# problem\_set\_6\_notebook

#### November 5, 2021

```
template=dataFile['template']
    th=template[0]
    tl=template[1]
    return th,tl
def read_file(filename):
    dataFile=h5py.File(filename, 'r')
    dqInfo = dataFile['quality']['simple']
    qmask=dqInfo['DQmask'][...]
    meta=dataFile['meta']
    #qpsStart=meta['GPSstart'].value
    gpsStart=meta['GPSstart'][()]
    #print meta.keys()
    #utc=meta['UTCstart'].value
    utc=meta['UTCstart'][()]
    #duration=meta['Duration'].value
    duration=meta['Duration'][()]
    #strain=dataFile['strain']['Strain'].value
    strain=dataFile['strain']['Strain'][()]
    dt=(1.0*duration)/len(strain)
```

```
dataFile.close()
return strain,dt,utc
```

# 1 Import Data

```
[3]: file_names = os.listdir(folder)
     # Save events from Hanford, Livingston and corresponding GW template
    L_events = []
    H_events = []
    for file in file_names:
        if file[0] == 'H':
            H_events.append(file)
            for j in file_names:
                if j[:] == 'L-L1_LOSC_4' + file[11:]:
                    L_events.append(j)
     # Save GW templates manually, making sure it corresponds to correct event
    GW = ['GW170104_4_template.hdf5','GW151226_4_template.
      →hdf5', 'LVT151012_4_template.hdf5', 'GW150914_4_template.hdf5']
[4]: print('''
    Handford Events:
     _____
    {0}
    Livinston Events:
     _____
    {1}
    GW Templates:
     '''.format(H_events, L_events, GW))
```

```
GW Templates:
-----
['GW170104_4_template.hdf5', 'GW151226_4_template.hdf5',
'LVT151012_4_template.hdf5', 'GW150914_4_template.hdf5']
```

### 2 PART A: ESTIMATING NOISE MODEL

To estimate the noise model, I'm going to average the PSD over all the events to average out the signal (and since the system noise is time invariant). I'll also make sure to window the data using the Tukey model, which is flat in the center where the signal is most likely to be.

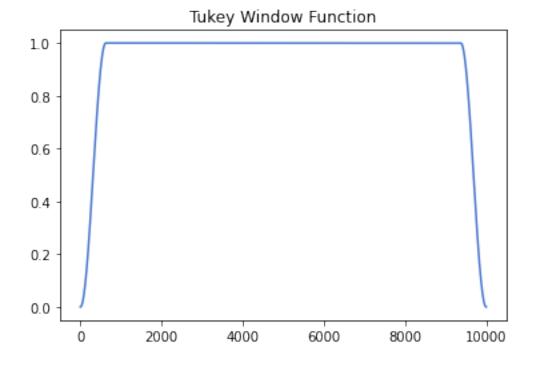
```
[5]: # To get an idea of the window function I am going to use, I've plotted an → example below.

tukey_window = sig.windows.tukey(10000, alpha=0.125)

plt.plot(tukey_window)

plt.title('Tukey Window Function')

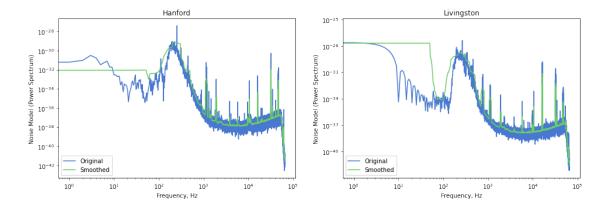
plt.show()
```



```
[6]: def get_noise_model(events):

    PSD = 0
    for i in events:
        strain, dt, utc = read_file(folder + i)
```

```
N = len(strain)
             # Window the function
             window = sig.windows.tukey(N, alpha = 0.2)
             # Get PSD
             psd = np.abs(np.fft.rfft(strain*window))**2
             PSD+=psd
         # Average
         noise_model = PSD/len(events)
         return noise_model
     # Get the noise model for each of the detectors
     raw_noise_model_H = get_noise_model(H_events)
     raw_noise_model_L = get_noise_model(L_events)
[7]: # Smooth the noise models
     def smooth_data(noise, n_smooth):
         smoothed = np.convolve(noise, np.ones(n_smooth)/n_smooth, mode = 'same')
         return smoothed
     smooth_noise_H = smooth_data(raw_noise_model_H, 100)
     smooth_noise_L = smooth_data(raw_noise_model_L, 100)
[8]: # Plot raw vs smoothed noise models
     # Plot
     fig, axs = plt.subplots(1,2, figsize = (16,5))
     axs[0].loglog((raw_noise_model_H), label = 'Original')
     axs[0].loglog(smooth_noise_H,label = 'Smoothed')
     axs[0].set_title('Hanford')
     leg0 = axs[0].legend(loc="lower left",
                      ncol=1, shadow=False, fancybox=False)
     axs[1].loglog(raw_noise_model_L, label = 'Original')
     axs[1].loglog(smooth_noise_L,label = 'Smoothed')
     axs[1].set_title('Livingston')
     leg1 = axs[1].legend(loc="lower left",
                      ncol=1, shadow=False, fancybox=False)
     for ax in axs.flat:
         ax.set(xlabel = 'Frequency, Hz', ylabel='Noise Model (Power Spectrum)')
```



