
pyLCSIM Documentation

Release 0.1

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DOCUMENTATION FOR PYLCSIM

1.1 Introduction

pyLCSIM is a python package to simulate X-ray lightcurves from coherent signals and power spectrum models.

Coherent signals can be specified as a sum of one or more sinusoids, each with its frequency, pulsed fraction and phase shift; or as a series of harmonics of a fundamental frequency (each with its pulsed fraction and phase shift).

Power spectra can be simulated from a model of the power spectrum density (PSD), using as a template one or more of the built-in library functions. The user can also define his/her custom models. Models are additive.

A PDF version of these notes is available [here](#).

Warning: the current release (0.x.y) is HIGHLY EXPERIMENTAL! Use at your own risk...

1.2 Prerequisites

pyLCSIM requires [Numpy](#) (at least v1.8) and [Astropy](#) (at least v0.3).

[Matplotlib](#) is highly recommended if you want to plot your simulations.

1.3 Installation

The package can be downloaded [here](#).

The installation follows the usual steps:

```
$ tar xzvf pyLCSIM-0.1.2.tar.gz
```

```
$ cd pyLCSIM-0.1.2
```

```
$ python setup.py install
```

The last step may require administrator privileges.

1.4 Example 1

Let's begin with a PSD model simulation.

We import the usual packages:

```
import matplotlib.pyplot as plt
import numpy as np
import pyLCSIM
```

We assume that our source has a rate of 30000 counts/s. Moreover, we have a 5000 counts/s background rate, and we have made a 50 s exposure. Our observation has a time resolution of 10 ms. We want to simulate a QPO at a frequency of 10 Hz, superimposed to a continuum modelled as a smoothly-varying broken power law (spectral indices 1 and 2, with a steepness change at 1 Hz). The required fractional RMS variation of the signal is 1%:

```
rate_src    = 30000.0
rate_bkg    = 5000.0
t_exp       = 50.0
dt          = 0.01
frms        = 0.01
```

The total bins of the lightcurve are therefore:

```
nbins = long(t_exp/dt)
```

The simulation follows as:

```
# Instantiate a simulation object
sim = pyLCSIM.Simulation()

# Add two PSD models: a smooth broken power law and a Lorentzian representing a QPO.
# See the documentation for details.
sim.addModel('smoothbknpo', [1., 1., 2., 1.])
sim.addModel('lorentzian', [10., 1., 100., 10])

# Run the simulation
sim.run(dt, nbins, rate_src, rms=frms)

# Add Poisson noise to the light curve
sim.poissonRandomize(dt, rate_bkg)

# Get lightcurve and power spectrum as 1-D arrays
time, rate = sim.getLightCurve()
f, psd = sim.getPowerSpectrum()
```

Done! We can save the results as FITS files:

```
# Save FITS files with lightcurve and spectrum
pyLCSIM.saveFITSLC("myLC.fits", time, rate)
pyLCSIM.saveFITSPSD("myPSD.fits", f, psd)
```

