

## Results of the dynamical core model intercomparison project (DCMIP): PUMA

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# Portable University Model of the Atmosphere (PUMA)

Dynamical core of the PlanetSimulator (Universität Hamburg)

Original version

- Global spectral general circulation model of the dry atmosphere (Hoskins-Simmons type)
- Numerical solution of hydrostatic EULER equations for an ideal gas on a rotating sphere
- Diabatic, dissipative processes: only Newtonian Cooling and Rayleigh friction

Simple physics extension

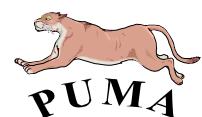
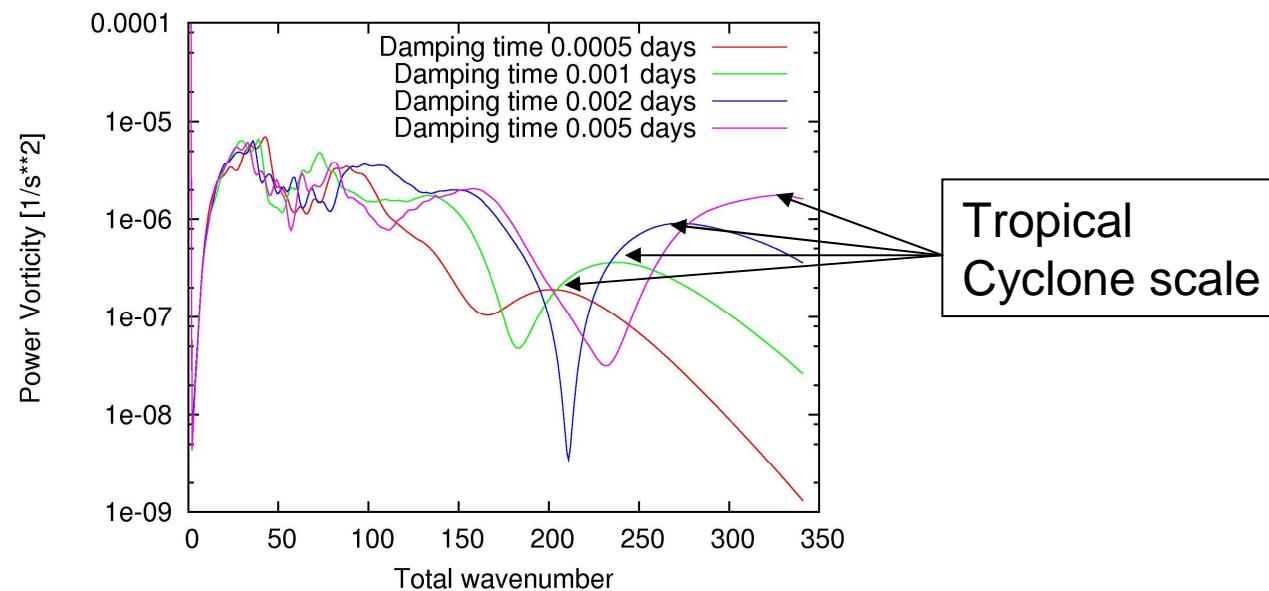
- Inclusion of a water vapor transport equation
- Boundary layer and condensation scheme by Reed and Jablonowski (2011)



## DCMIP experiments

- Specify horizontal hyperdiffusion (6th order) before the experiments.
- Diffusion avoids heap up of energy at the smallest resolved scale.
- Analysis of model spectra for different hyperdiffusion coefficients.
- The tropical cyclone experiment (51) sets the standard

Spectra of vorticity at the lowest model level for experiment 51 at day 1 for different hyperdiffusion decay time scales of the highest wave number



## DCMIP resolutions

The table gives information about the number of longitudes per latitude circle, spectral truncation, equatorial resolution and decay timescale of the highest total wavenumber mode by hyperdiffusion for the various DCMIP resolution configurations

| DCMIP denotation | Longitudes | Truncation | Resolution at equator [km] | Decay timescale for 6 <sup>th</sup> order hyperdiffusion [d] |
|------------------|------------|------------|----------------------------|--|
| LOW              | 256        | T85        | 156.4                      | 0.032  |
| MEDIUM           | 512        | T170       | 78.2                       | 0.008  |
| HIGH             | 1024       | T341       | 39.1                       | 0.002  |
| ULTRA            | 2048       | T682       | 19.6                       | 0.0005   |



# **Test Category 1-X: Tracer Transport Experiments**

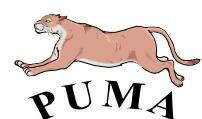
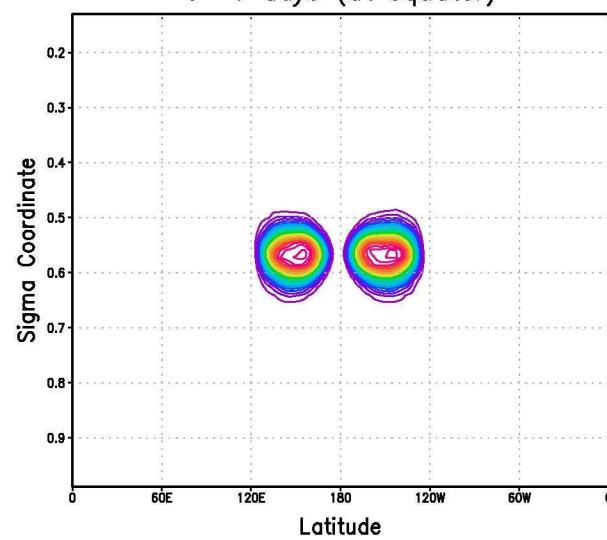
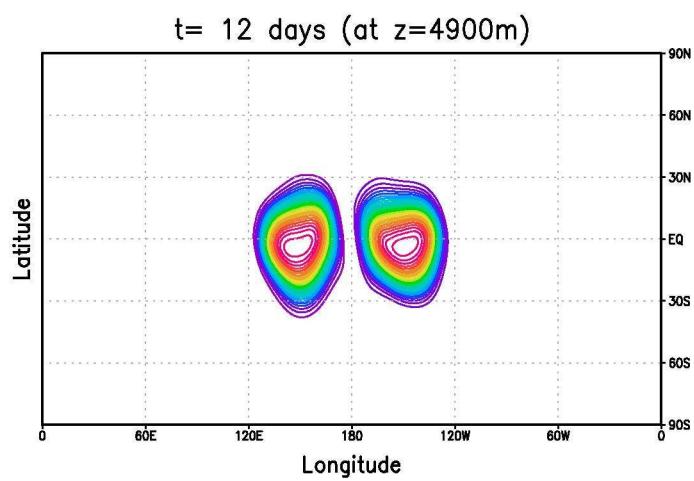
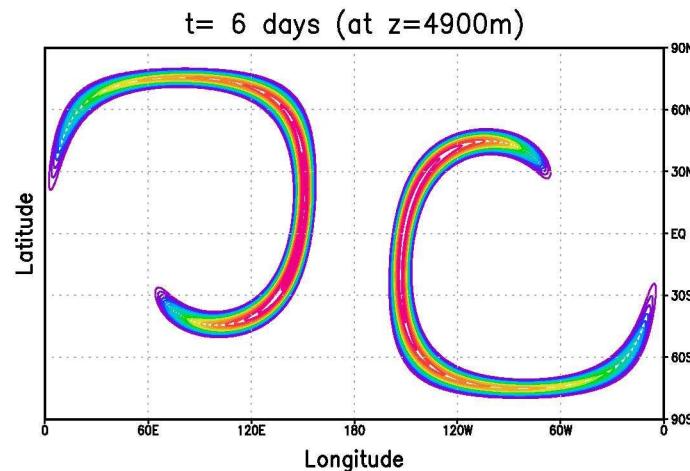
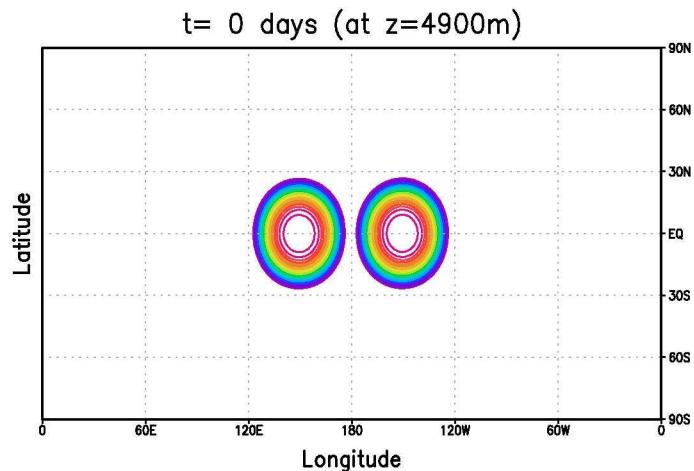
**Test 11:** 3D Deformational Flow

