

begin with calibrated
full-frame images



detect bright sources



calibrate astrometry



use WCS solution to project DR2
sources from (α, δ) to (x, y)



select photometric & astrometric
reference frames



transform each calibrated frame to
astrometric reference



convolve each resulting frame to
photometric reference; subtract.



do aperture photometry on
differenced images



remove trends from light curves