Parallactic Angle Calculation

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1 Introduction

The parallactic angle defines the rotation between RA/Dec and Az/ZA coordinates. It is calculated in FHD in the parallactic_angle.pro function and is used in rotate_jones_matrix.pro to construct the Jones matrix, which converts between instrumental and Stokes polarization modes.

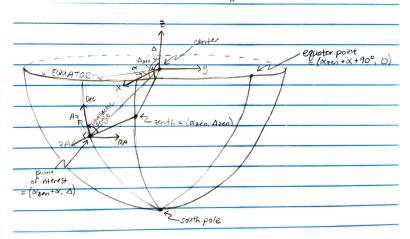
This calculation does not account for aberration.

2 Parallactic Angle Calculation

Here we consider a point with (RA, Dec) coordinates ($\alpha_{zen} + \alpha, \Delta$) and consider the angle that rotates a vector at that point between (Az, ZA) coordinates and (RA, Dec) coordinates.

The (RA, Dec) coordinates of zenith are $(\alpha_{\text{zen}}, \Delta_{\text{zen}})$. We can define a Cartesian coordinate system with the origin at the center of the sphere and the x-axis aligned with $\text{RA}=\alpha_{\text{zen}}$. We define the radius of the sphere to be 1.

We define our coordinates looking down, such that at zenith RA increases in the +y direction and Dec increases in the +z direction. ZA increases away from zenith and Az increases clockwise from the +y direction at zenith.



Converting (RA, Dec) to Cartesian coordinates, the vector to our point $(\alpha_{zen} + \alpha, \Delta)$ is

$$point = \cos \alpha \cos \Delta \hat{x} + \sin \alpha \cos \Delta \hat{y} + \sin \Delta \hat{z}. \tag{1}$$

and to zenith is

$$zenith = \cos \Delta_{zen} \hat{x} + \sin \Delta_{zen} \hat{z}. \tag{2}$$

At our point, the Az/ZA coordinate system is aligned with the plane that includes the center of the sphere, the point, and zenith. A vector normal to this plane can be calculated by taking the cross product between the vectors point and zenith:

zenith × point =
$$\sin \Delta_{\text{zen}} \cos \Delta \sin \alpha \hat{x} + (\sin \Delta_{\text{zen}} \cos \Delta \cos \alpha - \cos \Delta_{\text{zen}} \sin \Delta) \hat{y} + \cos \Delta_{\text{zen}} \cos \Delta \sin \alpha \hat{z}$$
(3)

To normalize, we find the amplitude of this vector:

$$|\text{zenith} \times \text{point}|^{2}$$

$$= \sin^{2} \alpha \cos^{2} \Delta + \sin^{2} \Delta_{\text{zen}} \cos^{2} \Delta \cos^{2} \alpha + \cos^{2} \Delta_{\text{zen}} \sin^{2} \Delta$$

$$- 2 \sin \Delta_{\text{zen}} \cos \Delta_{\text{zen}} \cos \alpha \sin \Delta \cos \Delta$$

$$= \cos^{2} \Delta_{\text{zen}} (\cos \Delta \tan \Delta_{\text{zen}} - \sin \Delta \cos \alpha)^{2} + \cos^{2} \Delta_{\text{zen}} \sin^{2} \alpha,$$
(4)

so the unit vector normal to the plane is

$$\frac{\frac{\mathrm{zenith} \times \mathrm{point}}{|\mathrm{zenith} \times \mathrm{point}|}}{\frac{\mathrm{sin} \, \Delta_{\mathrm{zen}} \cos \Delta \sin \alpha \hat{x} + (\sin \Delta_{\mathrm{zen}} \cos \Delta \cos \alpha - \cos \Delta_{\mathrm{zen}} \sin \Delta) \hat{y} + \cos \Delta_{\mathrm{zen}} \cos \Delta \sin \alpha \hat{z}}{\sqrt{\cos^2 \Delta_{\mathrm{zen}} (\cos \Delta \tan \Delta_{\mathrm{zen}} - \sin \Delta \cos \alpha)^2 + \cos^2 \Delta_{\mathrm{zen}} \sin^2 \alpha}}.$$
(5)

This vector is parallel to the -Az axis.

At our point, the RA/Dec coordinate system is aligned with a plane that includes the center of the sphere, the point, and a point we'll call "equator point" at

equator point =
$$\cos(\alpha + 90^{\circ})\hat{x} + \sin(\alpha + 90^{\circ})\hat{y} = -\sin\alpha\hat{x} + \cos\alpha\hat{y}$$
. (6)

A vector normal to this plane can be calculated by taking the cross product between the vectors point and equator point:

point × equator point =
$$-\sin \Delta \cos \alpha \hat{x} - \sin \Delta \sin \alpha \hat{y} + \cos \Delta \hat{z}$$
. (7)

To normalize, we find that the amplitude of this vector is

$$|point \times equator\ point|^2 = 1,$$
 (8)

so the unit vector normal to the plane is

$$\frac{\text{point} \times \text{equator point}}{|\text{point} \times \text{equator point}|} = -\sin \Delta \cos \alpha \hat{x} - \sin \Delta \sin \alpha \hat{y} + \cos \Delta \hat{z}. \tag{9}$$

This vector is parallel to the Dec axis.

