





QhX: A Python package for periodicity detection in red noise

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Summary

QhX is a Python package for detecting periodicity in red noise time series, developed as an in-kind contribution to the Vera C. Rubin Observatory Legacy Survey of Space and Time (LSST, [Ivezić et al., 2019](#)). Traditional Fourier-based methods often struggle with red noise, which is common in quasar light curves and other accreting objects. QhX addresses these challenges with its core 2D Hybrid method ([Kovačević et al., 2018](#)). Input data are mapped into a time-period plane via wavelet transforms, which are (auto)correlated to produce a correlation density map in a “period-period” plane. Statistical vetting incorporates significance, upper and lower error bounds, and the novel Intersection over Union (IoU) metric to evaluate the proximity and overlap of detected periods across bands and objects. In addition to a vetted numerical catalog, QhX dynamically visualizes periodicity across photometric bands and objects.

Statement of need

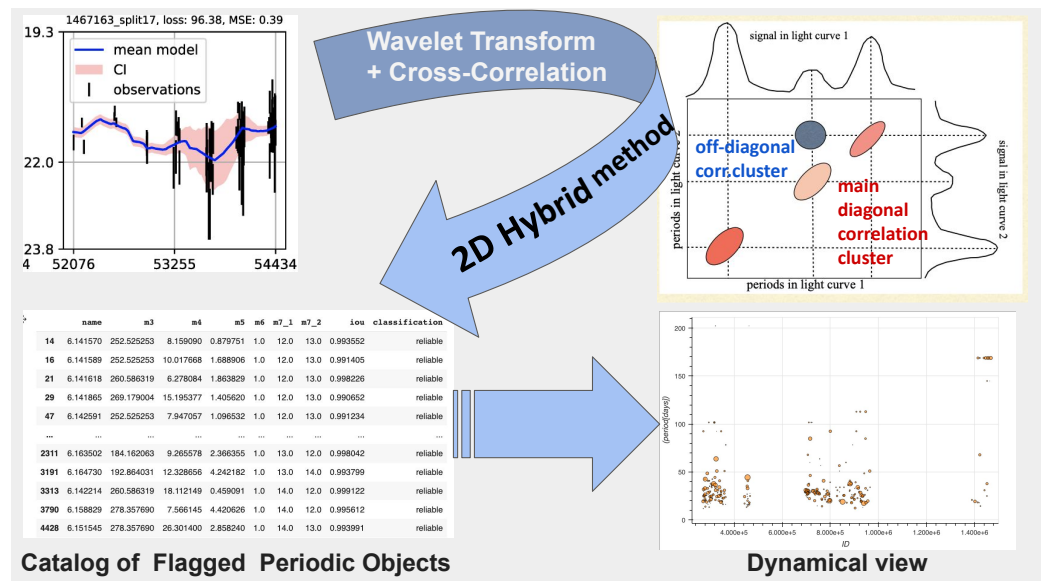


Figure 1: The left panel shows a 1D light curve with observational data (black error bars) and a model (blue line). QhX transforms the time-series into the time-frequency domain and cross-correlates wavelet matrices to produce a 2D period-period correlation map (right), where clusters indicate periodic signals. After map integration, statistical vetting generates a numerical catalog of flagged periodic objects (bottom left) and a dynamic view of detected periods across objects and bands (bottom right).

Periodic variability spans a range of astronomical objects, from asteroids to quasars. Identifying meaningful signals is complicated by red noise (see, e.g., Figure 1 in [Gaia Collaboration et al., 2019](#); [Kasliwal et al., 2015](#); [Kovačević, Radović, et al., 2022](#)), which exhibits fractal-like patterns across time scales ([Belete et al., 2018](#); [Vio et al., 1991](#)). Non-stationary signals and unfavorable sampling ([Brandt et al., 2018](#); [D’Orazio & Charisi, 2023](#)) further obscure coherent patterns. Traditional time-frequency methods, constrained by the Fourier uncertainty principle (i.e., Gabor limit, [Gabor, 1947](#)), often fail with such complex signals, highlighting the need for nonlinear approaches ([Abry et al., 1995](#); [Cohen, 1995](#)).

