

The MuLan Experiment

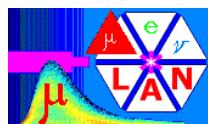
Measuring the muon lifetime to 1ppm

Kevin Lynch
MuLan Collaboration
Boston University

Outline:

- Motivate the measurement
- Describe the experiment
- Past and developing results

Berkeley, Boston, Illinois,
James Madison, Kentucky, KVI, PSI



Precision electroweak predictions rest on three parameters

Fine Structure Constant

$$\frac{\delta \alpha_{\text{em}}}{\alpha_{\text{em}}} \approx 0.37 \text{ ppb}$$

Gabrielse *et al*
2008

Mass of the neutral weak boson

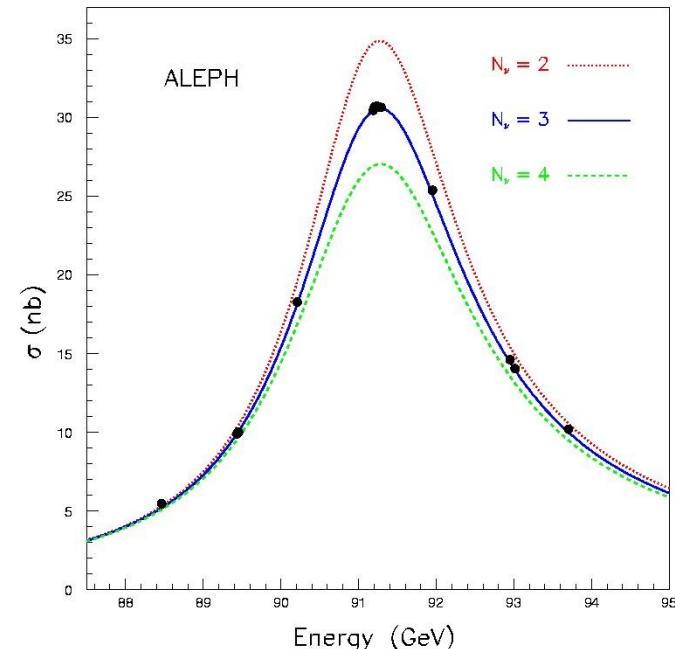
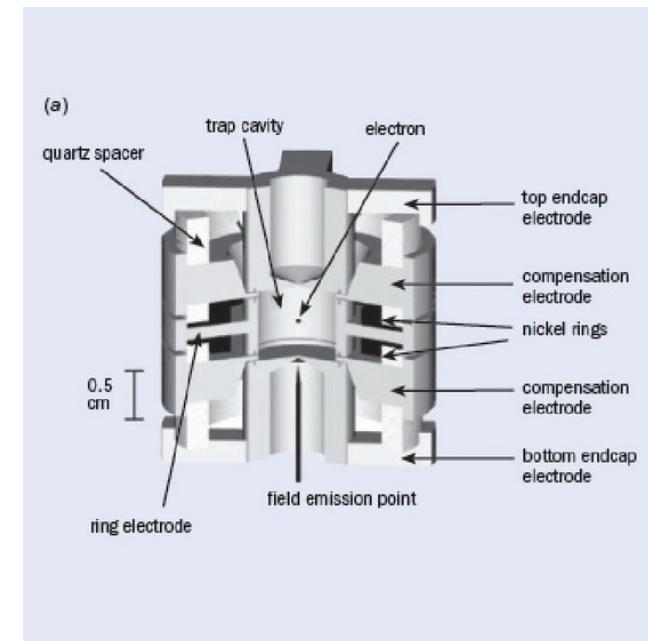
$$\frac{\delta M_{Z^0}}{M_{Z^0}} \approx 23 \text{ ppm}$$

LEP EWWG
2005

Fermi Constant

$$\frac{\delta G_F}{G_F} \approx 4 \text{ ppm}$$

Chitwood *et al*
2006

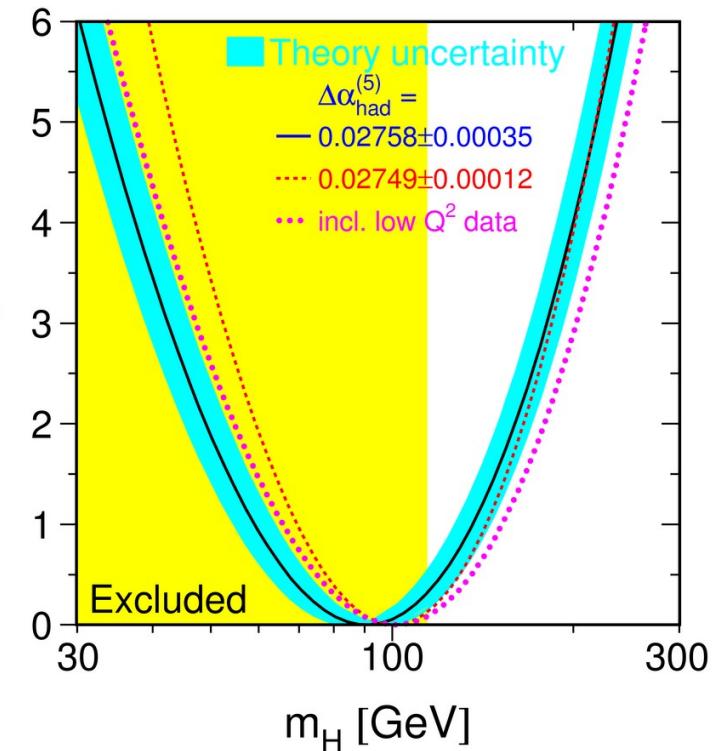
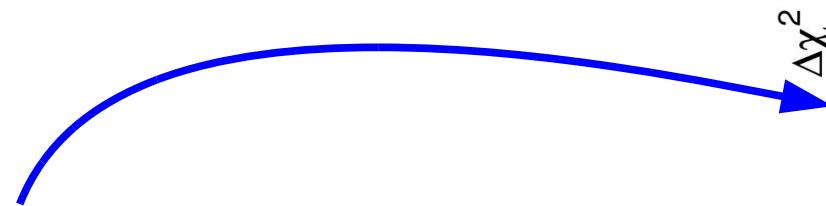


The Fermi constant is an implicit input to all precision electroweak studies

$$\frac{G_F}{\sqrt{2}} = \frac{g^2}{8M_W^2} (1 + \Delta r(m_t, m_H, \dots))$$

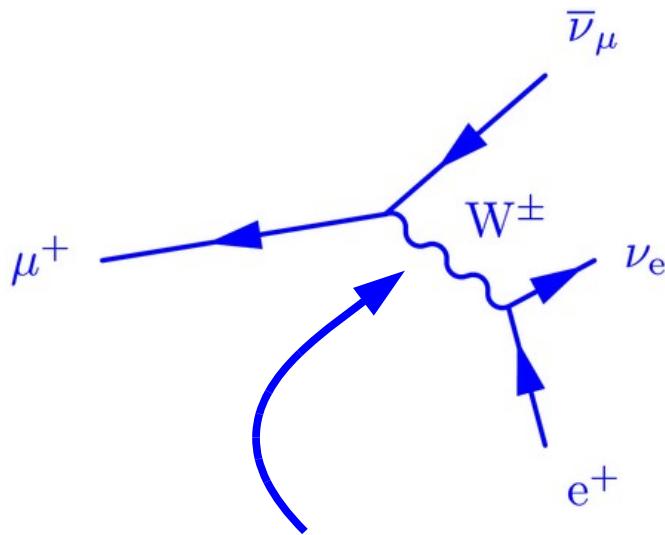
Contains all weak interaction loop corrections.

Example: the “blue band” Higgs limit plot.



Plot borrowed from LEP Electroweak Working Group publications

Muon decay gives us unique access to the electroweak scale

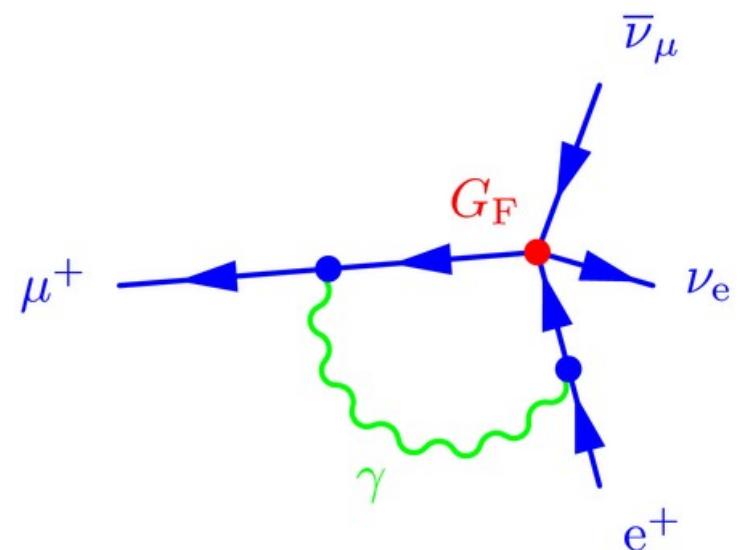


The muon *only* decays via the weak interaction, which gives it a very long lifetime.

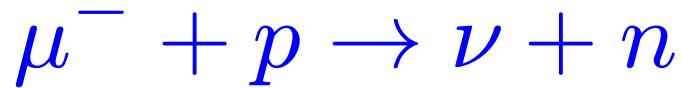
All relevant weak interaction physics confined to one easily measured parameter with a clean theoretical interpretation.

The V-A theory factorizes into a pure **weak** contribution, and **non-weak** corrections, essentially uncontaminated by hadronic uncertainties.

$$\frac{1}{\tau_{\mu^+}} = \frac{G_F^2 m_\mu^5}{192\pi^3} (1 + q)$$



Additionally, the free muon lifetime is a precision reference for nuclear capture measurements



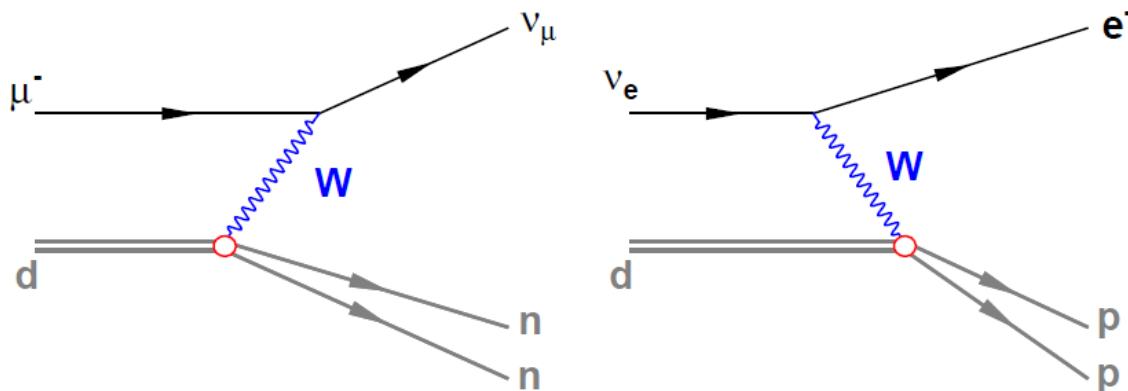
MuCap



MuSun

These are the simplest weak interaction processes in nuclei with precise theoretical predictions in QCD (EFT, χ PT, pQCD)

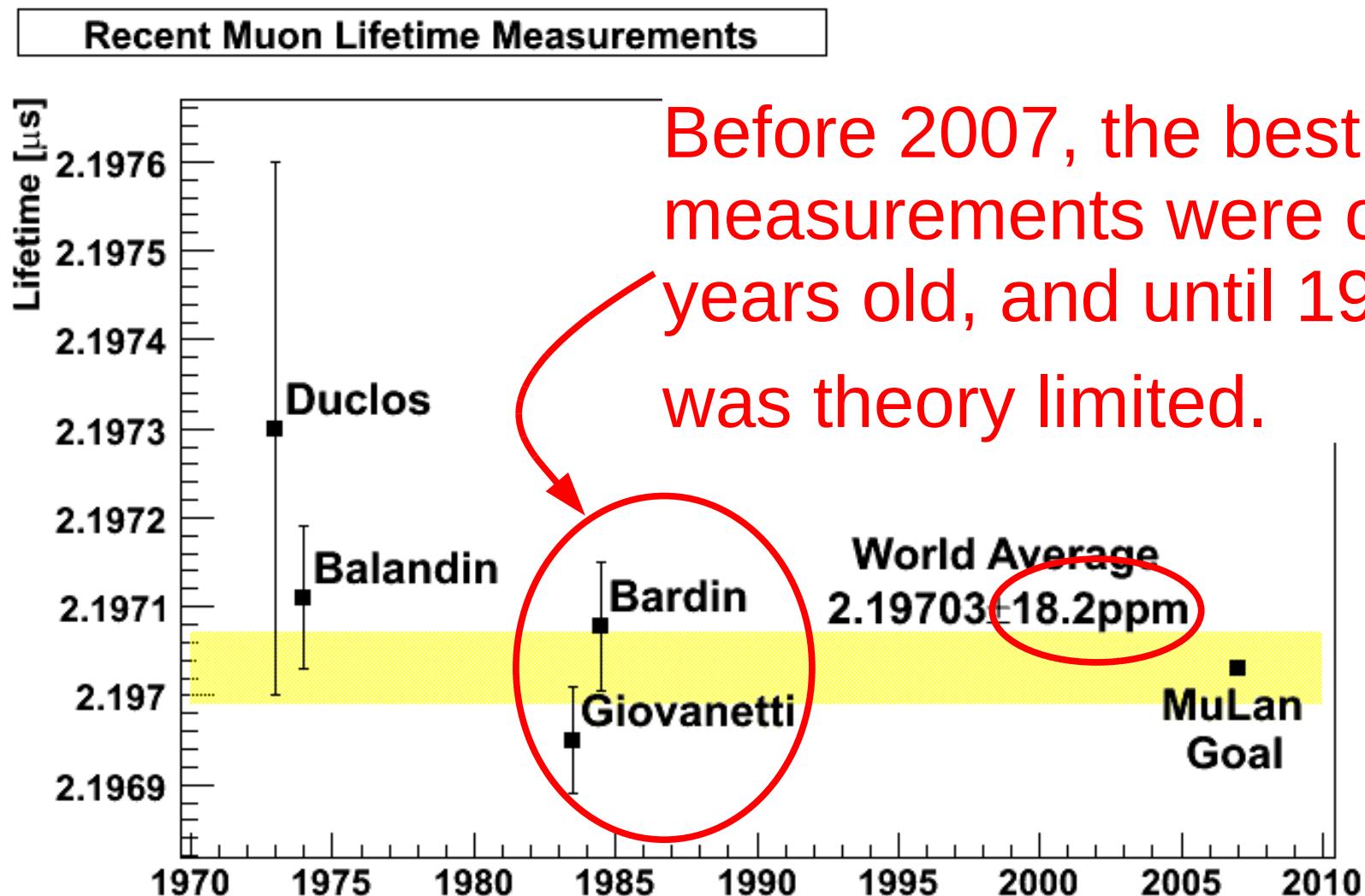
These measurements also calibrate some processes of astrophysical interest



Solar pp fusion cycle
 $p + p \rightarrow d + e^- + \bar{\nu}_e$

vd scattering in SNO
 $\nu_\mu + d \rightarrow \mu^- + n + n$

A brief history...