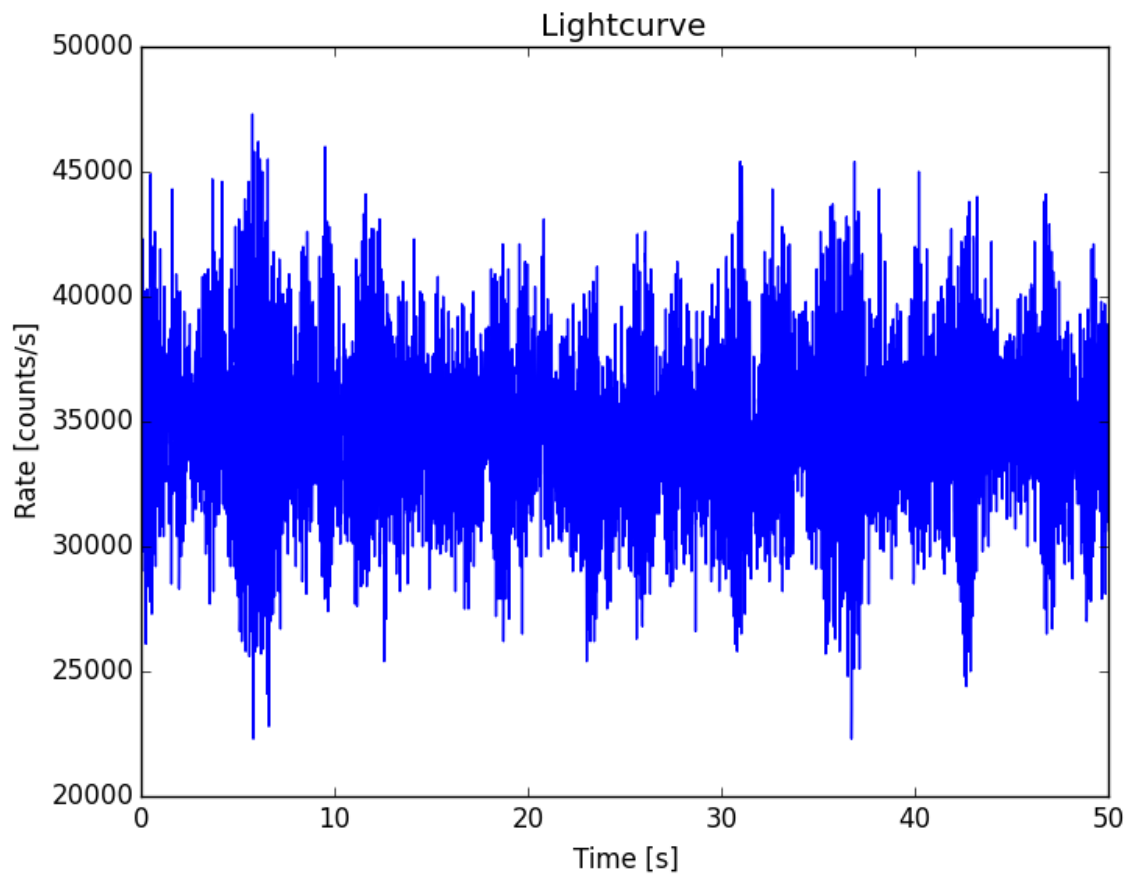
and view the results:

