoll, 2002].

Mars Daily Global Map (MDGM)

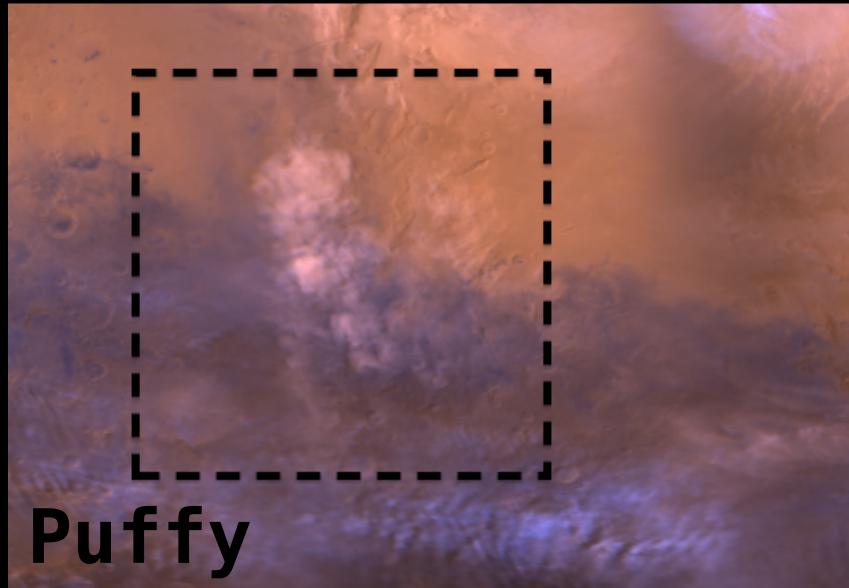


- One map for each day
- Each map is generated from 13 pairs of red and blue MOC wide angle images
- Restricted to 2 PM local time
- Occasional missing data (black and blue sections)
- Dust storms appear orange, condensate clouds appear blue in MDGM