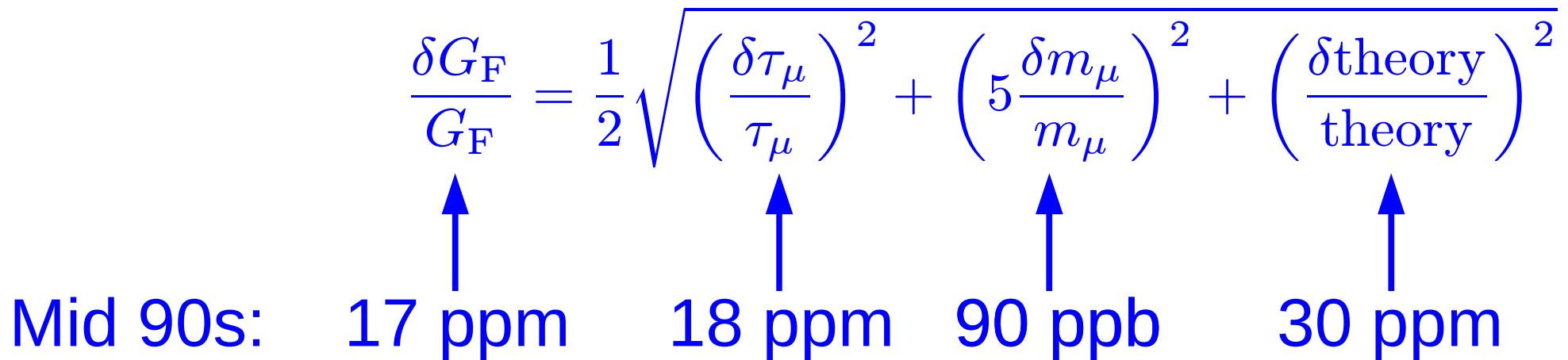
G. Bardin et al., Phys. Lett. B 137, 135 (1984)

K. Giovanetti et al., Phys. Rev. D 29, 343 (1984)

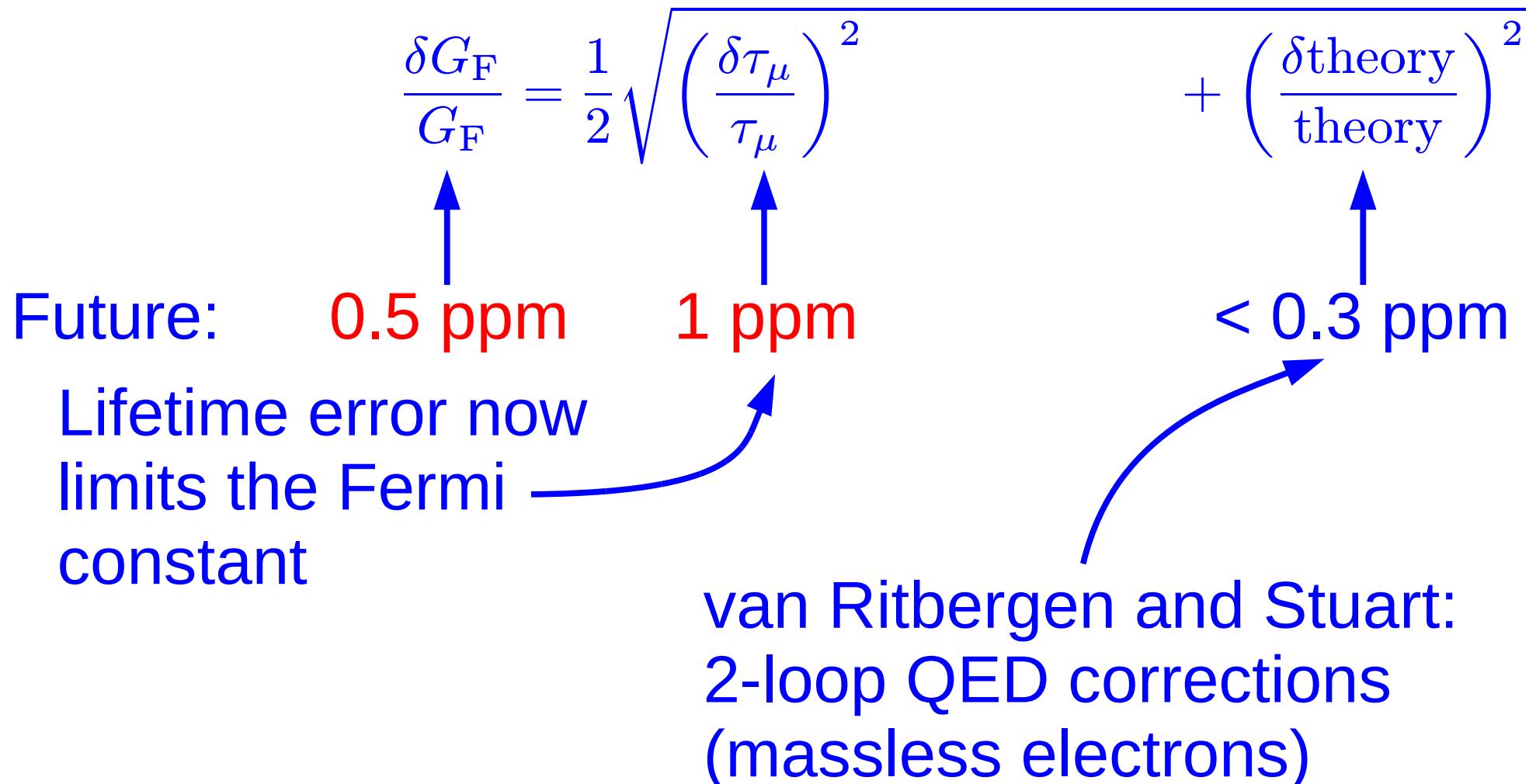
The Standard Model Fermi extraction is no longer theory limited

$$\frac{\delta G_F}{G_F} = \frac{1}{2} \sqrt{\left(\frac{\delta\tau_\mu}{\tau_\mu}\right)^2 + \left(5\frac{\delta m_\mu}{m_\mu}\right)^2 + \left(\frac{\delta\text{theory}}{\text{theory}}\right)^2}$$

Mid 90s: 17 ppm 18 ppm 90 ppb 30 ppm

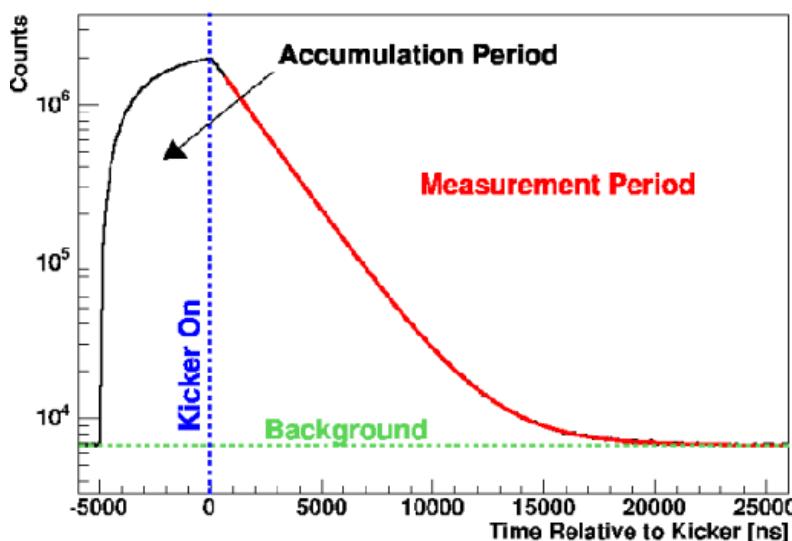


The Standard Model Fermi extraction is no longer theory limited



What exactly is the “lifetime” of a particle?

If an unstable particle exists at a certain time, then it has a fixed (history independent) probability of decaying in the next “clock tick”.



$$p_d(dt) = \frac{1}{\tau} dt \quad \text{Not normalized!}$$

$$p_s(t + dt) = p_s(t) (1 - p_d(dt))$$

$$\frac{dp_s(t)}{dt} = -\frac{1}{\tau} p_s(t)$$

$$p_s(t) = \frac{1}{\tau} e^{-t/\tau} \quad \text{Normalized!}$$

τ is called the lifetime.

How do you measure the muon lifetime?