QhX provides features specifically designed to address these challenges. The first feature is its core 2D Hybrid method (see Figure 1), detailed in ([Kovačević et al., 2018](#)), inspired by 2D Correlation Spectroscopy ([Kovačević, 2024](#); [Noda, 2018](#)). By applying wavelet transforms, QhX maps time-series data into the time-frequency domain and (auto)correlates it, generating a period-period correlation density that enhances signal detection. Secondly, QhX introduces an Intersection over Union (IoU) metric, combined with standard statistical measures (significance, upper and lower error bounds), to evaluate the overlap of detected periods across bands and objects. Each period is represented as the center of an “IoU ball,” with its radius reflecting relative error, calculated as the mean of the upper and lower error bounds—analogous to a circular aperture in photometry ([Saxena et al., 2024](#)). Thirdly, QhX enhances traditional analysis by generating numerical and interactive visual catalogs that rank periodicity candidates by reliability. These interactive catalogs enable detailed inspection of signal consistency, offering greater interpretability than traditional static plots.

across bands and objects. To our knowledge, this is the first application of the Intersection-over-Union (IoU) metric to quantify the overlap between detected and reference periods in astronomical time-series analysis.

- Statistical vetting categorizes detected periods for each object and band as reliable, medium, or poor.

3. Data Management:

- QhX assumes input time-series data in a simple tabular format containing time, flux (or magnitude), and associated uncertainties. Examples in the documentation illustrate how to map data from other commonly used formats into this structure.
- `data_manager` and `data_manager_dynamical` modules manage data flow, data loading, outlier removal, and format compatibility.
- `batch_processor` and `parallelization_solver` modules optimize task distribution across multiple processors.

4. Visualization and Output:

- `plots` module includes tools for creating interactive visualizations, such as `interactive_plot`, which allows for exploring detected periodicities across bands and objects. For large datasets, `interactive_plot_large_files` enables in-depth inspection of signal consistency.
- `output` and `output_parallel` modules handle result storage, supporting both single-threaded and parallelized workflows.

5. Testing:

- `tests` module, containing `test_parallel` and `test_integrated`, validates the functionality across various processing setups.

Representative Applications

QhX has been benchmarked with respect to widely used periodicity detection software across multiple domains. In Fatović et al. (2023), applying QhX to SDSS J23205+0024 yielded a period of $278.36^{+57.34}_{-25.21}$ days, with a significance above 99% measured via the shuffling method, and a 90% significance from the Generalized Extreme Value (GEV) approach. A Lomb–Scargle periodogram applied to the same dataset produced a consistent period of 278-days at the same significance level. In Kovačević et al. (2019), QhX detected periods of 1972 ± 254 -days (observed light curve) and 1873 ± 250 days (modeled light curve) for PG-1302-102, both within 1σ of the 1884 ± 88 day period reported by Graham et al. (2015) using generalized Lomb–Scargle, wavelet, and autocorrelation methods; a Bayesian reanalysis by Zhu & Thrane (2020) on an extended dataset for PG-1302-102 yielded a comparable quasi-period of 5.6 yr, interpreted as quasiperiodic oscillations. In the case of Mrk 231 (Kovačević, Yi, et al., 2020), the 2D hybrid method identified a characteristic period of 403 days with a significance greater than 99.7%, in agreement with a Lomb–Scargle periodogram result of 413 days at a significance above 95%; the slightly larger uncertainty in the QhX-derived period reflects the temporal variation of the periodicity, while the average oscillation power is comparable between the two methods. The method has also been validated in the context of damped oscillations in the changing-look quasar NGC 3516 (Kovačević, Popović, et al., 2020), where experimental results demonstrated robustness against the combined effects of red noise and complex time-series structure. Beyond astrophysical applications, QhX has been applied to Very Low Frequency (VLF) signal analysis for pre- and post-earthquake intervals (Kovačević, Nina, et al., 2022). In the no-earthquake scenario (same date one year earlier), the topology of QhX 2D hybrid maps exhibited distinct correlation cluster patterns compared to earthquake-day records, with detected periods below 111 s in most intervals and a ~ 140 s signal in the -2 h segment, closely matching a 147 s signal detected during the earthquake event. Comparison with Fast Fourier Transform (FFT) results (Nina et al., 2020) showed strong agreement before the earthquake for periods below 1.5 min, and convergence of both methods to similar values in subsequent intervals. Post-earthquake periods obtained with QhX were also consistent with the < 10 s to few-hundred-second range reported in (Ohya et al., 2018). QhX is the LSST

114 [directable software in-kind contribution.](#)

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