DM_halo_models

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Chapter 1

DM_halo_models

Testing DM models with SPARC

See *insert publication here* for more information.

1.1 Overview

This repository was created to test several galactic dark matter (DM) models against the Spitzer Photometry and Accurate Rotation Curves (SPARC) catalog data (see http://astroweb.case.edu/SPARC/). It can be used to obtain fit results for each of the DM models analyzed, as well as calculate various properties such as the galactic DM mass, etc. For more information pertaining to the math and physics behind the calculations, see *insert publication here*.

1.2 Requirements

See requirements.txt

1.3 Fitting

All results, including figures, in *insert publication here* can be reproduced using the Jupyter notebooks. The main results of the paper are included in the notebooks :

- fits_CDM_all.ipynb
- fits_Einasto.ipynb

Comparisons to previous studies can be found in the notebooks :

- checks_CDM_all.ipynb
- checks_DC14_NFW.ipynb
- checks_Einasto_NFW.ipynb

1.4 Acknowledgements

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Thanks to Joshua Eby and Peter Suranyi for valuable discussions and comments from proofreading of *insert publication here*.

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1.5 Citations

Below is the bibtex formatted citation for *insert publication here*,

insert bibtex citation here

Below is the bibtex formatted citation for this repository,

@misc{Street_DM_halo_models, author = {Street, Lauren and Gnedin, Nickolay Y. and Wijewardhana, L. C. R.}, title = {{DM halo models}}, url = { https://github.com/laurenstreet/DM_halo_models} }

1.6 Contact

streetlg@mail.uc.edu

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Chapter 4

Namespace Documentation

4.1 alp_funcs Namespace Reference

The package containing the functions for the ULDM halo models.

Classes

- class base_funcs

  The class containing base functions needed to describe the ULDM halo models.
- class soliton

  The class containing mass functions for the ULDM halo models.

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- def **rhocrit** = standard_consts_in['rhocrit']

  float
  Critical density of universe in GeV^4
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.1.1 Detailed Description

The package containing the functions for the ULDM halo models.

4.2 alp_params Namespace Reference

The package containing the parameters for the ULDM halo models.

Classes

- class soliton

  The class containing all parameters for fitting the ULDM halo models.

Functions

- def alphamatched (m22, msol, c200, V200)

  Define the value of α when using ULDM matched models.

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- def **rhoc** = standard_consts_in['rhocrit']

  float
  Critical density of universe in GeV^4
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.2.1 Detailed Description

The package containing the parameters for the ULDM halo models.

4.2.2 Function Documentation

4.2.2.1 alphamatched()

```
def alp_params.alphamatched (
              m22,
              msol,
              c200,
              V200 )
```

Define the value of α when using ULDM matched models.

This defines the value of α in the ULDM matched models. This value of is fixed from the condition $\rho_{\text{sol}}\left(3\,r_{\text{c,sol}}\right) = \rho_{\text{Einasto}}\left(3\,r_{\text{c,sol}}\right)$.

Parameters

m22	mass of ULDM in units of 10^{-22} eV
msol	mass of soliton in units of solar masses
c200	Einasto profile variable c_{200}
V200	Einasto profile variable V_{200}

Returns

float Value of α to be used for ULDM matched models

4.3 cdm_funcs Namespace Reference

The package containing the functions for the CDM halo models.

Classes

- class base_funcs

  The class containing base functions needed to describe the CDM halo models.
- class halo

  The class containing mass functions for the CDM halo models.

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- def **rhocrit** = standard_consts_in['rhocrit']

  float
  Critical density of universe in GeV^4
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.3.1 Detailed Description

The package containing the functions for the CDM halo models.

4.4 cdm_params Namespace Reference

The package containing the parameters for the CDM halo models.

Classes

- class base_funcs

  The class containing all base functions for the CDM halo parameters.
- class halo

  The class containing all parameters for fitting the CDM halo models.

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- def **rhoc** = standard_consts_in['rhocrit']

  float
  Critical density of universe in GeV^4
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.4.1 Detailed Description

The package containing the parameters for the CDM halo models.

4.5 constants Namespace Reference

The package containing constants, fitting routine definitions, and fitting routine parameter values.

Classes

- class fitting_dict

  The class containing dictionaries of fitting variables.

Functions

- def standard ()

  The dictionary containing standard constants.

4.5.1 Detailed Description

The package containing constants, fitting routine definitions, and fitting routine parameter values.

4.5.2 Function Documentation

4.5.2.1 standard()

```
def constants.standard ( )
```

The dictionary containing standard constants.

This defines the dictionary containing standard constants needed during fitting procedure.

Returns

dictionary
Dictionary of standard constants:

- kmTOGeV in units $[\text{km}^{-1}\text{GeV}^{-1}]$ = conversion from kilometer to GeV
- sTOGeV in units $[\text{s}^{-1}\ \text{GeV}^{-1}]$ = conversion from second to GeV
- kpcTOGeV in units $[\text{kpc}^{-1}\ \text{GeV}^{-1}]$ = conversion from kpc to GeV
- msun in units $[\text{GeV}]$ = mass of sun in GeV
- MP in units $[\text{GeV}]$ = Planck mass in GeV
- rhocrit in units $[\text{GeV}^4]$ = critical density of universe in GeV^4
- MOND_gdag in units $[\text{m}\,\text{s}^{-2}]$ = gravitational acceleration constant for gravitational acceleration relation in MOND theory

4.6 galaxy Namespace Reference

The package containing information about all SPARC catalog galaxies.

Classes

- class galaxy

  The class containing data for all SPARC catalog galaxies.

4.6.1 Detailed Description

The package containing information about all SPARC catalog galaxies.

All data has been taken from SPARC database : http://astroweb.case.edu/SPARC/
Data files can be found in:

- DM_halos_models/pyfiles/data/MassModels_Lelli2016c.txt
- DM_halos_models/pyfiles/data/SPARC_Lelli2016c.txt

4.7 halo Namespace Reference

The package containing the general information for all DM halo models.

Classes

- class halo

  The class containing definitions to describe a DM halo.

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.7.1 Detailed Description

The package containing the general information for all DM halo models.

4.8 model_fit Namespace Reference

The package containing the general fitting procedures.

Classes

- class grar_fit

  The class containing the fitting procedure for the gravitational acceleration relation.
- class model_fit

  The class containing the fitting procedure assuming a chosen halo model for a chosen galaxy.

Functions

- def conc_mass_rel_Du (m200)

  Define the concentration-mass relation (CMR).
- def conc_mass_rel_Wa (m200)

  Define the concentration-mass relation (CMR).
- def BTFR (Vf)

  Define the baryonic Tully-Fisher relation (BTFR).
- def abund_match_rel (m200)

  Define the abundance matching relation (AMR).

Variables

- **standard_consts_in** = constants.standard()

  dictionary
  Dictionary containing standard constants.
- def **kmTOGeV** = standard_consts_in['kmTOGeV']

  float
  Conversion from km to GeV
- def **sTOGeV** = standard_consts_in['sTOGeV']

  float
  Conversion from s to GeV
- def **kpcTOGeV** = standard_consts_in['kpcTOGeV']

  float
  Conversion from kpc to GeV
- def **msun** = standard_consts_in['msun']

  float
  Mass of sun in GeV
- def **MP** = standard_consts_in['MP']

  float
  Planck mass in GeV
- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.8.1 Detailed Description

The package containing the general fitting procedures.

4.8.2 Function Documentation

4.8.2.1 abund_match_rel()

```
def model_fit.abund_match_rel (
              m200 )
```

Define the abundance matching relation (AMR).

This defines the AMR for a given halo mass. See Peter S. Behroozi et al 2013 ApJ 770 57 : https://doi.org/10.1088/0004-637x/770/1/57 and Peter S. Behroozi et al 2013 ApJL 762 L31 : https://doi.org/10.1088/2041-8205/762/2/l31 for more information. The AMR is given by,

$$\frac{M_*}{M_{200}} = 2N \left[\left(\frac{M_{200}}{M_1\, M_\odot} \right)^{-\beta} + \left(\frac{M_{200}}{M_1\, M_\odot} \right)^{-\gamma} \right]^{-1},$$

where $N = 0.0351$, $\beta = 1.375$, $\gamma = 0.608$, and $\log_{10}(M_1) = 11.59$.

Parameters

m200	float Total mass of the DM halo in units of solar masses.

Returns

float
Total stellar mass of a galaxy in units of solar masses.

4.8.2.2 BTFR()

```
def model_fit.BTFR (
              Vf )
```

Define the baryonic Tully-Fisher relation (BTFR).

This defines the BTFR for a given maximum circular velocity of a galaxy. See Federico Lelli et al 2016 ApJL 816 L14 : https://doi.org/10.3847/2041-8205/816/1/l14 for more information. The BTFR is given by,

$$\log_{10} \frac{M_b}{M_\odot} = s \log_{10} \left(\frac{V_f}{\mathrm{km\,s^{-1}}} \right) + \log_{10} A,$$

where $M_\odot$ is the mass of sun, $s = 3.71 \pm 0.08$, and $\log_{10} A = 2.27 \pm 0.18$.

Parameters

Vf	float Maximum circular velocity in units of km s^{-1}.

Returns

float
Total baryonic mass of a galaxy in units of solar masses.

4.8.2.3 conc_mass_rel_Du()

```
def model_fit.conc_mass_rel_Du (
              m200 )
```

Define the concentration-mass relation (CMR).

This defines the CMR for a given halo mass. See Monthly Notices of the Royal Astronomical Society, Volume 441, Issue 4, 11 July 2014, Pages 3359–3374 : https://doi.org/10.1093/mnras/stu742 for more information. The CMR is given by,

$$\log_{10} c_{200} = 0.905 - 0.101 \log_{10} \left(\frac{M_{200}}{10^{12} h^{-1}\, M_\odot} \right),$$

where $h^{-1} = 0.73$ and $M_\odot$ is the mass of sun.

Parameters

m200	float Total mass of the DM halo in units of solar masses.

Returns

float
Concentration parameter assuming the CMR.

4.8.2.4 conc_mass_rel_Wa()

```
def model_fit.conc_mass_rel_Wa (
              m200 )
```

Define the concentration-mass relation (CMR).

This defines the CMR for a given halo mass. See Nature volume 585, pages 39–42 (2020) : https://doi.org/10.1038/s41586-020-2642-9 for more information. The CMR is given by,

$$c_{200} = \sum_{i=0}^{5} c_i \ln\left(\frac{M_{200}}{h^{-1}\,M_\odot}\right)^i,$$

where $h^{-1} = 0.6777$, $M_\odot$ is the mass of sun, $c_0 = 27.112$, $c_1 = -0.381$, $c_2 = -1.853 \times 10^{-3}$, $c_3 = -4.141 \times 10^{-4}$, and $c_5 = 3.208 \times 10^{-7}$.

Parameters

m200	float Total mass of the DM halo in units of solar masses.

Returns

float
Concentration parameter assuming the CMR.

4.9 results Namespace Reference

The package containing the procedures to obtain results for various cases.

Classes

- class plots

  The class to obtain rotation curve plots for given galaxies.
- class results_CDM_all

  The class to obtain fit results for all CDM halos analyzed.

- class results_CDM_check

  The class to obtain fit results to check all CDM model implementation.
- class results_DC14_check

  The class to obtain fit results to check the DC14 and NFW model implementation.
- class results_Einasto_check

  The class to obtain fit results to check the Einasto and NFW model implementation.
- class results_psi_multi_all

  The class to obtain fit results for all double flavored ULDM models analyzed.
- class results_psi_single_all

  The class to obtain fit results for all single flavored ULDM models analyzed.

Variables

- **fitting_dict_in** = constants.fitting_dict()

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class using all default values

4.9.1 Detailed Description

The package containing the procedures to obtain results for various cases.

Chapter 5

Class Documentation

5.1 alp_funcs.base_funcs Class Reference

The class containing base functions needed to describe the ULDM halo models.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)

  Define the constructor of the base_funcs class.
- def msol (self, params)

  Define the total soliton mass.
- def rc (self, params)

  Define the soliton profile variable rc.
- def rhoc (self, params)

  Define the soliton profile variable rhoc.
- def xc (self, params, r)

  Define the soliton profile variable xc.
- def mass_sol_init (self, params, r)

  Define the mass profile of the soliton.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**

  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- matched

  bool
  Denotes how to combine the soliton and outer halo.
- mfree

float
Denotes how to treat soliton particle mass in fitting procedure.

- cdmhalo

  float
  Denotes which CDM profile to use for outer halo in ULDM galactic structure.

- **m22**

  float
  Soliton particle mass in units of 10^{-22} *eV.*

- **m22_2**

  float
  Soliton particle mass two in units of 10^{-22} *eV.*

5.1.1 Detailed Description

The class containing base functions needed to describe the ULDM halo models.