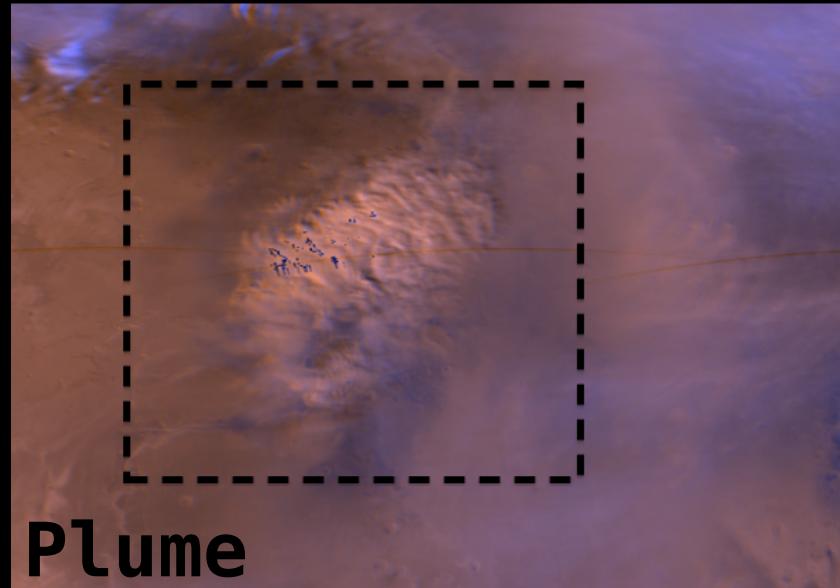
Identifying Dust Lifting



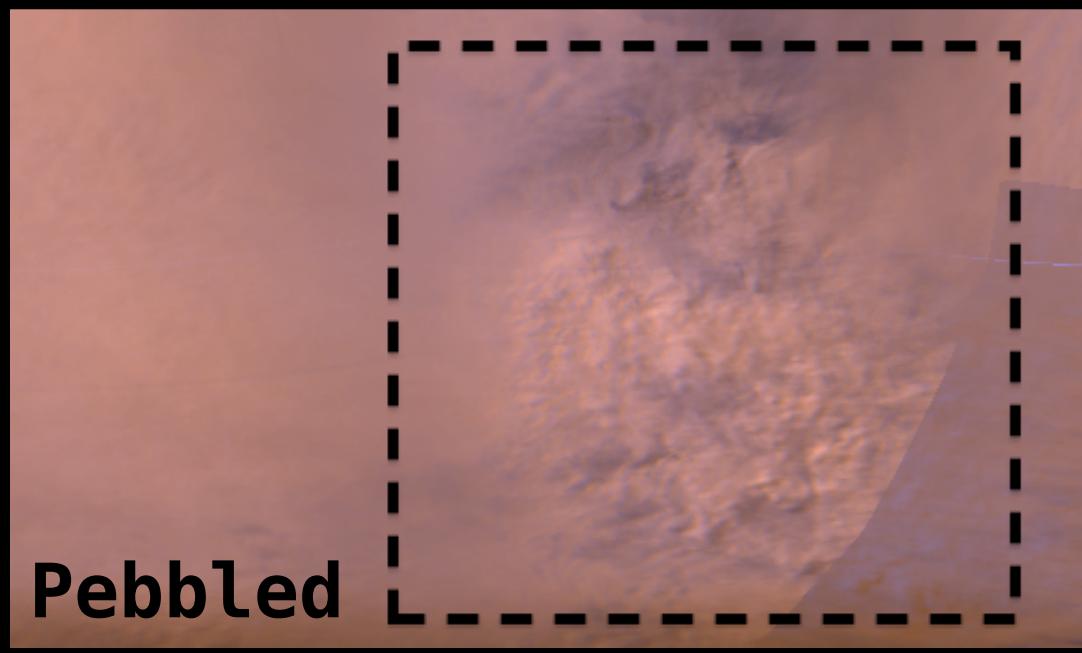
Morphological Classifications



Puffy



Plume



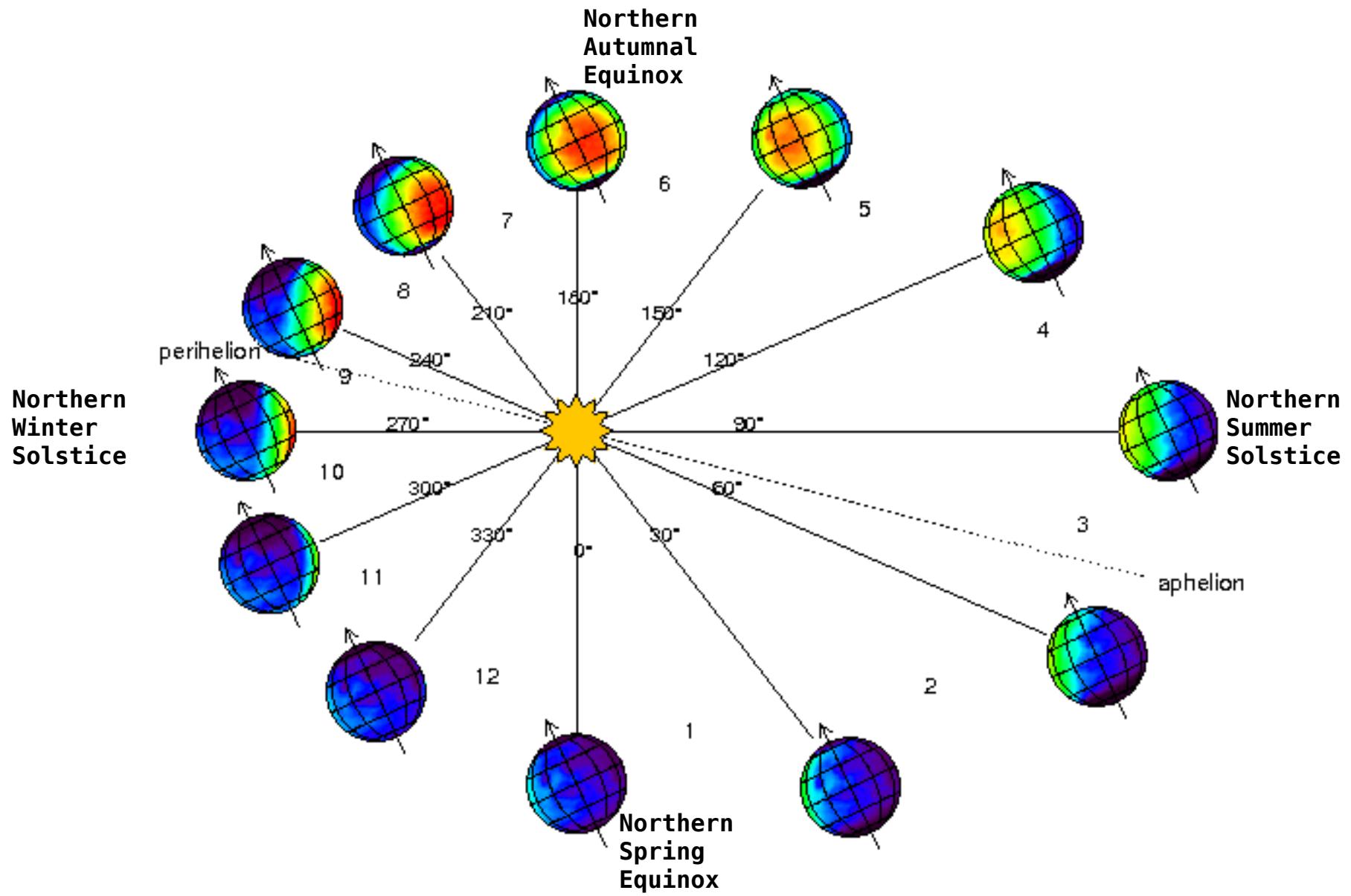
Pebbled

Analysis

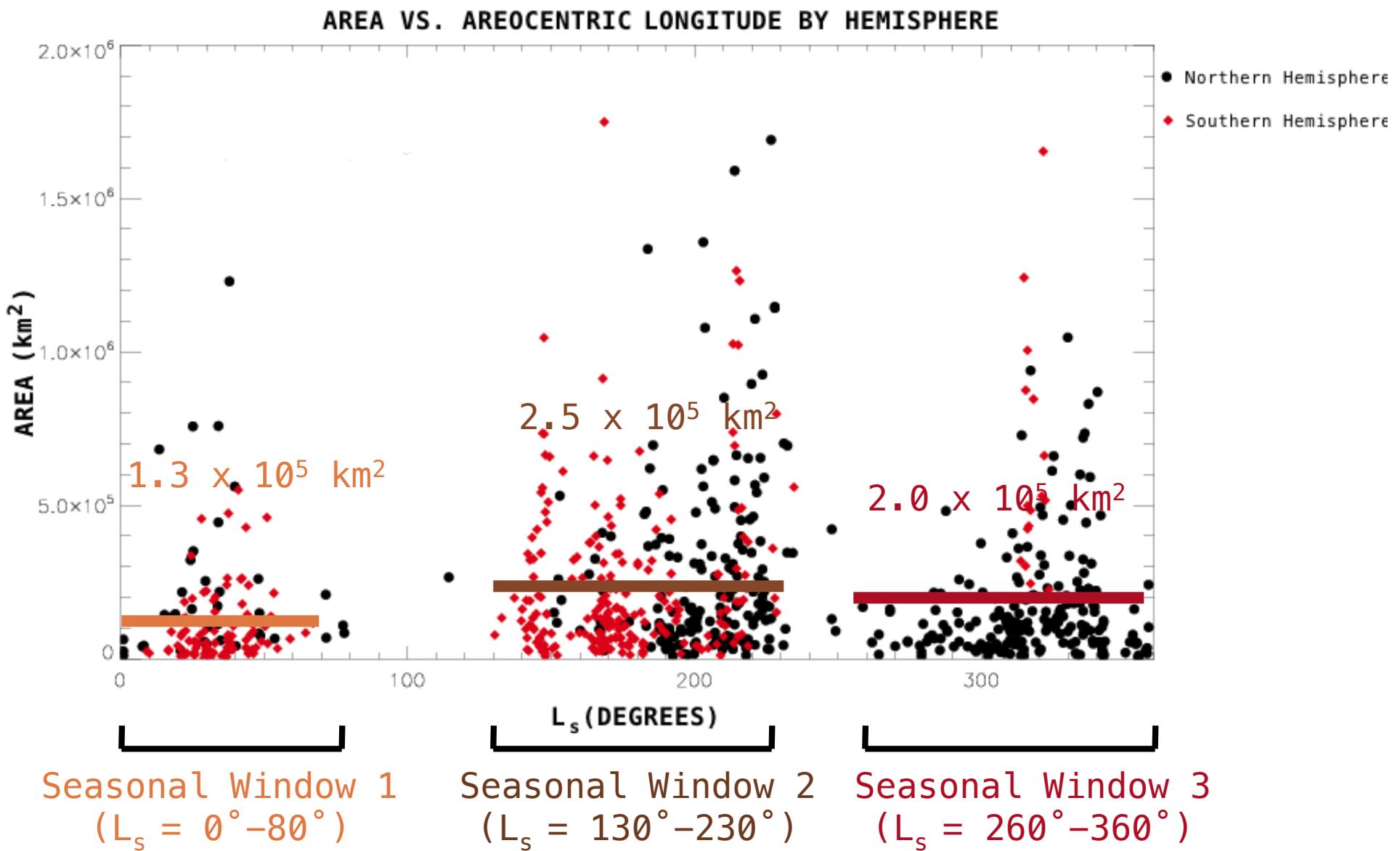
- I. Seasonal & Size Distributions
- II. Frequency Maps
- III. Morphological Properties