One-at-a-time



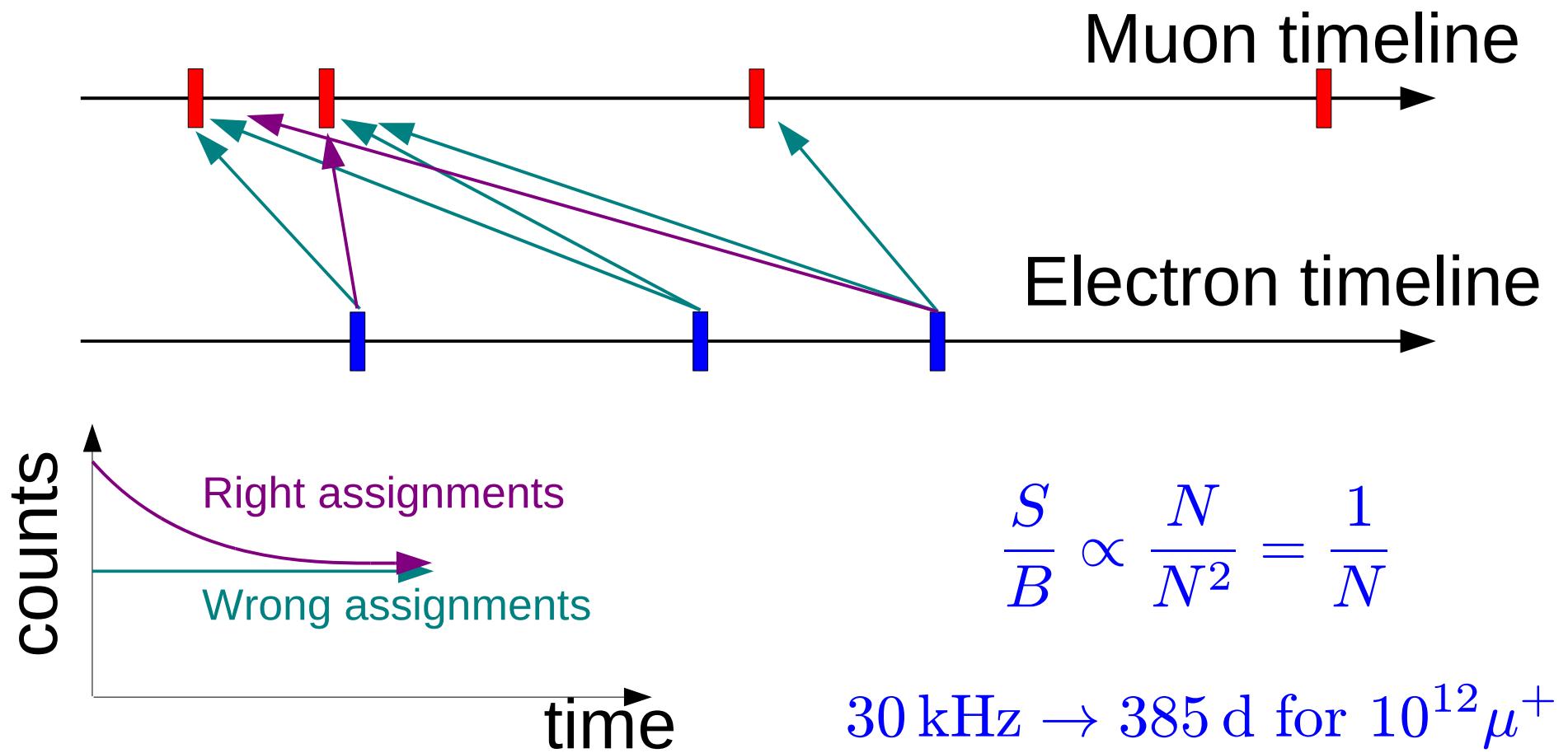
Many-at-once



$$\tau_\mu = 2.197 \mu\text{s}$$

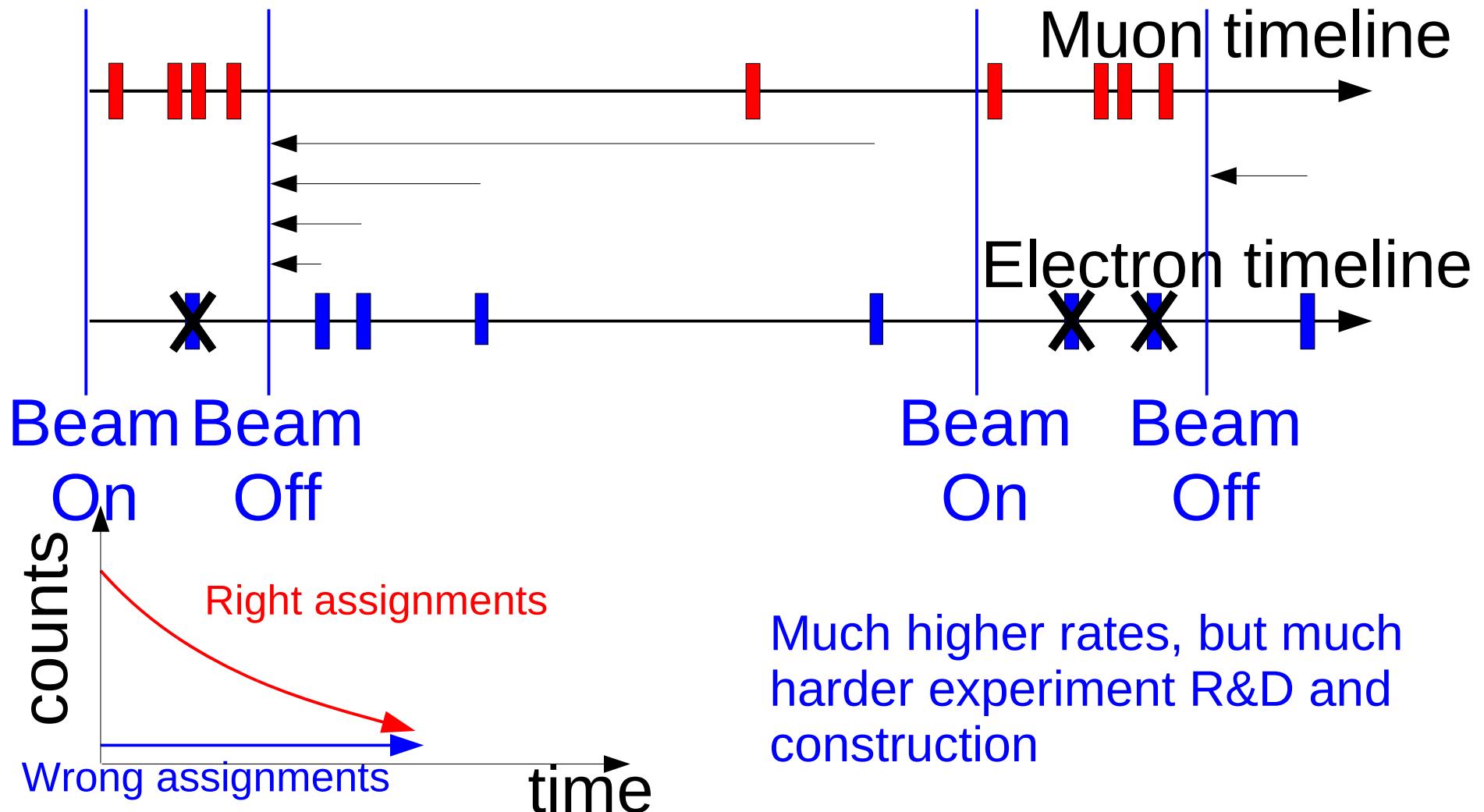
One-at-a-time

Can't really do one-at-a-time, the next best thing is a low rate, DC beam.