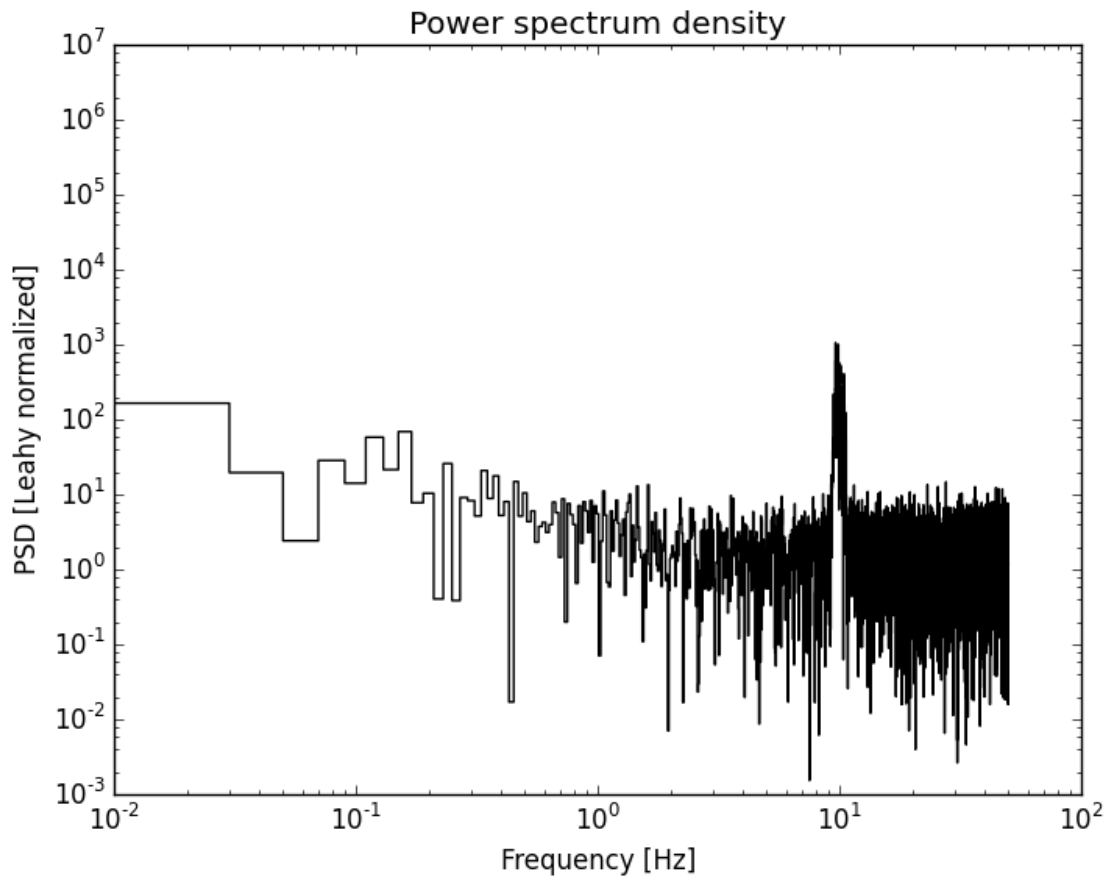
```
# Plot the lightcurve and power spectrum
fig0 = plt.figure()
plt.plot(time, rate)
plt.xlabel("Time [s]")
plt.ylabel("Rate [counts/s]")
plt.title("Lightcurve")

fig1 = plt.figure()
plt.loglog(f, psd, drawstyle='steps-mid', color='black')
plt.xlabel("Frequency [Hz]")
plt.ylabel("PSD [Leahy normalized]")
plt.title("Power spectrum density")
```



```
plt.show()
```





1.5 Example 2

The following example shows the simulation of a coherent signal using a sum of sinusoids:

```
import matplotlib.pyplot as plt
import numpy as np
import pyLCSIM
```

```
rate_src    = 30000.0
rate_bkg    = 5000.0
t_exp       = 1.0
dt          = 0.0001
nbins = long(t_exp/dt)
```