I. Seasonal and Size Distribution

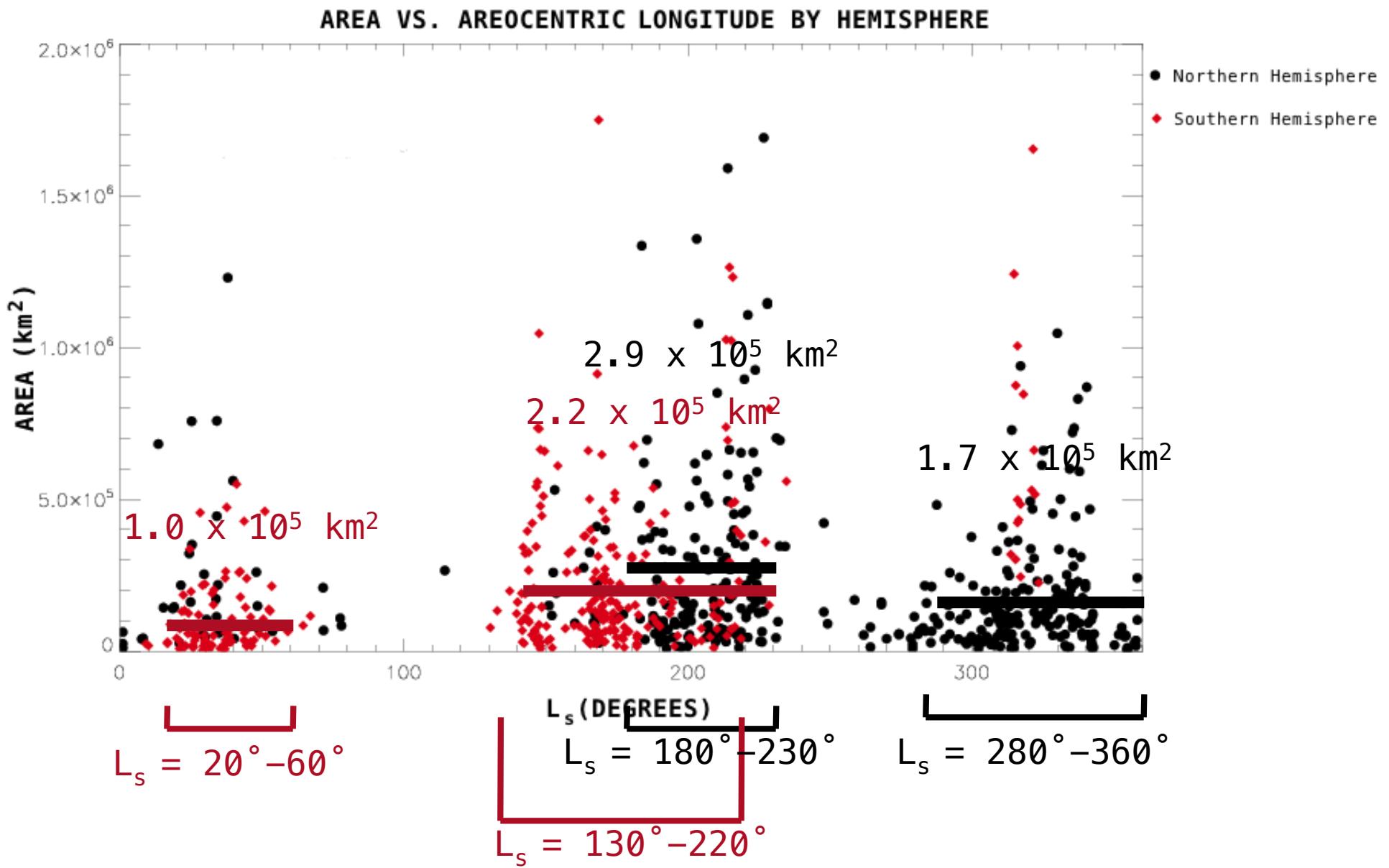
Areocentric Longitude (L_s)