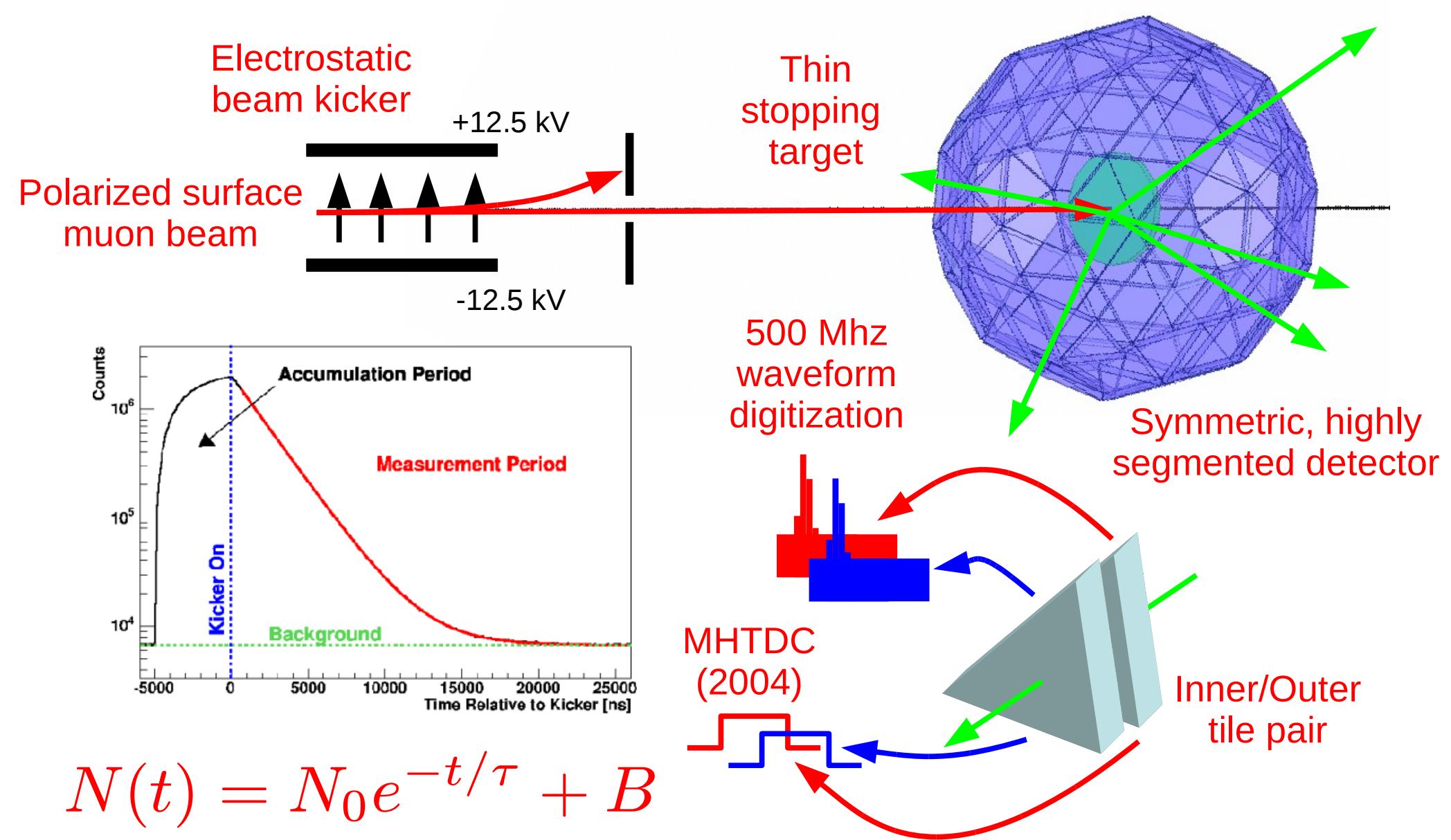


Many-at-once

Need time structured (AC) beam, not a continuous (DC) beam



We will reach our goal by running many muon decay experiments simultaneously

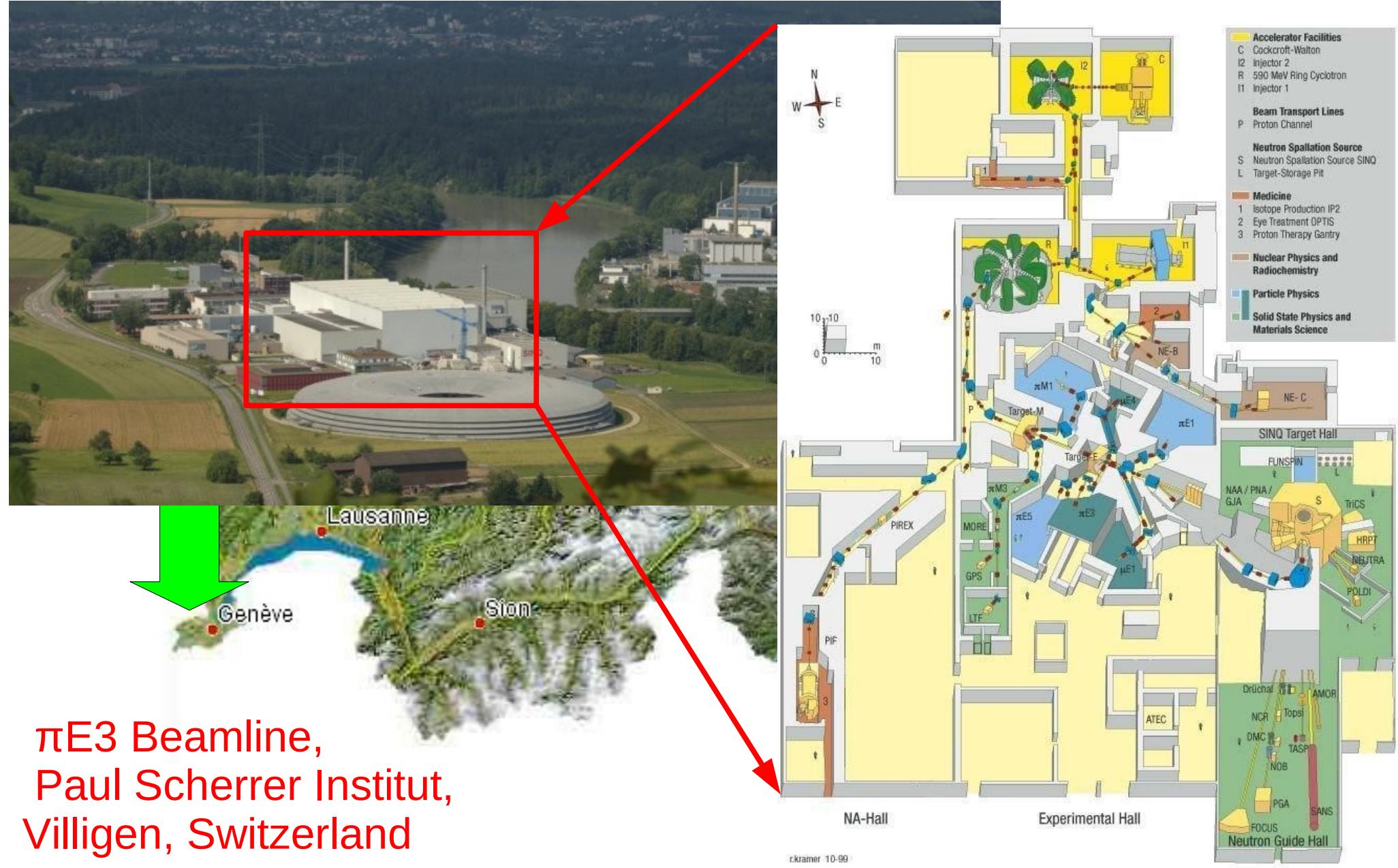


Finding muons isn't such a problem

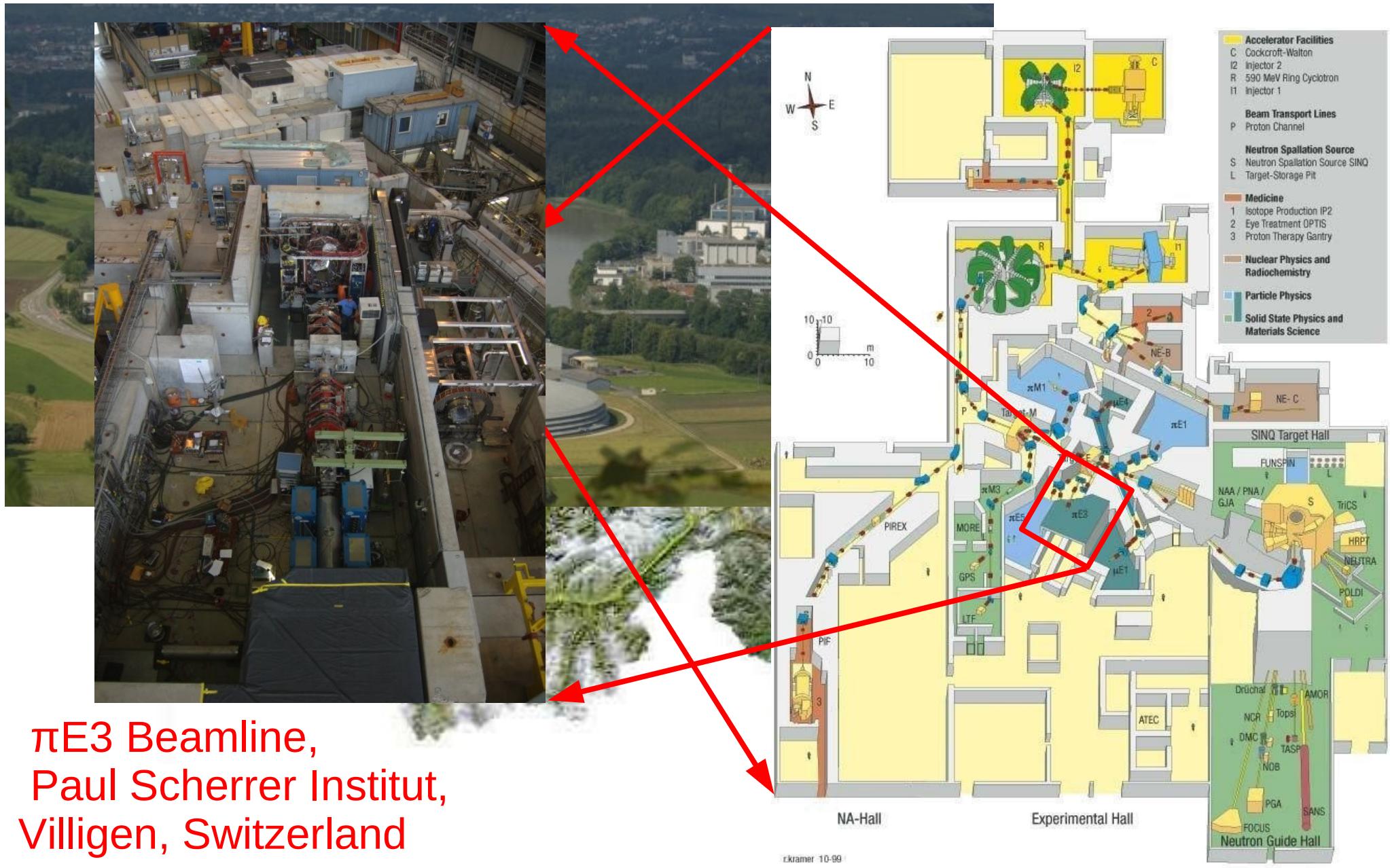


πE3 Beamlne,
Paul Scherrer Institut,
Villigen, Switzerland

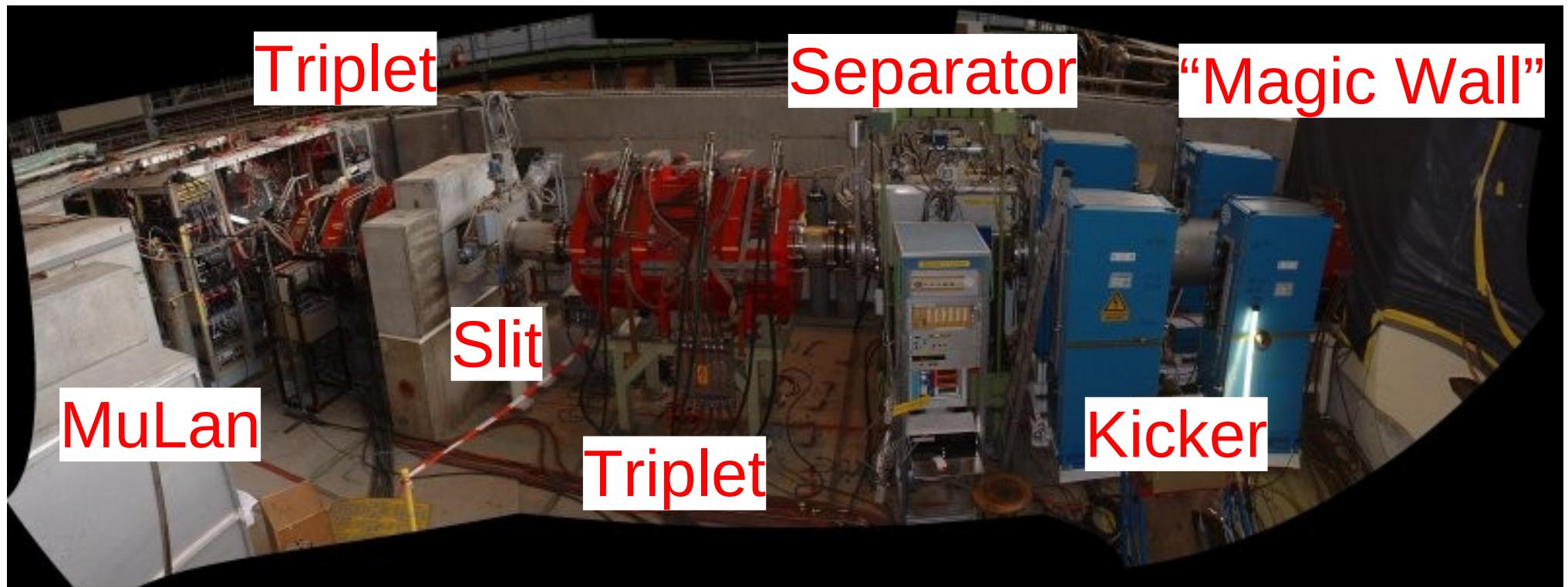
Finding muons isn't such a problem



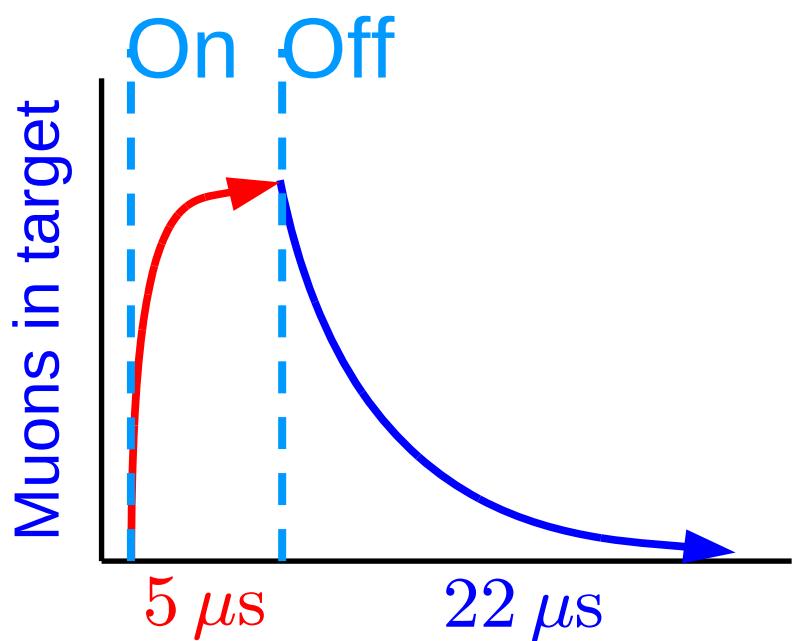
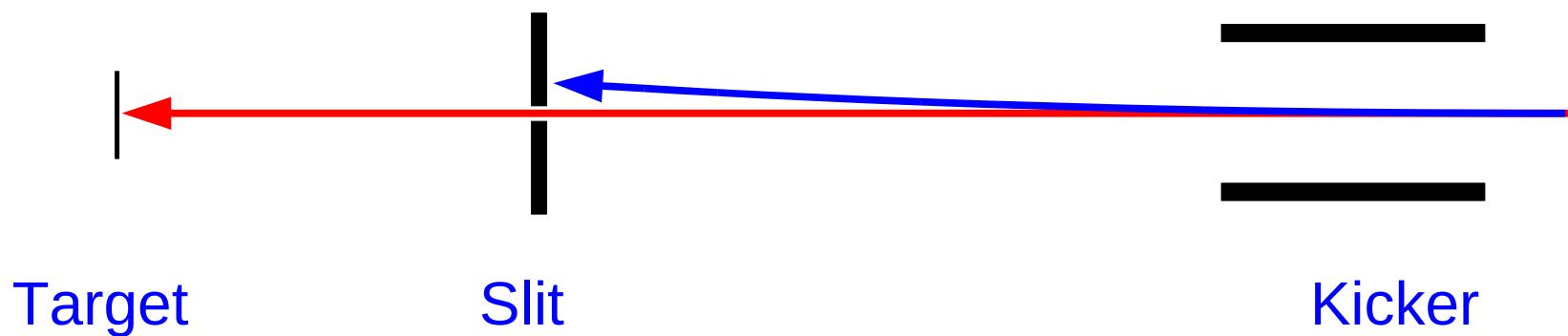
Finding muons isn't such a problem