**Test 12:** Hadley-like Merdional Circulation

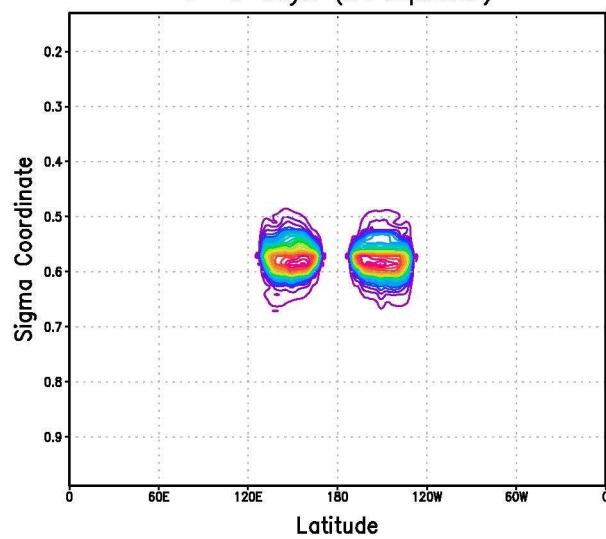
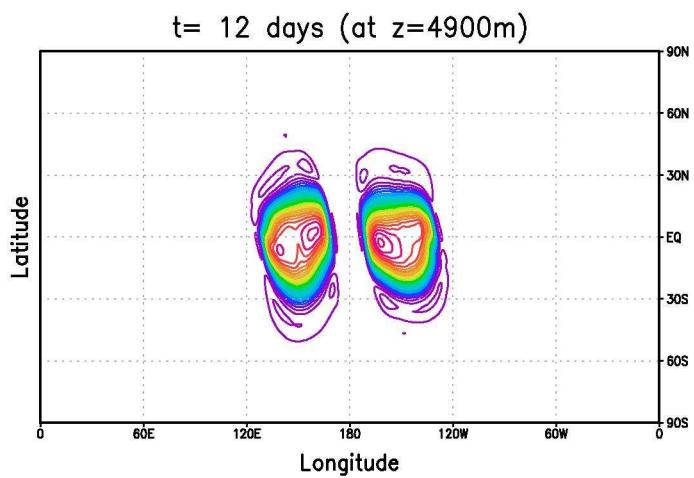
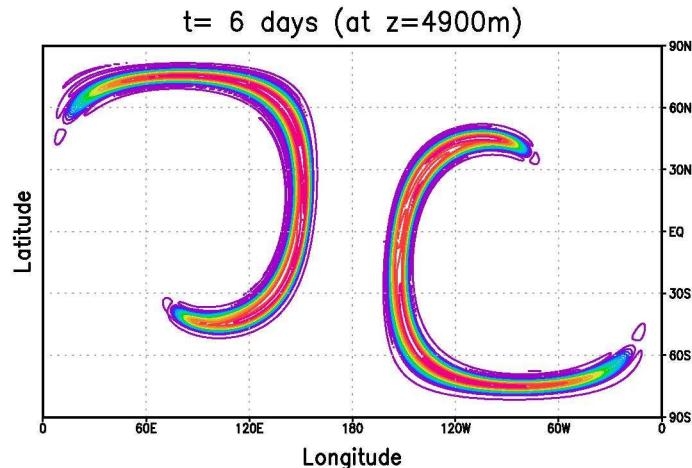
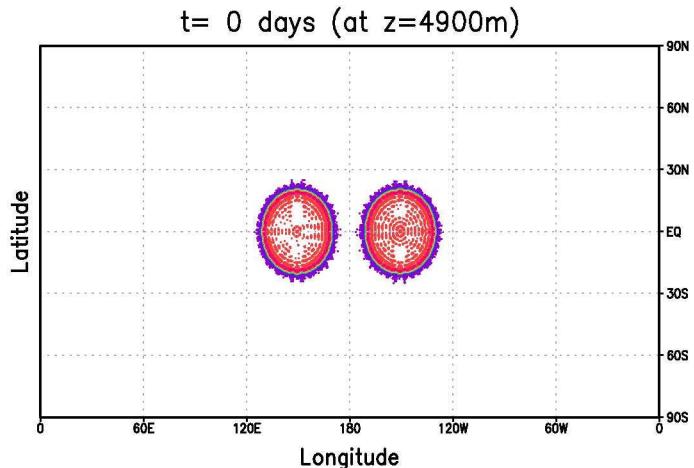
**Test 13:** Horizontal Advection of Thin Cloud-like Tracers in the  
Presence of orography