5.1.2 Constructor & Destructor Documentation

5.1.2.1 __init__()

```
def alp_funcs.base_funcs.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the base_funcs class.

This defines the constructor of the base_funcs class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • psi_single • psi_multi
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.1.3 Member Function Documentation

5.1.3.1 mass_sol_init()

```
def alp_funcs.base_funcs.mass_sol_init (
              self,
              params,
              r )
```

Define the mass profile of the soliton.

This defines the mass of the soliton at a given radius for the given model parameters.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N] (for single flavored models) or ndarray[2,N] (for double flavored models)
Numpy array of N soliton mass values (for single flavored models) or numpy array of [2,N] mass values (for double flavored models) in units of solar mass.

5.1.3.2 msol()

```
def alp_funcs.base_funcs.msol (
              self,
              params )
```

Define the total soliton mass.

This defines the total soliton mass for the given model parameters.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.

Returns

float (for single flavored models) or ndarray[2] (for double flavored models)
The total soliton mass (for single flavored models) or both total soliton masses, msol_1 and msol_2 (for double flavored models) in units of solar mass.

5.1.3.3 rc()

```
def alp_funcs.base_funcs.rc (
              self,
              params )
```

Define the soliton profile variable rc.

This defines the soliton profile variable rc for the given model parameters. The soliton profile variable rc is given by,

$$r_c \approx 0.228 \left(\frac{M_{\rm sol}}{10^9}\right)^{-1} m^{-2},$$

where $M_{\rm sol}$ = alp_funcs.base_funcs.msol and m = m22 in alp_params.soliton.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.

Returns

float (for single flavored models) or ndarray[2] (for double flavored models)
The halo profile variable rc in units of kpc. The soliton profile variable rc (for single flavored models) or both soliton profile variables, rc_1 and rc_2 (for double flavored models) in units of kpc.

5.1.3.4 rhoc()

```
def alp_funcs.base_funcs.rhoc (
              self,
              params )
```

Define the soliton profile variable rhoc.

This defines the soliton profile variable rhoc for the given model parameters. The soliton profile variables rhoc is given by,

$$\rho_c \approx 7 \times 10^9 \left(\frac{M_{\rm sol}}{10^9}\right)^{4} m^{6},$$

where $M_{\rm sol}$ = alp_funcs.base_funcs.msol and m = m22 in alp_params.soliton.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.

Returns

float (for single flavored models) or ndarray[2] (for double flavored models)
The soliton profile variable rhoc (for single flavored models) or both soliton profile variables, rhoc_1 and rhoc_2 (for double flavored models) in units of $M_\odot/\text{kpc}^3$.

5.1.3.5 xc()

```
def alp_funcs.base_funcs.xc (
              self,
              params,
              r )
```

Define the soliton profile variable xc.

This defines the soliton profile variable xc = r/rc for the given model parameters. The soliton profile variable xc is given by,

$$x_c = r/r_c,$$

where $r_c =$ alp_funcs.base_funcs.rc.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N] (for single flavored models) or ndarray[2,N] (for double flavored models)
Numpy array of N (for N given radii) soliton profile variables xc (for single flavored models) or [2,N] soliton profile variables, xc_1 and xc_2 (for double flavored models).

5.1.4 Member Data Documentation

5.1.4.1 cdmhalo

```
alp_funcs.base_funcs.cdmhalo
```

float
Denotes which CDM profile to use for outer halo in ULDM galactic structure.

Equal to constants.fitting_dict.sol_cdmhalo

5.1.4.2 matched

`alp_funcs.base_funcs.matched`

bool
Denotes how to combine the soliton and outer halo.

Equal constants.fitting_dict.sol_match

5.1.4.3 mfree

`alp_funcs.base_funcs.mfree`

float
Denotes how to treat soliton particle mass in fitting procedure.

Equal to constants.fitting_dict.sol_mfree

The documentation for this class was generated from the following file:

- pyfiles/models/alps/alp_funcs.py

5.2 cdm_funcs.base_funcs Class Reference

The class containing base functions needed to describe the CDM halo models.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)
  Define the constructor of the base_funcs class.
- def mass_frac_dc14 (self, params)
  Define the mass fraction for the DC14 model.
- def rc (self, params)
  Define the halo profile variable rc.
- def rhoc (self, params)
  Define the halo profile variable rhoc.
- def xc (self, params, r)
  Define the halo profile variable xc.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- mass_num
  int
  Used to differentiate between soliton 1 and soliton 2 in double flavored ULDM models.

5.2.1 Detailed Description

The class containing base functions needed to describe the CDM halo models.

5.2.2 Constructor & Destructor Documentation

5.2.2.1 __init__()

```
def cdm_funcs.base_funcs.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the base_funcs class.

This defines the constructor of the base_funcs class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.2.3 Member Function Documentation

5.2.3.1 mass_frac_dc14()

```
def cdm_funcs.base_funcs.mass_frac_dc14 (
              self,
              params )
```

Define the mass fraction for the DC14 model.

This defines the stellar to DM mass fraction for the DC14 model assuming the given model parameters. The stellar to DM mass fraction is given by,

$$X = \log_{10}\left(M_*/M_{\rm halo}\right),$$

where,

$$M_* \approx \left(\tilde{\Upsilon}_{\rm disk} + \tilde{\Upsilon}_{\rm bulge}\right) L,$$

with L = Luminosity in galaxy.galaxy.data, $M_{\rm halo}$ = cdm_funcs.halo.Mvir, and $\tilde{\Upsilon}_{\rm disk}$ = MLD and $\tilde{\Upsilon}_{\rm bulge}$ = MLB in cdm_params.halo.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.

Returns

float
The stellar to DM mass fraction in the DC14 model.

5.2.3.2 rc()

```
def cdm_funcs.base_funcs.rc (
              self,
              params )
```

Define the halo profile variable rc.

This defines the halo profile variable rc for the given model parameters. The halo profile variable rc is given by,

$$r_c = \sqrt{\frac{3}{2\pi\,\rho_{\rm crit}}}\,\frac{M_P\,V_{200}}{20\,c_{200}},$$

where $\rho_{\rm crit}$ = rhocrit and M_P = MP in constants.standard, while V_{200} = v200 and c_{200} = c200 in cdm_params.halo.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.

Returns

float
The halo profile variable rc in units of kpc.

5.2.3.3 rhoc()

```
def cdm_funcs.base_funcs.rhoc (
              self,
              params )
```

Define the halo profile variable rhoc.

This defines the halo profile variable rhoc for the given model parameters. The halo profile variables rhoc is given by,

$$\rho_c = \frac{M_{200}}{4\pi\, r_c^3 \left[\ln\left(1+c_{200}\right) - \frac{c_{200}}{1+c_{200}}\right]},$$

where $r_c =$ cdm_funcs.base_funcs.rc, $M_{200} =$ cdm_funcs.halo.Mvir, and $c_{200} =$ c200 in cdm_params.halo.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.

Returns

float
The halo profile variable rhoc in units of $M_\odot/\text{kpc}^3$.

5.2.3.4 xc()

```
def cdm_funcs.base_funcs.xc (
              self,
              params,
              r )
```

Define the halo profile variable xc.

This defines the halo profile variable xc for the given model parameters. The halo profile variable xc is given by,

$$x_c = r/r_c,$$

where $r_c =$ cdm_funcs.base_funcs.rc.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N]
Numpy array of N halo profile variables xc.

5.2.4 Member Data Documentation

5.2.4.1 mass_num

`cdm_funcs.base_funcs.mass_num`

int
Used to differentiate between soliton 1 and soliton 2 in double flavored ULDM models.

Outer CDM halo is halo 1 if mass_num = 0 and is halo 2 if mass_num = 1.

The documentation for this class was generated from the following file:

- pyfiles/models/cdm/cdm_funcs.py

5.3 cdm_params.base_funcs Class Reference

The class containing all base functions for the CDM halo parameters.

Public Member Functions

- def __init__ (self)
  Define the constructor of the base_funcs class.
- def v200min_frac_dc14check (self, mstar, mgas, vfac)
  Define the expression for V200 to be assumed in particular cases.
- def v200min_dc14 (self, mstar)
  Define the minimum allowed V200 for the DC14 model.

5.3.1 Detailed Description

The class containing all base functions for the CDM halo parameters.

This class contains functions necessary to describe CDM halo parameters for various cases.

5.3.2 Constructor & Destructor Documentation

5.3.2.1 __init__()

```
def cdm_params.base_funcs.__init__ (
            self )
```

Define the constructor of the base_funcs class.

This defines the constructor of the base_funcs class.

Parameters

self	object pointer

5.3.3 Member Function Documentation

5.3.3.1 v200min_dc14()

```
def cdm_params.base_funcs.v200min_dc14 (
              self,
              mstar )
```

Define the minimum allowed V200 for the DC14 model.

This defines the minimum allowed V200 for the DC14 model. This is assumed when using the fit routines :

- constants.fitting_dict.fit_routine = 'uni_priors',
- constants.fitting_dict.fit_routine = 'c200_priors_check',
- constants.fitting_dict.fit_routine = 'v200_priors_check',
- constants.fitting_dict.fit_routine = 'MLd_priors_check',
- constants.fitting_dict.fit_routine = 'MLb_priors_check',
- constants.fitting_dict.fit_routine = 'CDM_check' The allowed values for V200 are given by,

$$V_{200} \geq \left[10^{1.3} M_* \sqrt{\frac{2\pi\,\rho_{\rm crit}}{3}} \frac{20 \times 10^9}{M_P^3} \right]^{1/3},$$

where,

$$M_* \approx \left(\tilde{\Upsilon}_{\rm disk} + \tilde{\Upsilon}_{\rm bulge} \right) L.$$

Here, $\rho_{\rm crit}$ = rhocrit and M_P = MP in constants.standard, $M_{\rm gas}$ = Mgas and L = Luminosity in galaxy.galaxy.data, $\tilde{\Upsilon}_{\rm disk}$ = MLD and $\tilde{\Upsilon}_{\rm bulge}$ = MLB in cdm_params.halo.params.

Parameters

self	object pointer
mstar	float Total stellar mass in units of 10^9 solar masses.

Returns

float
Minimun allowed value for V200 for the DC14 model for various cases.

5.3.3.2 v200min_frac_dc14check()

```
def cdm_params.base_funcs.v200min_frac_dc14check (
              self,
              mstar,
              mgas,
              vfac )
```

Define the expression for V200 to be assumed in particular cases.

This defines the allowable values for V200 in $\log_{10}$ space when using the fit routine constants.fitting_dict.fit_routine = 'DC14_check'. The allowable values for V200 are given by,

$$V_{200} \geq \left[\left(M_* + M_{\rm gas} \right) \sqrt{\frac{2\pi\, \rho_{\rm crit}}{3}} \frac{20 \times 10^9}{0.2\, M_P^3} \right]^{1/3},$$

where,

$$M_* \approx \left(\tilde{\Upsilon}_{\rm disk} + \tilde{\Upsilon}_{\rm bulge} \right) L.$$

Here, $\rho_{\rm crit}$ = rhocrit and M_P = MP in constants.standard, $M_{\rm gas}$ = Mgas and L = Luminosity in galaxy.galaxy.data, $\tilde{\Upsilon}_{\rm disk}$ = MLD and $\tilde{\Upsilon}_{\rm bulge}$ = MLB in cdm_params.halo.params.

Parameters

self	object pointer
mstar	float Total stellar mass in units of 10^9 solar masses.
mgas	float Total HI gas mass in units of 10^9 solar masses.
vfac	float Factor to vary in fitting procedure.

Returns

float
Value for V200 in $\log_{10}$ space in units of $\log_{10}\left(\mathrm{km\,s^{-1}}\right)$.

The documentation for this class was generated from the following file:

- pyfiles/models/cdm/cdm_params.py

5.4 constants.fitting_dict Class Reference

The class containing dictionaries of fitting variables.

Public Member Functions

- def __init__ (self, fit_routine='uni_priors', sol_match=False, sol_mfree=False, sol_cdm_halo='Einasto', sol_m22=10 **(1.5), sol_m22_2=1, sol_m22_tab=np.float_power(10,(np.arange(0, 2, 0.15))), sol_m22_tab_prime=np.↩ float_power(10,(np.arange(-3, 4, 0.5))), sol_m22_2_tab=np.float_power(10,(np.arange(0, 2, 0.15))), sol_m22_2_tab_prime=np.float_power(10,(np.arange(-3, 4, 0.5))), sol_mass_ind=0)

Define the constructor of the fitting_dict class.

- def args_opts (self)

  Define the fitting rules.

- def params_vals (self)

  Define the fitting values.

Public Attributes

- **fit_routine**

  str
  Denotes the fitting routine to use.

- **sol_match**

  bool
  Denotes how to combine the soliton and outer halo.

- **sol_mfree**

  bool
  Denotes how to treat soliton particle mass in fitting procedure.

- **sol_cdm_halo**

  str
  Denotes which CDM profile to use for outer halo in ULDM galactic structure.

- **sol_m22**

  float
  Soliton particle mass (for single flavored models) and soliton 1 particle mass (for double flavored models) in units of 10^{-22} *eV.*

- **sol_m22_2**

  float
  Soliton 2 particle mass (for double flavored models) in units of 10^{-22} *eV.*

- **sol_m22_tab**

  ndarray[N]
  Numpy array of soliton particle masses (for single flavored, matched models) and soliton 1 particle masses (for double flavored, matched models) in units of 10^{-22} *eV.*

- **sol_m22_tab_prime**

  ndarray[N]
  Numpy array of soliton particle masses (for single flavored, summed models) and soliton 1 particle masses (for double flavored, summed models) in units of 10^{-22} *eV.*

- **sol_m22_2_tab**

  ndarray[N]
  Numpy array of soliton 2 particle masses (for double flavored, matched models) in units of 10^{-22} *eV.*

- **sol_m22_2_tab_prime**

  ndarray[N]
  Numpy array of soliton 2 particle masses (for double flavored, summed models) in units of 10^{-22} *eV.*

- **sol_mass_ind**

  int
  Used to differentiate between soliton 1 and soliton 2 in double flavored models.

