

Report page ExoTIC-ISM

W17_G141_lc_11395.txt - 11395

Input parameters:

Number of systematic models: 50
Wavelength mid point = 11380.141699602069
Wavelength half width = 90.80626964749172

Planet parameters:

$R_p/R^* = 0.1255$
Epoch (MJD) = 57957.97108811848
Inclination (deg) = 86.93051272857655
Eccentricity = 0.0
Omega (deg) = 0.0
Period (days) = 3.7354850226
 $a/R^* = 7.025$

Stellar parameters:

FeH (dex) = -0.25
Teff (K) = 6550.0
 $\log(g)$ (cgs) = 4.2

Output parameters:

Limb-darkening coefficients:

$C1 = 0.8529625463058947$
 $C2 = -0.8296916633180598$
 $C3 = 0.7767245286699795$
 $C4 = -0.2756867391195464$

Top five systematic models by their weight

Check the chi-squared values and the AIC evidence for reasonable fits.
If the chi-squared values far exceed the DOF then it is likely that the input data contains additional noise, double check the spectral extraction.
Model numbers = [18 19 23 17 43]
DOF = [45. 44. 44. 46. 44.]
Chi-squared = [55.37153692 54.68610039 55.18283248 57.30579833 55.3151131]
AIC evidence = [339.57891576 339.42163403 339.17326799 339.11178506 339.10712767]
Weights = [0.1042506758832423 0.08907837585340858 0.06948775612153642
0.06534413175378025 0.06504050656035855]
SDNR = [293.07907639 291.253453 292.56903461 298.23600113 292.94609316]

Top model Noise Statistics:

White noise = 0.0

Red noise = 0.0

Beta = 1.0

If the red-noise is significant it means the data is poorly fit by any of the systematic models. It is recommended that the input lightcurves are checked for additional noise sources.

Marginalised parameters:

If None, parameter was not fit for.