Note the different exposure time (1 s) and time resolution (100 us):

```
# Instantiate a simulation object, this time as coherent
sim = pyLCSIM.Simulation(kind='coherent')
```

```
# Run the simulation, using:
```

```
# four sinusoidal frequencies: 340, 550, 883, 1032 Hz;
```

```
# with pulsed fractions 10%, 5%, 7% and 15% respectively;
```

```
# the third frequency has a 35 degree phase shift with respect to the others
```

```
sim.run(dt, nbins, rate_src, freq=[340, 550, 883, 1032], amp=[0.1, 0.05, 0.07, 0.15], phi=[0., 0, 35
```

```

# Add Poisson noise to the light curve
sim.poissonRandomize(dt, rate_bkg)

# Get lightcurve and power spectrum as 1-D arrays
time, rate = sim.getLightCurve()
f, psd = sim.getPowerSpectrum()

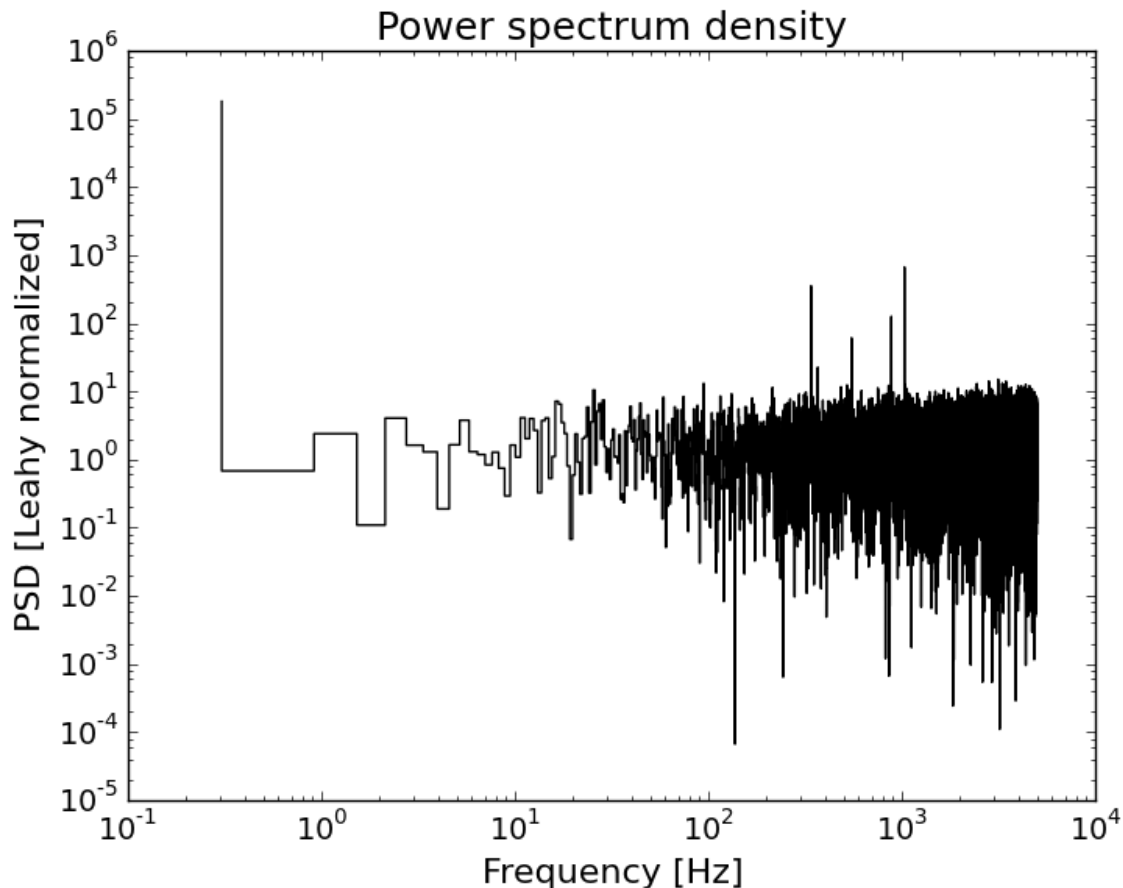
# Plot the lightcurve and power spectrum
fig0 = plt.figure()
plt.plot(time, rate)
plt.xlabel("Time [s]")
plt.ylabel("Rate [counts/s]")
plt.title("Lightcurve")

fig1 = plt.figure()
plt.loglog(f, psd, drawstyle='steps-mid', color='black')
plt.xlabel("Frequency [Hz]")
plt.ylabel("PSD [Leahy normalized]")
plt.title("Power spectrum density")

# Save FITS files with lightcurve and spectrum
pyLCSIM.saveFITS("myLC.fits", time, rate)
pyLCSIM.saveFITS("myPSD.fits", f, psd)

plt.show()