5.4.1 Detailed Description

The class containing dictionaries of fitting variables.

5.4.2 Constructor & Destructor Documentation

5.4.2.1 __init__()

```
def constants.fitting_dict.__init__ (
                self,
                fit_routine = 'uni_priors',
                sol_match = False,
                sol_mfree = False,
                sol_cdm_halo = 'Einasto',
                sol_m22 = 10**(1.5),
                sol_m22_2 = 1,
                sol_m22_tab = np.float_power(10, (np.arange(0, 2, 0.15))),
                sol_m22_tab_prime = np.float_power(10, (np.arange(-3, 4, 0.5))),
                sol_m22_2_tab = np.float_power(10, (np.arange(0, 2, 0.15))),
                sol_m22_2_tab_prime = np.float_power(10, (np.arange(-3, 4, 0.5))),
                sol_mass_ind = 0 )
```

Define the constructor of the fitting_dict class.

This defines the constructor of the fitting_dict class.

Parameters

self	object pointer
fit_routine	(optional) str Denotes the fitting routine to use. Can be equal to: • 'uni_priors' : uniform priors on all parameters (used for main results of paper) • 'c200_priors_check' : used to check affect of changing prior ranges for c200 • 'v200_priors_check' : used to check affect of changing prior ranges for v200 • 'MLd_priors_check' : used to check affect of changing prior ranges for MLd • 'MLb_priors_check' : used to check affect of changing prior ranges for MLb • 'CDM_check' : used to obtain results to be compared with Pengfei Li et al 2020 ApJS 247 31 : https://doi.org/10.3847/1538-4365/ab700e • 'DC14_check' : used to obtain results to be compared with Monthly Notices of the Royal Astronomical Society, Volume 466, Issue 2, April 2017, Pages 1648–1668 : https://doi.org/10.1093/mnras/stw3101 • 'Einasto_check' : used to obtain results to be compared with Nicolas Loizeau and Glennys R. Farrar 2021 ApJL 920 L10 : https://doi.org/10.3847/2041-8213/ac1bb7
sol_match	(optional) bool Denotes how to combine the soliton and outer halo. If equal to: • False : soliton and CDM halo to be summed • True : soliton and CDM halo to be matched
sol_mfree	(optional) bool Denotes how to treat soliton particle mass in fitting procedure. If equal to: • False : soliton particle mass set to be fixed • True : soliton particle mass to be free

Parameters

sol_cdm_halo	(optional) str Denotes which CDM profile to use for outer halo in ULDM galactic structure. Can be equal to: • 'Burkert' • 'DC14' • 'Einasto' • 'NFW'
sol_m22	(optional) float Soliton particle mass (for single flavored models) and soliton 1 particle mass (for double flavored models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_m22_2	(optional) float Soliton 2 particle mass (for double flavored models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_m22_tab	(optional) ndarray[N] Numpy array of soliton particle masses (for single flavored, matched models) and soliton 1 particle masses (for double flavored, matched models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_m22_tab_prime	(optional) ndarray[N] Numpy array of soliton particle masses (for single flavored, summed models) and soliton 1 particle masses (for double flavored, summed models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_m22_2_tab	(optional) ndarray[N] Numpy array of soliton 2 particle masses (for double flavored, matched models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_m22_2_tab_prime	(optional) ndarray[N] Numpy array of soliton 2 particle masses (for double flavored, summed models) in units of 10^{-22} eV. Only used for case in which soliton particle mass is fixed (i.e. constants.fitting_dict.sol_mfree = False)
sol_mass_ind	(optional) int Used to differentiate between soliton 1 and soliton 2 in double flavored models. Can be equal to 0 (to denote soliton 1) or 1 (to denote soliton 2)

5.4.3 Member Function Documentation

5.4.3.1 args_opts()

```
def constants.fitting_dict.args_opts (
            self )
```

Define the fitting rules.

This defines the rules to be assumed during fitting.
Many prior cases and fitting routines to compare to previous studies are taken into account.

Parameters

self	object pointer

Returns

dictionary
Dictionary of rules to be assumed during fitting:

- 'fitting_routine' : Dictionary of possible fitting routines
  Each of the following fitting routines can be equal to true or false. By default, the only fitting routine that is set to true is the one equal to fit_routine.
  - 'uni_priors' : bool
  - 'c200_priors_check' : bool
  - 'v200_priors_check' : bool
  - 'MLd_priors_check' : bool
  - 'MLb_priors_check' : bool
  - 'CDM_check' : bool
  - 'DC14_check' : bool
  - 'Einasto_check' : bool
- 'soliton' : Dictionary of soliton variables
  - 'matched' : bool (equal to constants.fitting_dict.sol_match)
  - 'mfree' : bool (equal to constants.fitting_dict.sol_mfree)
  - 'cdm_halo' : str (equal to constants.fitting_dict.sol_cdm_halo)
  - 'm22' : float (equal to constants.fitting_dict.sol_m22)
  - 'm22_2' : float (equal to constants.fitting_dict.sol_m22_2)
  - 'm22_tab' : ndarray[N] (equal to constants.fitting_dict.sol_m22_tab)
  - 'm22_tab_prime' : ndarray[N] (equal to constants.fitting_dict.sol_m22_tab_prime)
  - 'm22_2_tab' : ndarray[N] (equal to constants.fitting_dict.sol_m22_2_tab)
  - 'm22_2_tab_prime' : ndarray[N] (equal to constants.fitting_dict.sol_m22_2_tab_prime)
  - 'mass_ind' : int (equal to constants.fitting_dict.sol_mass_ind)

5.4.3.2 params_vals()

```
def constants.fitting_dict.params_vals (
              self )
```

Define the fitting values.

This defines the values for various variables used during the fitting procedure.
Many prior cases and fitting routines to compare to previous studies are taken into account. Values are setup to be utilized in lmfit.Parameters class.

Parameters

self	object pointer

Returns

dictionary
Possible dictionary of variables values to be used during fitting.
Variables values are setup in lists with format [starting value,min,max]
Dictionary to be used depends on value of fit_routine.
If:

- fit_routine = 'uni_priors', resulting dictionary is:
  - 'CDM' : values for variables describing CDM halos
    - 'c200' : $[3,\ 1,\ 100]$
    - 'v200' : $[100,\ 1,\ 1000]$ in units of km s^{-1}
    - 'v200_fac' : $[1.5,\ 1,\ \infty]$
    - 'MLd' : $[0.5,\ 0.01,\ 5]$ in units of $M_\odot/L_\odot$
    - 'MLb' : $[0.7,\ 0.01,\ 5]$ in units of $M_\odot/L_\odot$
    - 'alpha' : $[0.16,\ -\infty,\ \infty]$ (only used for Einasto profile)
  - 'soliton' : values for variables describing soliton
    - 'Msol' : $[10^9,\ 10^{4.5},\ 10^{12}]$ in units of $M_\odot$
    - 'Msol_2' : $[10^9,\ 10^{4.5},\ 10^{12}]$ in units of $M_\odot$
    - 'm22' : $[1,\ 10^{-3},\ 10^3]$ in units of 10^{-22} eV
    - 'm22_2' : $[1,\ 10^{-3},\ 10^3]$ in units of 10^{-22} eV
- fit_routine = 'c200_priors_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'c200' : $[3,\ 0,\ \infty]$
- fit_routine = 'v200_priors_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'v200' : $[100,\ 0,\ \infty]$ in units of km s^{-1}
- fit_routine = 'MLd_priors_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'MLd' : $[0.5,\ 0,\ \infty]$ in units of $M_\odot/L_\odot$
- fit_routine = 'MLb_priors_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'MLb' : $[0.7,\ 0,\ \infty]$ in units of $M_\odot/L_\odot$
- fit_routine = 'CDM_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'c200' : $[3,\ 10^{-1},\ 10^3]$
    - 'v200' : $[100,\ 10,\ 500]$ in units of km s^{-1}
- fit_routine = 'DC14_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'c200' : $[\log_{10}5,\ \log_{10}1,\ \log_{10}100]$
    - 'v200' : $[\log_{10}100,\ \log_{10}10,\ \log_{10}500]$ in units of $\log_{10}\left(\text{km s}^{-1}\right)$
    - 'v200_fac' : $[2,\ 1,\ \infty]$
    - 'MLd' : $[\log_{10}0.5,\ \log_{10}0.3,\ \log_{10}0.8]$ in units of $\log_{10}\left(M_\odot/L_\odot\right)$
- fit_routine = 'Einasto_check', resulting dictionary is same as 'uni_priors' with following changes:
  - 'CDM' :
    - 'v200' : $[100,\ 1,\ 500]$ in units of km s^{-1}
    - 'alpha' : $[0.16,\ 10^{-3},\ 10]$

The documentation for this class was generated from the following file:

- pyfiles/data_models/constants.py

5.5 galaxy.galaxy Class Reference

The class containing data for all SPARC catalog galaxies.

Public Member Functions

- def __init__ (self, name)

  Define the constructor of the galaxy class.

Public Attributes

- **name**

  str
  Galaxy name
- df

  Pandas dataframe
  Dataframe that contains data from SPARC catalog.
- df1

  Pandas dataframe
  Dataframe that contains data from SPARC catalog.
- **gal_init**

  Pandas dataframe
  Dataframe that contains data for an initialized galaxy.
- data

  dictionary Dictionary that contains data for the chosen galaxy (i.e.

5.5.1 Detailed Description

The class containing data for all SPARC catalog galaxies.

5.5.2 Constructor & Destructor Documentation

5.5.2.1 __init__()

```
def galaxy.galaxy.__init__ (
              self,
              name )
```

Define the constructor of the galaxy class.

This defines the constructor of the galaxy class.

Parameters

self	object pointer
name	str
	Can be equal to : • Any galaxy name in the SPARC catalog (e.g. 'CamB') The galaxy with galaxy name will be analyzed.

5.5.3 Member Data Documentation

5.5.3.1 data

`galaxy.galaxy.data`

dictionary Dictionary that contains data for the chosen galaxy (i.e.

either 'name' or 'random') :

- 'Radius' : ndarray[N]
  Numpy array of galactocentric radii in units of $[\text{kpc}]$
- 'Vobs' : ndarray[N]
  Numpy array of observed circular velocities at given radii in units of $[\text{km s}^{-1}]$
- 'eVobs' : ndarray[N]
  Numpy array of error in observed circular velocities in units of $[\text{km s}^{-1}]$
- 'Vgas' : ndarray[N]
  Numpy array of contribution from gas to circular velocities in units of $[\text{km s}^{-1}]$
- 'Vdisk' : ndarray[N]
  Numpy array of contribution from the disk to circular velocities in units of $[\text{km s}^{-1}]$
- 'Vbulge' : ndarray[N]
  Numpy array of contribution from the bulge to circular velocities in units of $[\text{km s}^{-1}]$
- 'Luminosity' : float
  Total luminosity at $[3.6\,\mu\text{m}]$ in units of $[10^9\,L_\odot]$
- 'Mgas': float
  Total mass of HI gas in units of $[10^9\,M_\odot]$
- 'Vf' : float
  Maximum circular velocity in units of $[\text{km s}^{-1}]$
- 'Inclination' : float
  Inclination of galaxy in units of $[\text{degrees}]$
- 'Quality' : int
  Quality flag of galaxy

5.5.3.2 df

`galaxy.galaxy.df`

Pandas dataframe
Dataframe that contains data from SPARC catalog.

Contains data from DM_halo_models/pyfiles/data/MassModels_Lelli2016c.txt

5.5.3.3 df1

`galaxy.galaxy.df1`

Pandas dataframe
Dataframe that contains data from SPARC catalog.

Contains data from DM_halo_models/pyfiles/data/SPARC_Lelli2016c.txt

The documentation for this class was generated from the following file:

- pyfiles/data_models/galaxy.py

5.6 model_fit.grar_fit Class Reference

The class containing the fitting procedure for the gravitational acceleration relation.

Public Member Functions

- def __init__ (self, model, ULDM_fits=False, fit_dict_in=fitting_dict_in)
  Define the constructor of the grar_fit class.
- def g_tot (self, params, data)
  Define the total radial gravitational acceleration of the galaxy.
- def g_bar (self, params, data)
  Define the radial gravitational acceleration due to baryonic matter.
- def g_rar (self, gdag, params, data)
  Define the radial gravitational acceleration relation.
- def grar_model (self, gbarin, gdag)
  Define the general model of the radial gravitational acceleration relation.
- def fit (self)
  Define the fit result for the modeled gravitational acceleration relation.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- **ULDM_fits**
  bool
  True if performing fits for the ULDM models

5.6.1 Detailed Description

The class containing the fitting procedure for the gravitational acceleration relation.