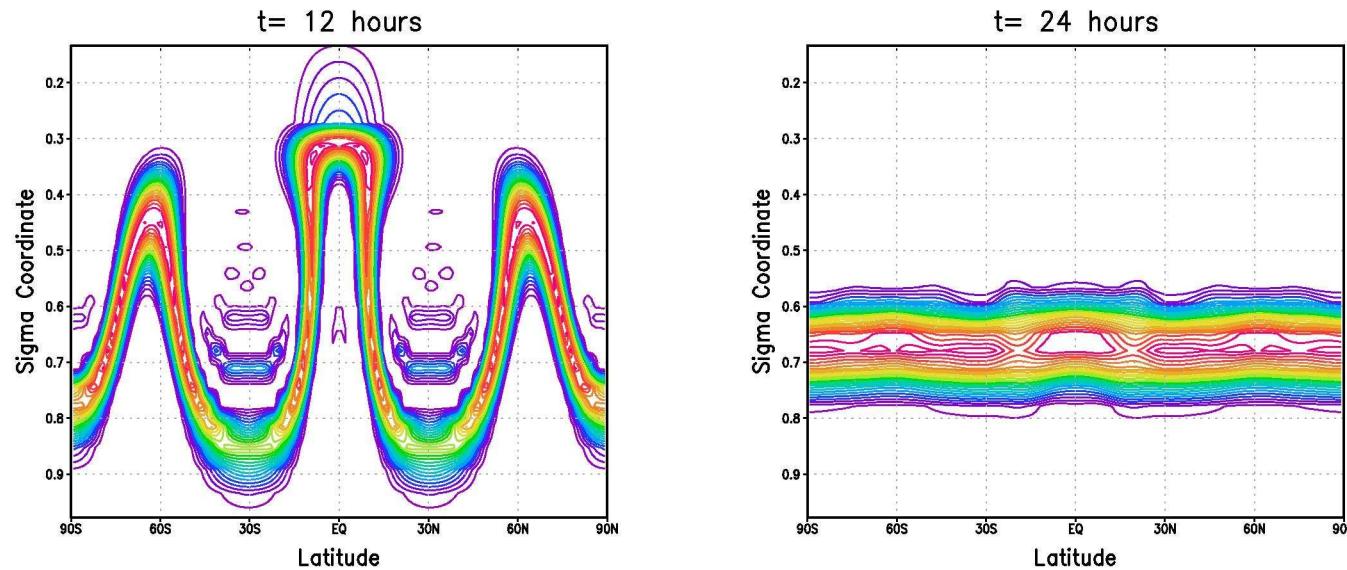
# Experiment 11, tracer q1



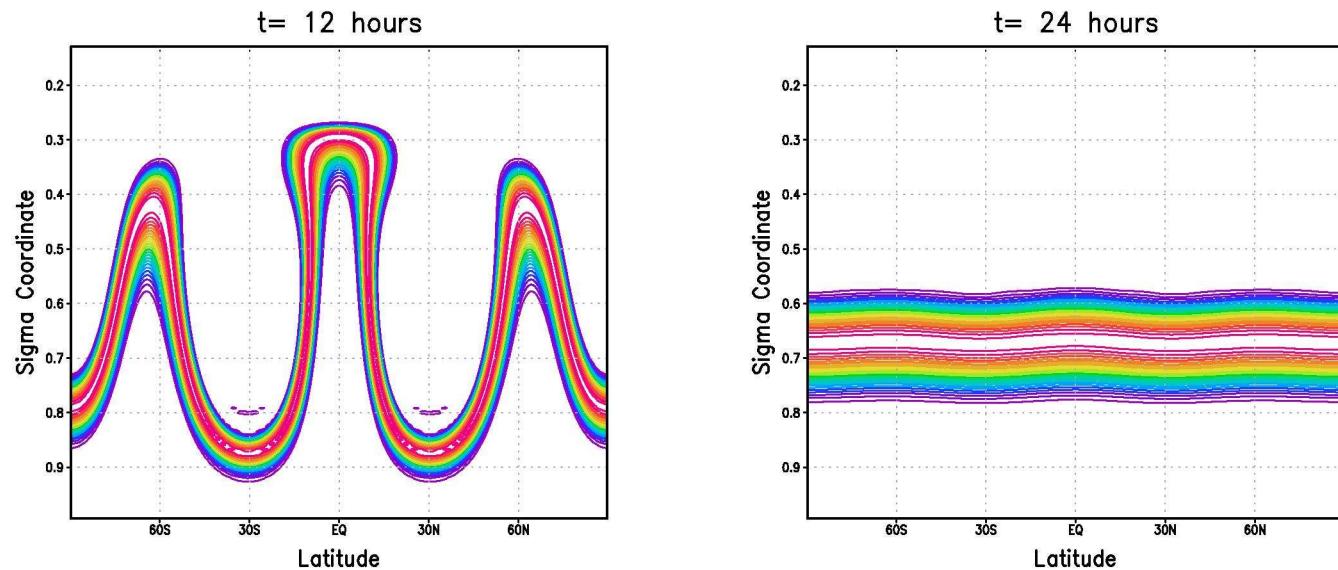
# Experiment 11, tracer q3



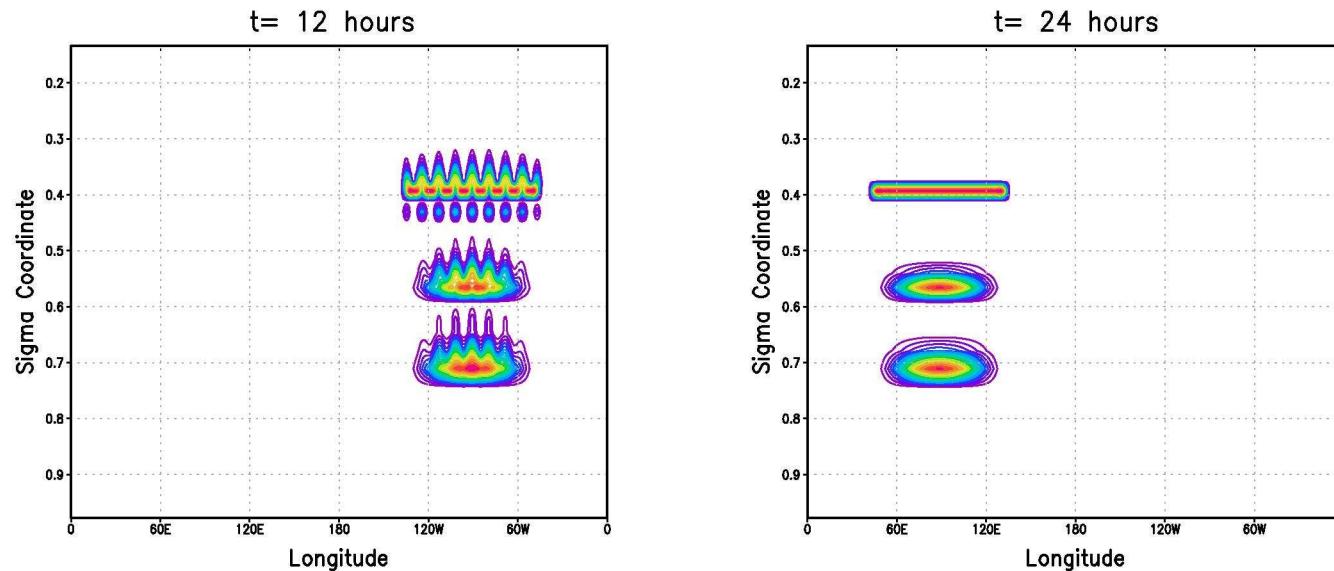
Experiment 12.low.L30, tracer q1



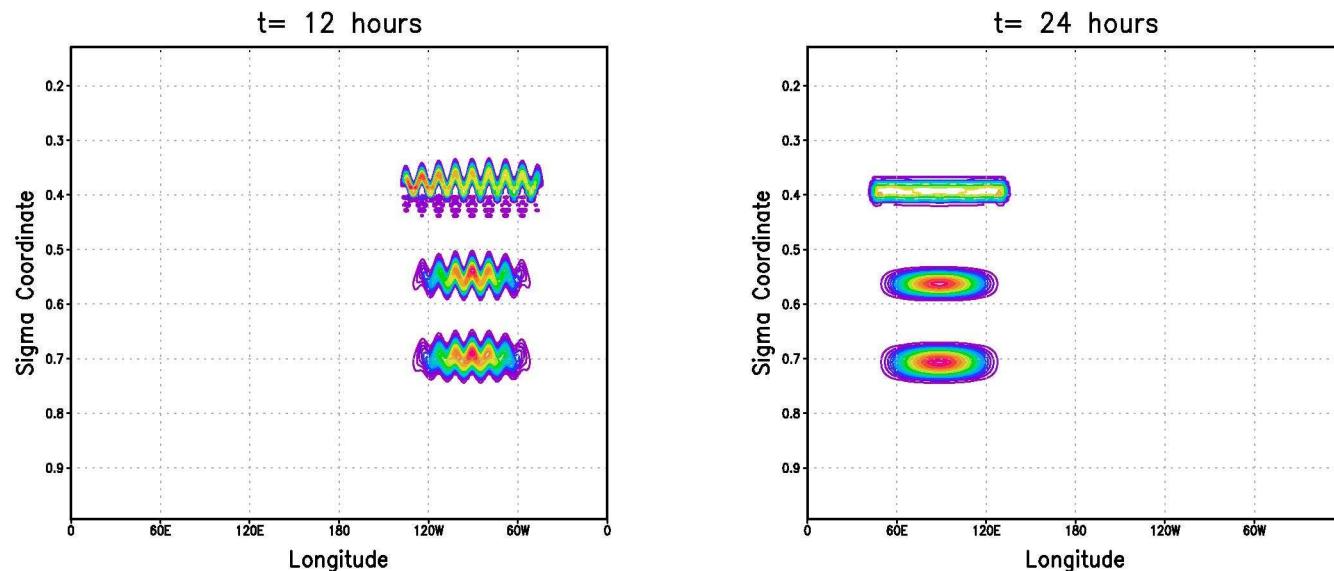
Experiment 12.high.L120, tracer q1



Experiment 13.medium.L30, tracer q4



Experiment 13.medium.L120, tracer q4



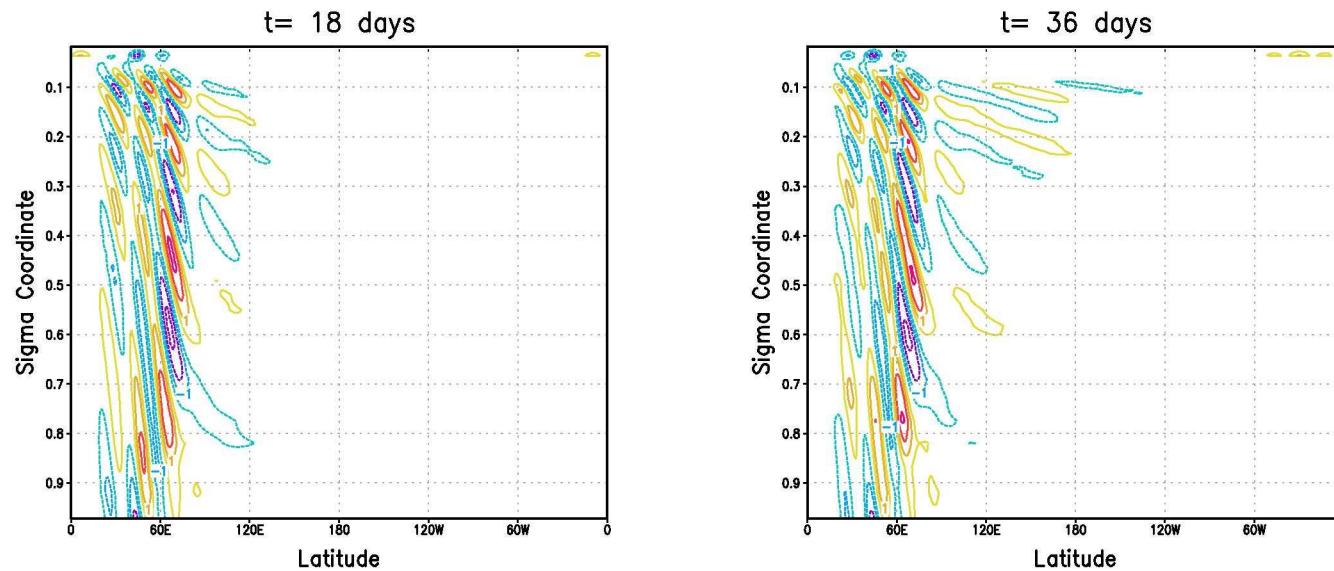
## **Test Category 2-X: Impact of Orography on a Non-Rotating Planet**

**Test 21:** Non-hydrostatic Mountain Waves over a Schär-type Mountain:  
Non-sheared Background Flow

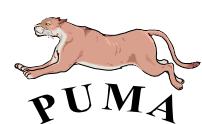
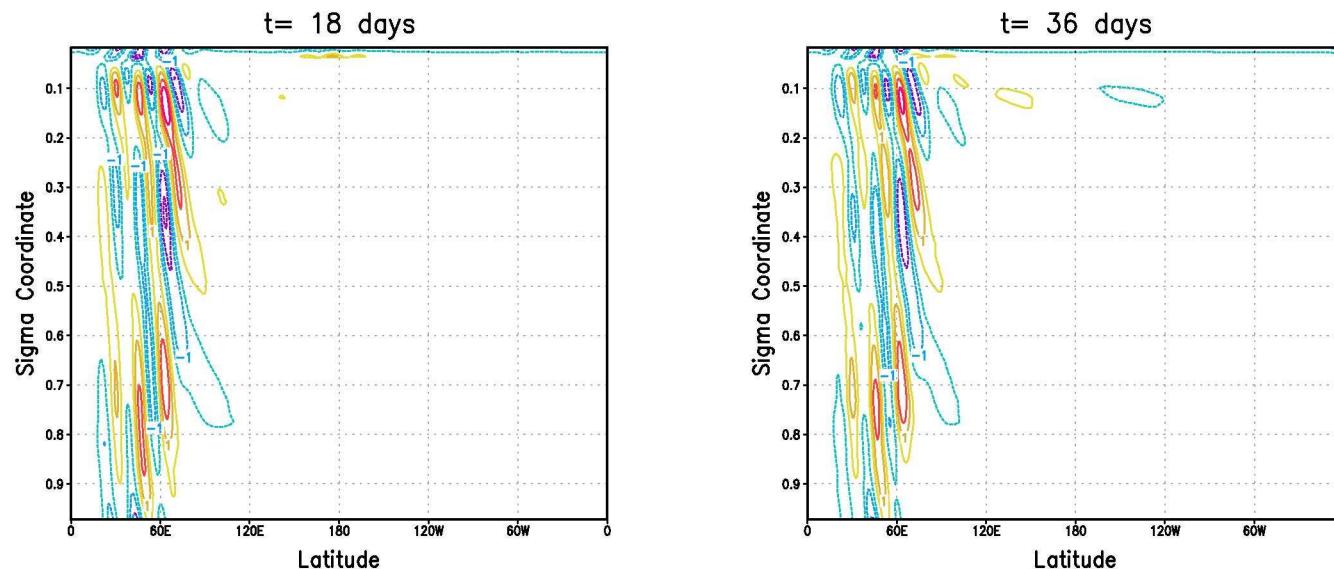
**Test 22:** Non-hydrostatic Mountain Waves over a Schär-type Mountain:  
Sheared Background Flow



