Here  $\alpha$  and  $G_F$ , respectively, are the fine-structure constant and the Fermi constant,  $\theta_i^2$  is the mixing angle defined by

$$\theta_i^2 = \frac{\left(m_D^{\dagger} m_D\right)_{ii}}{m_{\chi_i}^2} = \mathcal{O}\left[\left(\frac{M_N^D}{M_N^M}\right)^2\right] \frac{v_L^2}{v_R^2}.$$
 (43)

With the previous parameter choice, we can determine the mixing angle to be  $\theta_i^2 \simeq 10^{-22}$ . Therefore the decay into the electron-positron pairs (42b) can [26] provide a natural explanation for the flux of 511 keV photons from the galactic bulge observed by INTEGRAL [27] satellite,

$$\frac{\Phi_{511\gamma}}{\Phi_{\rm exp}} \simeq \Sigma_i \frac{\theta_i^2}{10^{-22}} \left(\frac{m_{\chi_i}}{1.3 \,\text{MeV}}\right)^4 \frac{\Omega_{\chi_i}}{\Omega_{\chi}}. \tag{44}$$

It is easy to check that our scenario is consistent with other astrophysical and cosmological constraints [28]. Alternatively, we may explain the observed cosmic positron/electron excess [29–33], which is probably from continuum distribution of pulsars [34, 35], by fine tuning the parameters.

## V. CONCLUSION

In summary we have shown the dark matter can be well determined by the neutrino masses and the baryon asymmetry in the left-right symmetric model with doublet and singlet fields. In this model, the SM fermions obtain masses by integrating out charged singlet fermions. In the neutrino sector, the right(left)-handed neutral fermions, associated with the left(right)-handed Higgs doublet, can generate the left(right)-handed neutrino masses through the seesaw. The mass matrices of the leftand right-handed neutrinos have a same structure as a result of the left-right symmetry. The neutral singlets are also responsible for the baryon asymmetry and the dark matter. Specifically their 2-body decays can produce a desired lepton asymmetry in the left-handed leptons and then the observed baryon asymmetry can be realized by the sphaleron induced lepton-to-baryon conversion. At the same time, the right-handed neutrinos can serve as the dark matter as they have a right relic density from the 3-body decays of the neutral singlets. The decays of these non-thermally produced right-handed neutrinos can easily induce the observed fluxes of 511 keV photons from the Galactic bulge. The attractive feature of our scenario is that the left-right symmetry can connect the properties of the dark matter to the neutrino masses and the baryon asymmetry.

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- J. Dunkley *et al.*, [WMAP Collaboration], Astrophys. J. Suppl. **180**, 306 (2009).
- V.A. Kuzmin, arXiv: hep-ph/9701269; R. Kitano and I. Low, Phys. Rev. D 71, 023510 (2005); N. Cosme, L. Lopez Honorez, and M.H.G. Tytgat, Phys. Rev. D 72, 043505 (2005); D.E. Kaplan, M.A. Luty, and K.M. Zurek, Phys. Rev. D 79, 115016 (2009); P.H. Gu, U. Sarkar, and X. Zhang, Phys. Rev. D 80, 076003 (2009); I.M. Shoemaker and A. Kusenko, Phys. Rev. D 80, 075021 (2009); P.H. Gu and U. Sarkar, arXiv:0909.5463 [hep-ph]; H. An, S.L. Chen, R.N. Mohapatra, and Y. Zhang, arXiv:0911.4463 [hep-ph].
- [3] T. Schwetz, M.A. Tórtola, J.W.F. Valle, New J. Phys. 10, 113011 (2008).
- [4] P. Minkowski, Phys. Lett. B 67, 421 (1977).
- [5] T. Yanagida, in *Proc. of the Workshop on Unified Theory* and the Baryon Number of the Universe, ed. O. Sawada and A. Sugamoto (KEK, Tsukuba, 1979), p. 95.
- [6] M. Gell-Mann, P. Ramond, and R. Slansky, in Supergravity, ed. F. van Nieuwenhuizen and D. Freedman (North Holland, Amsterdam, 1979), p. 315.
- [7] S.L. Glashow, in *Quarks and Leptons*, ed. M. Lévy *et al.* (Plenum, New York, 1980), p. 707.
- [8] R.N. Mohapatra and G. Senjanović, Phys. Rev. Lett. 44, 912 (1980).
- M. Magg and C. Wetterich, Phys. Lett. B 94, 61 (1980);
   J. Schechter and J.W.F. Valle, Phys. Rev. D 22, 2227 (1980);
   T.P. Cheng and L.F. Li, Phys. Rev. D 22, 2860

- (1980); G. Lazarides, Q. Shafi, and C. Wetterich, Nucl. Phys. B **181**, 287 (1981); R.N. Mohapatra and G. Senjanović, Phys. Rev. D **23**, 165 (1981).
- [10] M. Fukugita and T. Yanagida, Phys. Lett. B 174, 45 (1986).
- [11] P. Langacker, R.D. Peccei, and T. Yanagida, Mod. Phys. Lett. A 1, 541 (1986); M.A. Luty, Phys. Rev. D 45, 455 (1992); R.N. Mohapatra and X. Zhang, Phys. Rev. D 46, 5331 (1992).
- [12] M. Flanz, E.A. Paschos, and U. Sarkar, Phys. Lett. B
  345, 248 (1995); M. Flanz, E.A. Paschos, U. Sarkar, and J. Weiss, Phys. Lett. B 389, 693 (1996); L. Covi, E. Roulet, and F. Vissani, Phys. Lett. B 384, 169 (1996).
- [13] A. Pilaftsis, Phys. Rev. D 56, 5431 (1997).
- [14] E Ma and U. Sarkar, Phys. Rev. Lett. 80, 5716 (1998).
- [15] W. Buchmüller, R.D. Peccei, and T. Yanagida, Ann. Rev. Nucl. Part. Sci. 55, 311 (2005), and references therein.
- [16] J.C. Pati and A. Salam, Phys. Rev. D 10, 275 (1974);
  R.N. Mohapatra and J.C. Pati, Phys. Rev. D 11, 566 (1975);
  R.N. Mohapatra and J.C. Pati, Phys. Rev. D 11, 2558 (1975);
  R.N. Mohapatra and G. Senjanović, Phys. Rev. D 12, 1502 (1975).
- [17] B. Brahmachari, E Ma, and U. Sarkar, Phys. Rev. Lett. 91, 011801 (2003).
- [18] F. Siringo, Eur. Phys. J. C **32**, 555 (2004).
- [19] D. Chang, R.N. Mohapatra, and M.K. Parida, Phys. Rev. Lett. 52, 1072 (1984); D. Chang, R.N. Mohapatra, and M.K. Parida, Phys. Rev. D 30, 1052 (1984).