Documentation and report of Astrophysics with Artificial Intelligence(Astropy and AstroML) – UVES Spectroscopy with Astropy

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* **All the information is based on and originated from ‘astropy.org’, ‘atsroml.org’, and ‘wikipedia.org’**

<1> The fundamental knowledge to utilize astropy and astroML for astrophysics

1. Accretion Disk

The accretion disk which is also known as circumstellar disk is the structure which is created by diffuse material in orbital motion around a massive central body. Typically, the central body is the star such as protostar, white dwarf, neutron star, or black hole and also there are theoretical stars such as black dwarf and blue dwarf. The instabilities in the disk causes redistribution of the angular momentum and to move spiral the materials inward toward the central body. In this process, the gravitational energy and frictional forces are converted to the thermal energy by compressing and raising the temperature of the material and this occurs the emission of electromagnetic radiation on the surface of the disk. The range of the frequency of the emission of electromagnetic radiation depends on the central body. For instance, the accretion disk of the protostar radiates the infrared and the neutron star and black hole emit the X-ray.

The equation of the angular momentum

L = r × p = r × mv

(m : mass of the material,

P : linear momentum of the material,

r : position vector from origin 0 to material,

v : velocity of the material)

The

1. Protostar

Mj = (9/4) × (1/2𝝅n)1/2 × (1/m2 ) × (kT/G)3/2

(n : the particle number density,

m : average mass of the gas particle in the interstellar cloud,

T : temperature of the gas)