Experiment 21 ( $X=1.1574$ ) Temperature anomaly [K]



Experiment 22 ( $X=1.1574$ ) Temperature anomaly [K]



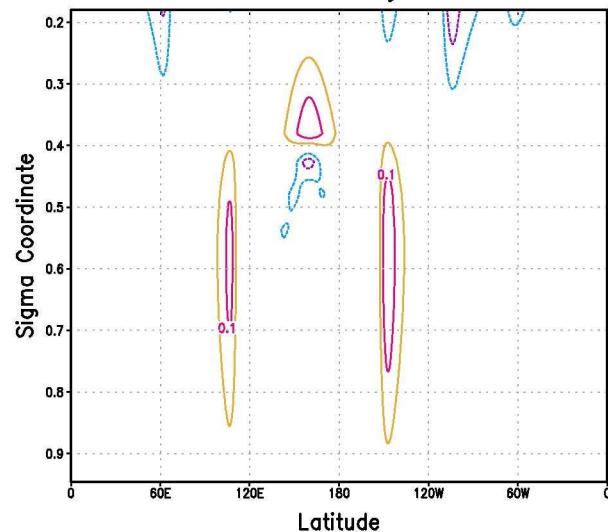
## **Test Category 3: Non-Hydrostatic Gravity Waves**

**Test 31:** Overlaid Potential Temperature Perturbation at the Equator

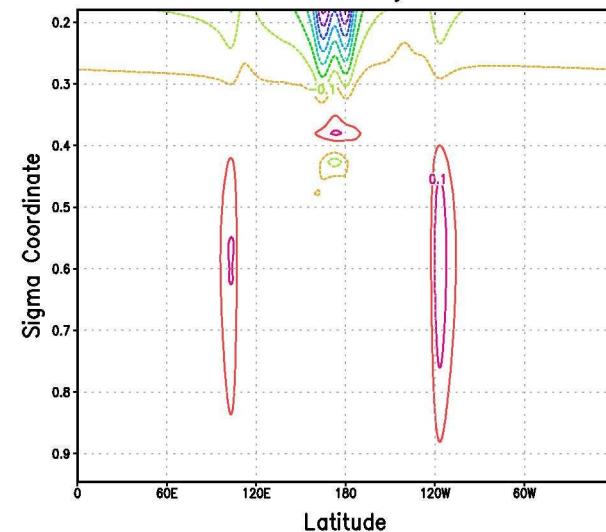


# Experiment 31 ( $X=1.1574$ ) potential temperature anomaly [K]

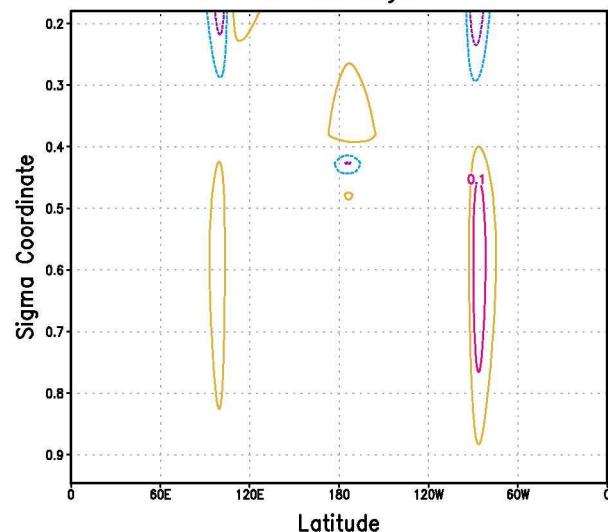
$t = 18$  days



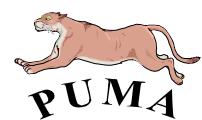
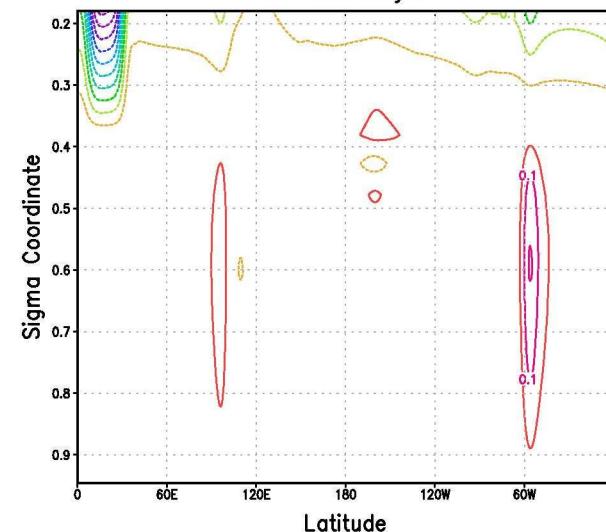
$t = 36$  days



$t = 18$  days



$t = 36$  days

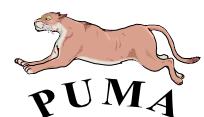


## **Test Category 4-X: Baroclinic Instability**

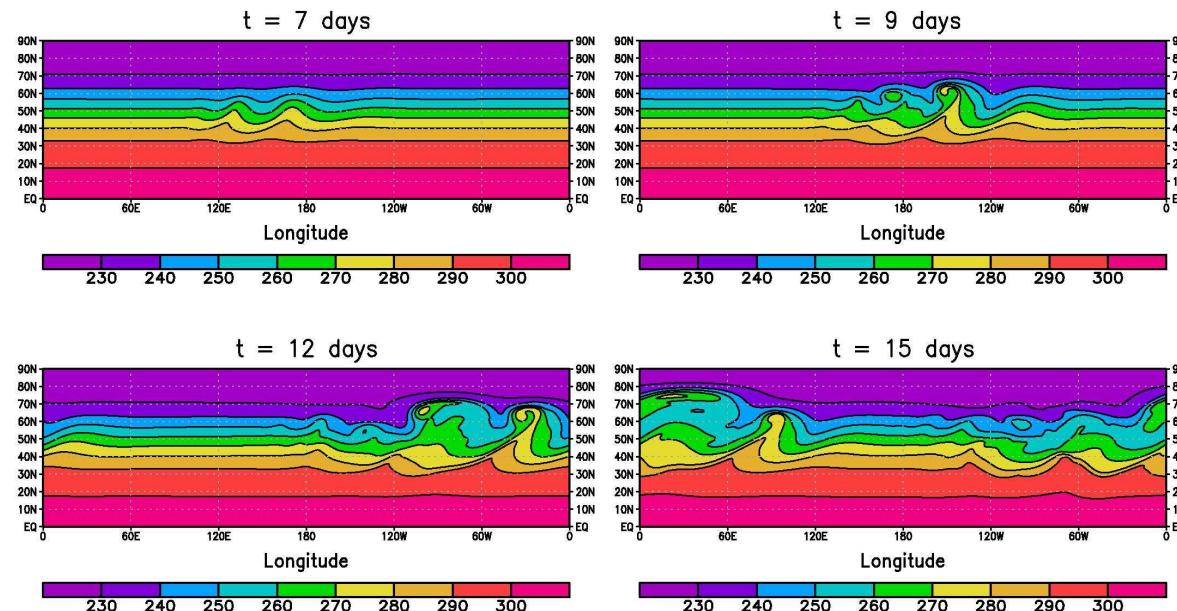
**Test 41-X:** Dry Baroclinic Instability on a Small Planet with Dynamic Tracers [Only 410 ( $X=1$ ) has been conducted]

**Test 42:** Moist Variant of the Baroclinic Wave Test Case with Large-Scale Condensation

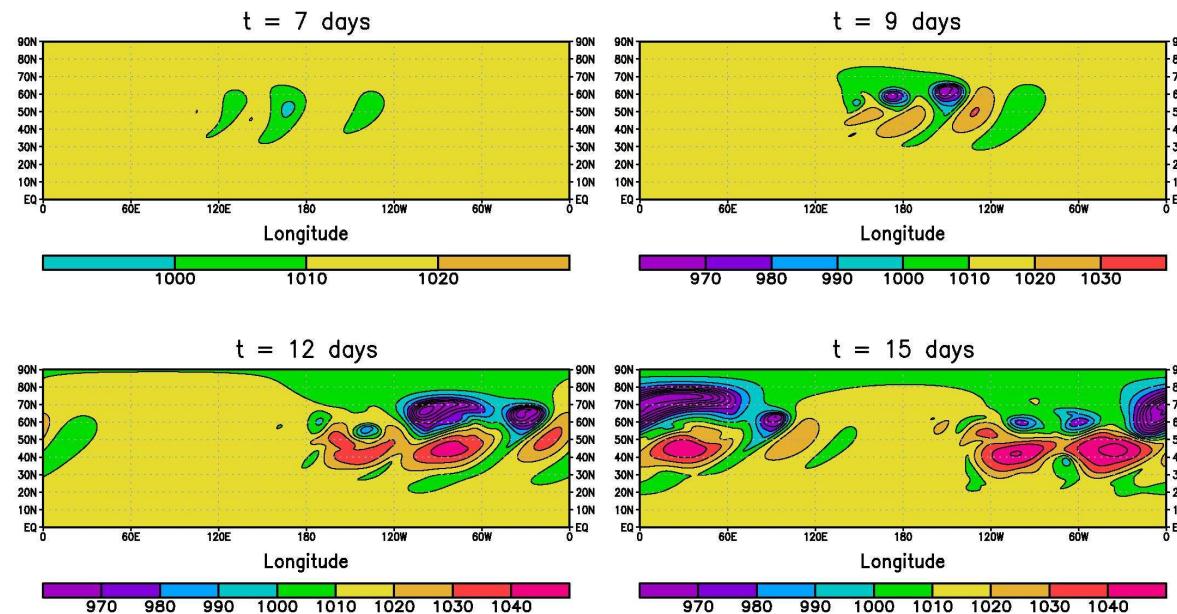
**Test 43:** Moist Variant of the Baroclinic Wave Test Case, driven by “Simple-Physics”