```



1.6 Example 3

Finally, an example with a fundamental frequency and two harmonics:

```
import matplotlib.pyplot as plt
import numpy as np
import pyLCSIM

rate_src      = 300000.0
rate_bkg      = 3000.0
t_exp         = 1.0
dt            = 0.0001
phase_shift   = 0.
nbins         = long(t_exp/dt)

print nbins

# Instantiate a simulation object, this time as coherent
sim = pyLCSIM.Simulation(kind='coherent')

# Run the simulation:
# Fundamental at 500 Hz, 3 harmonics (500, 1000, 1500 Hz)
# with pulsed fractions 10%, 5% and 15% respectively
sim.run(dt, nbins, rate_src, freq=500, nha=3, amp=[0.1, 0.05, 0.15])

# Add Poisson noise to the light curve
sim.poissonRandomize(dt, rate_bkg)

# Get lightcurve and power spectrum as 1-D arrays
time, rate = sim.getLightCurve()
f, psd = sim.getPowerSpectrum()

# Plot the lightcurve and power spectrum
fig0 = plt.figure()
plt.plot(time, rate)

fig1 = plt.figure()
plt.loglog(f, psd, drawstyle='steps-mid', color='black')

# Save FITS files with lightcurve and spectrum
pyLCSIM.saveFITSLC("myLC.fits", time, rate)
pyLCSIM.saveFITSPSD("myPSD.fits", f, psd)

plt.show()
```

1.7 Main module

A python package to simulate X-ray lightcurves from coherent signals and power spectrum models.

```
class pyLCSIM.Simulation (kind='psd')
    Main simulation class.
```

History: v1.2: Added reset method. Riccardo Campana, 2014. v1.1: Bugfix. Riccardo Campana, 2014. v1: Initial python implementation. Riccardo Campana, 2014.

addModel (*modelName, modelParams*)

Append simulation model to model dictionary

getLightCurve ()

Get lightcurve as time and rate arrays

getPowerSpectrum ()

Get power spectrum as frequency and power arrays

info ()

Prints simulation informations

reset ()

Reset the simulation, emptying models array and other members

run (*dt, nbins_old, mean, rms=None, freq=None, nha=None, amp=None, phi=None, verbose=False*)

Run the simulation

`pyLCSIM.lcsinusoid` (*dt=1.0, nbins=65536, mean=0.0, freq=None, nha=1, amp=None, phi=None*)

Generate coherent signals as a sequence of sinusoids (if `len(freq) > 1`) or of the fundamental frequency plus `nha-1` harmonics (if `len(freq) == 1`), each with normalized pulsed fraction `amp[i]`.

Kwargs: `dt`: time resolution of the lightcurve to be simulated (default: 1.0).

`nbins`: Number of bins of the simulated lightcurve (default: 65536).

`freq`: if float: frequency of the fundamental harmonic, if array: frequencies of sinusoids.

`nha`: number of harmonics (>1)

`amp`: array with `nha/nfreq` elements; pulsed fraction for each frequency

`phi`: array with `nha/nfreq` elements; phases (in degrees!) for each frequency

Returns: `time`: time array

`rate`: array of count rates

History:

v1: Initial python implementation, from the IDL procedure `lcharmonics.pro v0.0.3` by I. Donnarumma & R. Campana. Riccardo Campana, 2014.

`pyLCSIM.lcpsd` (*dt=1.0, nbins=65536, mean=0.0, rms=1.0, seed=None, models=None, phase_shift=None, verbose=False*)

Simulate a light-curve with a general power spectrum shape. For the underlying algorithm see: J. Timmer & M. Koenig, "On generating power law noise", A&A, 300, 707-710 (1995).

Kwargs: `dt`: time resolution of the lightcurve to be simulated

`nbins`: Number of bins of the simulated lightcurve (default:65536). Should be power of two for optimum performance (FFT...)

`mean`: Mean count rate of the simulated lightcurve (default: 0.).

`rms`: Total fractional RMS of the simulated lightcurve.

`seed`: Seed for the random number generator.

`models`: List of tuples, each containing:

1. Name of a function returning the desired PSD shape,

2. Parameters of the model (argument to model). Total model is the sum of these tuples.

`phase_shift`: Constant phase shift (in degrees) to the FFT.

`verbose`: If True, prints some debugging information.

Returns: time: time array.

rate: array of count rates.

History: v1: Initial python implementation, based on AITLIB IDL procedure timmerlc.pro. Riccardo Campana, 2014.

