

where Φ_f and Φ_w are the $\underline{24}$ representation of $SU(5)$ group. The breaking component is

$$\Phi_i = \text{Diagonal} \left(1, 1, 1, -\frac{3}{2}, -\frac{3}{2} \right) (\sigma_i + \varphi_i), \quad (12)$$

where the subscript $i = f, w$. From (11) – (12) we obtain

$$V_{fw} = -A \left(2\sigma_f \sigma_w \varphi_f \varphi_w + \sigma_f \varphi_f \varphi_w^2 + \sigma_w \varphi_w \varphi_f^2 + \frac{1}{2} \varphi_f^2 \varphi_w^2 \right). \quad (13)$$

$|\sigma_w| = |\sigma_f|$ because of the symmetry of s – *matter* and f – *matter*. Both σ_i and $m(\varphi_i)$ are functions of temperature T . When $T \geq T_{cr}$, $\sigma_i = m(\varphi_i) = 0$. Consequently f – *particles* and w – *particles* can easily transform from one to another so that $\rho_F = \rho_W$. When $T \sim 0$, both $|\sigma_i|$ and $m(\varphi_i)$ are large enough. Consequently interaction between f – *particles* and w – *particles* by the scalar bosons may be ignored. Thus there is only the gravitation between f – *matter* and w – *matter* when temperature is low.

There are the couplings of fermions (and gauge particles) and scalar bosons^[7]. Hence there are the interactions of f – *fermions* and w – *fermions* shown in figures 1-3 and the interactions of f – *gauge* bosons and w – *gauge* bosons via the scalar bosons φ_f and φ_w . In the figures the dotted lines with arrows denote W – *fermion* field ψ_w , the dotted lines without arrow denote W – *scalar* field φ_w , the lines with arrows denote F – *fermion* field ψ_f , the lines without arrow denote F – *scalar* field φ_f , $M^2 = -2A\sigma_f\sigma_w$, $R_f = -A\sigma_f$, $R_w = -A\sigma_w$ and $S = -A/2$.

It can be seen from figure 1 and (13) that when $-A\sigma_f\sigma_w > 0$ and $k^2 - m^2 < 0$ or $-A\sigma_f\sigma_w < 0$ and $k^2 - m^2 > 0$, f – *fermions* and w – *fermions* are repulsive each other; when $-A\sigma_f\sigma_w > 0$ and $k^2 - m^2 > 0$ or $-A\sigma_f\sigma_w < 0$ and $k^2 - m^2 < 0$, f – *fermions* and w – *fermions* are attractive each other.

V. FEATURES AND OBSERVATION OF DARK MATTER IN PRESENT MODEL

According to the present model^[3,4], v – *UFM* and v – *UWM* cannot form cluster, loosely distribute in space and have positive gravitational masses, hence both should be identified as cold dark matter. From (10) we have

$$\begin{aligned} \frac{\rho_{vFu} + \rho_{vWu}}{\rho_v} &= \frac{0.095 \times 2}{0.27} = \frac{19}{27}, \\ \frac{\rho_{vFu} + \rho_{vWu}}{\rho_{vD}} &= \frac{0.095 \times 2}{0.23} = \frac{19}{23}. \end{aligned} \quad (14)$$