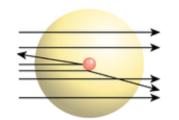
RUTHERFORD'S ATOM MODEL

- i) Majority of α particles passed without any deviation.
- ii) Some are scattered at small angle θ (impact parameter is equal to that of nuclear radius)
- iii) Only few alpha particle retrace the path (impact parameter = 0)



BOHR ATOM MODEL

First postulate

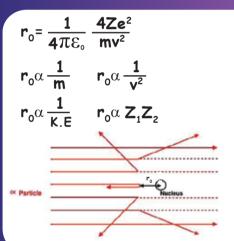
$$F = \frac{1}{4\pi\epsilon_0} \frac{Ze \times e}{r^2}$$

Second postulate

$$mvr = \frac{nh}{2\pi}$$

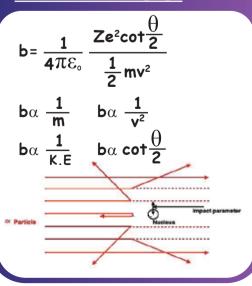
$$\frac{mv^2}{r} = \frac{1}{4\pi\epsilon_o} \frac{Ze^2}{r^2}$$

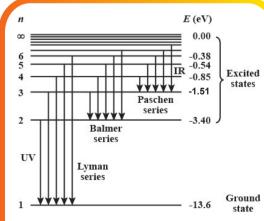
DISTANCE OF CLOSEST APPROACH OF ∝-PARTICLES



- RADIUS OF ORBIT $r_n = 0.53 \frac{n^2}{7}$
- VELOCITY OF ELECTRON Vna Zn
- TIME PERIOD $T\alpha \frac{n^3}{z^2}$
- FREQUENCY = $\frac{1}{T} \alpha \frac{z^2}{n^3}$
- CURRENT = $\frac{e}{T} \alpha \frac{z^2}{n^3}$
- MAGNETIC FIELD B $\alpha \frac{v}{r^2} \Rightarrow B \alpha \frac{z^3}{n^5}$
- MAGNETIC DIPOLE MOMENT Man

IMPACT PARAMETER





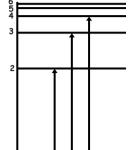
As the n value increases, the energy difference betwen adjacent level decreases 5 \rightarrow 1 > 4 \rightarrow 1 > 3 \rightarrow 1 > 2 \rightarrow 1 > > 4 \rightarrow 2 > 3 \rightarrow 2 \to 2 \to 2 \to 3 \to 2 \to 2 \to 3 \to 3 \to 3 \to 3 \to 4 \to 5 \to 6 \to 9 \to 1 \to

ATOMS



HYDROGEN SPECTRUM

Absorption spectrum



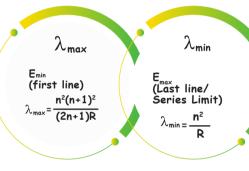
Electrons absorb only those photons whose energy
-Energy difference of 2 shall

=Energy difference of 2 shells

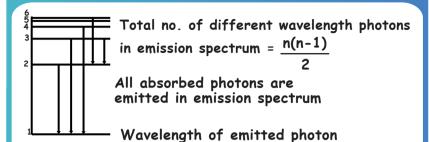
If atomic excitation takes place upto nth shell starting from ground state then (n-1)different photons are absorbed

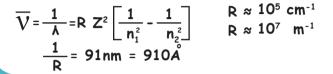
ENERGY

Total energy = -13.6 $\frac{z^2}{n^2}$ eV K.E = -T.E = +13.6 $\frac{z^2}{n^2}$ eV P.E = 2T.E

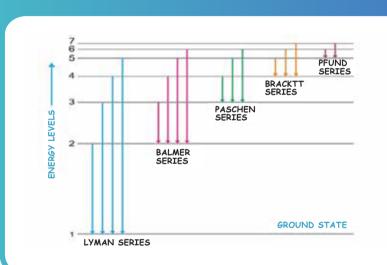


EMISSION SPECTRUM





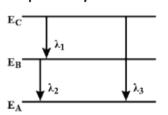
LINE SPECTRUM OF HYDROGEN ATOM



LINE SPECTRUM OF HYDROGEN ATOM

Spectral series	n ₁	n ₂	Wavelength	λ_{max} $(n_2=n_1+1)$	λ_{\min} (n ₂ = ∞)	$\frac{\lambda_{\text{max}}}{\lambda_{\text{min}}}$	Region	Range
Lyman	1	2,3,4	$\frac{1}{\lambda_{Ly}} = R\left(\frac{1}{1^2} - \frac{1}{n^2}\right)$	4 3R	1 R	4 3	Ultra - violet	911.6 Å to 1216 Å
Balmer	2	3,4,5	$\frac{1}{\lambda_{\rm B}} = R \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$	36 5R	4 R	9 5	Visible	3646 Å to 6563 Å
Paschen	3	4,5,6	$\frac{1}{\lambda_p} = R \left(\frac{1}{3^2} - \frac{1}{n^2} \right)$	144 7R	9 R	16 7	Near infar-red	8204 Å to 18753 Å
Brackett	4	5,6,7	$\frac{1}{\lambda_{Br}} = R \left(\frac{1}{4^2} - \frac{1}{n^2} \right)$	400 9R	16 R	<u>25</u> 9	Middle infra-red	14585 Å to 40515 Å
Pfund	5	6,7,8	$\frac{1}{\lambda_{\rm Pf}} = R \left(\frac{1}{5^2} - \frac{1}{n^2} \right)$	900 11R	25 R	36 11	Far infra-red	22790 Å to 74583 Å

Energy levels A,B & C of a certain atom correspond to increasing values of energy, i.e. $E_A < E_B < E_C$. If $\lambda_1, \lambda_2, \lambda_3$ are the wavelengths of radiations corresponding to transitions C to B,B to A and C to A respectively then



- a) $\lambda_3 = \lambda_1 + \lambda_2$ c) $\lambda_1 + \lambda_2 + \lambda_3 = 0$
- **b)** $\lambda_3 = \frac{\lambda_1 \lambda_2}{\lambda_1 + \lambda_2}$ **d)** $\lambda_3^2 = \lambda_1^2 + \lambda_2^2$