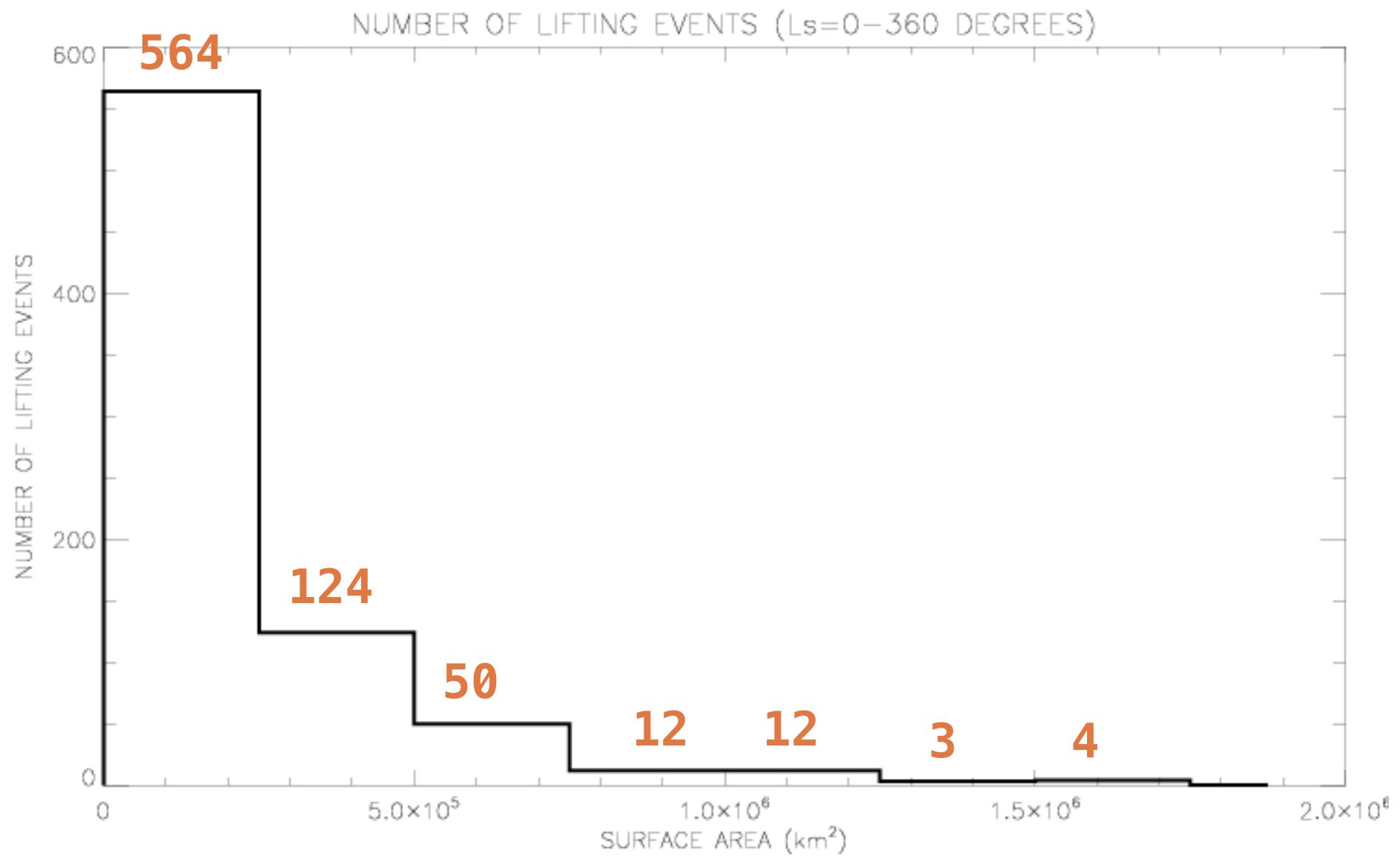
Size Distribution



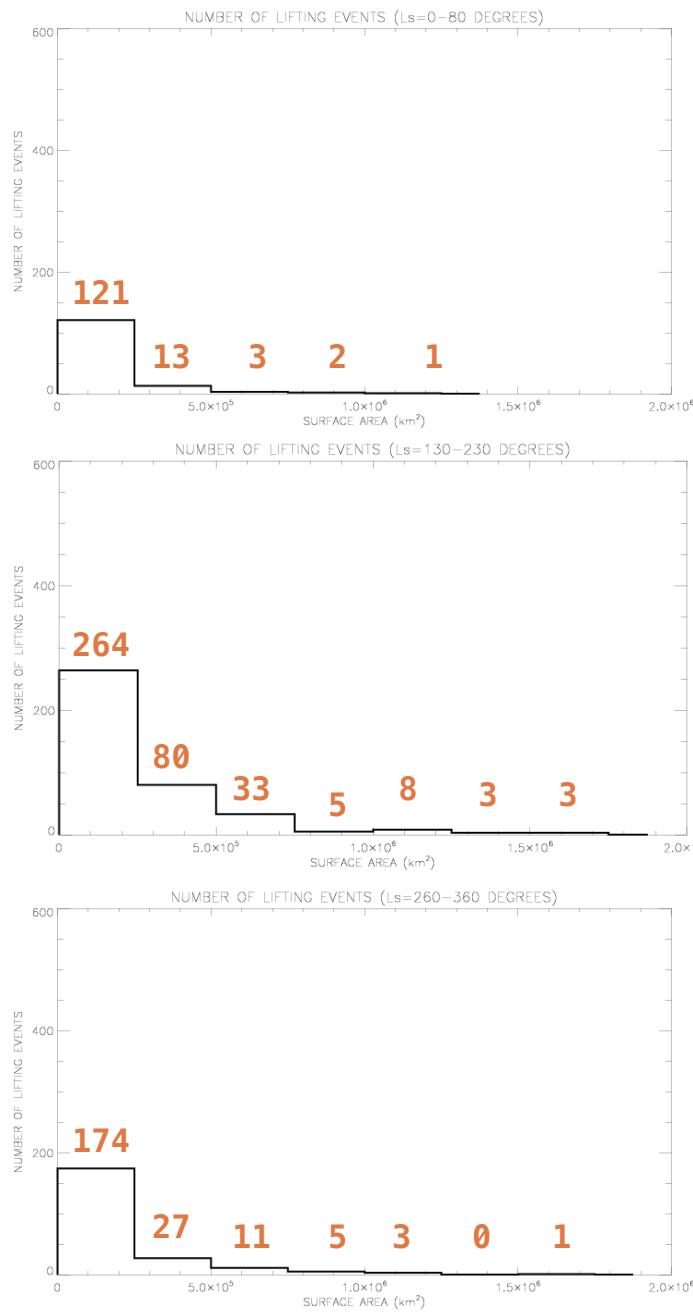
Size Distribution



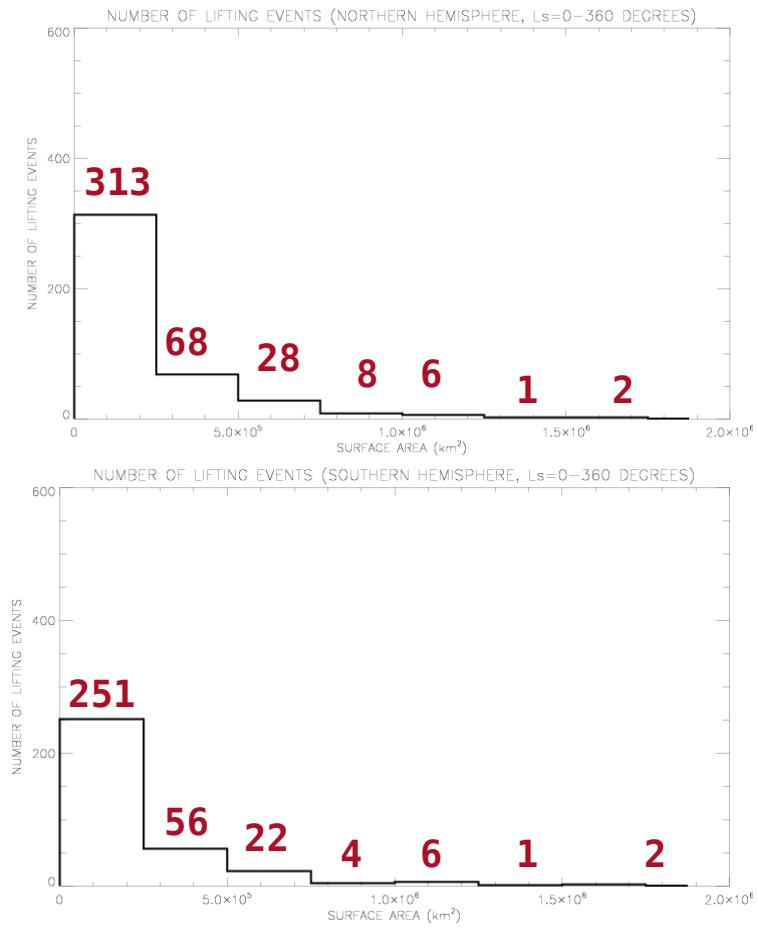
Size Distribution



Seasonal Windows



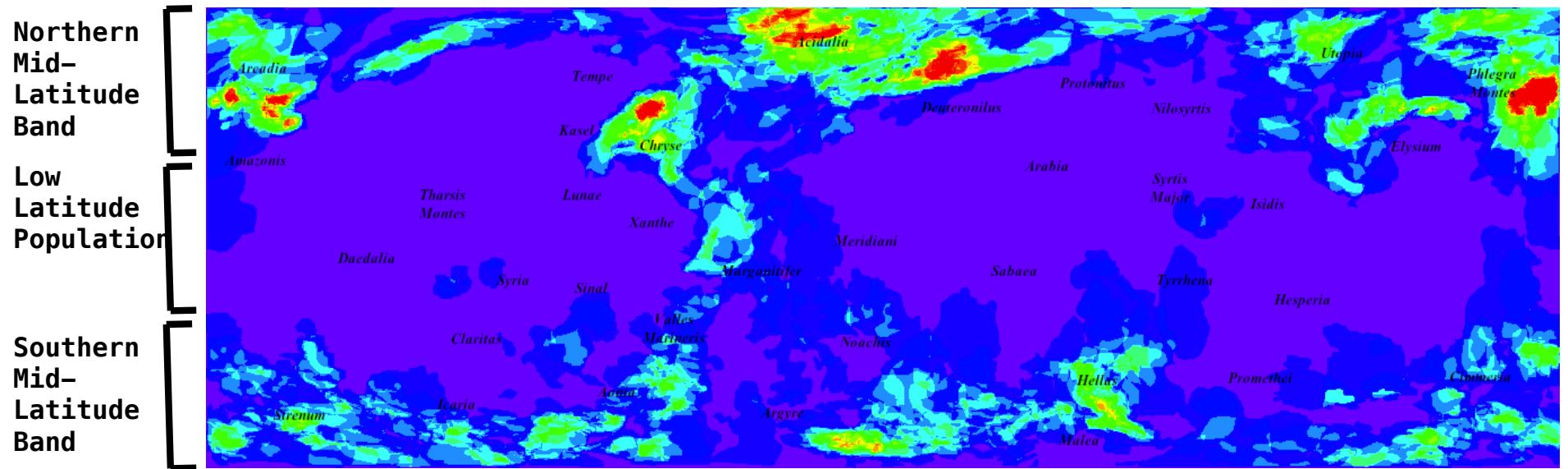
Northern and Southern Hemispheres