See See Stacy S. McGaugh, Federico Lelli, and James M. Schombert Phys. Rev. Lett. 117, 201101 : https://doi.org/10.1103/PhysRevLett.117.201101 for more information

5.6.2 Constructor & Destructor Documentation

5.6.2.1 __init__()

```
def model_fit.grar_fit.__init__ (
              self,
              model,
              ULDM_fits = False,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the grar_fit class.

This defines the constructor of the grar_fit class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW • psi_single • psi_multi
ULDM_fits	(optional) bool True if performing fits for the ULDM models
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.6.3 Member Function Documentation

5.6.3.1 fit()

```
def model_fit.grar_fit.fit (
              self )
```

Define the fit result for the modeled gravitational acceleration relation.

This defines the fit result for the modeled gravitational acceleration relation.

Parameters

self	object pointer

Returns

lmfit.Minimizer.minimize fit result
Fit results contain the optimized parameters and several goodness-of-fit statistics.

5.6.3.2 g_bar()

```
def model_fit.grar_fit.g_bar (
              self,
              params,
              data )
```

Define the radial gravitational acceleration due to baryonic matter.

This defines the radial gravitational acceleration due to baryonic matter assuming the given model parameters and galaxy data. The radial gravitational acceleration due to baryonic matter is given by,

$$g_{\text{bar}}(r) = V_{\text{bar}}(r)^2/r,$$

where,

$$V_{\text{bar}}(r) = \sqrt{|V_{\text{gas}}(r)|\, V_{\text{gas}}(r) + \tilde{\Upsilon}_{\text{disk}}\, |V_{\text{disk}}(r)|\, V_{\text{disk}}(r) + \tilde{\Upsilon}_{\text{bulge}}\, \left|V_{\text{bulge}}(r)\right|\, V_{\text{bulge}}(r)}.$$

Here, $V_{\text{gas}}(r)$, $V_{\text{disk}}(r)$, and $V_{\text{bulge}}(r)$ can be found in galaxy.galaxy.data. Finally, $\tilde{\Upsilon}_{\text{disk}} =$ MLD and $\tilde{\Upsilon}_{\text{bulge}} =$ MLB in cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models).

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.

Returns

ndarray[N]
Numpy array of N values of the radial gravitational acceleration due to baryons in units of m s^{-2}.

5.6.3.3 g_rar()

```
def model_fit.grar_fit.g_rar (
              self,
              gdag,
              params,
              data )
```

Define the radial gravitational acceleration relation.

This defines the radial gravitational acceleration relation assuming the given relation constant, gdag, model parameters and galaxy data. The radial gravitational acceleration relation is given by,

$$g_{\text{tot}}(r) = \frac{g_{\text{bar}}(r)}{1 - \exp\left(-\sqrt{g_{\text{bar}}(r)/g_\dagger}\right)},$$

where $g_{\text{bar}}(r) =$ model_fit.grar_fit.g_bar and $g_\dagger$ is some constant.

Parameters

self	object pointer
gdag	float Constant in the gravitational acceleration relation. Must be in units of m s^{-2}.
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.

Returns

ndarray[N]
Numpy array of N values of the total radial gravitational acceleration relation in units of m s^{-2}.

5.6.3.4 g_tot()

```
def model_fit.grar_fit.g_tot (
              self,
              params,
              data )
```

Define the total radial gravitational acceleration of the galaxy.

This defines the total radial gravitational acceleration of the galaxy assuming the given model parameters and galaxy data. The total radial gravitational acceleration is given by,

$$g_{\text{tot}}(r) = V_{\text{tot}}(r)^2/r,$$

where $V_{\text{tot}}(r) =$ model_fit.model_fit.velocity_tot.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.

Returns

ndarray[N]
Numpy array of N values of the total radial gravitational acceleration in units of m s^{-2}.

5.6.3.5 grar_model()

```
def model_fit.grar_fit.grar_model (
              self,
              gbarin,
              gdag )
```

Define the general model of the radial gravitational acceleration relation.

This defines the general model of the radial gravitational acceleration relation for the model parameters, gbarin and gdag. This is the form of the radial gravitational acceleration relation to be used in fitting and is given by,

$$g_{\text{tot}}(r) = \frac{g_{\text{bar}}(r)}{1 - \exp\left(-\sqrt{g_{\text{bar}}(r)/g_\dagger}\right)}.$$

Here, the fit parameter $g_\dagger$ can be found by fitting this model using $g_{\text{tot}}(r) =$ model_fit.grar_fit.g_tot and $g_{\text{bar}} =$ model_fit.grar_fit.g_bar.

Parameters

self	object pointer
gbarin	ndarray[N] Numpy array of N values of baryonic gravitational accerlation in units of m s^{-2}.
gdag	ndarray[N] Constant in the gravitational acceleration relation in units of m s^{-2}.

Returns

ndarray[N]
Numpy array of N values of the total modeled radial gravitational acceleration in units of m s^{-2}.

The documentation for this class was generated from the following file:

- pyfiles/fitting/model_fit.py

5.7 cdm_funcs.halo Class Reference

The class containing mass functions for the CDM halo models.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)
  Define the constructor of the halo class.
- def mass (self, params, r)
  Define the mass profile of the CDM galactic structure.
- def Mvir (self, params)
  Define the total galactic DM halo mass.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- mass_num
  int
  Used to differentiate between soliton 1 and soliton 2 in double flavored ULDM models.
- **base_funcs_in**
  cdm_funcs.base_funcs instance
  Instance of the cdm_funcs.base_funcs class

5.7.1 Detailed Description

The class containing mass functions for the CDM halo models.

5.7.2 Constructor & Destructor Documentation

5.7.2.1 __init__()

```
def cdm_funcs.halo.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the halo class.

This defines the constructor of the halo class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.7.3 Member Function Documentation

5.7.3.1 mass()

```
def cdm_funcs.halo.mass (
              self,
              params,
              r )
```

Define the mass profile of the CDM galactic structure.

This defines the mass of the CDM galactic structure at a given radius given the model parameters. The mass profile depends on the model assumed. The mass profile for each model is given by :

- Burkert,

$$M_{\text{Burkert}}(r) = \pi\,\rho_c\,r_c^3\left\{-2\arctan(x_c(r)) + \log\left[(1+x_c(r))^2\left(1+x_c(r)^2\right)\right]\right\},$$

- DC14,

$$M_{\text{DC14}}(r) = 4\pi\,\rho_c\,r_c^3\,x_c(r)^{3-\gamma}\,{}_2\text{F}_1\left(\frac{3-\gamma}{\alpha}, \frac{\beta-\gamma}{\alpha}, \frac{3+\alpha-\gamma}{\alpha}, -x_c(r)^\alpha\right)(3-\gamma)^{-1},$$

where,

$$\alpha = 2.94 - \log_{10}\left[\left(10^{X+2.33}\right)^{-1.08} + \left(10^{X+2.33}\right)^{2.99}\right],$$

$$\beta = 4.23 + 1.34X + 0.26X^2,$$

$$\gamma = -0.06 - \log_{10}\left[\left(10^{X+2.56}\right)^{-0.68} + 10^{X+2.56}\right].$$

Here, $X =$ cdm_funcs.base_funcs.mass_frac_dc14, and ${}_2\mathrm{F}_1 =$ Gaussian hypergeometric function.

- Einasto,

$$M_{\mathrm{Einasto}}(r) = 4\pi\,\rho_c\,r_c^3\exp\left(2/\alpha\right)\left(2/\alpha\right)^{-3/\alpha}\,\Gamma\left(3/\alpha, 2x_c^{\alpha}/\alpha\right)/\alpha,$$

where $\alpha =$ alpha in cdm_params.halo.params and $\Gamma\left(a,x\right) =$ incomplete Gamma function.

- NFW,

$$M_{\mathrm{NFW}}(r) = 4\pi\,\rho_c\,r_c^3\left[\ln(1+x_c(r)) - \frac{x_c(r)}{1+x_c(r)}\right].$$

For each of the above masses, $\rho_c =$ cdm_funcs.base_funcs.rhoc, $r_c =$ cdm_funcs.base_funcs.rc, and $x_c(r) =$ cdm_funcs.base_funcs.xc.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N]
Numpy array of N mass values in units of solar mass.

5.7.3.2 Mvir()

```
def cdm_funcs.halo.Mvir (
              self,
              params )
```

Define the total galactic DM halo mass.

This defines the total mass of the galactic DM given the model parameters. The total galactic DM mass is given by,

$$M_{200} = \sqrt{\frac{3}{2\pi\rho_{\mathrm{crit}}}}\frac{M_P^3 V_{200}^3}{20},$$

where $\rho_{\mathrm{crit}} =$ rhocrit and M_P = MP in constants.standard, while $V_{200} =$ v200 in cdm_params.halo.params.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.

Returns

float
Total mass of the galactic DM in units of solar mass.

5.7.4 Member Data Documentation

5.7.4.1 mass_num

`cdm_funcs.halo.mass_num`

int
Used to differentiate between soliton 1 and soliton 2 in double flavored ULDM models.

Outer CDM halo is halo 1 if mass_num = 0 and is halo 2 if mass_num = 1.

The documentation for this class was generated from the following file:

- pyfiles/models/cdm/cdm_funcs.py

5.8 cdm_params.halo Class Reference

The class containing all parameters for fitting the CDM halo models.

Public Member Functions

- def __init__ (self, model, data, fit_dict_in=fitting_dict_in)
  Define the constructor of the halo class.
- def params (self)
  Define the parameters to be assumed.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- data
  dictionary
  Dictionary containing all the necessary data for a given galaxy.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- params_vals

dictionary
Dictionary of variable values to be used during fitting.

- **c200**

  list
  Values for c200 in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **v200**

  list
  Values for V200 in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **v200fac**

  list
  Values for v200fac in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **MLD**

  list
  Values for MLd in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **MLB**

  list
  Values for MLb in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **alpha**

  list
  Values for alpha in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals

5.8.1 Detailed Description

The class containing all parameters for fitting the CDM halo models.

This class contains all parameters necessary to perform the fitting procedures for all CDM halo models.

5.8.2 Constructor & Destructor Documentation

5.8.2.1 __init__()

```
def cdm_params.halo.__init__ (
              self,
              model,
              data,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the halo class.

This defines the constructor of the halo class.

Parameters

self	object pointer

Parameters

model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.8.3 Member Function Documentation

5.8.3.1 params()

```
def cdm_params.halo.params (
              self )
```

Define the parameters to be assumed.

This defines the parameters to be assumed during the fitting procedures for all CDM halos.

Parameters

self	object pointer

Returns

lmfit.Parameters instance Instance of the lmfit.Parameters class

5.8.4 Member Data Documentation

5.8.4.1 data

```
cdm_params.halo.data
```

dictionary
Dictionary containing all the necessary data for a given galaxy.

This can be created using galaxy.galaxy.

5.8.4.2 params_vals

`cdm_params.halo.params_vals`

dictionary
Dictionary of variable values to be used during fitting.

Instance of constants.fitting_dict.params_vals

The documentation for this class was generated from the following file:

- pyfiles/models/cdm/cdm_params.py

5.9 halo.halo Class Reference

The class containing definitions to describe a DM halo.

Public Member Functions

- def __init__ (self, model, data, fit_dict_in=fitting_dict_in)
  Define the constructor of the halo class.
- def mass (self, params, r)
  Define the DM halo mass within some radius assuming the chosen model.
- def velocity (self, params, r)
  Define the total circular velocity of the DM halo at some radius assuming the chosen model.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- data
  dictionary
  Dictionary containing all the necessary data for a given galaxy.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- params_vals
  dictionary
  Dictionary of variable values to be used during fitting.
- params
  lfmit.Parameters instance
  Instance of the lmfit.Parameters class All model parameters contained here.
- halo_init
  Can be :

5.9.1 Detailed Description

The class containing definitions to describe a DM halo.

5.9.2 Constructor & Destructor Documentation

5.9.2.1 __init__()

```
def halo.halo.__init__ (
              self,
              model,
              data,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the halo class.

This defines the constructor of the halo class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW • psi_single • psi_multi
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.9.3 Member Function Documentation

5.9.3.1 mass()

```
def halo.halo.mass (
              self,
```

```
                params,
                r )
```

Define the DM halo mass within some radius assuming the chosen model.

This defines the total DM halo mass within the given radius assuming the given model parameters. Particular equations for mass depends on chosen model.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N]
Numpy array of N mass values in units of solar mass.

5.9.3.2 velocity()

```
def halo.halo.velocity (
                self,
                params,
                r )
```

Define the total circular velocity of the DM halo at some radius assuming the chosen model.

This calculates the total circular velocity of the DM halo at the given radius assuming the given model parameters. Total circular velocity is given by,

$$V(r) = \sqrt{\frac{M(r)}{M_P^2\, r}}$$

where $M(r)$ = halo.halo.mass, and M_P is the Planck mass (see constants.standard)

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N]
Numpy array of N circular velocity values in units of km/s.

5.9.4 Member Data Documentation

5.9.4.1 data

`halo.halo.data`

dictionary
Dictionary containing all the necessary data for a given galaxy.

This can be created using galaxy.galaxy.

5.9.4.2 halo_init

`halo.halo.halo_init`

Can be :

- cdm_funcs.halo instance
- alp_funcs.soliton instance
  Instance of cdm_funcs.halo or alp_funcs.soliton class depending on chosen model.

5.9.4.3 params

`halo.halo.params`

lfmit.Parameters instance
Instance of the lmfit.Parameters class All model parameters contained here.

This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)

5.9.4.4 params_vals

`halo.halo.params_vals`

dictionary
Dictionary of variable values to be used during fitting.

Instance of constants.fitting_dict.params_vals

The documentation for this class was generated from the following file:

- pyfiles/data_models/halo.py

5.10 model_fit.model_fit Class Reference

The class containing the fitting procedure assuming a chosen halo model for a chosen galaxy.

Public Member Functions

- def __init__ (self, halo, ULDM_fits=False)

  Define the constructor of the model_fit class.
- def velocity_tot (self, params, r, data)

  Define the total circular velocity of a galaxy at some radius for the given halo model.
- def residual (self, params, r, data)

  Define the residual of the modeled total circular velocity from the data for a chosen galaxy.
- def fit (self, params, data, calc_covar_in=True)

  Define the fit result for the modeled total circular velocity for a galaxy.
- def bic (self, fit)

  Define the BIC value for a given fit.
- def plot (self, halo, gal, axs)

  Define the plot of the model of the total circular velocity for a galaxy.