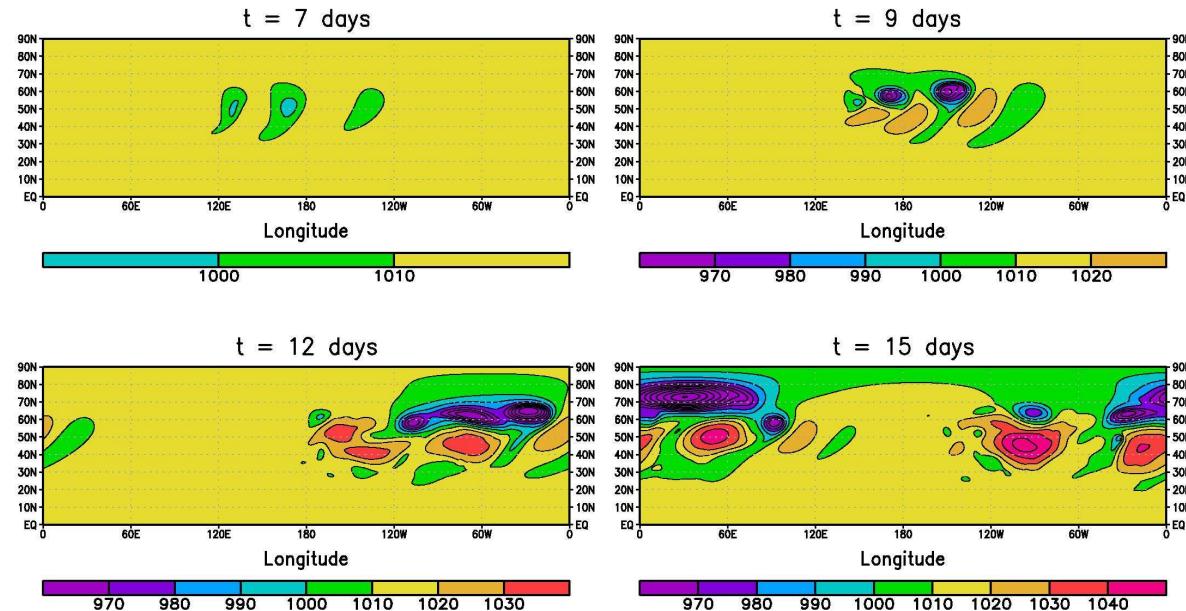
## Experiment 410.medium: 850hPa Temperature (Contour interval 5K)



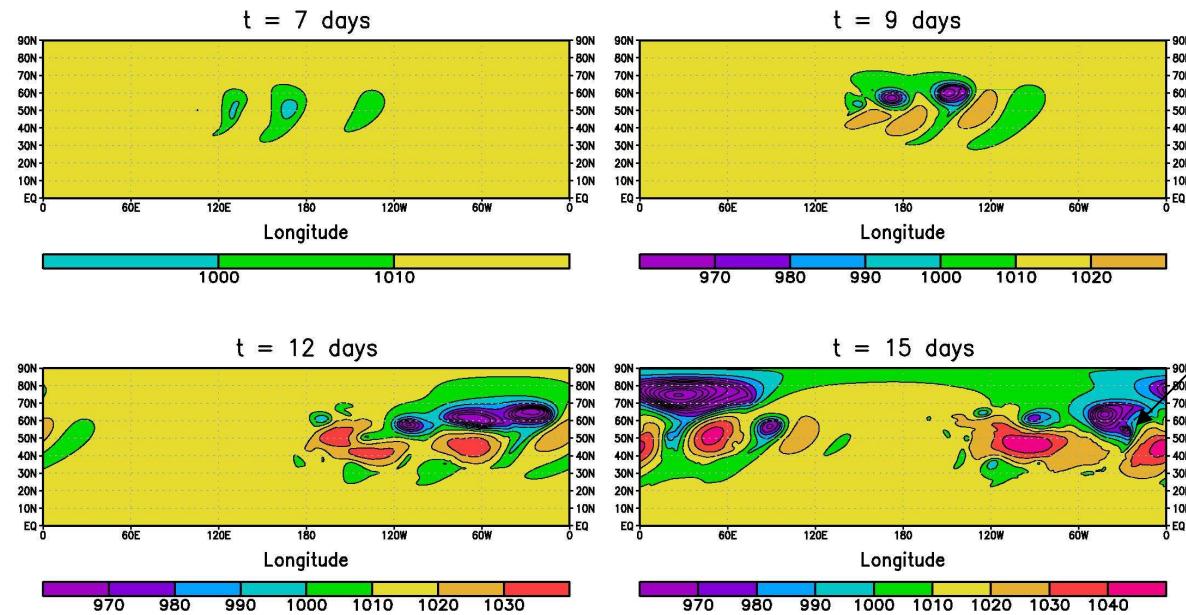
## Experiment 410: Surface pressure (Contour interval 10hPa)



## Experiment 42.medium: Surface pressure (Contour interval 10hPa)



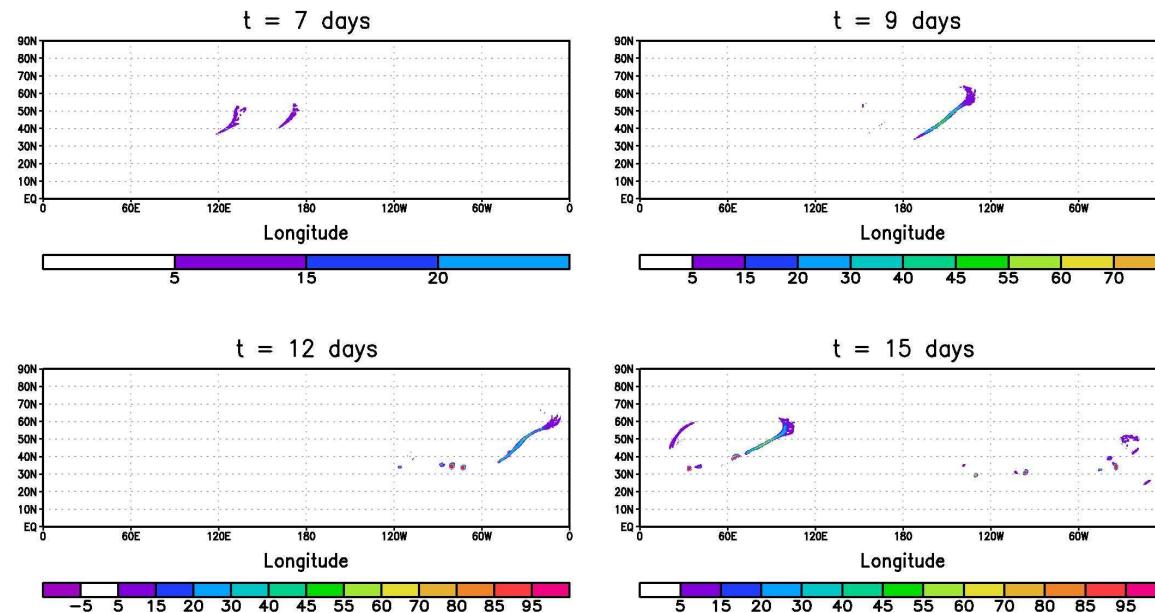
## Experiment 42.high: Surface pressure (Contour interval 10hPa)



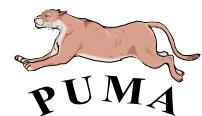
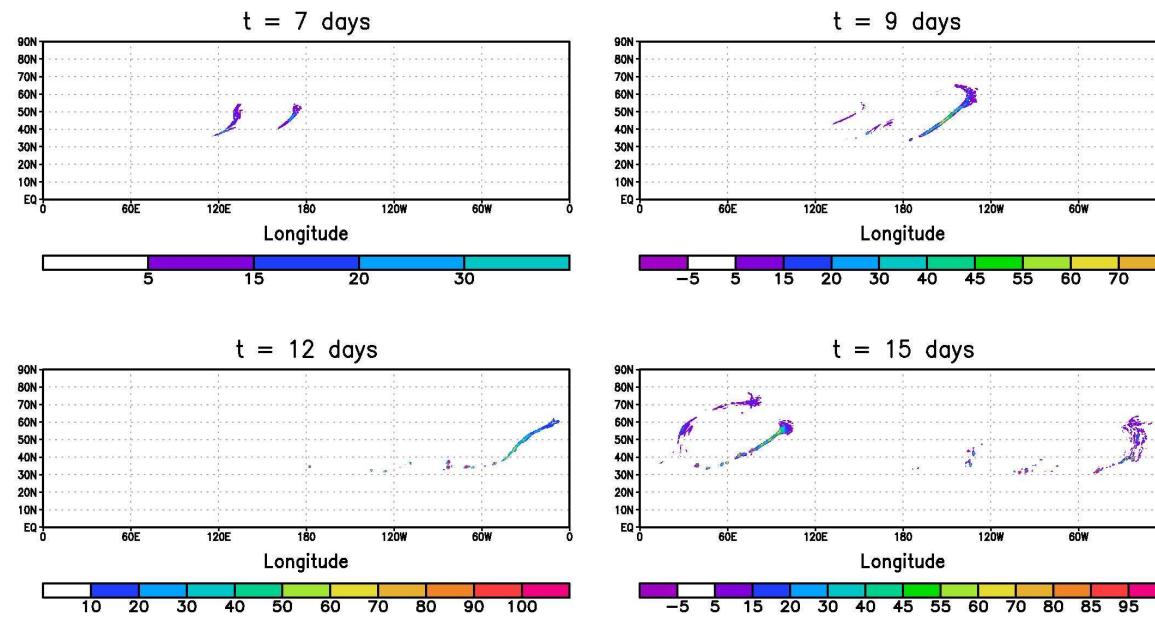
Intense  
Mesocyclone!