$R_p/R^* = 0.12299246762814105 \pm 0.0004816974733537$

Epoch (MJD) = 57957.9702101891 $\pm 0.0006150565073334445$

Inclination (rad) = None \pm None

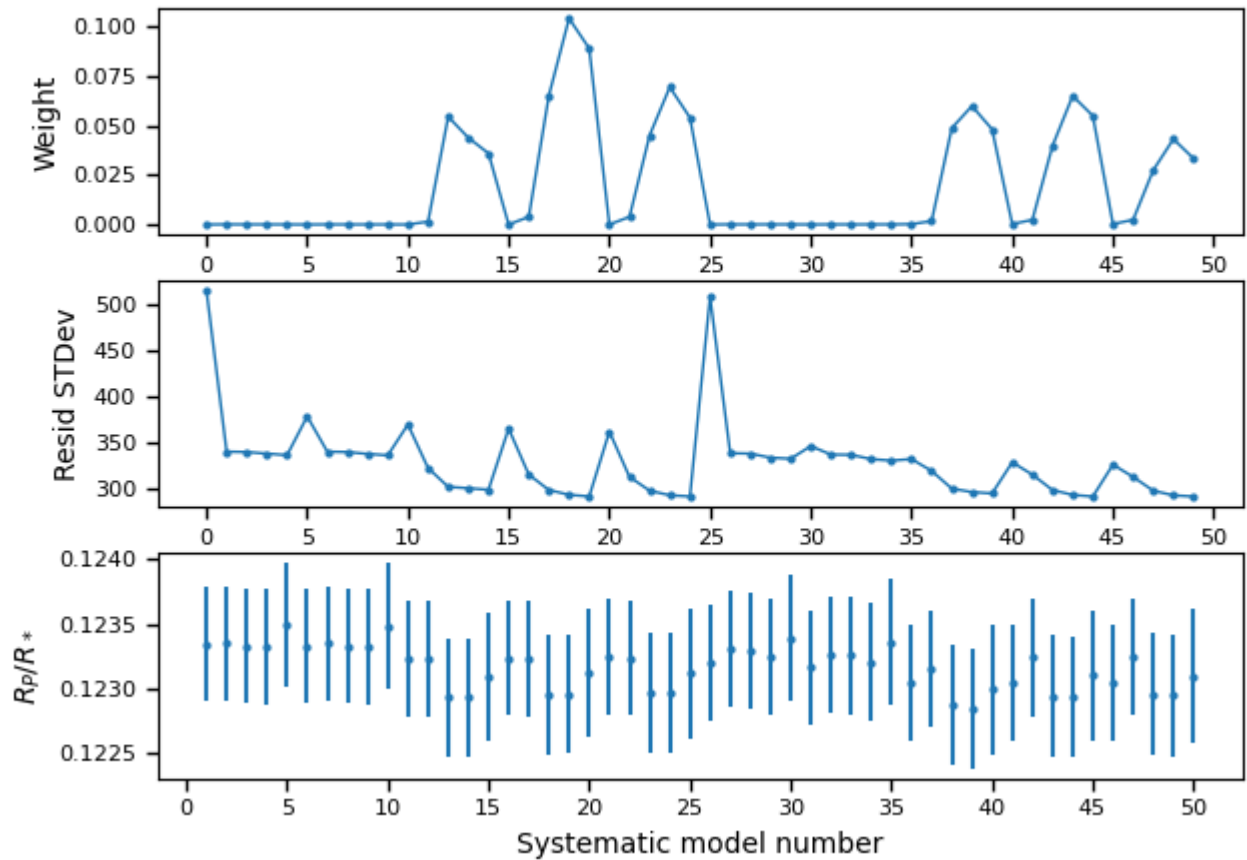
Inclination (deg) = None \pm None

System density ($M_s + M_p/R^3$) = None \pm None

$a/R^* =$ None \pm None

Systematics

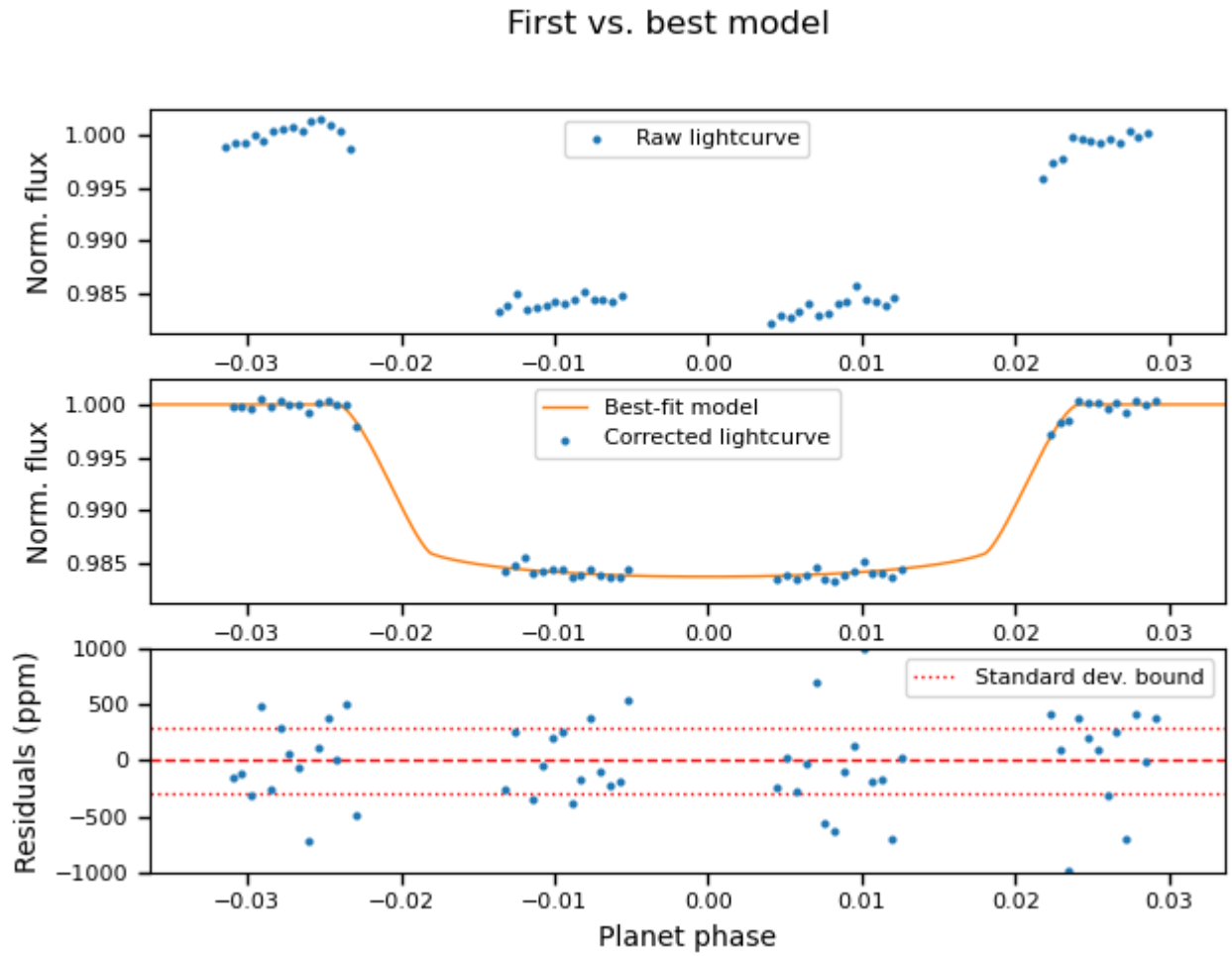
Marginalisation results



Top: Evidence-based weight associated with each systematic model when fit with the data. *Middle:* Standard deviation of the residuals after correcting for each systematic model. *Bottom:* Radius ratio

measured from the transit depth when the light curve has been corrected using each systematic model. *If present, grey crosses mark discarded systematic models (poor AIC evidence).*

Lightcurves



Top: Input lightcurve with no systematic model correction applied. *Middle:* Lightcurve corrected by highest weight systematic model plotted with the smooth planetary transit model centred on the mid-transit time. *Bottom:* Residuals and uncertainties associated with the middle panel lightcurve. The upper and lower standard deviation bounds are shown in dotted lines relative to zero.