Public Attributes

- **halo**

  halo.halo instance
  Instance of the halo.halo class.
- args_opts

  dictionary
  Dictionary of rules to be assumed during fitting.
- **ULDM_fits**

  bool
  True if performing fits for the ULDM models

5.10.1 Detailed Description

The class containing the fitting procedure assuming a chosen halo model for a chosen galaxy.

5.10.2 Constructor & Destructor Documentation

5.10.2.1 __init__()

```
def model_fit.model_fit.__init__ (
              self,
              halo,
              ULDM_fits = False )
```

Define the constructor of the model_fit class.

This defines the constructor of the model_fit class.

Parameters

self	object pointer
halo	halo.halo instance Instance of the halo.halo class. Assumed halo model and proper definitions contained in halo class instance.
ULDM_fits	(optional) bool True if performing fits for the ULDM models

5.10.3 Member Function Documentation

5.10.3.1 bic()

```
def model_fit.model_fit.bic (
              self,
              fit )
```

Define the BIC value for a given fit.

This defines the BIC value for a given fit.

Parameters

self	object pointer
fit	model_fit.model_fit.fit instance Instance of the model_fit.model_fit.fit function.

Returns

float
BIC value for a given fit.

5.10.3.2 fit()

```
def model_fit.model_fit.fit (
              self,
              params,
              data,
              calc_covar_in = True )
```

Define the fit result for the modeled total circular velocity for a galaxy.

This defines the fit result for the modeled total circular velocity assuming the given model parameters and galaxy data.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.
calc_covar_in	(optional) bool True(False) if covariance matrix is to be(is not to be) calculated in fitting procedure

Returns

lmfit.Minimizer.minimize fit result
Fit result contains the optimized parameters and several goodness-of-fit statistics.

5.10.3.3 plot()

```
def model_fit.model_fit.plot (
              self,
              halo,
              gal,
              axs )
```

Define the plot of the model of the total circular velocity for a galaxy.

This defines the plot of the model of total circular velocity for a galaxy assuming the best fit parameters.

Parameters

self	object pointer
halo	halo.halo instance Instance of the halo.halo class. Assumed halo model and proper definitions contained in halo class instance.
gal	galaxy.galaxy instance Instance of the galaxy.galaxy class. Contains all necessary galaxy data.
axs	matplotlib axis Axis in which matplotlib plot will be contained

Returns

matplotlib plot
Plot of the circular velocity of the different galaxy contributions including the DM halo given the assumed model and best fit parameters.

5.10.3.4 residual()

```
def model_fit.model_fit.residual (
                self,
                params,
                r,
                data )
```

Define the residual of the modeled total circular velocity from the data for a chosen galaxy.

This calculates the residual of the modeled total circular velocity from the data for a galaxy at a given radius assuming the given model parameters and galaxy data. This is used as the fitting residual in the lmfit.Minimizer class and is given by,

$$\text{res} = \frac{V_{\text{obs}}(r) - V_{\text{model}}(r)}{V_{\text{obs,err}}(r)},$$

where $V_{\text{obs}}(r)$ and $V_{\text{obs,err}}(r)$ can be found in galaxy.galaxy.data and $V_{\text{model}}(r) =$ model_fit.model_fit.velocity_tot.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
r	ndarray[N] Numpy array of N radii values in units of kpc.
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.

Returns

ndarray[N]
Numpy array of N residual values.

5.10.3.5 velocity_tot()

```
def model_fit.model_fit.velocity_tot (
                self,
                params,
                r,
                data )
```

Define the total circular velocity of a galaxy at some radius for the given halo model.

This defines the total circular velocity of a galaxy at a given radius assuming the given model parameters and galaxy data. The total circular velocity is given by,

$$V(r) = \sqrt{V_{\text{bar}}(r)^2 + V_{\text{DM}}(r)^2},$$

where,

$$V_{\text{bar}}(r) = \sqrt{\left|V_{\text{gas}}(r)\right| V_{\text{gas}}(r) + \tilde{\Upsilon}_{\text{disk}} \left|V_{\text{disk}}(r)\right| V_{\text{disk}}(r) + \tilde{\Upsilon}_{\text{bulge}} \left|V_{\text{bulge}}(r)\right| V_{\text{bulge}}(r)}.$$

Here, $V_{\text{DM}}(r) =$ halo.halo.velocity, $V_{\text{gas}}(r)$, $V_{\text{disk}}(r)$, and $V_{\text{bulge}}(r)$ can be found in galaxy.galaxy.data. Finally, $\tilde{\Upsilon}_{\text{disk}} =$ MLD and $\tilde{\Upsilon}_{\text{bulge}} =$ MLB in cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models).

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params (for CDM models) or alp_params.soliton.params (for ULDM models)
r	ndarray[N] Numpy array of N radii values in units of kpc.
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.

Returns

ndarray[N]
Numpy array of N circular velocity values in units of km/s.

5.10.4 Member Data Documentation

5.10.4.1 args_opts

`model_fit.model_fit.args_opts`

dictionary
Dictionary of rules to be assumed during fitting.

Instance of constants.fitting_dict.args_opts

The documentation for this class was generated from the following file:

- pyfiles/fitting/model_fit.py

5.11 results.plots Class Reference

The class to obtain rotation curve plots for given galaxies.

Public Member Functions

- def __init__ (self)

  Define the constructor of the plots class.
- def rotcurves_CDM_all (self, galaxies, fit_dict_in=fitting_dict_in, size=(20, 20), save_file=None)

  Define the plot of the rotation curves for given galaxies and for all CDM models.
- def rotcurves_psi_all (self, galaxies, fit_dict_in=fitting_dict_in, size=(20, 20), save_file=None)

  Define the plot of the rotation curves for given galaxies and for the single ULDM and double ULDM models.
- def rotcurves_CDM_check (self, galaxies, fit_dict_in=fitting_dict_in, size=(20, 20))

  Define the plot of the rotation curves for given galaxies and for all CDM models.
- def rotcurves_DC14_check (self, galaxies, fit_dict_in=fitting_dict_in, size=(20, 20))

  Define the plot of the rotation curves for given galaxies and for the DC14 and NFW models.
- def rotcurves_Einasto_check (self, galaxies, fit_dict_in=fitting_dict_in, size=(20, 20))

  Define the plot of the rotation curves for given galaxies and for the Einasto and NFW models.
- def BIC_diffs_CDM (self, dict_in, bins=250, lwin=3, save_file=None)

  Define the plot of the BIC differences for different priors using CDM models.
- def BIC_CDM (self, dict_in, bins=250, lwin=3, save_file=None)

  Define the plot of the BIC differences for the CDM models.
- def chi_CDM (self, dict_in, splt=50, lwin=3, save_file=None)

  Define the plot of the reduced chi-square comparisons for the CDM models.
- def chi_dist_CDM (self, dict_in, bins=50, save_file=None)

  Define the plot of the reduced chi-square distributions for the CDM models.
- def params_dist_CDM (self, dict_in, bins=50, save_file=None)

  Define the plot of the parameter distributions for the CDM models.
- def MLd_degen_CDM (self, save_file=None)

  Define the plot of the reduced chi-square contours for the NFW model.
- def relations_CDM (self, dict_in, save_file=None)

  Define the plot of the empirical relations for the CDM models.
- def BIC_psi_mfree (self, psis_dict_in, psim_dict_in, CDM_dict_in, splt=50, lwin=3, save_file=None)

  Define the plot of the BIC differences between the ULDM (particle mass free) and Einasto models.
- def chi_psi_mfree (self, psis_dict_in, psim_dict_in, CDM_dict_in, splt=50, lwin=3, save_file=None)

  Define the plot of the reduced chi-square for the ULDM (particle mass free) and Einasto models.
- def Msol_psi_mfree (self, psis_dict_in, psim_dict_in, save_file=None)

  Define the plot of the soliton-halo (SH) relation for the ULDM (particle mass free) models.
- def chi_dist_psi_mfree (self, psis_dict_in, psim_dict_in, bins=50, save_file=None)

  Define the plot of the reduced chi-square distributions for the ULDM (particle mass free) models.
- def params_dist_psi_mfree (self, psis_dict_in, psim_dict_in, bins=50, save_file=None)

  Define the plot of the parameter distributions for the ULDM (particle mass free) models.
- def relations_psi_mfree (self, psis_dict_in, psim_dict_in, save_file=None)

  Define the plot of the empirical relations for the ULDM (particle mass free) models.
- def chi_psi_mfix (self, psis_dict_in, psim_dict_in, CDM_dict_in, lwin=4, fit_dict_in=fitting_dict_in, save_↩ file=None)

  Define the plot of the reduced chi-square for the ULDM (particle mass fixed and scanned) and Einasto models.
- def chi_gal_psi_mfix (self, psis_dict_in, CDM_dict_in, lwin=4, fit_dict_in=fitting_dict_in, save_file=None)

  Define the plot of the cumulative reduced chi-square for single flavor ULDM (particle mass fixed and scanned) and Einasto models.
- def Msol_psi_mfix (self, psis_dict_in, psim_dict_in, bins=50, fit_dict_in=fitting_dict_in, save_file=None)

  Define the plot of the soliton-halo (SH) relation for the ULDM (particle mass fixed and scanned) models.
- def BIC_psi_mfix_ex (self, psis_dict_in, psim_dict_in, CDM_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), bins=50, lwin=3, save_file=None)

  Define the plot of the BIC differences distributions between the ULDM (particle mass fixed) and Einasto models.

- def chi_psi_mfix_ex (self, psis_dict_in, psim_dict_in, CDM_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), splt=50, lwin=3, save_file=None)

  Define the plot of the reduced chi-square for the ULDM (particle mass fixed) and Einasto models.
- def Msol_psi_mfix_ex (self, psis_dict_in, psim_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), save_↩ file=None)

  Define the plot of the soliton-halo (SH) for the ULDM (particle mass fixed) and Einasto models.
- def chi_dist_psi_mfix_ex (self, psis_dict_in, psim_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), bins=50, save_file=None)

  Define the plot of the reduced chi-square distributions for the ULDM (particle mass fixed) models.
- def params_dist_psi_mfix_ex (self, psis_dict_in, psim_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), bins=50, save_file=None)

  Define the plot of the parameter distributions for the ULDM (particle mass fixed) models.
- def relations_psi_mfix_ex (self, psis_dict_in, psim_dict_in, m22_ex=10 ∗∗(1.5), m22_2_ex=10 ∗∗(1.8), save_file=None)

  Define the plot of the empirical relations for the ULDM (particle mass fixed) models.
- def chi_dist_CDM_checks (self, dict_in, bins=50)

  Define the plot of the reduced chi-square distributions for the CDM models (checks).
- def params_dist_CDM_checks (self, dict_in, bins=50)

  Define the plot of the parameter distributions for the CDM models (checks).
- def chi_dist_DC14_checks (self, dict_in, bins=50)

  Define the plot of the reduced chi-square distributions for the DC14 and NFW models (checks).
- def params_dist_DC14_checks (self, dict_in, bins=50)

  Define the plot of the parameter distributions for the DC14 and NFW models (checks).
- def chi_box_Einasto_checks (self, dict_in)

  Define the plot of the reduced chi-square for the Einasto and NFW models (checks).
- def chi_dist_Einasto_checks (self, dict_in, bins=50)

  Define the plot of the reduced chi-square distributions for the Einasto and NFW models (checks).
- def params_dist_Einasto_checks (self, dict_in, bins=50)

  Define the plot of the parameters distributions for the Einasto and NFW models (checks).
- def params_scatter_Einasto_checks (self, dict_in, fit_dict_in=fitting_dict_in)

  Define the plot of the halo parameters for the Einasto model (checks).

Public Attributes

- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **size**

  tuple
  Size of figure.

5.11.1 Detailed Description

The class to obtain rotation curve plots for given galaxies.

5.11.2 Constructor & Destructor Documentation

5.11.2.1 __init__()

```
def results.plots.__init__ (
              self )
```

Define the constructor of the plots class.

This defines the constructor of the plots class.

Parameters

self	object pointer
galaxies	str[N] List of galaxy names (or 'random') to plot the given galaxies (or random galaxies) in the SPARC catalog.

5.11.3 Member Function Documentation

5.11.3.1 BIC_CDM()

```
def results.plots.BIC_CDM (
              self,
              dict_in,
              bins = 250,
              lwin = 3,
              save_file = None )
```

Define the plot of the BIC differences for the CDM models.

This defines the plot of the differences in BIC between each CDM model analyzed.

Parameters

self	object pointer
dict_in	dictionary See fits_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_all.fit.
bins	(optional) int Number of bins for histogram plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,3]
Figure of the differences in BIC between each CDM model.

5.11.3.2 BIC_diffs_CDM()

```
def results.plots.BIC_diffs_CDM (
              self,
              dict_in,
              bins = 250,
              lwin = 3,
              save_file = None )
```

Define the plot of the BIC differences for different priors using CDM models.

This defines the plot of the differences in BIC for different prior cases for each CDM model analyzed.

Parameters

self	object pointer
dict_in	dictionary See fits_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'uni_priors' : {...} • 'c200_priors_check' : {...} • 'v200_priors_check' : {...} • 'MLd_priors_check' : {...} • 'MLb_priors_check' : {...} }, where each {...} must be of the form: {...} = • 'Burkert' : {....} • 'DC14' : {....} • 'Einasto' : {....} • 'NFW' : {....} Finally, each {....} must be an instance of results.results_CDM_all.fit.
bins	(optional) int Number of bins for histogram plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the BIC differences for different priors using all CDM models.