Some additional notes on the above: - I've smoothed out the noise model: however, I have not removed the big spikes since they are intrinsic to the actual noise. I have, however, removed small scale oscilations that were most likely just white noise. - The data we care about lies in the range of  $\sim$ 20 Hz to  $\sim$ 2000 Hz (according to LIGO). Consequently, it doesn't matter that the noise model for frequencies lower than this don't follow the data that closely.

### 3 PART B: MATCHED FILTERS

#### 3.0.1 Whitening the data

Using the noise model, we can now whiten our data and look for the signal.

```
[9]: def whiten(signal, noise_model, fs = 4096):
    spectr = np.fft.rfft(signal)
    white_ft = spectr / np.sqrt(noise_model) / fs
    return white_ft

def xcorr(strain, template, noise_model, fs = 4096):

# Get window
    window = sig.windows.tukey(len(strain), alpha = 0.125)

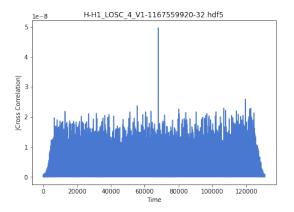
# Get whitened data in Fourier space, making sure to window the data and_unitemplate
    strain_ft_whitened = whiten(strain*window, noise_model, fs = fs)
    template_ft_whitened = whiten(template*window, noise_model, fs = fs)

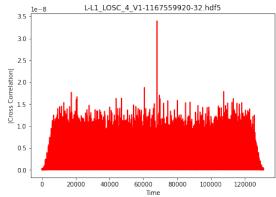
# Get matched filter
    xcorr = np.fft.irfft(strain_ft_whitened*np.conj(template_ft_whitened))
```

return xcorr

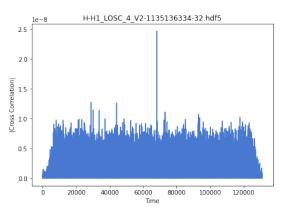
```
[10]: # Let's see how we did!
      xcorrs_H = np.empty([4,131072], dtype = 'complex')
      xcorrs_L = np.empty([4,131072], dtype = 'complex')
      for i in range(4):
          # Load in each event
          H_i, dt, utc = read_file(folder + H_events[i])
          L_i, dt, utc = read_file(folder + L_events[i])
          th_i, tl_i = read_template(folder + GW[i])
          # Get xcorr for each event
          xcorr_H_i = xcorr(H_i, th_i, smooth_noise_H)
          xcorr_L_i = xcorr(L_i, tl_i, smooth_noise_L)
          # Save data for analysis later
          xcorrs_H[i,:] = xcorr_H_i
          xcorrs_L[i,:] = xcorr_L_i
          N = len(xcorr_H_i)
          # Plot
          print('EVENT: {0}'.format(i+1))
          print('____')
          fig, axs = plt.subplots(1,2, figsize = (16,5))
          axs[0].plot(np.abs(np.roll(xcorr_H_i, N//2)), label = H_events[i])
          axs[0].set_title(H_events[i])
          \#leg0 = axs[0].legend(loc = 'upper right', ncol=1, shadow=False, 
       \rightarrow fancybox=False)
          axs[1].plot(np.abs(np.roll(xcorr_L_i, N//2)), label = L_events[i], color = __
       →'red')
          axs[1].set_title(L_events[i])
          \#leg1 = axs[1].legend(loc = 'upper right', ncol=1, shadow=False, 
       \rightarrow fancybox=False)
          for ax in axs.flat:
              ax.set(xlabel = 'Time', ylabel= '|Cross Correlation|')
          plt.show()
```