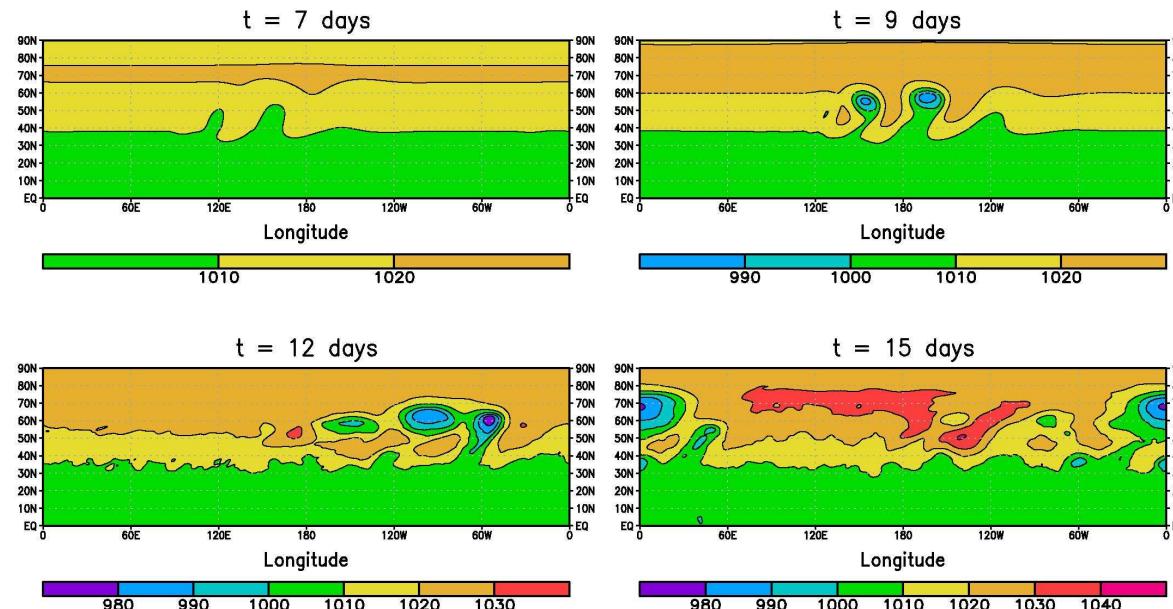
## Experiment 42.medium: Precipitation (Contour interval 5mm/h)



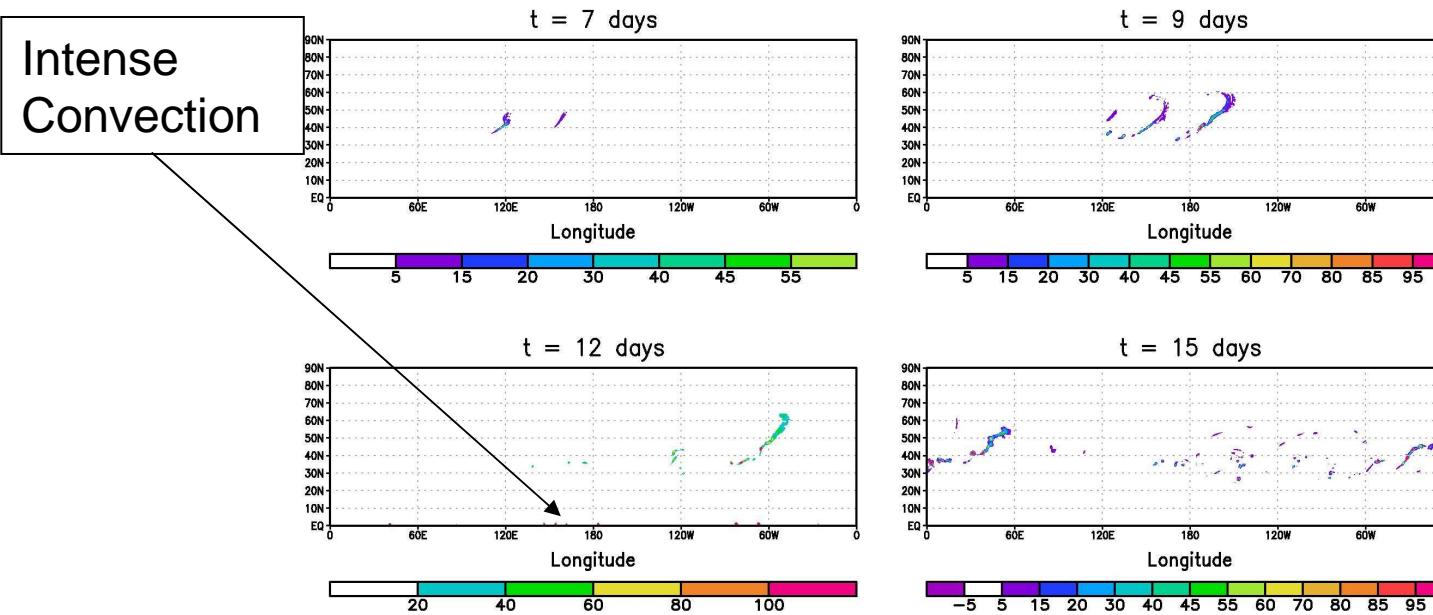
## Experiment 42.high: Precipitation (Contour interval 5mm/h)



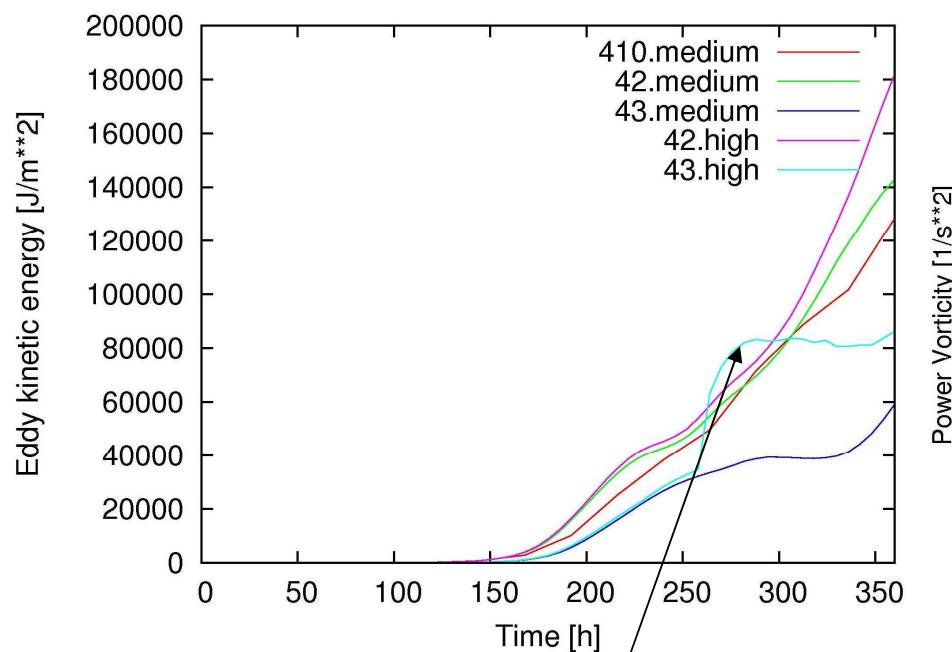
Experiment 43.high: Surface pressure (Contour interval 10hPa)



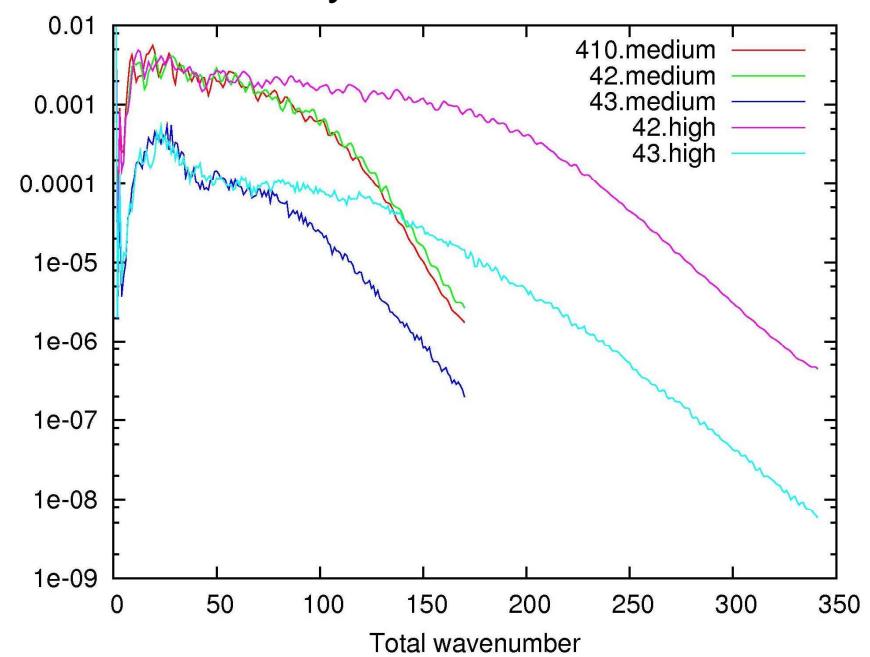
Experiment 43.high: Precipitation (Contour interval 5mm/h)



Time development of eddy kinetic energy



Spectra of vorticity at the lowest level  
at day 15



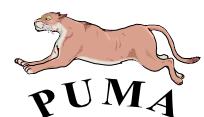
Development of convection at the  
Equator (ITCZ-development?)