Filling the bucket

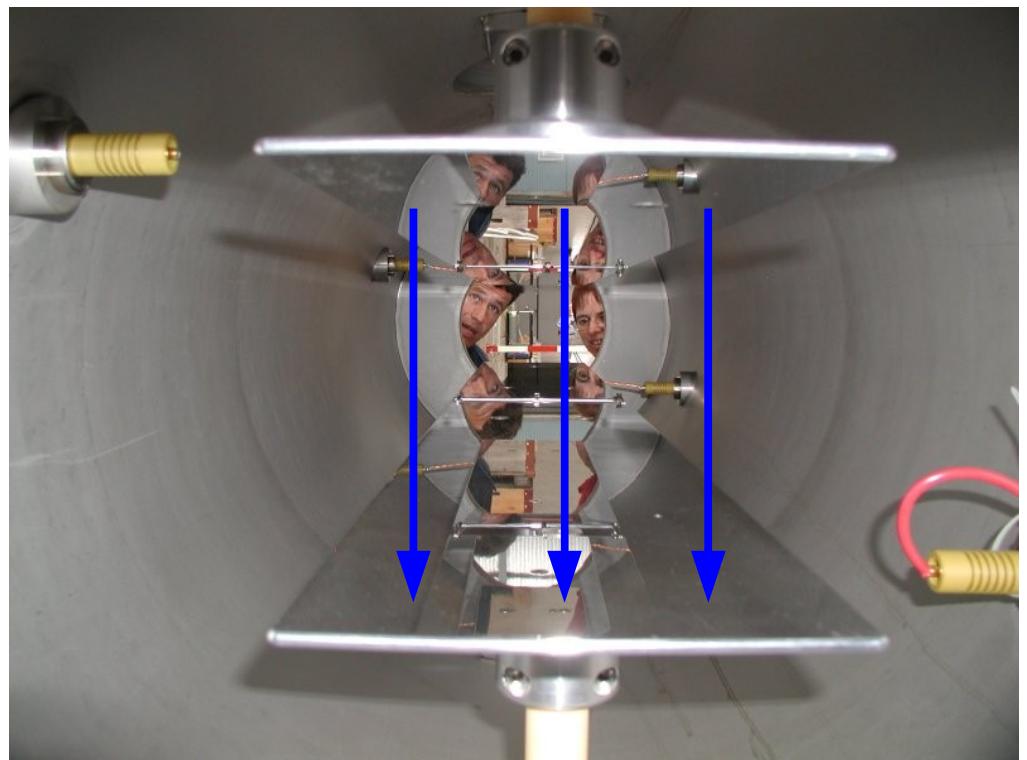


Time structuring the beam

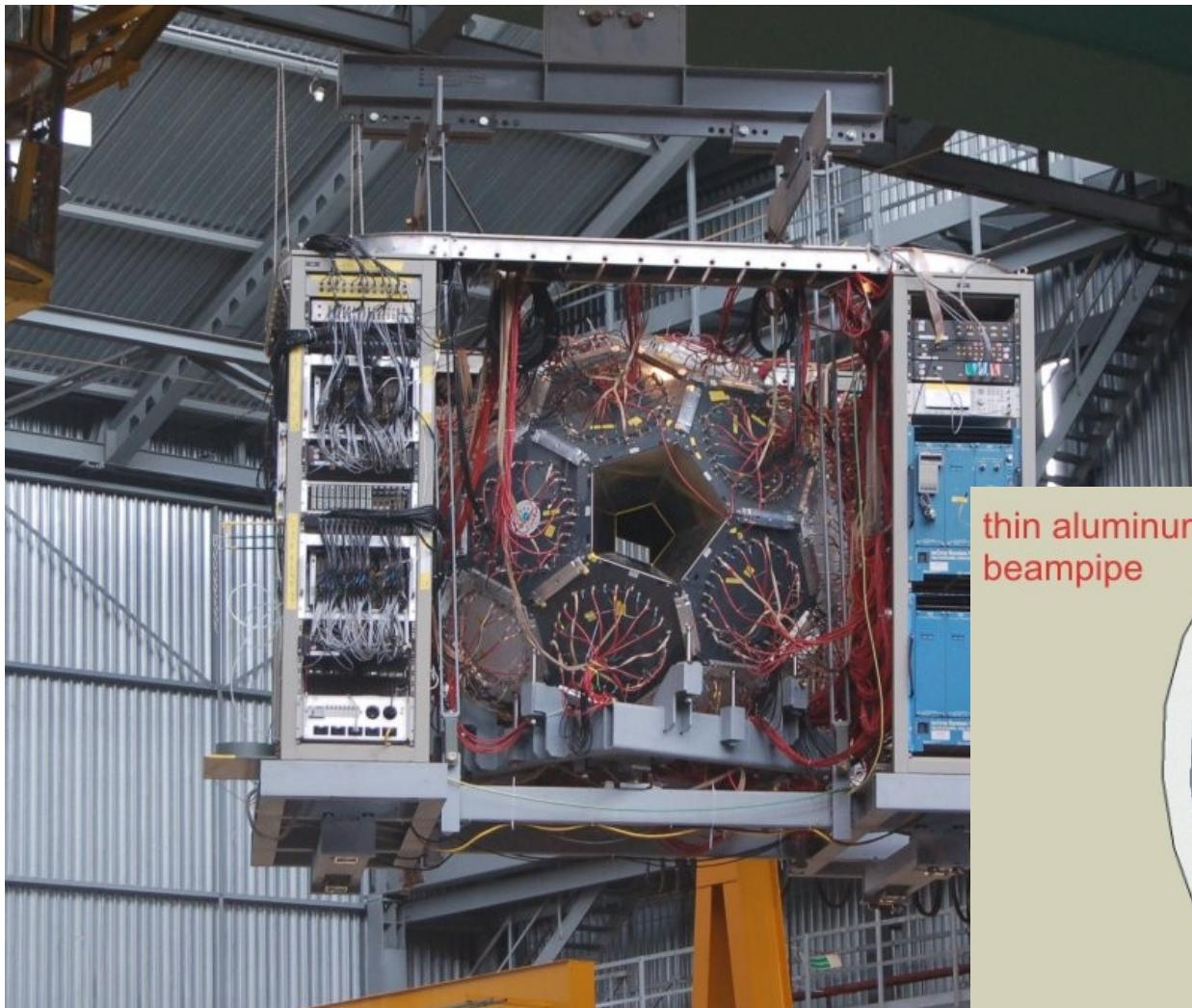


$$N_{\text{in}}(t) = R_\mu \tau \left(1 - e^{-t/\tau}\right)$$

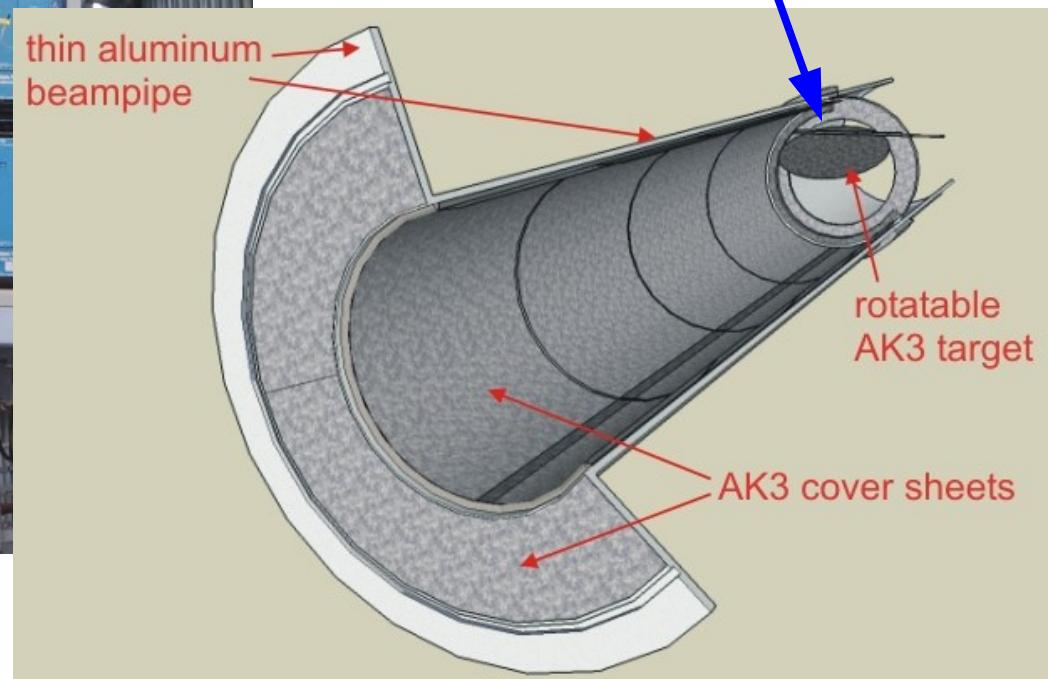
$$N_{\text{out}}(t) = N_{\text{in}}(t_c) e^{-t/\tau}$$



Where exactly is the bucket?

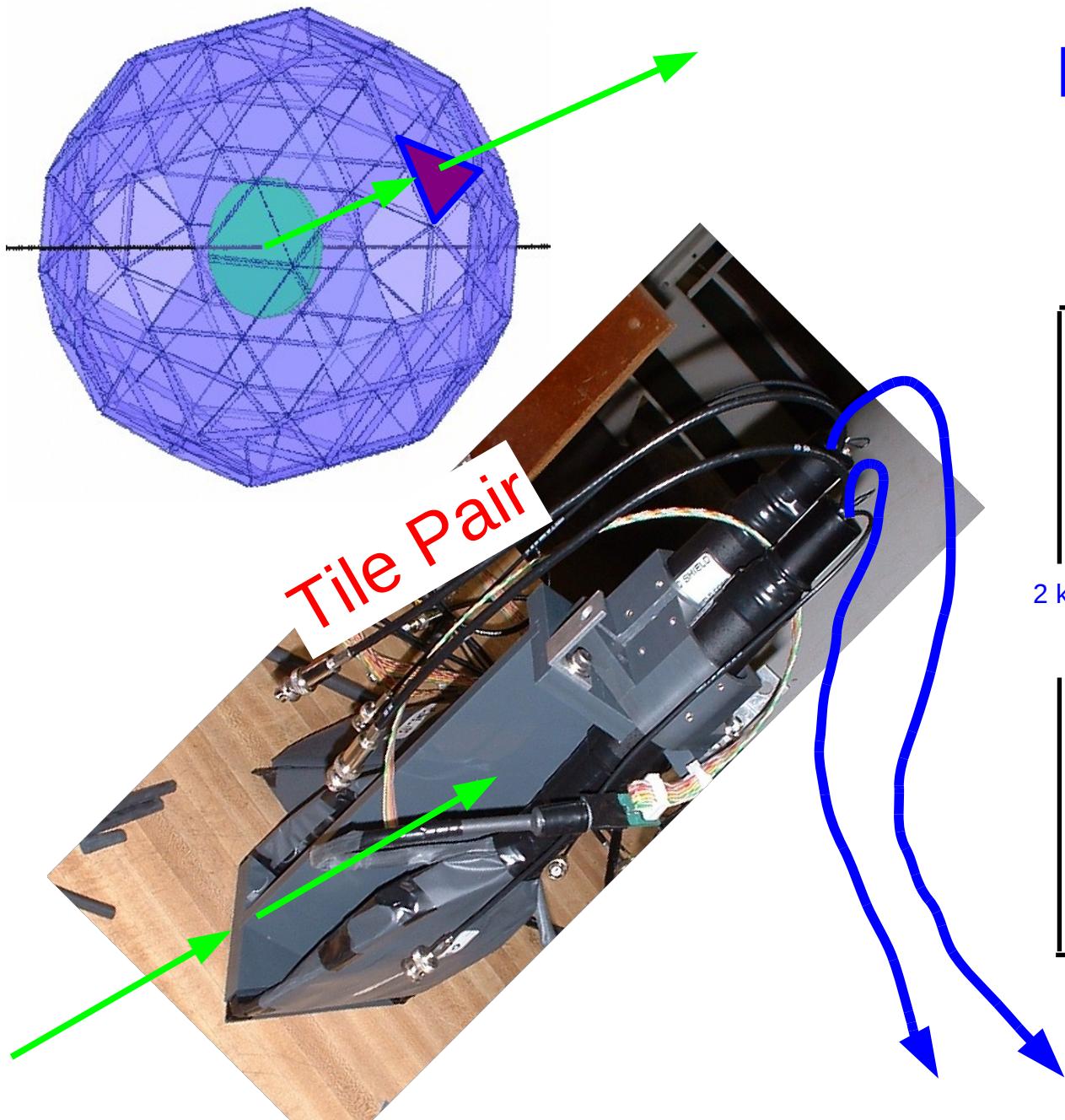


Right there!

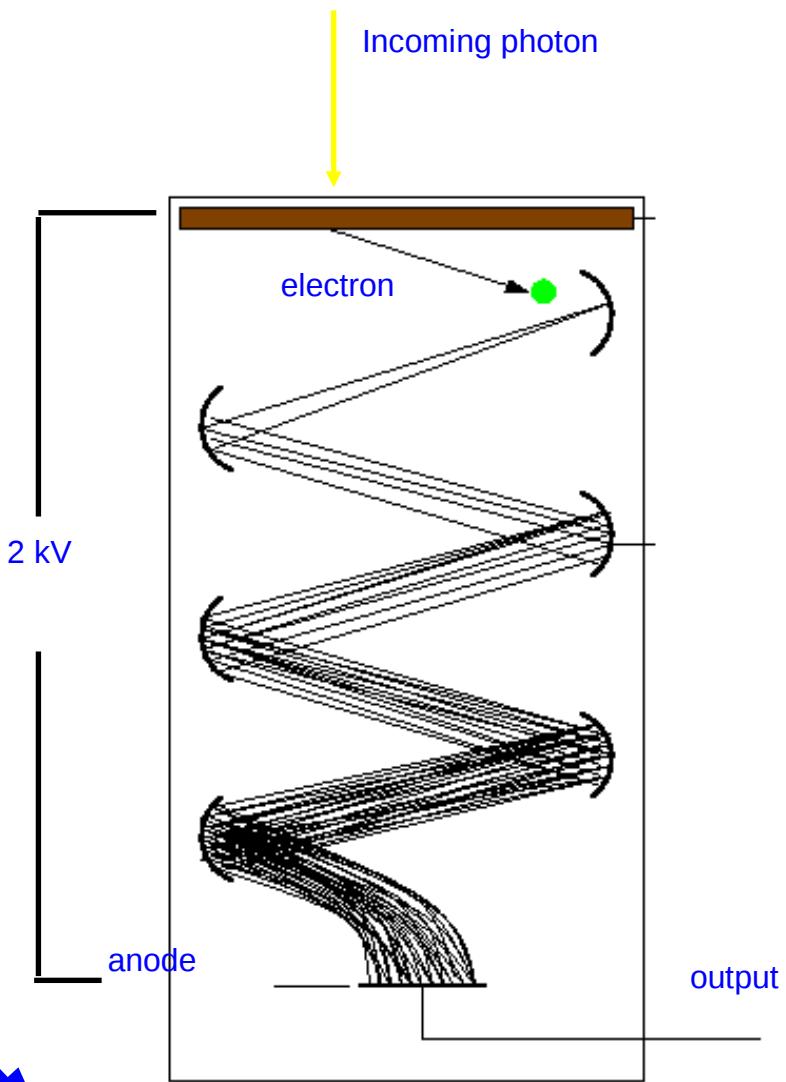


Rob Carey: “The world's largest research grade soccer ball”

Watching the bucket empty

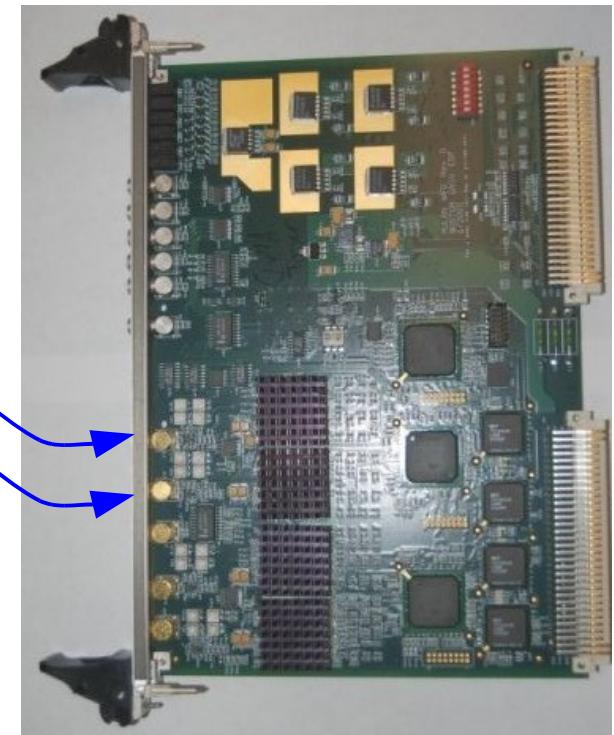
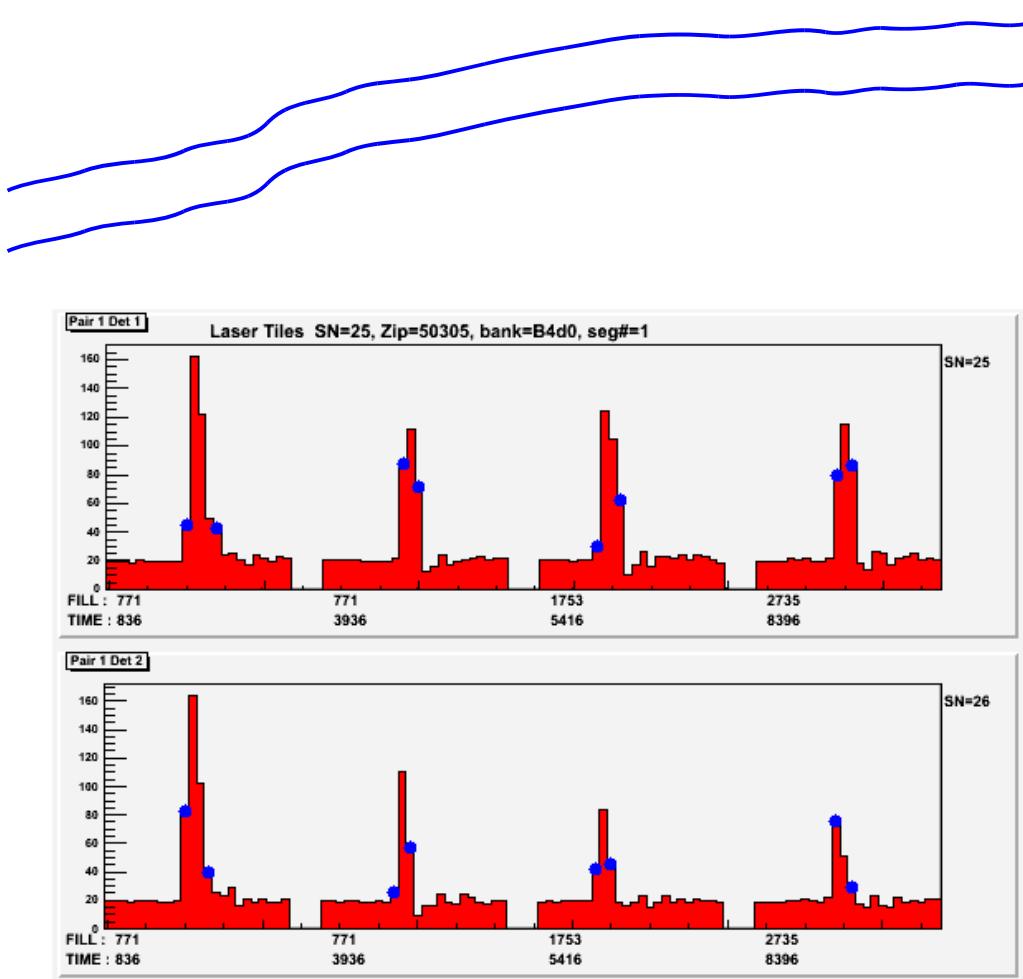