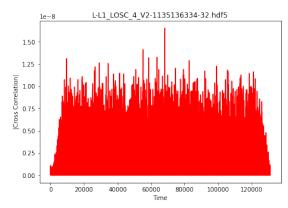
EVENT: 1



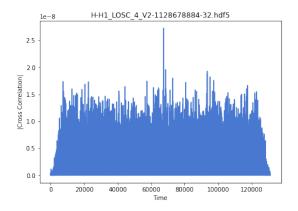


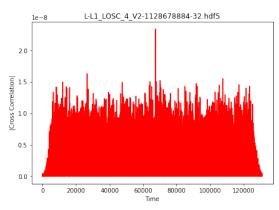
-----



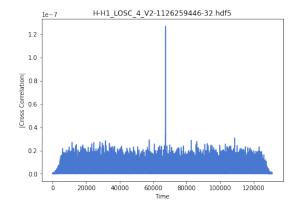


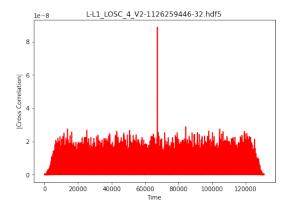
## EVENT: 3





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# 4 PART C: SIGNAL TO NOISE

**NOTE: SEE LEGENDS WITHIN GRAPHS FOR OBTAINED SNR AND NOISE VALUES** In this section, we estimate the signal-to-noise ratio of the detection.

We can estimate the noise by evaluating the standard deviation of the signal in an area that isn't affected by the window nor where a strong detection is. In all of the plots, a reasonable spot where the noise looks random is around 20000 - 40000 on the 'time' axis.

#### 4.0.1 Individual SNRs

```
[52]: # Get noise of each event by estimating it from 20000-40000
N = xcorrs_H.shape[1]

# Roll to center
rolled_H = np.roll(xcorrs_H, N//2, axis = 1)
abs_rolled_H = np.abs(rolled_H)

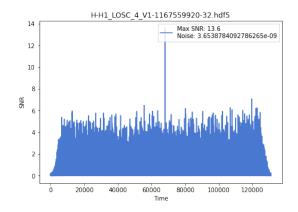
rolled_L = np.roll(xcorrs_L, N//2, axis = 1)
abs_rolled_L = np.abs(rolled_L)

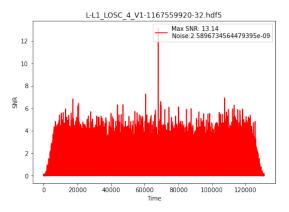
# Get standard deviation of area that looks like random noise
noise_H_events = np.abs(np.std(abs_rolled_H[:,20000:40000], axis = 1))
noise_L_events = np.abs(np.std(abs_rolled_L[:,20000:40000], axis = 1))

indmax_SNR_H = []
indmax_SNR_H = []
max_SNR_H_list = []
max_SNR_L_list = []
```

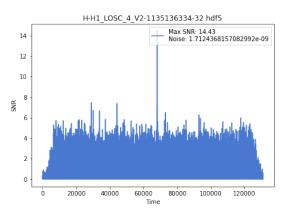
```
for i in range(4):
    # Plot SNR Graphs
    SNR_H_i = abs_rolled_H[i,:] / noise_H_events[i]
    SNR_L_i = abs_rolled_L[i,:] / noise_L_events[i]
    max_SNR_H = np.argmax(SNR_H_i)
    max_SNR_L = np.argmax(SNR_L_i)
    indmax_SNR_H.append(max_SNR_H)
    indmax_SNR_L.append(max_SNR_L)
    max_SNR_H_list.append(SNR_H_i[max_SNR_H])
    max_SNR_L_list.append(SNR_L_i[max_SNR_L])
    print('EVENT: {0}'.format(i+1))
    print('____')
    fig, axs = plt.subplots(1,2, figsize = (16,5))
    axs[0].plot(SNR_H_i, label = 'Max SNR: {0}\nNoise: {1}'.

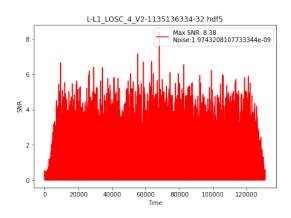
→format(round(SNR_H_i[max_SNR_H], 2), noise_H_events[i]))
    axs[0].set_title(H_events[i])
    leg0 = axs[0].legend(loc = 'upper right', ncol=1, shadow=False,__
 →fancybox=False)
    axs[1].plot(SNR_L_i, color = 'red', label = 'Max SNR: {0}\nNoise:{1}'.
 →format(round(SNR_L_i[max_SNR_L], 2), noise_L_events[i]))
    axs[1].set_title(L_events[i])
    leg1 = axs[1].legend(loc = 'upper right', ncol=1, shadow=False,__
 →fancybox=False)
    for ax in axs.flat:
        ax.set(xlabel = 'Time', ylabel= 'SNR')
   plt.show()
```