- The rapid decay of the number of events with increasing area is true for all seasonal windows and both hemispheres
- The northern and southern hemispheres exhibit similar rates of decay
- Large scale storms tend to occur during the second seasonal window

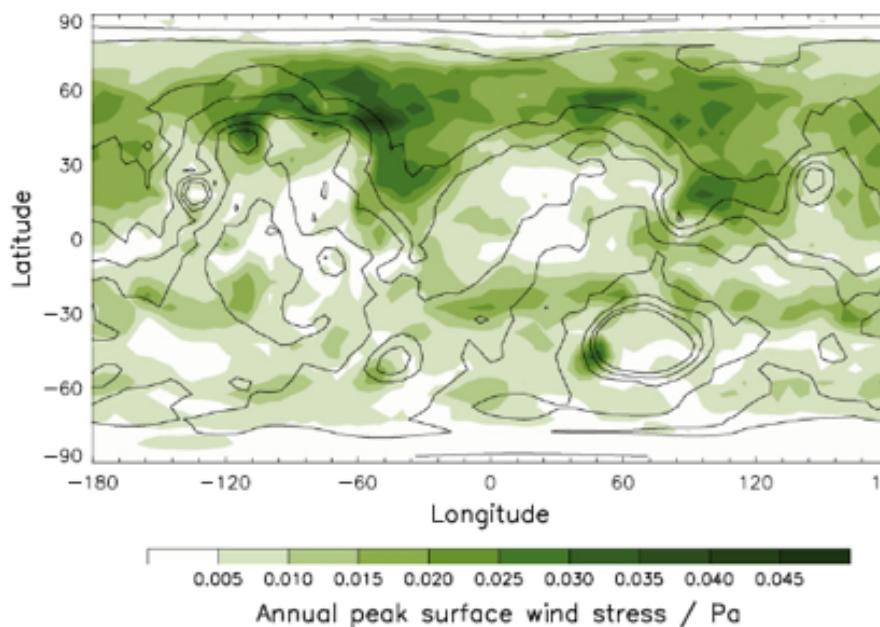
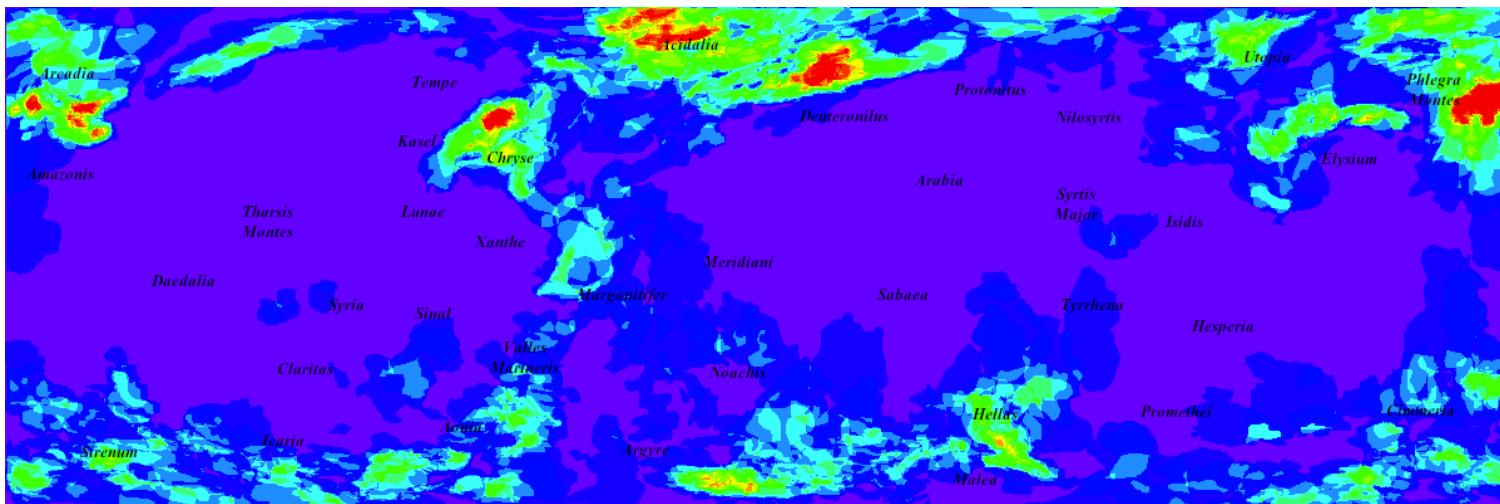
II. Frequency Distribution