Photomultiplier tube



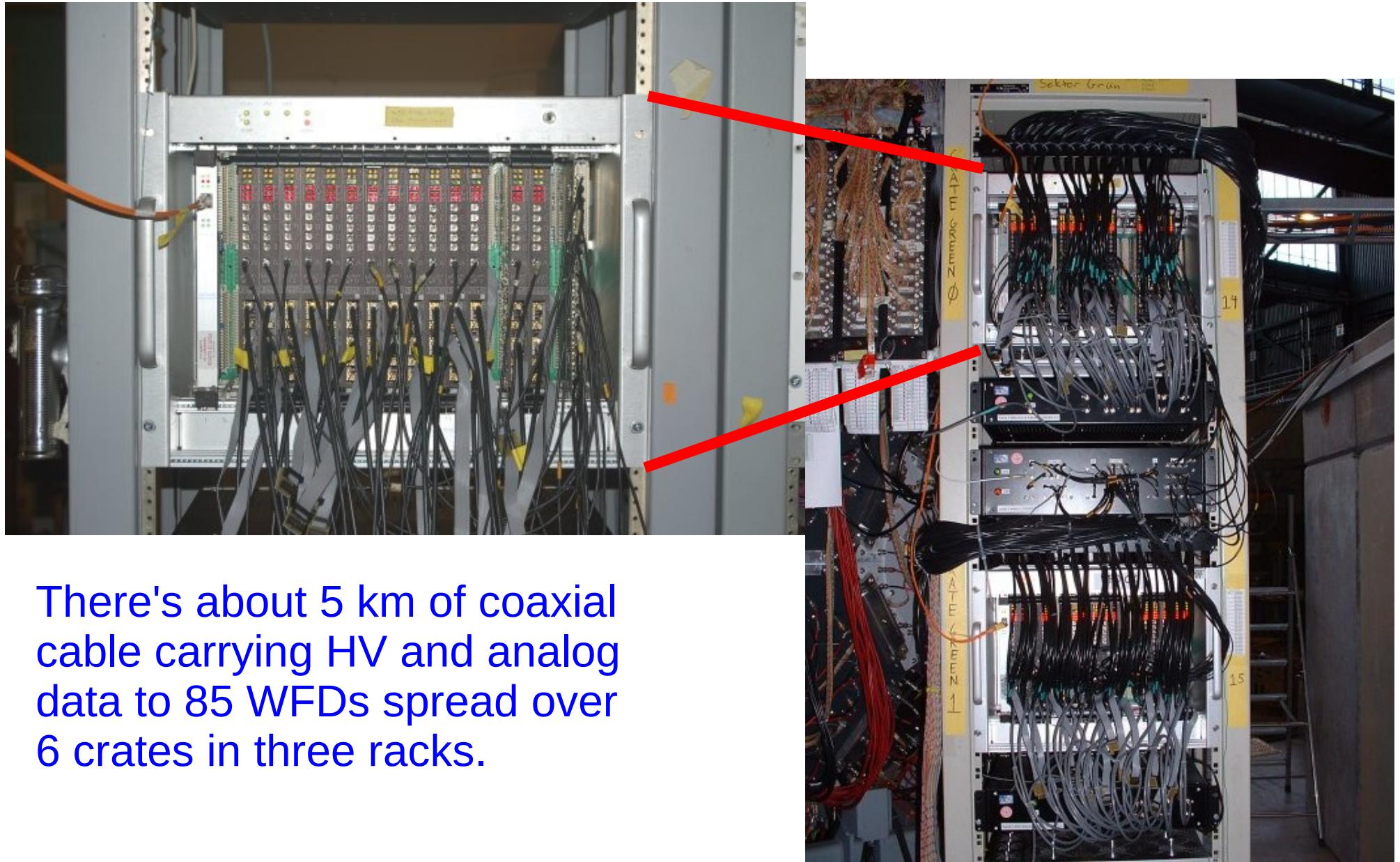
Watching the bucket empty

The PMTs feed the WFDs



This gives about 40 MB/s of data that has to be stored!

Where does all that data go?

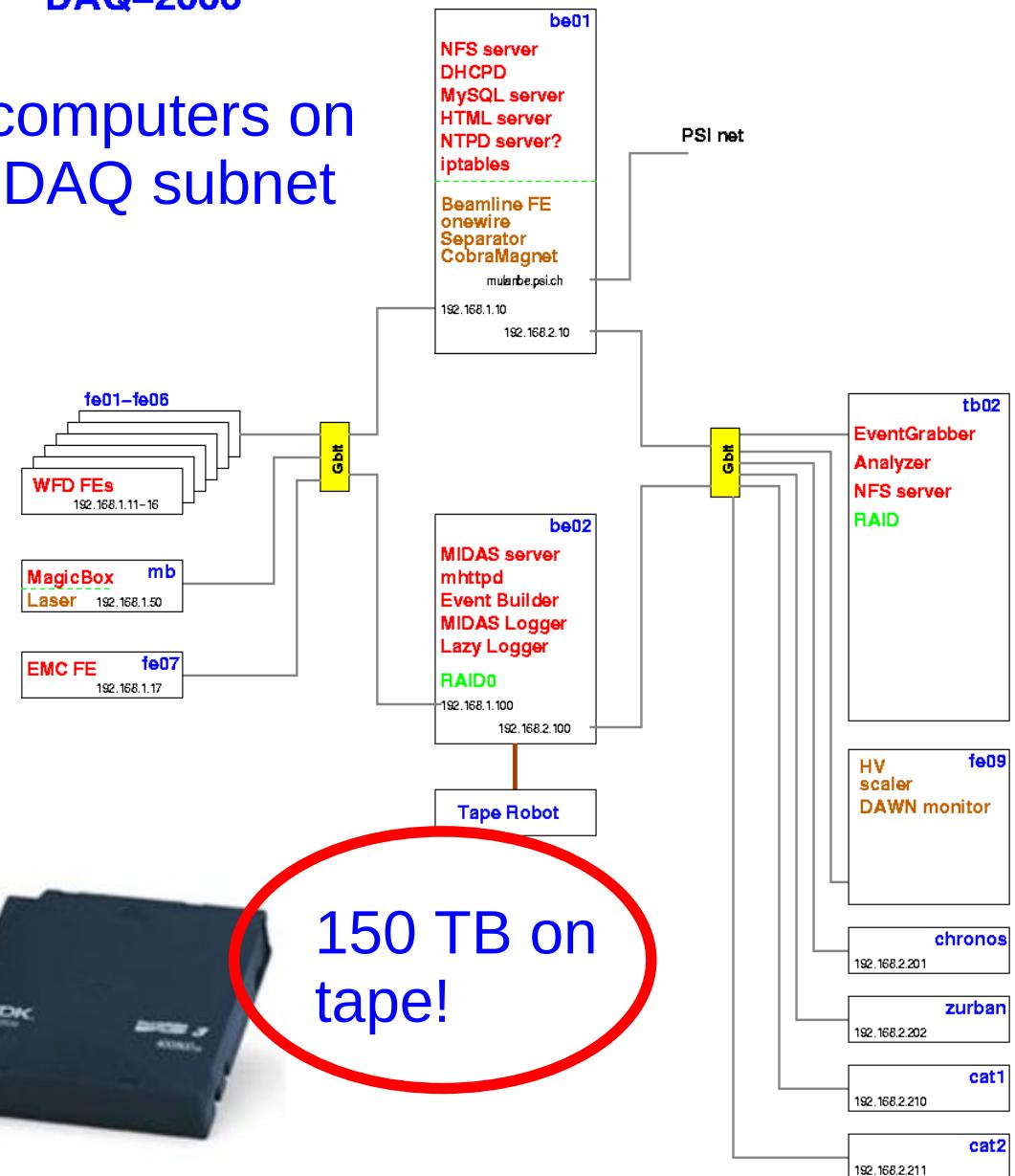


Computers and tapes galore!



DAQ-2006

16 computers on
the DAQ subnet



150 TB on
tape!

Time-dependent systematics are the core concern for a 10^{12} data set

Early-to-late changes, for instance:

Instrumental issues

PMT gains

Discriminator threshold

walk

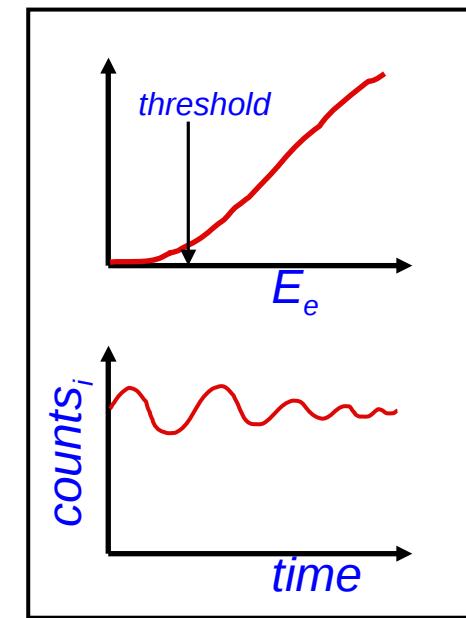
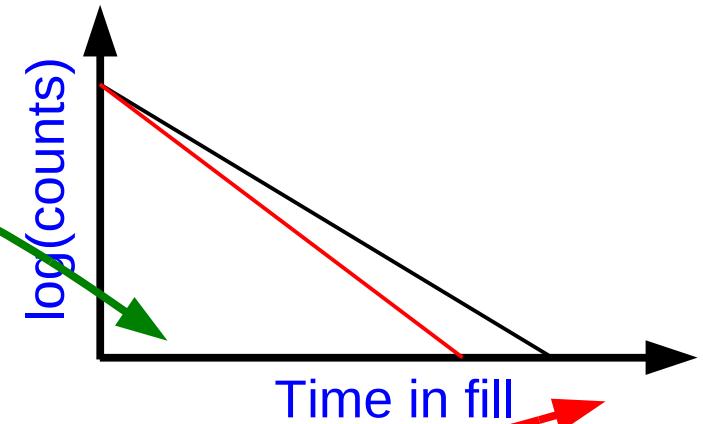
Kicker voltage sag