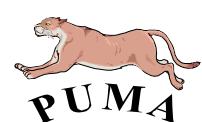
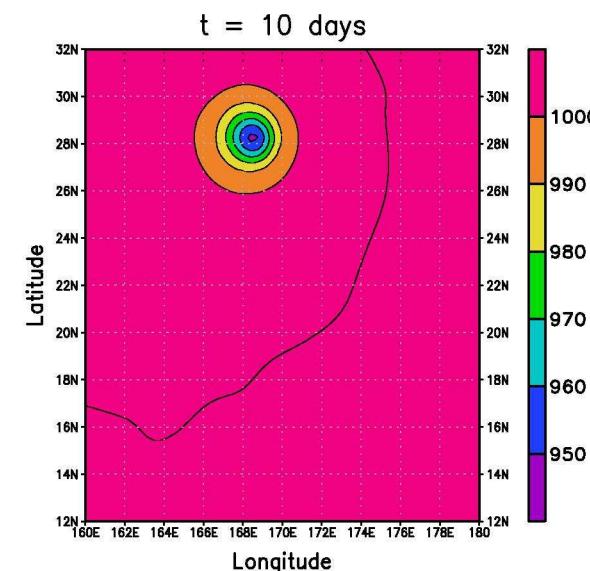
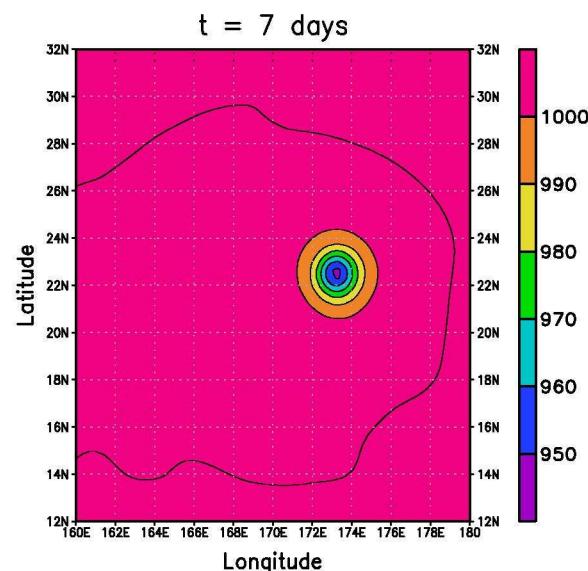
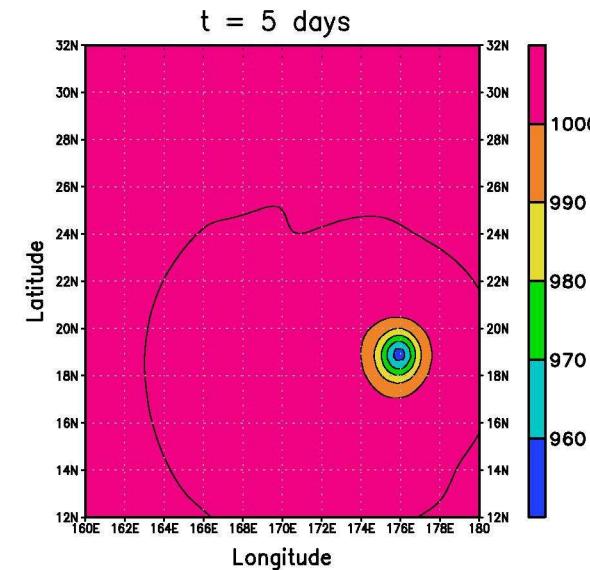
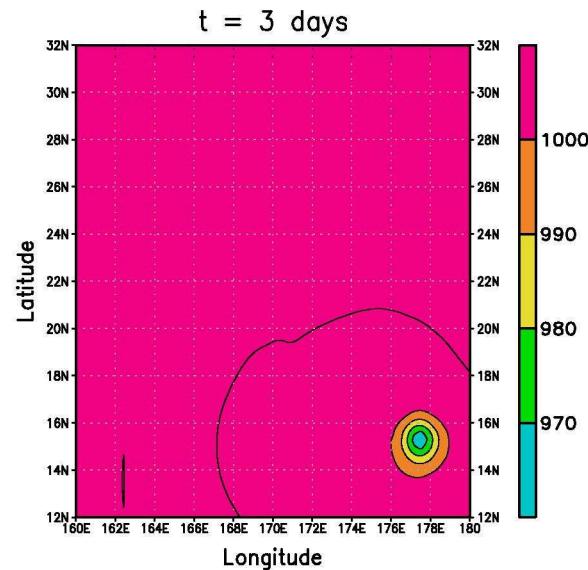
# **Test Category 5-X: Idealized Tropical Cyclone Experiments**

**Test 51:** Coupling to the “Simple-Physics” Physical Parameterization

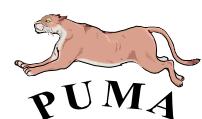
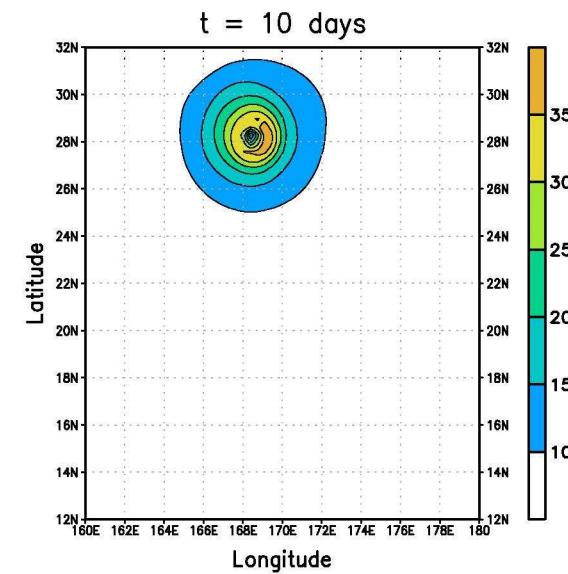
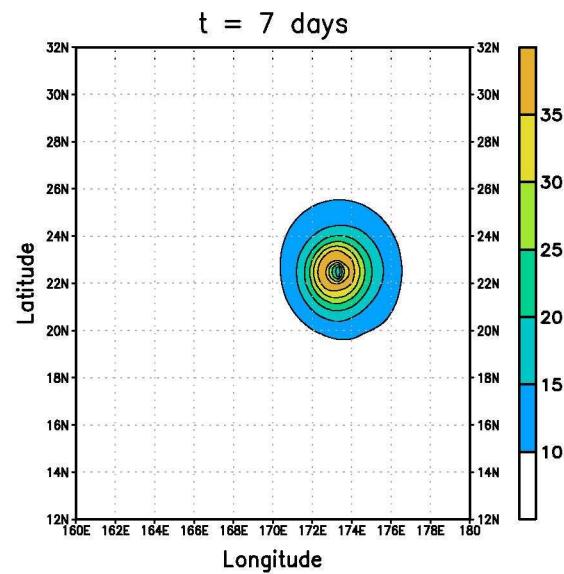
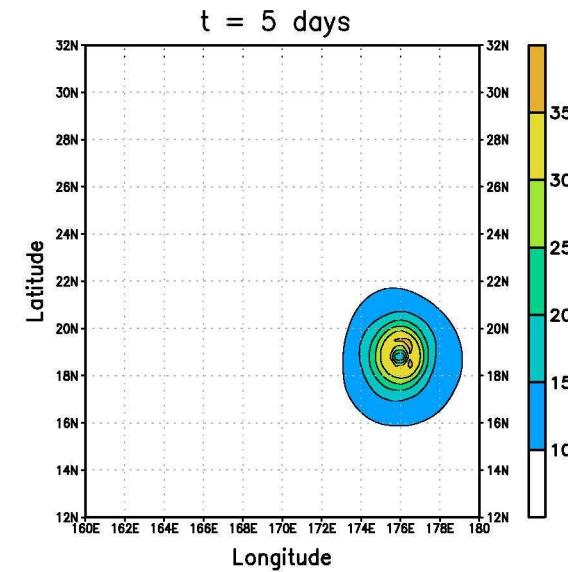
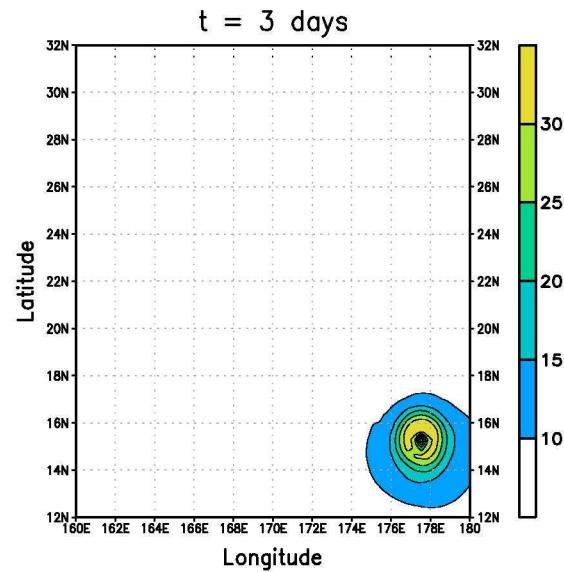
**Test 52:** Using the Model’s Full-Physics Aqua-Planet Mode with a constant SST of 28°C (not conducted)