Frequency Distribution

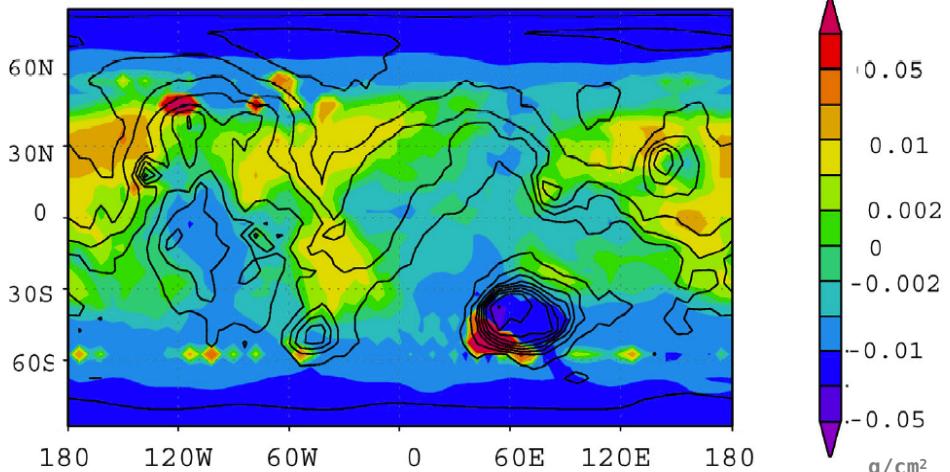


- Non-uniform dust distribution
- Northern band includes regions with the highest dust lifting frequency (Arcadia, Acidalia, Cebrenia)
- Low latitude population has a highly asymmetric distribution
- Southern band has more zonal symmetry

Frequency Distribution



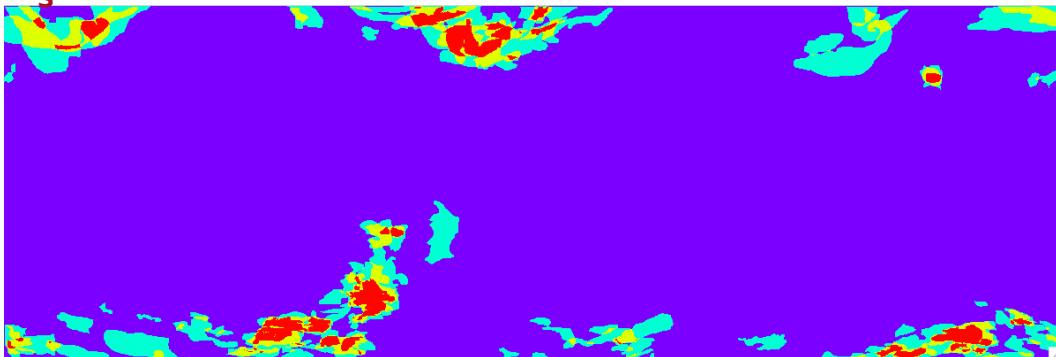
Peak annual wind stress predicted by GCM [Mulholland et al., 2013]



Net annual erosion predicted by the GCM [Basu et al., 2006]

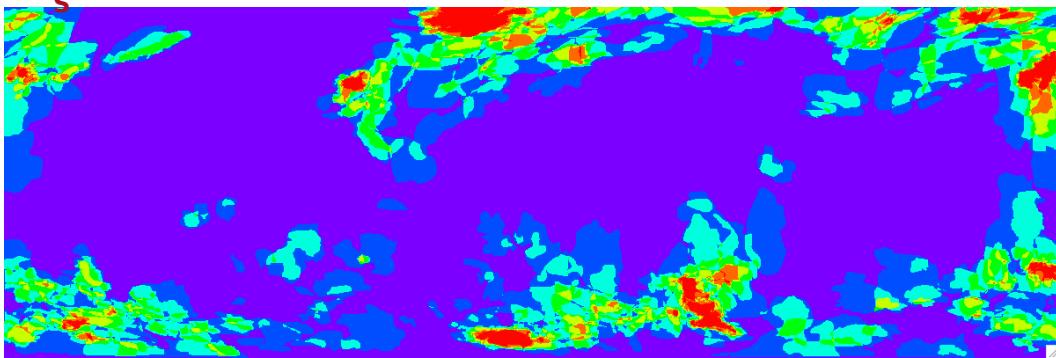
Frequency Distribution

$L_s = 0^\circ - 80^\circ$



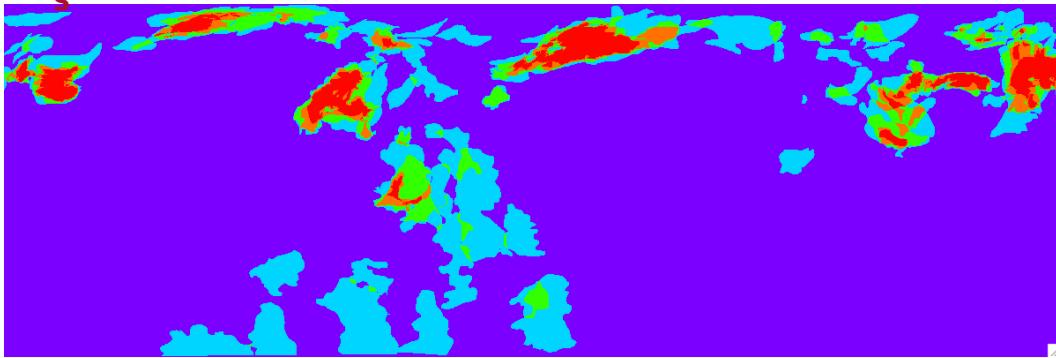
- Northern mid-latitude band and southern mid-latitude band are most active
- The low latitude population is relatively inactive