Pileup

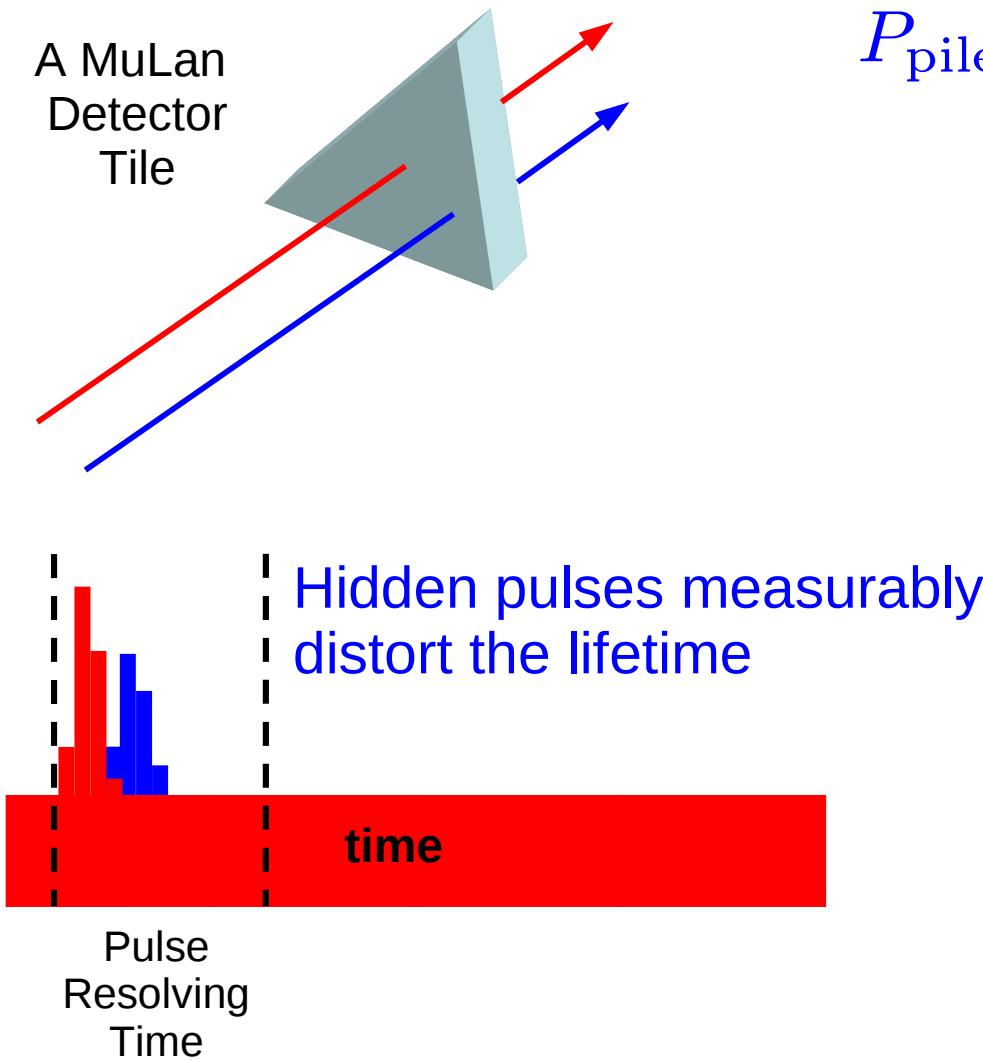
Physics issues

Spin polarization

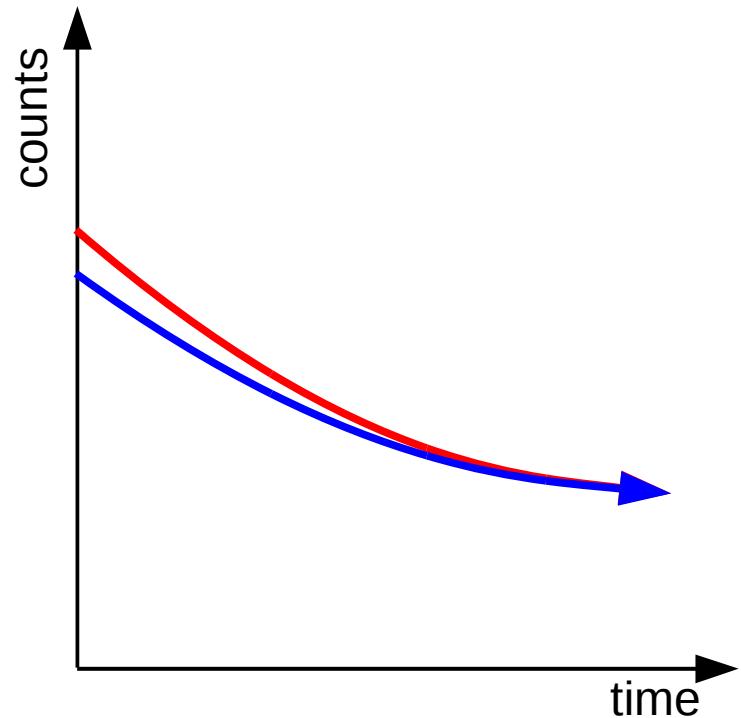
Non-flat background sources



What's pileup?

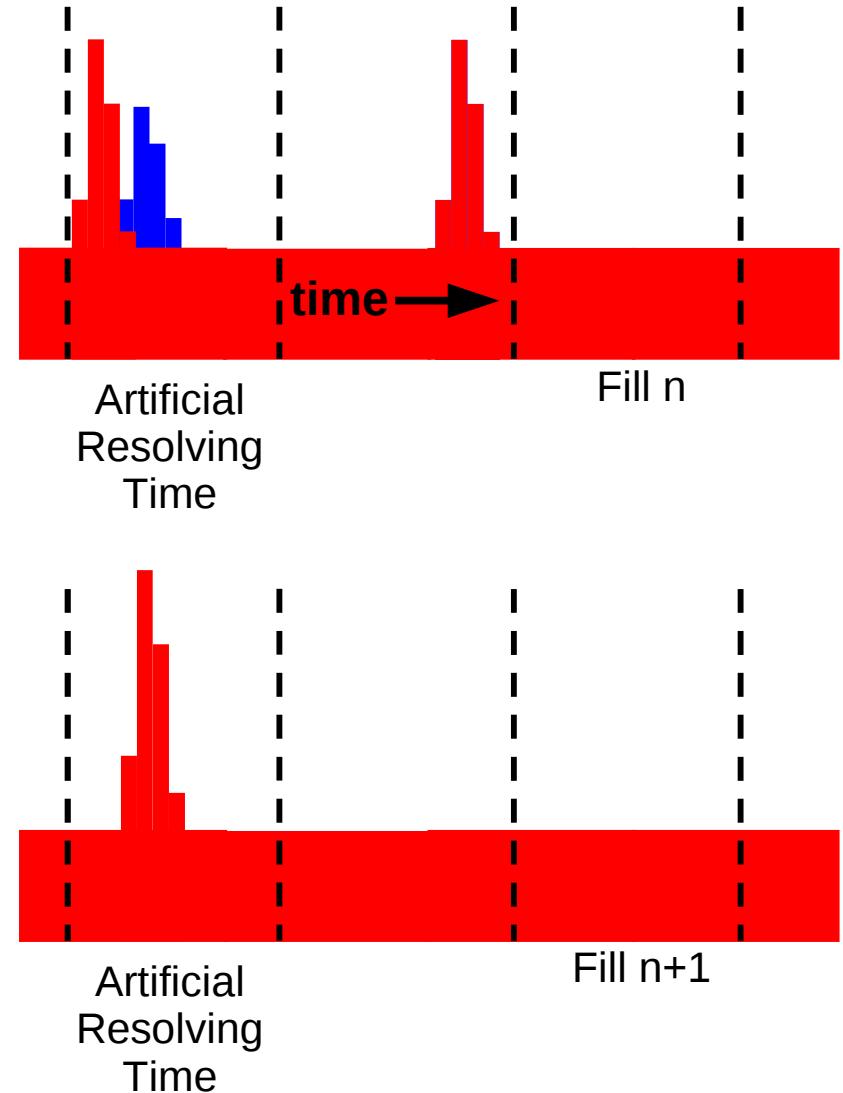
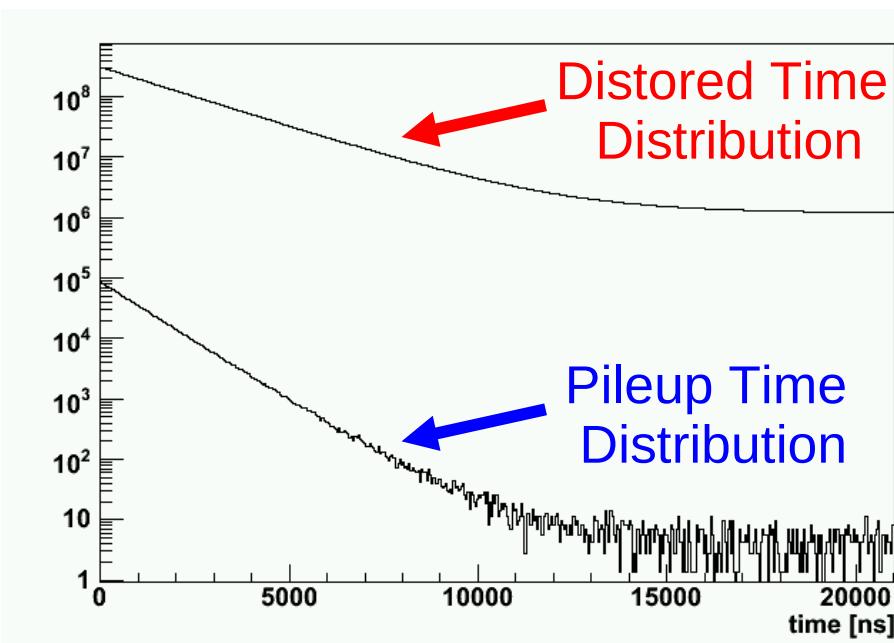
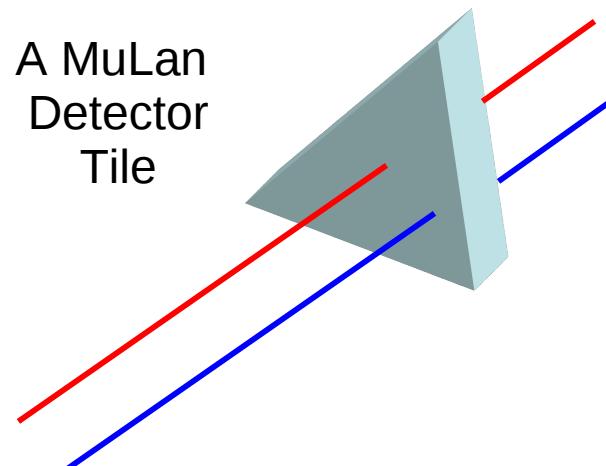


$$P_{\text{pileup}} \propto \int_0^{t_r} P(t)P(t+t')dt' \propto e^{-2t/\tau}$$



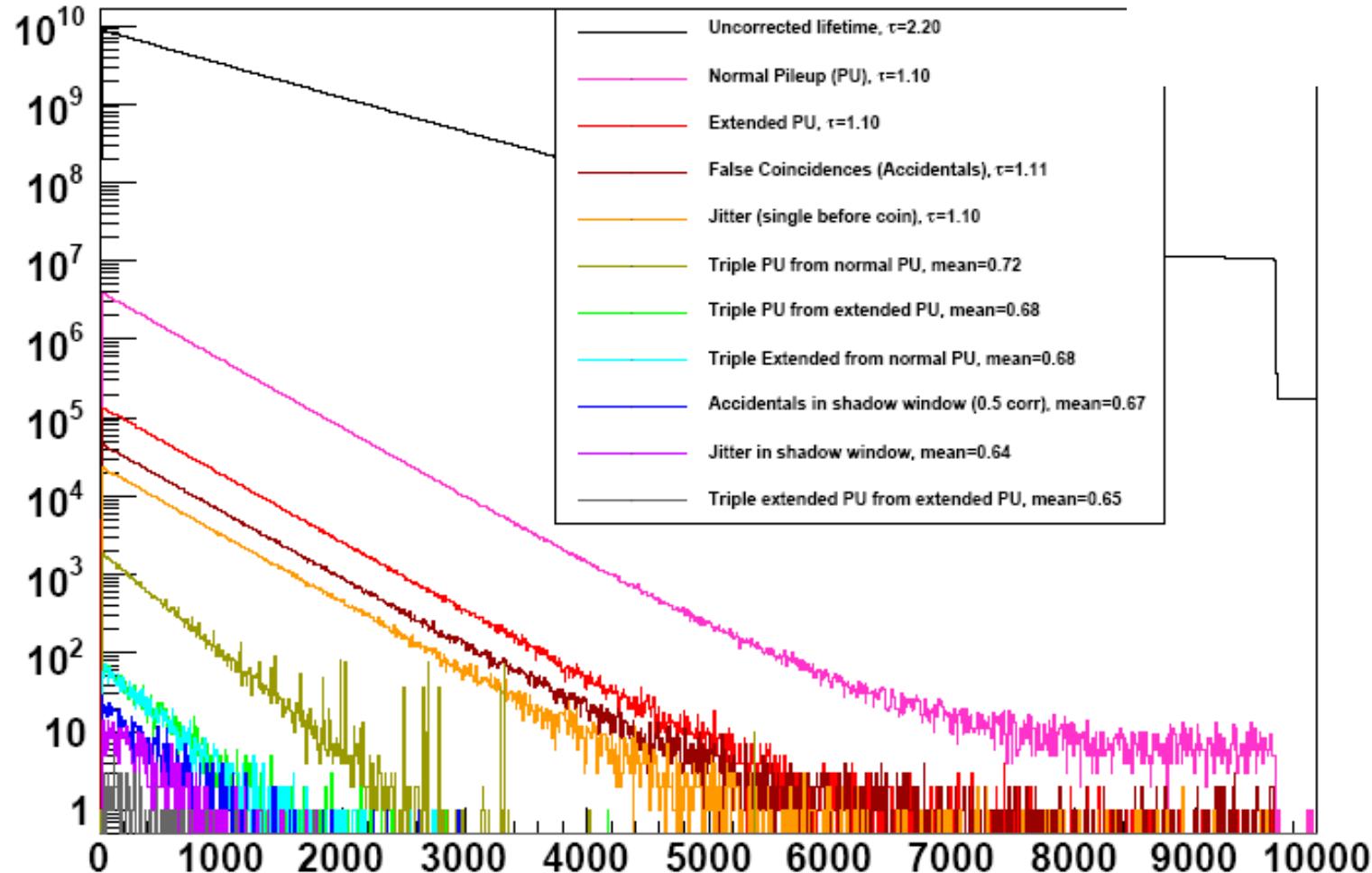
We could fit for this, at a significant cost in statistical error ... but we can actually use the data itself to construct a correction function!

Here's how it's done



Adding the pileup distribution to the normal distribution (statistically) corrects for what's killed by the imposed deadtime!