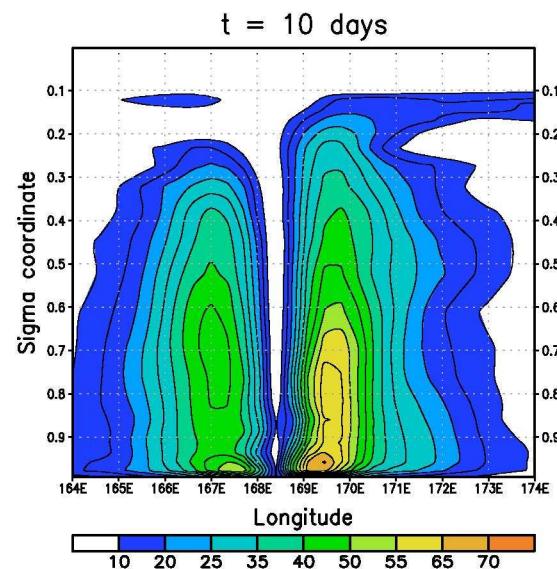
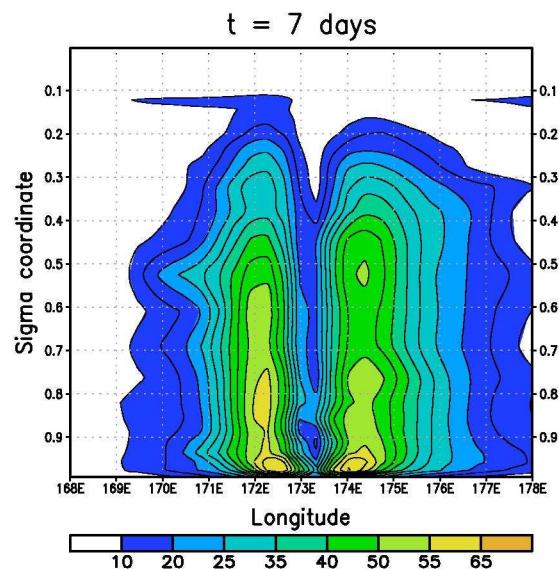
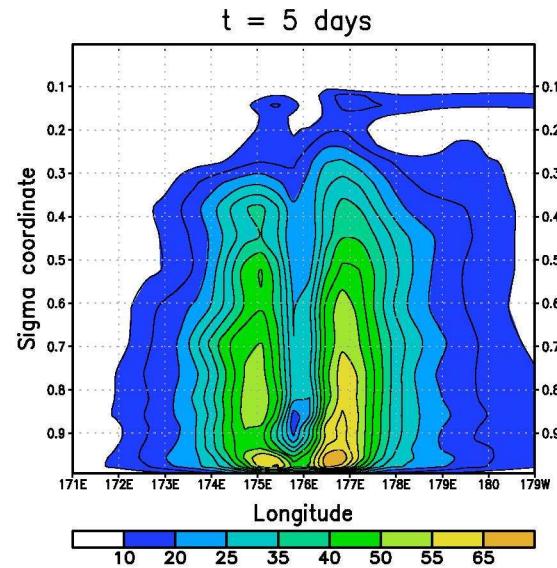
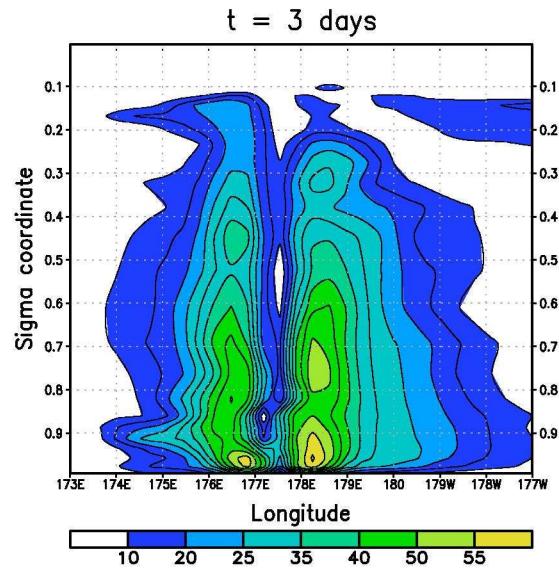
# Experiment 51.high: Surface pressure (Contour interval 10hPa)



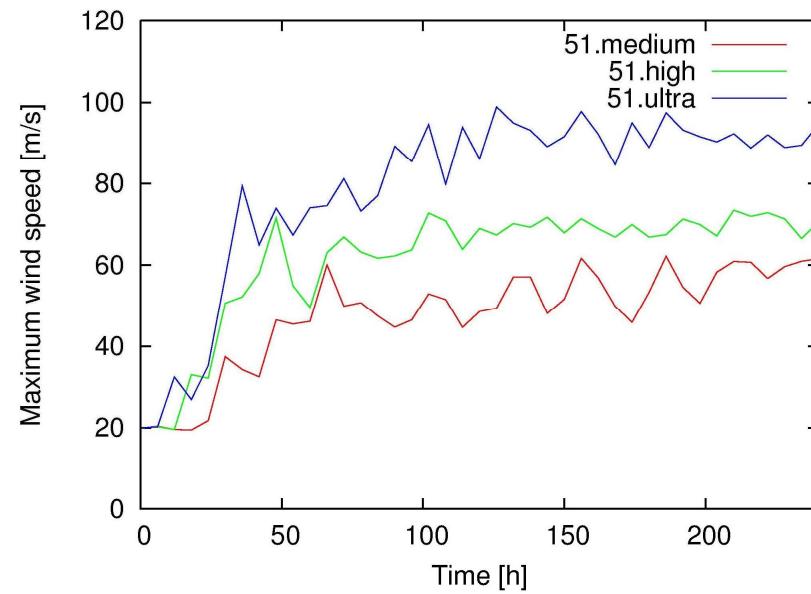
# Experiment 51.high: Wind speed lowest level (Contour interval 5m/s)



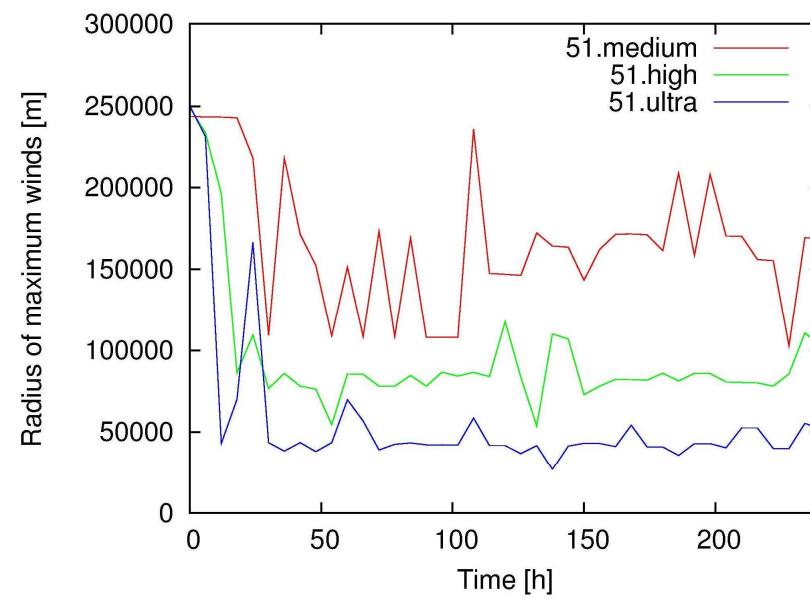
## Experiment 51.high: Wind speed (Contour interval 5m/s)



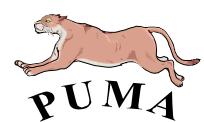
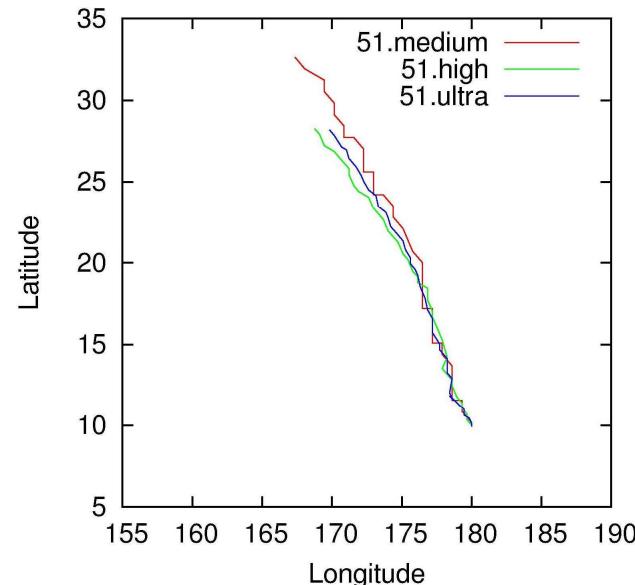
### Development of maximum wind speed



### Development of the radius of maximum wind



### Cyclone tracks



## Conclusion

- For a large enough resolution the transport of the tracer within PUMA is reasonable. Difficulties arise for discontinuous fields since the spectral representation of such fields is problematic.
- The pure Sigma-coordinate in PUMA is possibly not a good choice for simulation of mountain induced gravity waves. Furthermore, the upper boundary condition of a rigid lid can hardly be mimicked within PUMA.
- Moist and dry baroclinic instability can reasonably be simulated by PUMA. Some phenomena (intense mesocyclones, ITCZ formation) arise only for the high resolution.
- The idealized tropical cyclogenesis experiments appear plausible. With increasing resolution the intensity increases and the radius of maximum winds decreases.