$L_s = 130^\circ - 230^\circ$



- Most active period of dust lifting
- Lifting occurs in northern and southern bands and low latitude population
- Hellas basin becomes active
- Southern activity has greater zonal symmetry than the mid-latitude northern band or the low latitude population

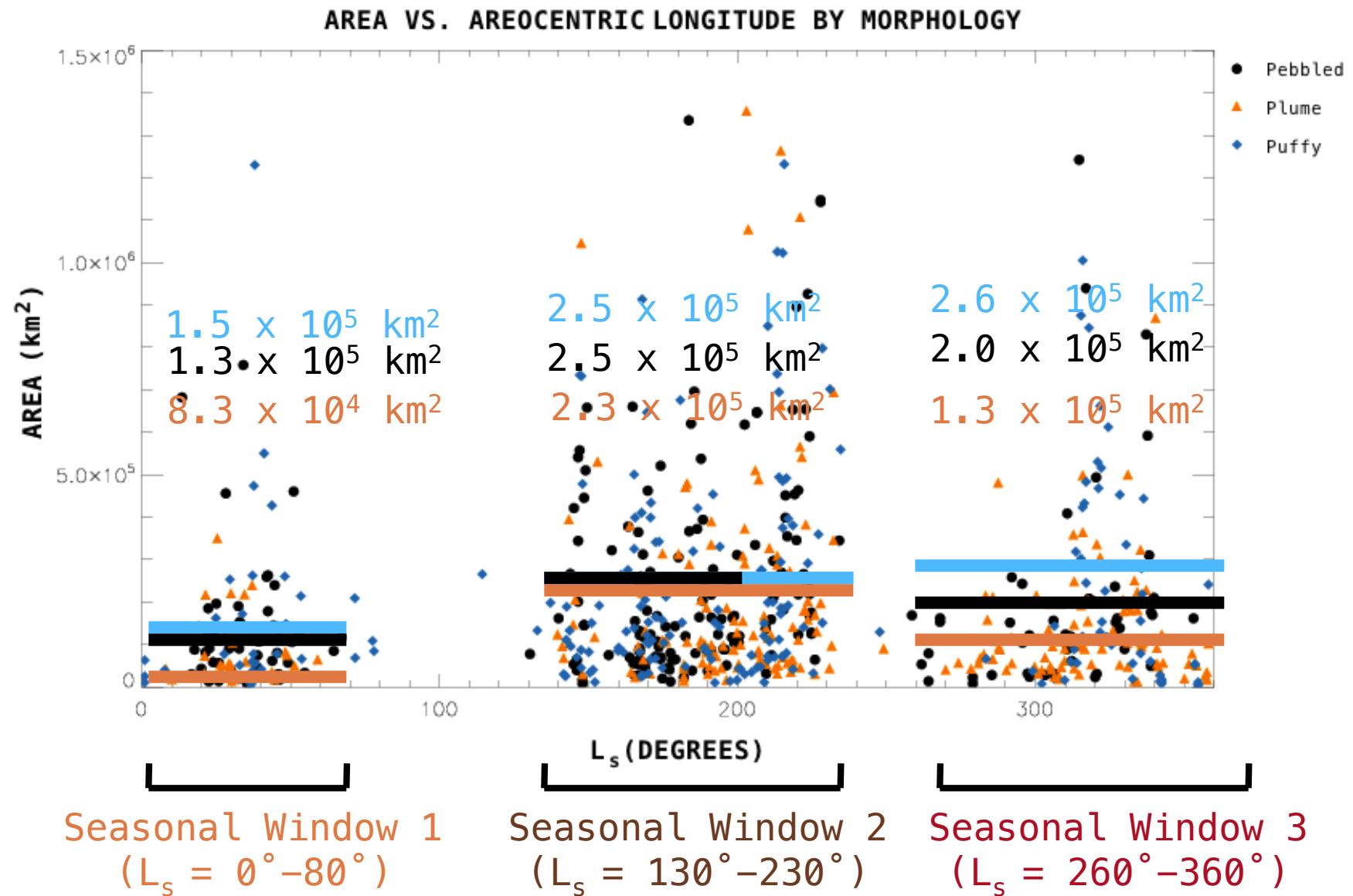
$L_s = 260^\circ - 360^\circ$



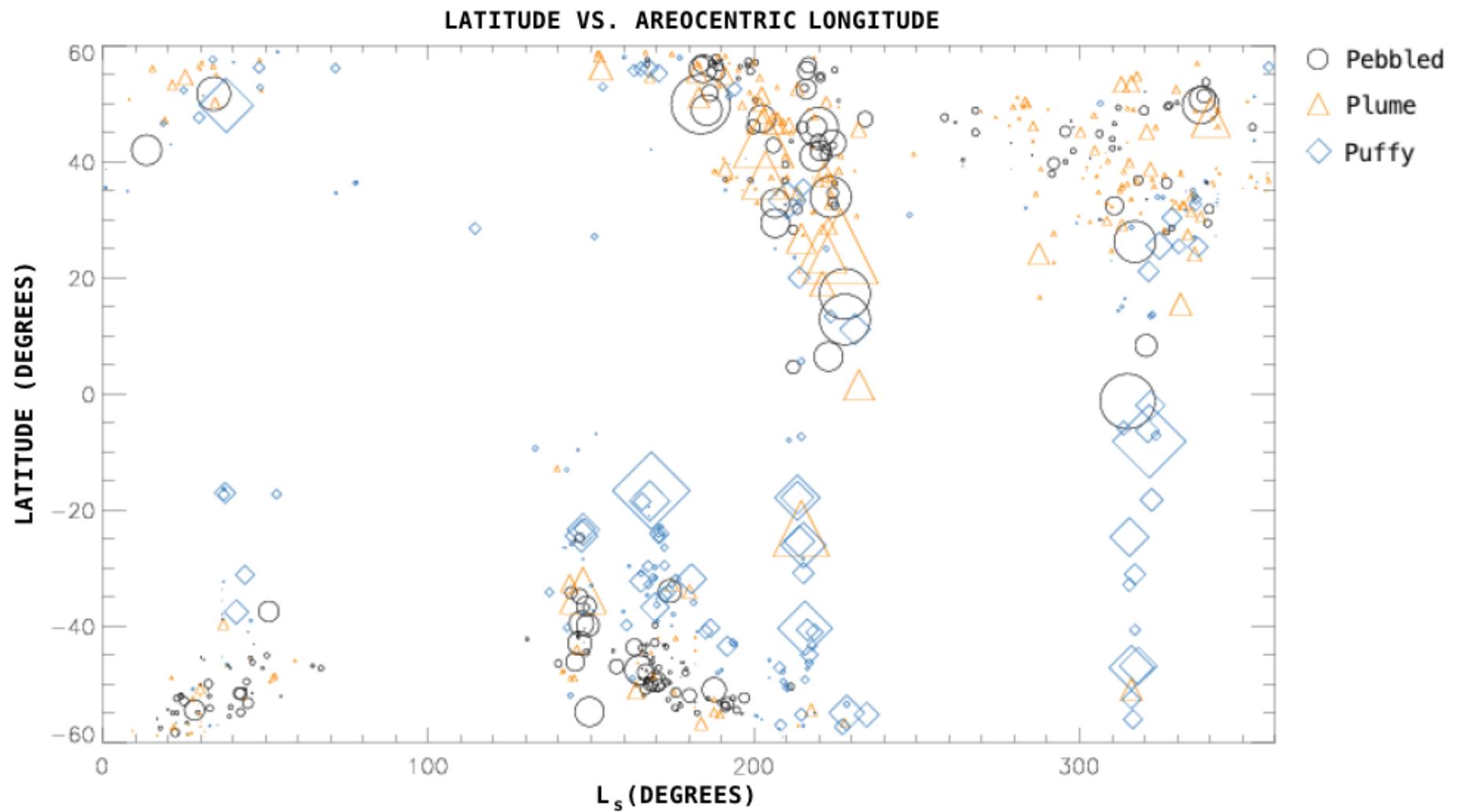
- Northern mid-latitude band similar to second seasonal window
- South is much less active and restricted in zonal expanse
- Hellas is no longer active