Taking the dot product of these two vectors gives

$$\left(\frac{\text{zenith} \times \text{point}}{|\text{zenith} \times \text{point}|}\right) \cdot \left(\frac{\text{point} \times \text{equator point}}{|\text{point} \times \text{equator point}|}\right) \\
= \frac{\cos \Delta_{\text{zen}} \sin \alpha}{\sqrt{\cos^2 \Delta_{\text{zen}} (\cos \Delta \tan \Delta_{\text{zen}} - \sin \Delta \cos \alpha)^2 + \cos^2 \Delta_{\text{zen}} \sin^2 \alpha}} \tag{10}$$

so the angle between the planes is

$$\cos^{-1} \left[\frac{\cos \Delta_{\text{zen}} \sin \alpha}{\sqrt{\cos^2 \Delta_{\text{zen}} (\cos \Delta \tan \Delta_{\text{zen}} - \sin \Delta \cos \alpha)^2 + \cos^2 \Delta_{\text{zen}} \sin^2 \alpha}} \right]. \tag{11}$$

This corresponds to the angle that rotates the -Az axis counterclockwise into the +Dec axis. The parallactic angle is defined as the angle that rotates the -ZA axis into the +Dec, so

parallactic angle

$$= \sin^{-1} \left[\frac{-\cos \Delta_{\rm zen} \sin \alpha}{\sqrt{\cos^2 \Delta_{\rm zen} (\cos \Delta \tan \Delta_{\rm zen} - \sin \Delta \cos \alpha)^2 + \cos^2 \Delta_{\rm zen} \sin^2 \alpha}} \right]. \tag{12}$$

To simplify, we can use some handy trig identities to rewrite

parallactic angle =
$$\tan^{-1} \left(\frac{-\sin \alpha}{\cos \Delta \tan \Delta_{\text{zen}} - \sin \Delta \cos \alpha} \right)$$
. (13)

The parallactic angle is defined as the counterclockwise angle between the -ZA axis and the +Dec axis. It lies in quadrants III and IV when $\alpha > 0$ and quadrants I and II when $\alpha < 0$.

3 Implementation in FHD

The parallactic angle is calculated in FHD in parallactic_angle.pro. In that function, latitude = Δ_{zen} , dec = Δ , and hour_angle = $-\alpha$ as defined in this memo (note that the hour angle is defined as increasing to the West, the opposite of the Right Ascension).

The IDL inverse tangent function at an has a range of $(-\pi/2, \pi/2]$ when only one argument is provided. When two arguments are provided, the range is $(-\pi, \pi]$: the range is $(0, \pi]$ when the first argument is positive and $(-\pi, 0)$ when the first argument is negative.

parallactic_angle.pro is called by projection_slant_orthographic.pro and by rotate_jones_matrix.pro, which is used in constructing the Jones matrix for converting between Stokes and instrumental polarization modes. In FHD, the Jones matrix is defined such that

$$J\begin{bmatrix} RA \\ Dec \end{bmatrix} = \begin{bmatrix} p \\ q \end{bmatrix}, \tag{14}$$

where p is the East-West dipole orientation, q is the North-South dipole orientation, and

$$J = \begin{bmatrix} J[0,0] & J[1,0] \\ J[0,1] & J[1,1] \end{bmatrix}$$
 (15)

Before rotation, the function $fhd_struct_init_antenna.pro$ returns a Jones matrix (that we will call J') that converts from Az/ZA to instrumental polarizations:

$$J' \begin{bmatrix} ZA \\ Az \end{bmatrix} = \begin{bmatrix} p \\ q \end{bmatrix}. \tag{16}$$

rotate_jones_matrix.pro applies a rotation to calculate J = J'R where R is the

rotation

$$R \begin{bmatrix} RA \\ Dec \end{bmatrix} = \begin{bmatrix} ZA \\ Az \end{bmatrix}. \tag{17}$$

The parallactic angle is the angle between the -ZA axis (pointing towards zenith) and the Dec axis. The rotation matrix is

$$R = \begin{bmatrix} R[0,0] & R[1,0] \\ R[0,1] & R[1,1] \end{bmatrix} = \begin{bmatrix} \sin \delta & -\cos \delta \\ -\cos \delta & -\sin \delta \end{bmatrix}.$$
 (18)

Note that $R^{-1} = R$. Under IDL convention, the matrix product AB is given by matrix multiply (B, A).