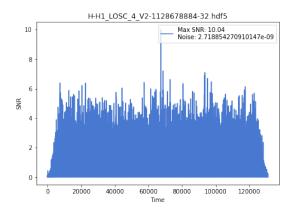


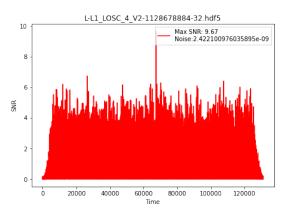
-----



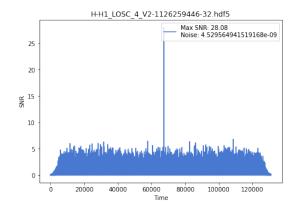


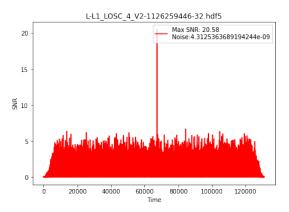
### EVENT: 3





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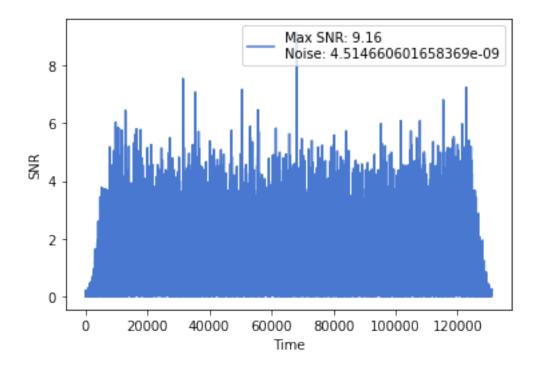


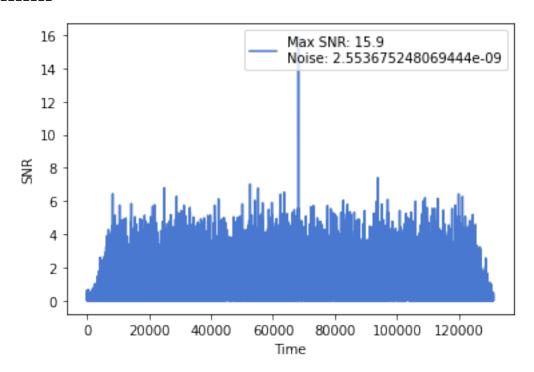


#### 4.0.2 Combined SNR

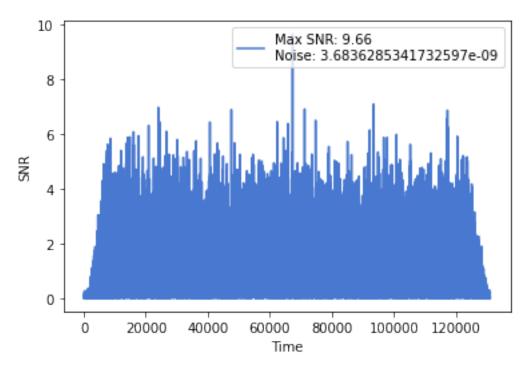
```
[78]: # Combine Cross Correlations
      xcorr_combined = xcorrs_H + xcorrs_L
      # Roll
      rolled\_comb = np.roll(xcorr\_combined, N//2, axis = 1)
      # Get absolute value
      abs_rolled_comb = np.abs(rolled_comb)
      # Get noise estimate
      noise_combined = np.abs(np.std(abs_rolled_comb[:,20000:40000], axis = 1))
      max_SNR_list = []
      max_SNR_index = []
      for i in range(4):
          SNR_i = abs_rolled_comb[i,:] / noise_combined[i]
          max_SNR = np.argmax(SNR_i)
          max_SNR_index.append(max_SNR)
          max_SNR_list.append(SNR_i[max_SNR])
          print('EVENT: {0}'.format(i+1))
          print('____')
          fig = plt.figure(figsize = (6,4))
          plt.plot(SNR_i, label= 'Max SNR: {0}\nNoise: {1}'.
       →format(round(SNR_i[max_SNR], 2), noise_combined[i]) )
          plt.legend()
          plt.xlabel('Time')
          plt.ylabel('SNR')
          plt.show()
```