III. Morphological Properties

Morphological Properties

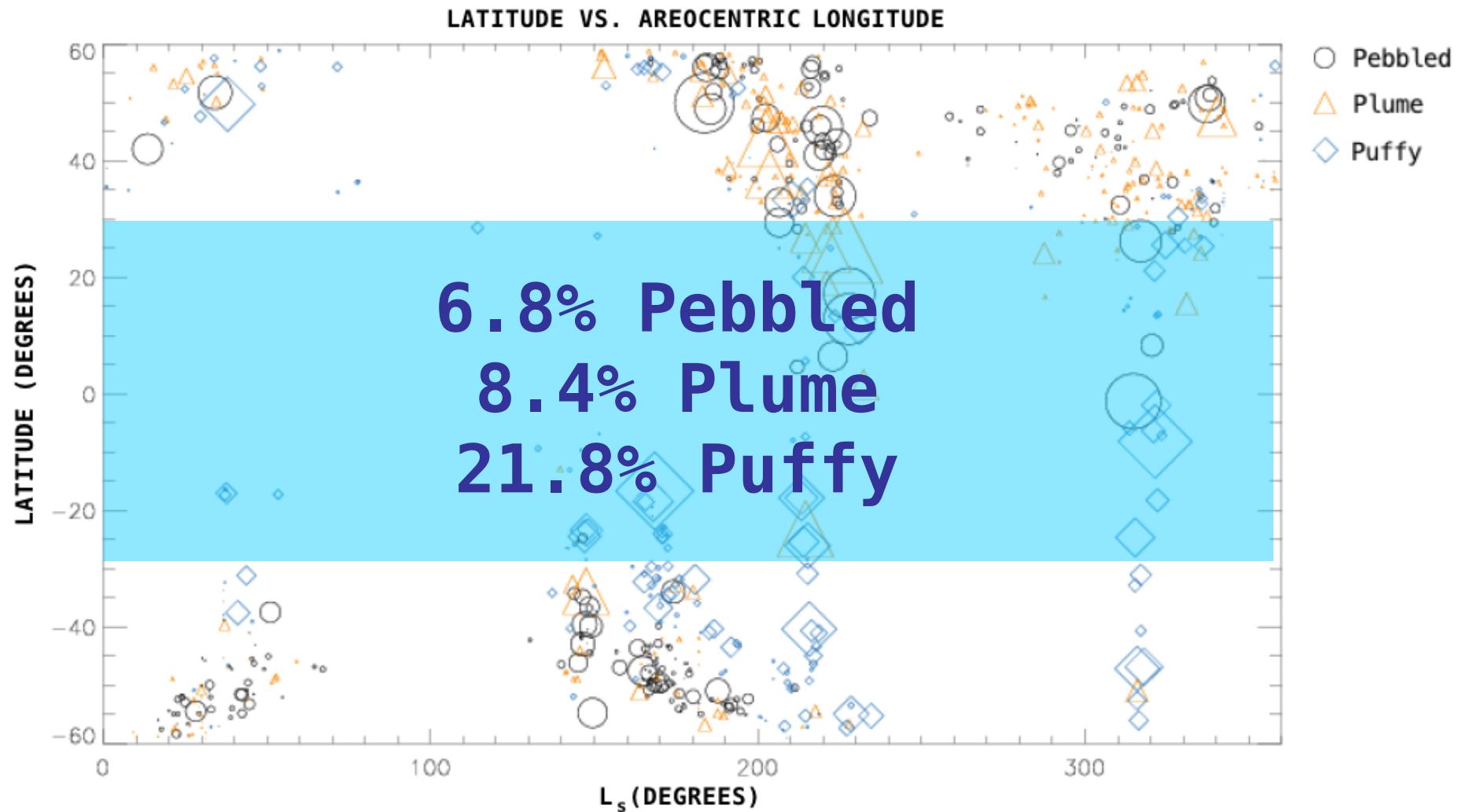


Morphological Properties



Each symbol is proportional to the area of the lifting event

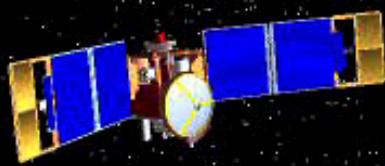
Morphological Properties



Concluding Remarks

- A definite seasonal pattern of dust lifting has been observed ($L_s = 0^\circ - 80^\circ$, $L_s = 130^\circ - 230^\circ$, and $L_s = 260^\circ - 360^\circ$)
- The average size of dust lifting varies with season (largest during the second season window)
- Small scale lifting is more common than large scale lifting
- The north claimed more high frequency lifting areas than the south

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- Lifting frequency appeared to be correlated with wind stress predicted by GCMs
- Pebbled and plume structures are concentrated in higher latitudes while puffy lifting structures are also found in lower latitudes.

L. Kulowski was supported by the National Science Foundation Research Experiences for Undergraduates (REU) and Department of Defense ASSURE programs under NSF Grant no. 1262851 and by the Smithsonian Institution.