5.11.3.3 BIC_psi_mfix_ex()

```
def results.plots.BIC_psi_mfix_ex (
                self,
                psis_dict_in,
                psim_dict_in,
                CDM_dict_in,
                m22_ex = 10**(1.5),
                m22_2_ex = 10**(1.8),
                bins = 50,
                lwin = 3,
                save_file = None )
```

Define the plot of the BIC differences distributions between the ULDM (particle mass fixed) and Einasto models.

This defines the plot of the distributions of the differences in BIC between each of the ULDM (particle mass fixed) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
lwin	(optional) int Width of lines for plot.
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the BIC differences distributions between each ULDM (particle mass fixed) model and the Einasto model.

5.11.3.4 BIC_psi_mfree()

```
def results.plots.BIC_psi_mfree (
              self,
              psis_dict_in,
              psim_dict_in,
              CDM_dict_in,
              splt = 50,
              lwin = 3,
              save_file = None )
```

Define the plot of the BIC differences between the ULDM (particle mass free) and Einasto models.

This defines the plot of the differences in BIC vs. ULDM particle mass between each of the ULDM (particle mass free) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
splt	(optional) int Point size for plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[3,2]
Figure of the differences in BIC vs. particle mass between each ULDM (particle mass free) model and the Einasto model.

5.11.3.5 chi_box_Einasto_checks()

```
def results.plots.chi_box_Einasto_checks (
              self,
              dict_in )
```

Define the plot of the reduced chi-square for the Einasto and NFW models (checks).

This defines the plot of the reduced chi-square comparison between the Einasto and NFW models (checks).

Parameters

self	object pointer
dict_in	dictionary See checks_Einasto_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_Einasto_check.fit.

Returns

matplotlib.pyplot.figure[1,2]
Figure of the reduced chi-square comparisons between the Einasto and NFW models (checks).

5.11.3.6 chi_CDM()

```
def results.plots.chi_CDM (
              self,
              dict_in,
              splt = 50,
              lwin = 3,
              save_file = None )
```

Define the plot of the reduced chi-square comparisons for the CDM models.

This defines the plot of the reduced chi-square comparisons between each CDM model analyzed.

Parameters

self	object pointer
dict_in	dictionary See fits_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_all.fit.

Parameters

splt	(optional) int Point size for plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,3]
Figure of the reduced chi-square comparisons between each CDM model.

5.11.3.7 chi_dist_CDM()

```
def results.plots.chi_dist_CDM (
              self,
              dict_in,
              bins = 50,
              save_file = None )
```

Define the plot of the reduced chi-square distributions for the CDM models.

This defines the plot of the reduced chi-square distribution for each CDM model analyzed.

Parameters

self	object pointer
dict_in	dictionary See fits_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_all.fit.
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square distributions for each CDM model.

5.11.3.8 chi_dist_CDM_checks()

```
def results.plots.chi_dist_CDM_checks (
               self,
               dict_in,
               bins = 50 )
```

Define the plot of the reduced chi-square distributions for the CDM models (checks).

This defines the plot of the reduced chi-square distribution for each CDM model analyzed (checks).

Parameters

self	object pointer
dict_in	dictionary See checks_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[1,4]
Figure of the reduced chi-square distributions for each CDM model (checks).

5.11.3.9 chi_dist_DC14_checks()

```
def results.plots.chi_dist_DC14_checks (
               self,
               dict_in,
               bins = 50 )
```

Define the plot of the reduced chi-square distributions for the DC14 and NFW models (checks).

This defines the plot of the reduced chi-square distribution for the DC14 and NFW models (checks).

Parameters

self	object pointer

Parameters

dict↩ _in	dictionary See checks_DC14_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'DC14' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_DC14_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[1,2]
Figure of the reduced chi-square distributions for the DC14 and NFW models (checks).

5.11.3.10 chi_dist_Einasto_checks()

```
def results.plots.chi_dist_Einasto_checks (
              self,
              dict_in,
              bins = 50 )
```

Define the plot of the reduced chi-square distributions for the Einasto and NFW models (checks).

This defines the plot of the reduced chi-square distribution for the Einasto and NFW models (checks).

Parameters

self	object pointer
dict↩ _in	dictionary See checks_Einasto_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_Einasto_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[1,2]
Figure of the reduced chi-square distributions for the Einasto and NFW models (checks).

5.11.3.11 chi_dist_psi_mfix_ex()

```
def results.plots.chi_dist_psi_mfix_ex (
              self,
              psis_dict_in,
              psim_dict_in,
              m22_ex = 10**(1.5),
              m22_2_ex = 10**(1.8),
              bins = 50,
              save_file = None )
```

Define the plot of the reduced chi-square distributions for the ULDM (particle mass fixed) models.

This defines the plot of the reduced chi-square distributions for each of the ULDM (particle mass fixed) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square distributions for each of the ULDM (particle mass fixed) models.

5.11.3.12 chi_dist_psi_mfree()

```
def results.plots.chi_dist_psi_mfree (
              self,
```

```
                psis_dict_in,
                psim_dict_in,
                bins = 50,
                save_file = None )
```

Define the plot of the reduced chi-square distributions for the ULDM (particle mass free) models.

This defines the plot of the reduced chi-square distributions for each of the ULDM (particle mass free) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square distributions for each of the ULDM (particle mass free) models.

5.11.3.13 chi_gal_psi_mfix()

```
def results.plots.chi_gal_psi_mfix (
                self,
                psis_dict_in,
                CDM_dict_in,
                lwin = 4,
                fit_dict_in = fitting_dict_in,
                save_file = None )
```

Define the plot of the cumulative reduced chi-square for single flavor ULDM (particle mass fixed and scanned) and Einasto models.

This defines the plot of the cumulative reduced chi-square comparisons vs. galaxy between each of the single flavor ULDM (particle mass fixed and scanned) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
lwin	(optional) int Width of lines for plot.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains the mass ranges to be used.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,1]
Figure of the cumulative reduced chi-square vs. galaxy for each of the single flavor ULDM (particle mass fixed and scanned) models and the Einasto model.

5.11.3.14 chi_psi_mfix()

```
def results.plots.chi_psi_mfix (
              self,
              psis_dict_in,
              psim_dict_in,
              CDM_dict_in,
              lwin = 4,
              fit_dict_in = fitting_dict_in,
              save_file = None )
```

Define the plot of the reduced chi-square for the ULDM (particle mass fixed and scanned) and Einasto models.

This defines the plot of the reduced chi-square comparisons vs. particle mass between each of the ULDM (particle mass fixed and scanned) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
lwin	(optional) int Width of lines for plot.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains the mass ranges to be used.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square comparisons vs. particle mass between each of the ULDM (particle mass fixed and scanned) models and the Einasto model.

5.11.3.15 chi_psi_mfix_ex()

```
def results.plots.chi_psi_mfix_ex (
                self,
                psis_dict_in,
                psim_dict_in,
                CDM_dict_in,
                m22_ex = 10**(1.5),
                m22_2_ex = 10**(1.8),
                splt = 50,
                lwin = 3,
                save_file = None )
```

Define the plot of the reduced chi-square for the ULDM (particle mass fixed) and Einasto models.

This defines the plot of the reduced chi-square comparisons between each of the ULDM (particle mass fixed) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
splt	(optional) int Point size for plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square comparisons between each ULDM (particle mass fixed) model and the Einasto model.

5.11.3.16 chi_psi_mfree()

```
def results.plots.chi_psi_mfree (
              self,
              psis_dict_in,
              psim_dict_in,
              CDM_dict_in,
              splt = 50,
              lwin = 3,
              save_file = None )
```

Define the plot of the reduced chi-square for the ULDM (particle mass free) and Einasto models.

This defines the plot of the reduced chi-square comparisons between each of the ULDM (particle mass free) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
CDM_dict_in	dictionary The dictionary for the Einasto. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be an instance of results.results_CDM_all.fit.
splt	(optional) int Point size for plot.
lwin	(optional) int Width of lines for plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square comparisons between each of the ULDM (particle mass free) models and the Einasto model.

5.11.3.17 MLd_degen_CDM()

```
def results.plots.MLd_degen_CDM (
              self,
              save_file = None )
```

Define the plot of the reduced chi-square contours for the NFW model.

This defines the plot of the reduced chi-square contours in the $\tilde{\Upsilon}_d-c_{200}$ plane for the NFW model.

Parameters

self	object pointer
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[2,2]
Figure of the reduced chi-square contours in the $\tilde{\Upsilon}_d-c_{200}$ plane for the NFW model.

5.11.3.18 Msol_psi_mfix()

```
def results.plots.Msol_psi_mfix (
              self,
              psis_dict_in,
              psim_dict_in,
              bins = 50,
              fit_dict_in = fitting_dict_in,
              save_file = None )
```

Define the plot of the soliton-halo (SH) relation for the ULDM (particle mass fixed and scanned) models.

This defines the plot of the SH relation for each of the ULDM (particle mass fixed and scanned) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
bins	(optional) int Number of bins for histogram plot.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains the mass ranges to be used.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[3,2]
Figure of the SH relation for each of the ULDM (particle mass fixed and scanned) models.

5.11.3.19 Msol_psi_mfix_ex()

```
def results.plots.Msol_psi_mfix_ex (
              self,
              psis_dict_in,
              psim_dict_in,
              m22_ex = 10**(1.5),
              m22_2_ex = 10**(1.8),
              save_file = None )
```

Define the plot of the soliton-halo (SH) for the ULDM (particle mass fixed) and Einasto models.

This defines the plot of the SH relation for each of the ULDM (particle mass fixed) models and the Einasto model.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[3,2]
Figure of the SH relation for each ULDM (particle mass fixed) model and the Einasto model.

5.11.3.20 Msol_psi_mfree()

```
def results.plots.Msol_psi_mfree (
              self,
              psis_dict_in,
```

```
        psim_dict_in,
        save_file = None )
```

Define the plot of the soliton-halo (SH) relation for the ULDM (particle mass free) models.

This defines the plot of the SH relation vs. particle mass for each of the ULDM (particle mass free) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[3,2]
Figure of the SH relation vs. particle mass for each of the ULDM (particle mass free) models.

5.11.3.21 params_dist_CDM()

```
def results.plots.params_dist_CDM (
              self,
              dict_in,
              bins = 50,
              save_file = None )
```

Define the plot of the parameter distributions for the CDM models.

This defines the plot of the parameter distributions for each CDM model analyzed.

Parameters

self	object pointer
dict_in	dictionary See fits_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_all.fit.
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the parameter distributions for each CDM model.

5.11.3.22 params_dist_CDM_checks()

```
def results.plots.params_dist_CDM_checks (
              self,
              dict_in,
              bins = 50 )
```

Define the plot of the parameter distributions for the CDM models (checks).

This defines the plot of the parameter distributions for each CDM model analyzed (checks).

Parameters

self	object pointer
dict_in	dictionary See checks_CDM_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Burkert' : {...} • 'DC14' : {...} • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_CDM_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the parameter distributions for each CDM model (checks).

5.11.3.23 params_dist_DC14_checks()

```
def results.plots.params_dist_DC14_checks (
              self,
              dict_in,
              bins = 50 )
```

Define the plot of the parameter distributions for the DC14 and NFW models (checks).

This defines the plot of the parameter distributions for the DC14 and NFW models (checks).

Parameters

self	object pointer
dict_in	dictionary See checks_DC14_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'DC14' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_DC14_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[1,2]
Figure of the reduced chi-square distributions for the DC14 and NFW models (checks).

5.11.3.24 params_dist_Einasto_checks()

```
def results.plots.params_dist_Einasto_checks (
              self,
              dict_in,
              bins = 50 )
```

Define the plot of the parameters distributions for the Einasto and NFW models (checks).

This defines the plot of the parameter distributions for the Einasto and NFW models (checks).

Parameters

self	object pointer
dict_in	dictionary See checks_Einasto_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_Einasto_check.fit.
bins	(optional) int Number of bins for histogram plot.

Returns

matplotlib.pyplot.figure[3,2]
Figure of the parameter distributions for the Einasto and NFW models (checks).

5.11.3.25 params_dist_psi_mfix_ex()

```
def results.plots.params_dist_psi_mfix_ex (
              self,
              psis_dict_in,
              psim_dict_in,
              m22_ex = 10**(1.5),
              m22_2_ex = 10**(1.8),
              bins = 50,
              save_file = None )
```

Define the plot of the parameter distributions for the ULDM (particle mass fixed) models.

This defines the plot of the parameter distributions for each of the ULDM (particle mass fixed) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
bins	(optional) int Number of bins for histogram plot.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the parameter distributions for each of the ULDM (particle mass fixed) models.

5.11.3.26 params_dist_psi_mfree()

```
def results.plots.params_dist_psi_mfree (
              self,
```

```
                psis_dict_in,
                psim_dict_in,
                bins = 50,
                save_file = None )
```

Define the plot of the parameter distributions for the ULDM (particle mass free) models.

This defines the plot of the parameter distributions for each of the ULDM (particle mass free) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
bins	(optional) int Number of bins for histogram plot.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the parameter distributions for each of the ULDM (particle mass free) models.