EVENT: 1

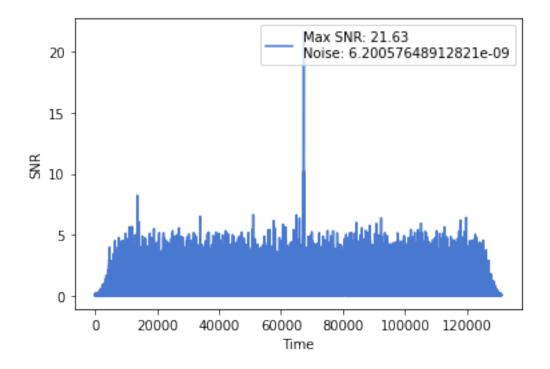




EVENT: 3



EVENT: 4

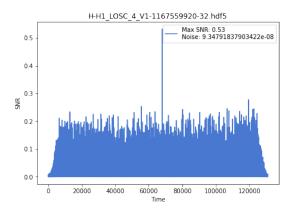


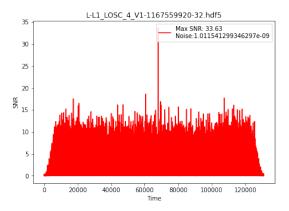
### 5 PART D: COMPARING SNR TO THEORY

**NOTE: SEE LEGENDS WITHIN GRAPHS FOR OBTAINED SNR AND NOISE VALUES** The theoretical noise is just the matched filter of the template with itself, as given in the tutorial. We then just sum over the cross correlation, take the absolute value and square root.

```
[71]: for i in range(4):
          th, tl = read_template(folder + GW[i])
          xcorr_th = xcorr(th, th, smooth_noise_H)
          xcorr_tl = xcorr(tl, tl, smooth_noise_L)
          # This is where we calculate the new noise
          sig_H = np.sqrt(np.abs(np.sum(xcorr_th)))
          sig_L = np.sqrt(np.abs(np.sum(xcorr_tl)))
          # Plot SNR Graphs
          SNR_H_i = abs_rolled_H[i,:] / sig_H
          SNR_L_i = abs_rolled_L[i,:] / sig_L
          max_SNR_H = np.argmax(SNR_H_i)
          max_SNR_L = np.argmax(SNR_L_i)
          print('EVENT: {0}'.format(i+1))
          print('____')
          fig, axs = plt.subplots(1,2, figsize = (16,5))
          axs[0].plot(SNR_H_i, label = 'Max SNR: {0}\nNoise: {1}'.
       →format(round(SNR_H_i[max_SNR_H], 2), sig_H))
          axs[0].set_title(H_events[i])
          leg0 = axs[0].legend(loc = 'upper right', ncol=1, shadow=False,__
       →fancybox=False)
          axs[1].plot(SNR_L_i, color = 'red', label = 'Max SNR: {0}\nNoise:{1}'.
       →format(round(SNR_L_i[max_SNR_L], 2), sig_L))
          axs[1].set_title(L_events[i])
          leg1 = axs[1].legend(loc = 'upper right', ncol=1, shadow=False,__
       →fancybox=False)
          for ax in axs.flat:
              ax.set(xlabel = 'Time', ylabel= 'SNR')
          plt.show()
```