In $fhd_struct_init_jones.pro$, Jmat is the tensor product of J with its conjugate:

$$\mathtt{Jmat} = J \otimes J^* = \begin{bmatrix} |J[0,0]|^2 & |J[1,0]|^2 & J[0,0]J^*[1,0] & J[1,0]J^*[0,0] \\ |J[0,1]|^2 & |J[1,1]|^2 & J[0,1]J^*[1,1] & J[1,1]J^*[0,1] \\ J[0,0]J^*[0,1] & J[1,0]J^*[1,1] & J[0,0]J^*[1,1] & J[1,0]J^*[0,1] \\ J[0,1]J^*[0,0] & J[1,1]J^*[1,0] & J[0,1]J^*[1,0] & J[1,1]J^*[0,0] \end{bmatrix} \quad (19)$$

where

$$\mathtt{Jmat} = \begin{bmatrix} \mathtt{Jmat}[0,0] & \mathtt{Jmat}[1,0] & \mathtt{Jmat}[2,0] & \mathtt{Jmat}[3,0] \\ \mathtt{Jmat}[0,1] & \ddots & & & \\ \mathtt{Jmat}[0,2] & & & & \\ \mathtt{Jmat}[0,3] & & & & \\ \end{bmatrix}. \tag{20}$$

Jmat converts from RA/Dec coordinates to instrumental polarization:

$$\operatorname{Jmat}\begin{bmatrix} \alpha \alpha^* \\ \Delta \Delta^* \\ \alpha \Delta^* \\ \Delta \alpha^* \end{bmatrix} = \begin{bmatrix} pp^* \\ qq^* \\ pq^* \\ qp^* \end{bmatrix}, \tag{21}$$

and its inverse, Jinv, does the opposite transformation:

$$\operatorname{Jinv} \begin{bmatrix} pp^* \\ qq^* \\ pq^* \\ qp^* \end{bmatrix} = \begin{bmatrix} \alpha\alpha^* \\ \Delta\Delta^* \\ \alpha\Delta^* \\ \Delta\alpha^* \end{bmatrix}. \tag{22}$$

4 Conversion to Stokes

The conversion between sky coordinates and Stokes parameters is done by stokes_cnv.pro in FHD. Stokes parameters are inherently tied to a sky coordinate system, preferably

RA/Dec. The parameters are defined as follows:

$$\begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix} = \begin{bmatrix} 1 & 1 & 0 & 0 \\ 1 & -1 & 0 & 0 \\ 0 & 0 & 1 & 1 \\ 0 & 0 & i & -i \end{bmatrix} \begin{bmatrix} \alpha \alpha^* \\ \Delta \Delta^* \\ \alpha \Delta^* \\ \Delta \alpha^* \end{bmatrix}$$
(23)

and the inverse

$$\begin{bmatrix} \alpha \alpha^* \\ \Delta \Delta^* \\ \alpha \Delta^* \\ \Delta \alpha^* \end{bmatrix} = \begin{bmatrix} 1/2 & 1/2 & 0 & 0 \\ 1/2 & -1/2 & 0 & 0 \\ 0 & 0 & 1/2 & -i/2 \\ 0 & 0 & 1/2 & i/2 \end{bmatrix} \begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix}.$$
 (24)

$$\begin{bmatrix} \alpha \alpha^* \\ \Delta \Delta^* \\ \alpha \Delta^* \\ \Delta \alpha^* \end{bmatrix} = \begin{bmatrix} \sin^2 \delta & \cos^2 \delta & -\sin \delta \cos \delta & -\sin \delta \cos \delta \\ \cos^2 \delta & \sin^2 \delta & \sin \delta \cos \delta & \sin \delta \cos \delta \\ -\sin \delta \cos \delta & \sin \delta \cos \delta & -\sin^2 \delta & \cos^2 \delta \\ -\sin \delta \cos \delta & \sin \delta \cos \delta & \cos^2 \delta & -\sin^2 \delta \end{bmatrix} \begin{bmatrix} |Z|^2 \\ |A|^2 \\ ZA^* \\ AZ^* \end{bmatrix}$$
(25)

where Z is the ZA coordinate and A is the Az coordinate. For an unpolarized source, such that Q = U = V = 0, we find that

$$\begin{bmatrix} |Z|^2 \\ |A|^2 \\ ZA^* \\ AZ^* \end{bmatrix} = \begin{bmatrix} \alpha \alpha^* \\ \Delta \Delta^* \\ \alpha \Delta^* \\ \Delta \alpha^* \end{bmatrix} = \begin{bmatrix} I/2 \\ I/2 \\ 0 \\ 0 \end{bmatrix}, \tag{26}$$

where I is the source flux, so the Az/ZA vector components have no dependence on the the parallactic angle.