`pyLCSIM.poisson_randomization` (*rate*, *dt=1.0*, *bkg=0.0*, *seed=None*)

Poisson randomization of the rate array.

Args: rate: input array of count rates (in cts/s).

Kwargs: dt: time resolution of the lightcurve to be simulated.

bkg: Mean count rate of the simulated lightcurve (default: 0.0).

seed: Seed for the random number generator.

Returns: newrate: Array of Poisson randomized count rates of length n.

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

`pyLCSIM.psd` (*time*, *rate*, *norm='leahy'*)

Returns power spectral density from a (real) time series with Leahy normalization.

Args: time: array of times (evenly binned).

rate: array of rate in counts/s.

Kwargs: norm: Normalization (only Leahy for now).

Returns: f: array of frequencies.

p: power spectrum.

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

`pyLCSIM.rebin` (*x*, *y*, *nbins*, *mode='rate'*)

Linearly rebins the (x, y) 1-D arrays using #nbins new bins, taking for each new element the mean of the corresponding x elements, and the sum (if mode=='counts') or the mean (if mode=='rate') of the corresponding y elements.

Args: x: x array (same length as y)

y: y array (same length as x)

nbins: Number of new bins. Should be a factor of len(x) == len(y).

Kwargs: mode: 'counts' or 'rate'; returns sum or mean of y-array elements

Returns: xreb: rebinned x array

yreb: rebinned y array

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

`pyLCSIM.logrebin` (*x*, *y*, *nbins*, *mode='rate'*)

Logarithmically rebins the (x, y) 1-D arrays using #nbins new bins, taking for each new element the logarithmic mean of the corresponding x elements, and the sum (if mode=='counts') or the mean (if mode=='rate') of the corresponding y elements.

Args: x: x array (same length as y)

y: y array (same length as x)

nbins: Number of new bins.

Kwargs: mode: 'counts' or 'rate'; returns sum or mean of y-array elements

Returns: xreb: rebinned x array

yreb: rebinned y array

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

`pyLCSIM.saveFITS LC (outfile name, time, rate, clobber=True)`

Produce an output FITS file containing a lightcurve.

Args: outfile name: Name of the output FITS file

time: Array of times

rate: Array of count rates

Kwargs: clobber: if True, overwrites existing files with same name

Returns: none

History: v2: OGIP-compliance (OGIP 93-003). Riccardo Campana, 2014 v1: Initial python implementation. Riccardo Campana, 2014

`pyLCSIM.saveFITS PSD (outfile name, freq, psd, clobber=True)`

Produce an output FITS file containing a power spectrum.

Args: outfile name: Name of the output FITS file

freq: Array of frequencies

psd: Array of power spectrum

Kwargs: clobber: if True, overwrites existing files with same name

Returns: none

History: v1: Initial python implementation. Riccardo Campana, 2014

1.8 Submodule: psd_models

Contains the analytic models for the power spectrum density.

`pyLCSIM.psd_models.lorentzian (x, p)`

Generalized Lorentzian function.

(**WARNING:** for $n \neq 2$ the function is no more normalized!)

Args: x: (non-zero) frequencies.

p[0]: x0 = peak central frequency.

p[1]: gamma = FWHM of the peak.

p[2]: value of the peak at $x = x_0$.

p[3]: n = power coefficient.

The quality factor is given by $Q = x_0/\text{gamma}$.

Returns: f: psd model.

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

`pyLCSIM.psd_models.smoothbknp (x, p)`

Smooth broken power law.

Args: x: (non-zero) frequencies.

p[0]: Normalization.

p[1]: power law index for $f \rightarrow 0$.

p[2]: power law index for $f \rightarrow \infty$.

p[3]: break frequency.

Returns: f: psd model.

History: v0.1: Riccardo Campana, 2014. Initial python implementation.

1.9 Changelog

v0.1.2: Added reset method to Simulation. Lightcurve FITS output is now OGIP-compliant.

v0.1.1: Bugfix. Modified Simulation class.

v0.1: Initial release.

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