5.11.3.27 params_scatter_Einasto_checks()

```
def results.plots.params_scatter_Einasto_checks (
                self,
                dict_in,
                fit_dict_in = fitting_dict_in )
```

Define the plot of the halo parameters for the Einasto model (checks).

This defines the plot of the halo parameters for the Einasto model (checks).

Parameters

self	object pointer

Parameters

dict_in	dictionary See checks_Einasto_NFW.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: dict_in = { • 'Einasto' : {...} • 'NFW' : {...} }, where each {...} must be an instance of results.results_Einasto_check.fit.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains the necessary parameters for fitting.

Returns

matplotlib.pyplot.figure[1,2]
Figure of the halo parameters for the Einasto model (checks).

5.11.3.28 relations_CDM()

```
def results.plots.relations_CDM (
              self,
              dict_in,
              save_file = None )
```

Define the plot of the empirical relations for the CDM models.

This defines the plot of the empirical relations for each CDM models analyzed.

Parameters

self	object pointer
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the empirical relations for each CDM model.

5.11.3.29 relations_psi_mfix_ex()

```
def results.plots.relations_psi_mfix_ex (
              self,
              psis_dict_in,
              psim_dict_in,
```

```
                m22_ex = 10**(1.5),
                m22_2_ex = 10**(1.8),
                save_file = None )
```

Define the plot of the empirical relations for the ULDM (particle mass fixed) models.

This defines the plot of the empirical relations for each of the ULDM (particle mass fixed) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
m22_ex	(optional) float Value for particle mass (single flavor models) or particle mass one (double flavor models).
m22_2_ex	(optional) float Value for particle mass two (double flavor models).
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the empirical relations for each of the ULDM (particle mass fixed) models.

5.11.3.30 relations_psi_mfree()

```
def results.plots.relations_psi_mfree (
                self,
                psis_dict_in,
                psim_dict_in,
                save_file = None )
```

Define the plot of the empirical relations for the ULDM (particle mass free) models.

This defines the plot of the empirical relations for each of the ULDM (particle mass free) models.

Parameters

self	object pointer
psis_dict_in	dictionary The dictionary for the single flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psis_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_single_all.fit.
psim_dict_in	dictionary The dictionary for the double flavor models. See fits_Einasto_all.ipynb for an example of formatting for the dictionary. Dictionary must be of the form: psim_dict_in = { • 'Summed' : {...} • 'Matched' : {...} }, where each {...} must be an instance of results.results_psi_multi_all.fit.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[4,4]
Figure of the empirical relations for each of the ULDM (particle mass free) models.

5.11.3.31 rotcurves_CDM_all()

```
def results.plots.rotcurves_CDM_all (
              self,
              galaxies,
              fit_dict_in = fitting_dict_in,
              size = (20,20),
              save_file = None )
```

Define the plot of the rotation curves for given galaxies and for all CDM models.

This defines the plot of the rotation curves for all given galaxies and for all CDM models.

Parameters

self	object pointer
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.
size	(optional) tuple Size of figure.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[N,4]
Figure of rotation curves for all given galaxies (N galaxies in total) and for all CDM halos (4 in total).

5.11.3.32 rotcurves_CDM_check()

```
def results.plots.rotcurves_CDM_check (
              self,
              galaxies,
              fit_dict_in = fitting_dict_in,
              size = (20,20) )
```

Define the plot of the rotation curves for given galaxies and for all CDM models.

This defines the plot of the rotation curves for all given galaxies and for all CDM models. For comparison with Pengfei Li et al 2020 ApJS 247 31 : https://doi.org/10.3847/1538-4365/ab700e

Parameters

self	object pointer
fit_dict↩ _in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.
size	(optional) tuple Size of figure.

Returns

matplotlib.pyplot.figure[N,4]
Figure of rotation curves for all given galaxies (N galaxies in total) and for all CDM halos (4 in total).

5.11.3.33 rotcurves_DC14_check()

```
def results.plots.rotcurves_DC14_check (
              self,
              galaxies,
              fit_dict_in = fitting_dict_in,
              size = (20,20) )
```

Define the plot of the rotation curves for given galaxies and for the DC14 and NFW models.

This defines the plot of the rotation curves for all given galaxies and for the DC14 and NFW models. For comparison with Monthly Notices of the Royal Astronomical Society, Volume 466, Issue 2, April 2017, Pages 1648–1668 : https://doi.org/10.1093/mnras/stw3101

Parameters

self	object pointer
fit_dict↩ _in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.
size	(optional) tuple Size of figure.

Returns

matplotlib.pyplot.figure[N,2]
Figure of rotation curves for all given galaxies (N galaxies in total) and for the DC14 and NFW models.

5.11.3.34 rotcurves_Einasto_check()

```
def results.plots.rotcurves_Einasto_check (
              self,
              galaxies,
              fit_dict_in = fitting_dict_in,
              size = (20,20) )
```

Define the plot of the rotation curves for given galaxies and for the Einasto and NFW models.

This defines the plot of the rotation curves for all given galaxies and for the Einasto and NFW models. For comparison with Nicolas Loizeau and Glennys R. Farrar 2021 ApJL 920 L10 : https://doi.org/10.↩ 3847/2041-8213/ac1bb7

Parameters

self	object pointer
fit_dict↩ _in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.
size	(optional) tuple Size of figure.

Returns

matplotlib.pyplot.figure[N,2]
Figure of rotation curves for all given galaxies (N galaxies in total) and for the Einasto and NFW models.

5.11.3.35 rotcurves_psi_all()

```
def results.plots.rotcurves_psi_all (
              self,
```

```
galaxies,
fit_dict_in = fitting_dict_in,
size = (20,20),
save_file = None )
```

Define the plot of the rotation curves for given galaxies and for the single ULDM and double ULDM models.

This defines the plot of the rotation curves for all given galaxies and for the single ULDM and double ULDM models.

Parameters

self	object pointer
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.
size	(optional) tuple Size of figure.
save_file	(optional) str File path to save plot. Default is None, in which case file is not saved.

Returns

matplotlib.pyplot.figure[N,2]
Figure of rotation curves for all given galaxies (N galaxies in total) and for the single and double ULDM (will correspond to either summed or matched model depending on value for constants.fitting_dict.sol_match).

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.12 results.results_CDM_all Class Reference

The class to obtain fit results for all CDM halos analyzed.

Public Member Functions

- def __init__ (self, model, ULDM_fits=False, fit_dict_in=fitting_dict_in)

  Define the constructor of the results_CDM_all class.
- def fit (self)

  Define the fit results.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**

  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- **matched**

  bool
  True if performing ULDM fits and ULDM matched models
- **ULDM_fits**

  bool
  True if performing ULDM fits

5.12.1 Detailed Description

The class to obtain fit results for all CDM halos analyzed.

This class can be used to obtain results for all CDM halos analyzed.
120 galaxies in the SPARC catalog are used for CDM only fits, 93 galaxies used for ULDM fits with Einasto halo.

5.12.2 Constructor & Destructor Documentation

5.12.2.1 __init__()

```
def results.results_CDM_all.__init__ (
              self,
              model,
              ULDM_fits = False,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_CDM_all class.

This defines the constructor of the results_CDM_all class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW
	Generated by Doxygen
ULDM_fits	(optional) bool True if performing fits for the ULDM models.
fit_dict_in	(optional) constants.fitting_dict instance

5.12.3 Member Function Documentation

5.12.3.1 fit()

```
def results.results_CDM_all.fit (
              self )
```

Define the fit results.

This defines the fit results for the assumed model for 120 galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}}
where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : ndarray[120],
  'fit' : ndarray[120],
  'Chi_sq' : ndarray[120],
  'BIC' : ndarray[120],
  'Mvir' : ndarray[120] (unumpy.uarray),
  'params' : dictionary,
  }

- fit['Vbulge'] = same as fit['Vbulge_none'] Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'fit' = model_fit.model_fit.fit object
- 'Chi_sq' = reduced chi-squared
- 'BIC' = Bayesian information criterion
- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'params' = {
  'c200' : ndarray[120] (unumpy.uarray) = halo concentration,
  'v200' : ndarray[120] (unumpy.uarray) = halo virial velocity,
  'MLd' : ndarray[120] (unumpy.uarray) = mass-to-light ratio of disk,
  'MLb' : ndarray[120] (unumpy.uarray) = mass-to-light ratio of bulge,
  'alpha' : ndarray[120] (unumpy.uarray) = Einasto halo parameter,
  'mstar' : ndarray[120] (unumpy.uarray) = mass of stellar component,
  'Vf' : ndarray[120] = maximum circular velocity,
  'mgas' : ndarray[120] = mass of gas component,
  }

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.13 results.results_CDM_check Class Reference

The class to obtain fit results to check all CDM model implementation.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)

  Define the constructor of the results_CDM_check class.
- def fit (self)

  Define the fit results.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.

5.13.1 Detailed Description

The class to obtain fit results to check all CDM model implementation.

The class to obtain fit results for comparison with Pengfei Li et al 2020 ApJS 247 31 : https://doi.org/10.3847/1538-4365/ab700e

5.13.2 Constructor & Destructor Documentation

5.13.2.1 __init__()

```
def results.results_CDM_check.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_CDM_check class.

This defines the constructor of the results_CDM_check class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Burkert • DC14 • Einasto • NFW

5.13.3 Member Function Documentation

5.13.3.1 fit()

```
def results.results_CDM_check.fit (
              self )
```

Define the fit results.

This defines the fit results for the assumed model for all galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}} where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : str[175],
  'Chi_sq' : ndarray[175],
  'params' : ndarray[175, 3 or 4],
  'Mvir' : float,
  'fit' : model_fit.model_fit.fit object,
  }

- fit['Vbulge'] = same as 'Vbulge_none' with the change :
  'params' : ndarray[175, 4 or 5]
  Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'Chi_sq' = reduced chi-squared
- 'params' = fit parameters
  - for Burkert, DC14, and NFW without bulge component: [c200, v200, MLd]
  - for Burkert, DC14, and NFW with bulge component: [c200, v200, MLd, MLb]
  - for Einasto without bulge component: [c200, v200, MLd, alpha]
  - for Einasto with bulge component: [c200, v200, MLd, MLb, alpha]
- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'fit' = model_fit.model_fit.fit object

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.14 results.results_DC14_check Class Reference

The class to obtain fit results to check the DC14 and NFW model implementation.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)

  Define the constructor of the results_DC14_check class.
- def fit (self)

  Define the fit results.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.

5.14.1 Detailed Description

The class to obtain fit results to check the DC14 and NFW model implementation.

The class to obtain fit results for comparison with Monthly Notices of the Royal Astronomical Society, Volume 466, Issue 2, April 2017, Pages 1648–1668 : https://doi.org/10.1093/mnras/stw3101

5.14.2 Constructor & Destructor Documentation

5.14.2.1 __init__()

```
def results.results_DC14_check.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_DC14_check class.

This defines the constructor of the results_DC14_check class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • DC14 • NFW
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used

5.14.3 Member Function Documentation

5.14.3.1 fit()

```
def results.results_DC14_check.fit (
              self )
```

Define the fit results.

This defines the fit results for the assumed model for 120 galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}} where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : str[149],
  'Chi_sq' : ndarray[149],
  'params' : ndarray[149, 3],
  'Mvir' : float,
  'BIC' : float,
  'fit' : model_fit.model_fit.fit object,
  }

- fit['Vbulge'] = same as 'Vbulge_none' with the change :
  'params' : ndarray[149, 4]
  Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'Chi_sq' = reduced chi-squared
- 'params' = fit parameters
  - for DC14, and NFW without/with bulge component: [c200, v200, MLd]
- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'BIC' = Bayesian information criterion (using lmfit definition)
- 'fit' = model_fit.model_fit.fit object

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.15 results.results_Einasto_check Class Reference

The class to obtain fit results to check the Einasto and NFW model implementation.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)

  Define the constructor of the results_Einasto_check class.
- def fit (self)

  Define the fit results.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.

5.15.1 Detailed Description

The class to obtain fit results to check the Einasto and NFW model implementation.

The class to obtain fit results for comparison with Nicolas Loizeau and Glennys R. Farrar 2021 ApJL 920 L10 : https://doi.org/10.3847/2041-8213/ac1bb7

5.15.2 Constructor & Destructor Documentation

5.15.2.1 __init__()

```
def results.results_Einasto_check.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_Einasto_check class.

This defines the constructor of the results_Einasto_check class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • Einasto • NFW
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.15.3 Member Function Documentation

5.15.3.1 fit()

```
def results.results_Einasto_check.fit (
              self )
```

Define the fit results.

This defines the fit results for the assumed model for 120 galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}} where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : str[121],
  'Chi_sq' : ndarray[121],
  'params' : ndarray[121, 3 or 4],
  'Mvir' : float,
  'fit' : model_fit.model_fit.fit object,
  }

- fit['Vbulge'] = same as 'Vbulge_none' with the change :
  'params' : ndarray[121, 4 or 5]
  Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'Chi_sq' = reduced chi-squared
- 'params' = fit parameters
  - for NFW without bulge component: [c200, v200, MLd]
  - for NFW with bulge component: [c200, v200, MLd, MLb]
  - for Einasto without bulge component: [c200, v200, MLd, alpha]
  - for Einasto with bulge component: [c200, v200, MLd, MLb, alpha]
- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'fit' = model_fit.model_fit.fit object

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.16 results.results_psi_multi_all Class Reference

The class to obtain fit results for all double flavored ULDM models analyzed.

Public Member Functions

- def __init__ (self, fit_dict_in=fitting_dict_in)

  Define the constructor of the results_psi_multi_all class.
- def fit (self)

  Define the fit results.