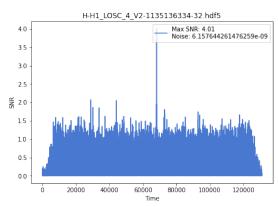
-----

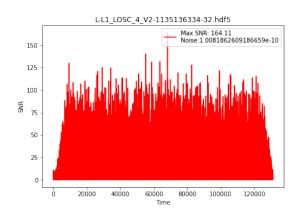




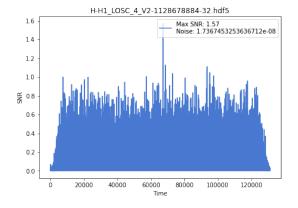
### EVENT: 2

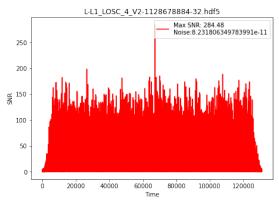
-----



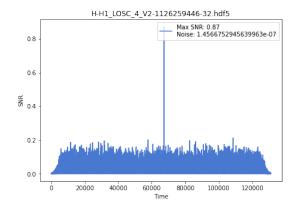


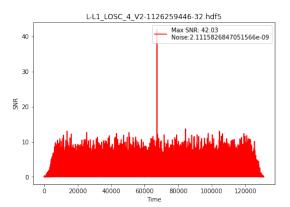
#### EVENT: 3





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Comparing the noise from the theoretical to the noise I obtained, we get a larger noise value for Hanford but smaller for Livingston. I cannot think of a reason as to why this is the case, but I imagine it has something to do with the templates and doing the cross correlation with itself. For some reason, I don't really trust this theoretical result that much given how different the answer is to what I got in PART C.

#### 6 PART E

The way I interpreted this question was that since it wanted a frequency, it made most sense to work in frequency space. Also, by whitening our data and looking at the power spectrum in fourier space, we can determine how much power is contained in the frequencies by summing up the power until we reach half of the total power. The results printed below highlight that most of the power is contained in the lower range of frequencies (which makes sense given the template).

```
[45]: half_freq_H = []
    half_freq_L = []

for i in range(4):

    # Load in template
    th, tl = read_template(folder + GW[i])
    freqs = np.fft.rfftfreq(len(th), 1/4096)

#------#
# Hanford #
#------#
h_fft_white = np.abs(whiten(th, smooth_noise_H))
```

```
tot_h = np.sum(h_fft_white) # Get combined power
   sum_h = 0
   for j, freq in enumerate(freqs):
       sum_h+=h_fft_white[j]
       # Check if power exceeds 1/2 of total power.
       if sum_h >= tot_h/2:
           half_freq = freq
           break
   # Append corresponding frequency to list to save data.
   half_freq_H.append(half_freq)
   #----#
   # Livingston #
    #----#
   # Repeat as above, but for Livingston template
   1_fft_white = np.abs(whiten(tl, smooth_noise_L))
   tot_1 = np.sum(l_fft_white)
   sum_1 = 0
   for j, freq in enumerate(freqs):
       sum_l+=l_fft_white[j]
       if sum_1 >= tot_1/2:
           half_freq = freq
           break
   half_freq_L.append(half_freq)
   print('EVENT: {0}'.format(i+1))
   print('____')
   print("""
_H = \{0\} Hz
_L = \{1\} Hz
   """.format(half_freq_H[i], half_freq_L[i]))
   print()
```

```
_H = 139.625 Hz
_L = 138.375 Hz
```

```
EVENT: 2

_H = 151.21875 Hz

_L = 167.9375 Hz

EVENT: 3

_H = 130.65625 Hz

_L = 142.6875 Hz

EVENT: 4

_H = 135.3125 Hz

_L = 145.90625 Hz
```

# 7 PART F

To calculate the difference in arrival time, all we need to do is take the difference in the location of the indices where the SNR is maximum and multiply that by dt.

EVENT: 1
----t\_arrival = 0.004638671875 s

```
EVENT: 2
------
t_arrival = 0.00048828125 s

EVENT: 3
-----
t_arrival = 0.001953125 s

EVENT: 4
-----
t_arrival = 0.009033203125 s
```

This question is extremely confusing and I wasn't sure how to interpret what it wanted, so I've given my interpretation below. Given that gravitational waves travel at the speed of light, the precision of the localization of the time arrival is highly dependent on the SNR of the detection, as we'd expect. For high SNR detection, we can expect roughly 1e2 km precision in positional uncertainty, whereas for low SNR bursts, we get roughly ~1e3 km uncertainty. This is highlighted in the plot I generated below where I compute the distance between the two detectors using  $d = c^*\delta t$ , gettheerrorbysubtractingthem from the true distance between the two detectors (3002km), and plotting as a function of S

```
[80]: distances = 299792*np.asarray(time)
d_true = 3002 #km, from https://www.ligo.caltech.edu/page/facilities
error = np.abs(d_true - distances)
plt.figure(figsize = (10,8))
plt.scatter(max_SNR_L_list, error)
plt.xlabel('Max_SNR_ of Detection')
plt.ylabel('Uncertainty in Distance Between Detectors')
plt.show()
```

