

## Homework 7

Astronomers use several different units to measure brightnesses: fluxes per unit wavelength ( $F_\lambda$ ), fluxes per unit frequency ( $F_\nu$ ), and magnitudes, both integrated and per unit wavelength and frequency ( $m = -2.5\log F$ ). Since magnitudes are logarithmic units, differences in magnitude correspond to ratios in fluxes.

- 1 A Jansky is a unit used to measure flux density, most often in the radio; one Jansky is  $10^{-26} \text{Wm}^{-2}\text{Hz}^{-1}$  (note that this is an  $F_\nu$  quantity). How bright is Vega at 5500 in Janskys, using the fact that the flux density of Vega at 5500 is  $3.6 \times 10^{-9} \text{ergs/cm}^2/\text{sec}/$  (note that this is an  $F_\lambda$  quantity)?
- 2 If a star has a flux density of  $3.6 \times 10^{-9} \text{ergs/cm}^2/\text{sec}/$  at 8500, how bright is it in Janskys?
- 3 If a star has a flux density of  $7.2 \times 10^{-14} \text{ergs/cm}^2/\text{sec}/$  at 5500, how much fainter is it than Vega in magnitudes?