Public Attributes

- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**

  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- **mfree**

  bool
  Equal to true/false if soliton particle mass free/fixed in fitting procedure
- cdmhalo

  str
  Can be equal to :
- **matched**

  bool
  True if using ULDM matched models

5.16.1 Detailed Description

The class to obtain fit results for all double flavored ULDM models analyzed.

This class can be used to obtain results for all double flavored ULDM models analyzed.
93 galaxies in the SPARC catalog are used.

5.16.2 Constructor & Destructor Documentation

5.16.2.1 __init__()

```
def results.results_psi_multi_all.__init__ (
              self,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_psi_multi_all class.

This defines the constructor of the results_psi_multi_all class.

Parameters

self	object pointer
fit_dict↩ _in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.16.3 Member Function Documentation

5.16.3.1 fit()

```
def results.results_psi_multi_all.fit (
              self )
```

Define the fit results.

This defines the fit results for the double flavor ULDM models for 93 galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}} where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : ndarray[93],
  'fit' : ndarray[93],
  'Chi_sq' : ndarray[93],
  'BIC' : ndarray[93],
  'Mhalo_1' : ndarray[93],
  'Mhalo_2' : ndarray[93],
  'Mvir' : ndarray[93],
  'params' : dictionary,
  }

- fit['Vbulge'] = same as fit['Vbulge_none'] Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'fit' = model_fit.model_fit.fit object
- 'Chi_sq' = reduced chi-squared
- 'BIC' = Bayesian information criterion
- 'Mhalo_1' = mass of ULDM halo one
- 'Mhalo_2' = mass of ULDM halo two

- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'params' = {
  'c200' : ndarray[93] (unumpy.uarray) = halo one concentration,
  'c200_2' : ndarray[93] (unumpy.uarray) = halo two concentration,
  'v200' : ndarray[93] (unumpy.uarray) = halo one virial velocity,
  'v200_2' : ndarray[93] (unumpy.uarray) = halo two virial velocity,
  'MLd' : ndarray[93] (unumpy.uarray) = mass-to-light ratio of disk,
  'MLb' : ndarray[93] (unumpy.uarray) = mass-to-light ratio of bulge,
  'alpha' : ndarray[93] (unumpy.uarray) = Einasto halo parameter,
  'm22' : ndarray[93] (unumpy.uarray) = ULDM particle one mass,
  'm22_2' : ndarray[93] (unumpy.uarray) = ULDM particle two mass,
  'Msol' : ndarray[93] (unumpy.uarray) = soliton one mass,
  'Msol_2' : ndarray[93] (unumpy.uarray) = soliton two mass,
  'mstar' : ndarray[93] (unumpy.uarray) = mass of stellar component,
  'Vf' : ndarray[93] = maximum circular velocity,
  'mgas' : ndarray[93] = mass of gas component,
  }

5.16.4 Member Data Documentation

5.16.4.1 cdmhalo

`results.results_psi_multi_all.cdmhalo`

str
Can be equal to :

- 'Burkert'
- 'DC14'
- 'Einasto'
- 'NFW'

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.17 results.results_psi_single_all Class Reference

The class to obtain fit results for all single flavored ULDM models analyzed.

Public Member Functions

- def __init__ (self, fit_dict_in=fitting_dict_in)
  Define the constructor of the results_psi_single_all class.
- def fit (self)
  Define the fit results.

Public Attributes

- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**

  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- **mfree**

  bool
  Equal to true/false if soliton particle mass free/fixed in fitting procedure
- cdmhalo

  str
  Can be equal to :
- **matched**

  bool
  True if using ULDM matched models

5.17.1 Detailed Description

The class to obtain fit results for all single flavored ULDM models analyzed.

This class can be used to obtain results for all single flavored ULDM models analyzed.
93 galaxies in the SPARC catalog are used.

5.17.2 Constructor & Destructor Documentation

5.17.2.1 __init__()

```
def results.results_psi_single_all.__init__ (
              self,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the results_psi_single_all class.

This defines the constructor of the results_psi_single_all class.

Parameters

self	object pointer
fit_dict↩ _in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.17.3 Member Function Documentation

5.17.3.1 fit()

```
def results.results_psi_single_all.fit (
              self )
```

Define the fit results.

This defines the fit results for the single flavor ULDM models for 93 galaxies in the SPARC catalog.

Parameters

self	object pointer

Returns

dictionary
Dictionary of resulting fit parameters. Results are as follows : fit = {'Vbulge_none' : {...}, 'Vbulge' : {...}}
where 'Vbulge_none'/'Vbulge' corresponds to galaxies without/with a bulge component.

- fit['Vbulge_none'] = {
  'Name' : ndarray[93],
  'fit' : ndarray[93],
  'Chi_sq' : ndarray[93],
  'BIC' : ndarray[93],
  'Mhalo' : ndarray[93],
  'Mvir' : ndarray[93],
  'params' : dictionary,
  }

- fit['Vbulge'] = same as fit['Vbulge_none'] Each of the components of fit['Vbulge_none'] and fit['Vbulge'] are :
- 'Name' = name of galaxy
- 'fit' = model_fit.model_fit.fit object
- 'Chi_sq' = reduced chi-squared
- 'BIC' = Bayesian information criterion
- 'Mhalo' = mass of ULDM halo
- 'Mvir' = M200 (mass enclosed within radius R200) in units of solar masses
- 'params' = {
  'c200' : ndarray[93] (unumpy.uarray) = halo concentration,
  'v200' : ndarray[93] (unumpy.uarray) = halo virial velocity,
  'MLd' : ndarray[93] (unumpy.uarray) = mass-to-light ratio of disk,
  'MLb' : ndarray[93] (unumpy.uarray) = mass-to-light ratio of bulge,
  'alpha' : ndarray[93] (unumpy.uarray) = Einasto halo parameter,
  'm22' : ndarray[93] (unumpy.uarray) = ULDM particle mass,
  'Msol' : ndarray[93] (unumpy.uarray) = soliton mass,
  'mstar' : ndarray[93] (unumpy.uarray) = mass of stellar component,
  'Vf' : ndarray[93] = maximum circular velocity,
  'mgas' : ndarray[93] = mass of gas component,
  }

5.17.4 Member Data Documentation

5.17.4.1 cdmhalo

`results.results_psi_single_all.cdmhalo`

str
Can be equal to :

- 'Burkert'
- 'DC14'
- 'Einasto'
- 'NFW'

The documentation for this class was generated from the following file:

- pyfiles/fitting/results.py

5.18 alp_funcs.soliton Class Reference

The class containing mass functions for the ULDM halo models.

Public Member Functions

- def __init__ (self, model, fit_dict_in=fitting_dict_in)
  Define the constructor of the soliton class.
- def mass (self, params, r)
  Define the mass profile of the ULDM galactic structure.
- def Mhalo (self, params)
  Define the total mass of the outer halo.
- def Mvir (self, params)
  Define the total galactic DM halo mass.

Public Attributes

- **model**

  str
  Model to assume for DM halo.
- **fit_dict_in**

  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**

  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- matched

  bool
  Denotes how to combine the soliton and outer halo.
- mfree

  float
  Denotes how to treat soliton particle mass in fitting procedure.
- **m22**

  float
  Soliton particle mass in units of 10^{-22} eV.
- **m22_2**

  float
  Soliton particle mass two in units of 10^{-22} eV.
- cdmhalo

  float
  Denotes which CDM profile to use for outer halo in ULDM galactic structure.
- **base_funcs_in**

  alp_funcs.base_funcs instance
  Instance of the alp_funcs.base_funcs class
- **halo_init**

  cdm_funcs.halo instance
  Instance of the cdm_funcs.halo class

5.18.1 Detailed Description

The class containing mass functions for the ULDM halo models.

5.18.2 Constructor & Destructor Documentation

5.18.2.1 __init__()

```
def alp_funcs.soliton.__init__ (
              self,
              model,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the soliton class.

This defines the constructor of the soliton class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • psi_single • psi_multi
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.18.3 Member Function Documentation

5.18.3.1 mass()

```
def alp_funcs.soliton.mass (
              self,
              params,
              r )
```

Define the mass profile of the ULDM galactic structure.

This defines the mass of the ULDM galactic structure (soliton and outer halo) at a given radius given the model parameters.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.
r	ndarray[N] Numpy array of N radii values in units of kpc.

Returns

ndarray[N]
Numpy array of N mass values in units of solar mass.

5.18.3.2 Mhalo()

```
def alp_funcs.soliton.Mhalo (
              self,
              params )
```

Define the total mass of the outer halo.

This defines the total mass of the outer halo assuming the model parameters.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using alp_params.soliton.params.

Returns

float
Total mass of the outer halo in units of solar mass.

5.18.3.3 Mvir()

```
def alp_funcs.soliton.Mvir (
              self,
              params )
```

Define the total galactic DM halo mass.

This defines the total mass of the galactic DM given the model parameters.

Parameters

self	object pointer
params	lmfit.Parameters instance Instance of the lmfit.Parameters class. All model parameters contained here. This can be created using cdm_params.halo.params.

Returns

float
Total mass of the galactic DM in units of solar mass.

5.18.4 Member Data Documentation

5.18.4.1 cdmhalo

```
alp_funcs.soliton.cdmhalo
```

float
Denotes which CDM profile to use for outer halo in ULDM galactic structure.

Equal to constants.fitting_dict.sol_cdmhalo

5.18.4.2 matched

`alp_funcs.soliton.matched`

bool
Denotes how to combine the soliton and outer halo.

Equal constants.fitting_dict.sol_match

5.18.4.3 mfree

`alp_funcs.soliton.mfree`

float
Denotes how to treat soliton particle mass in fitting procedure.

Equal to constants.fitting_dict.sol_mfree

The documentation for this class was generated from the following file:

- pyfiles/models/alps/alp_funcs.py

5.19 alp_params.soliton Class Reference

The class containing all parameters for fitting the ULDM halo models.

Public Member Functions

- def __init__ (self, model, data, fit_dict_in=fitting_dict_in)
  Define the constructor of the soliton class.
- def params (self)
  Define the parameters to be assumed.

Public Attributes

- **model**
  str
  Model to assume for DM halo.
- data
  dictionary
  Dictionary containing all the necessary data for a given galaxy.
- **fit_dict_in**
  constants.fitting_dict instance
  Instance of the constants.fitting_dict class.
- **args_opts**
  dictionary
  Dictionary of rules to be assumed during fitting Instance of constants.fitting_dict.args_opts
- params_vals

dictionary
Dictionary of variable values to be used during fitting.

- **c200**

  list
  Values for c200 in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **v200**

  list
  Values for V200 in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **v200fac**

  list
  Values for v200fac in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **MLD**

  list
  Values for MLd in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **MLB**

  list
  Values for MLb in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **alpha**

  list
  Values for alpha in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **msol**

  float
  Values for msol in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **m22**

  float
  Values for m22 in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **msol_2**

  float
  Values for msol two in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- **m22_2**

  float
  Values for m22 two in the form [starting value, min, max] Contained in constants.fitting_dict.params_vals
- matched

  bool
  Denotes how to combine the soliton and outer halo.
- mfree

  float
  Denotes how to treat soliton particle mass in fitting procedure.
- cdmhalo

  float
  Denotes which CDM profile to use for outer halo in ULDM galactic structure.

5.19.1 Detailed Description

The class containing all parameters for fitting the ULDM halo models.

This class contains all parameters necessary to perform the fitting procedures for all ULDM halo models.

5.19.2 Constructor & Destructor Documentation

5.19.2.1 __init__()

```
def alp_params.soliton.__init__ (
              self,
              model,
              data,
              fit_dict_in = fitting_dict_in )
```

Define the constructor of the soliton class.

This defines the constructor of the soliton class.

Parameters

self	object pointer
model	str Model to assume for DM halo. Can be equal to : • psi_single • psi_multi
data	dictionary Dictionary containing all the necessary data for a given galaxy. This can be created using galaxy.galaxy.
fit_dict_in	(optional) constants.fitting_dict instance Instance of the constants.fitting_dict class. Contains all the necessary rules and values to be used in fitting.

5.19.3 Member Function Documentation

5.19.3.1 params()

```
def alp_params.soliton.params (
              self )
```

Define the parameters to be assumed.

This defines the parameters to be assumed during the fitting procedures for all ULDM halos.

Parameters

self	object pointer

Returns

lmfit.Parameters instance Instance of the lmfit.Parameters class

5.19.4 Member Data Documentation

5.19.4.1 cdmhalo

`alp_params.soliton.cdmhalo`

float
Denotes which CDM profile to use for outer halo in ULDM galactic structure.

Equal to constants.fitting_dict.sol_cdmhalo

5.19.4.2 data

`alp_params.soliton.data`

dictionary
Dictionary containing all the necessary data for a given galaxy.

This can be created using galaxy.galaxy.

5.19.4.3 matched

`alp_params.soliton.matched`

bool
Denotes how to combine the soliton and outer halo.

Equal constants.fitting_dict.sol_match

5.19.4.4 mfree

`alp_params.soliton.mfree`

float
Denotes how to treat soliton particle mass in fitting procedure.

Equal to constants.fitting_dict.sol_mfree

5.19.4.5 params_vals

`alp_params.soliton.params_vals`

dictionary
Dictionary of variable values to be used during fitting.

Instance of constants.fitting_dict.params_vals

The documentation for this class was generated from the following file:

- pyfiles/models/alps/alp_params.py

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