In practice, there are many different pileup correction terms



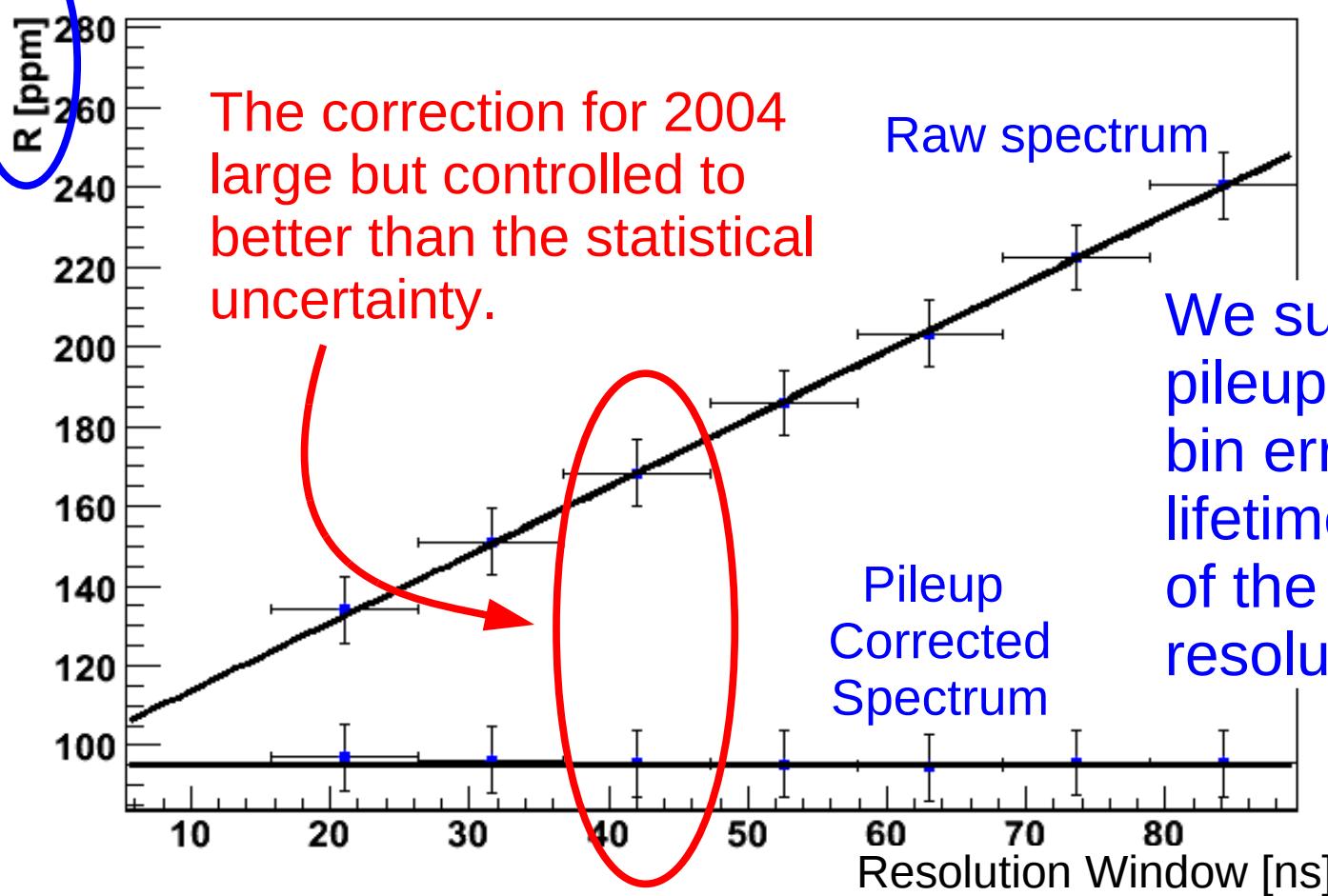
Left uncorrected, these terms shift the lifetime fit by hundreds of ppm at large resolving times.

How well does this method correct pileup?

Blind analysis!

Measured τ vs Pulse Resolution

2004

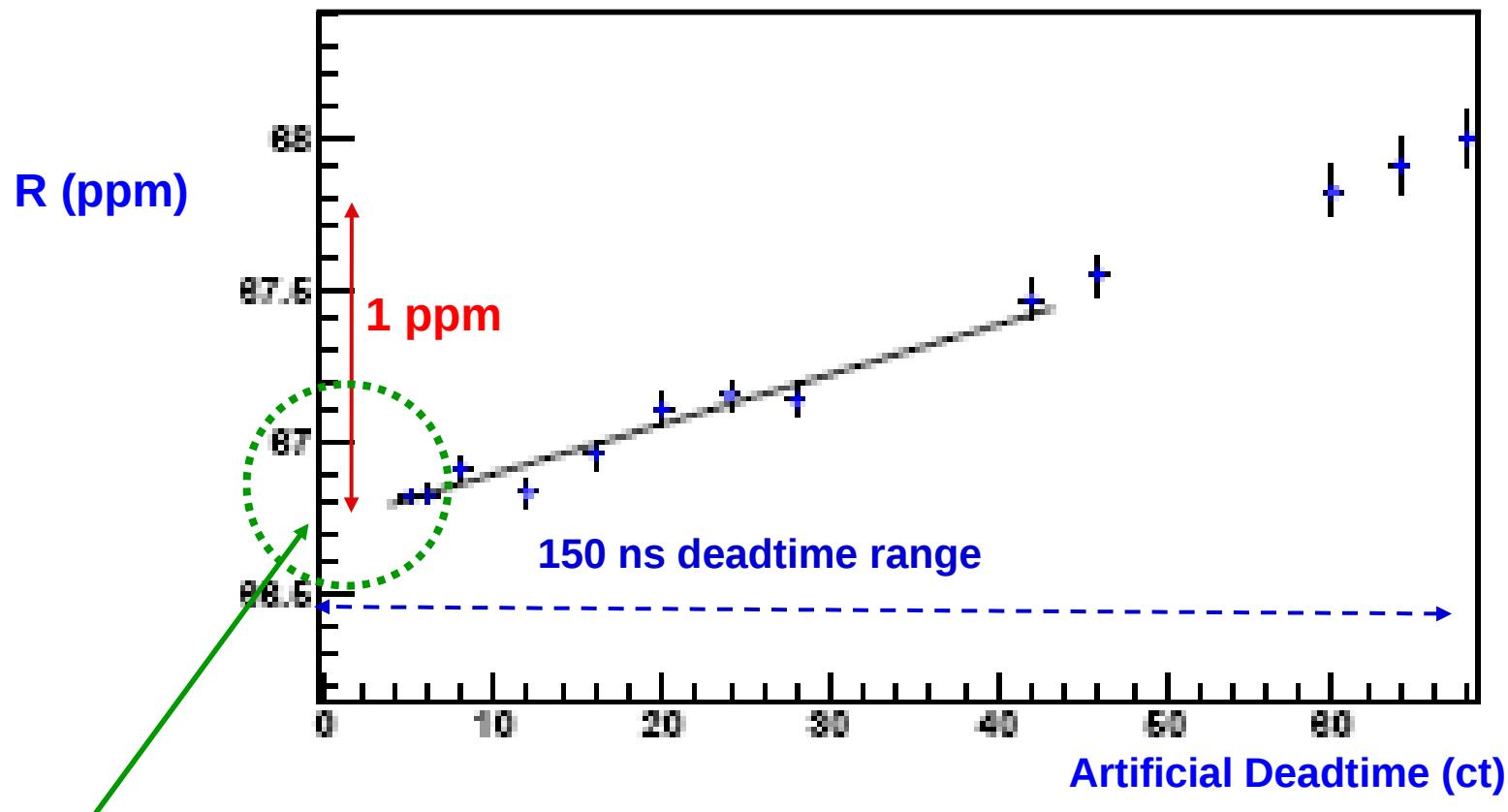


We sum the normal and pileup spectra (and correct bin errors!); the corrected lifetime fit is independent of the width of the time resolution window.

(\pm) ppm

How well does this method correct pileup?

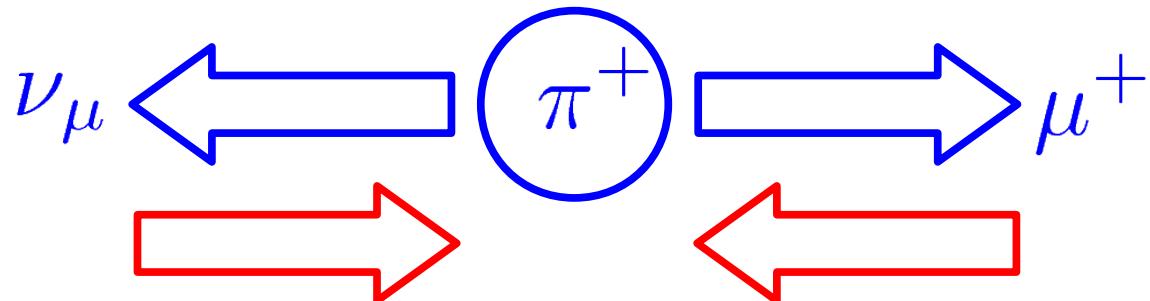
2006/2007



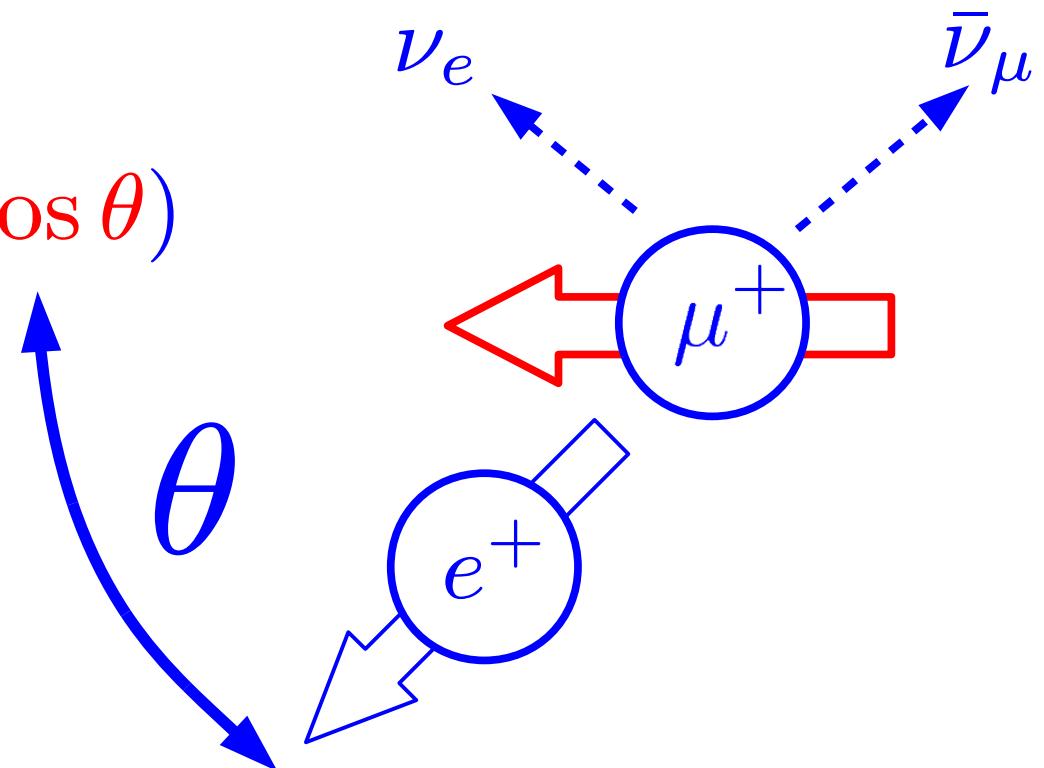
Extrapolation to 0 deadtime should be correct answer and our indications are that this extrapolation is right, but we continue to investigate the source of this shallow slope.

Muon beams are naturally polarized, and the Michel electron is not produced isotropically!

Parity violation in weak decays
requires left-handed neutrinos



$$\frac{d^2\Gamma_\mu^\pm}{dyd\theta} = n(y) (1 \mp a(y)\cos\theta)$$



Add in spin precession in magnetic fields and material based spin exchange interactions, and things can get complicated very quickly!

$$f(t) = N \left[1 + \frac{1}{3} \vec{P}_1 \cdot \hat{r}_D e^{-t/T_1} + \frac{1}{3} P_2 \sin(\omega t + \phi) e^{-t/T_2} \right] e^{-t/\tau_\mu} + B$$

Longitudinal component

Transverse component

Muon Lifetime

Flat background

The diagram illustrates the decomposition of the function $f(t)$ into several components. A blue arrow points from the term $\vec{P}_1 \cdot \hat{r}_D e^{-t/T_1}$ to the text "Longitudinal component". Another blue arrow points from the term $P_2 \sin(\omega t + \phi) e^{-t/T_2}$ to the text "Transverse component". A third blue arrow points from the term e^{-t/τ_μ} to the text "Muon Lifetime". A fourth blue arrow points from the constant term B to the text "Flat background".

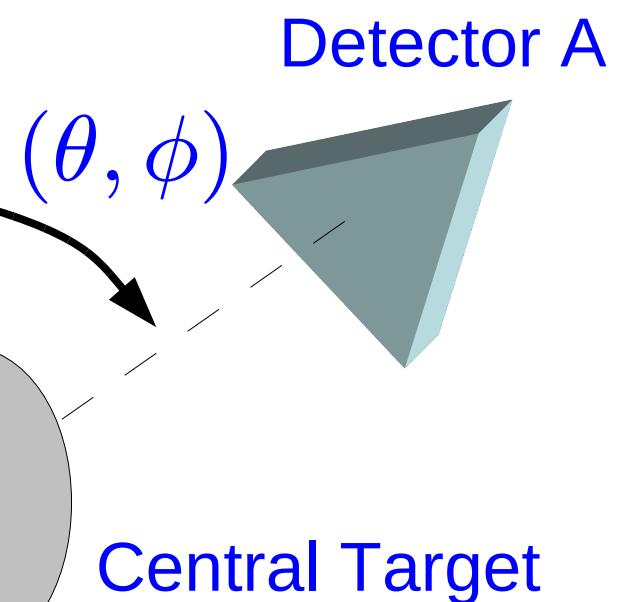
Any mismeasured polarization terms can have a large impact on the lifetime measurement

Since we start with nearly 100% polarized beam, how do we control polarization issues?

Point symmetry of the detector cancels polarization asymmetries in sum over symmetric tiles, up to source centrality and acceptance differences.

$$(\pi - \theta, \pi + \phi)$$

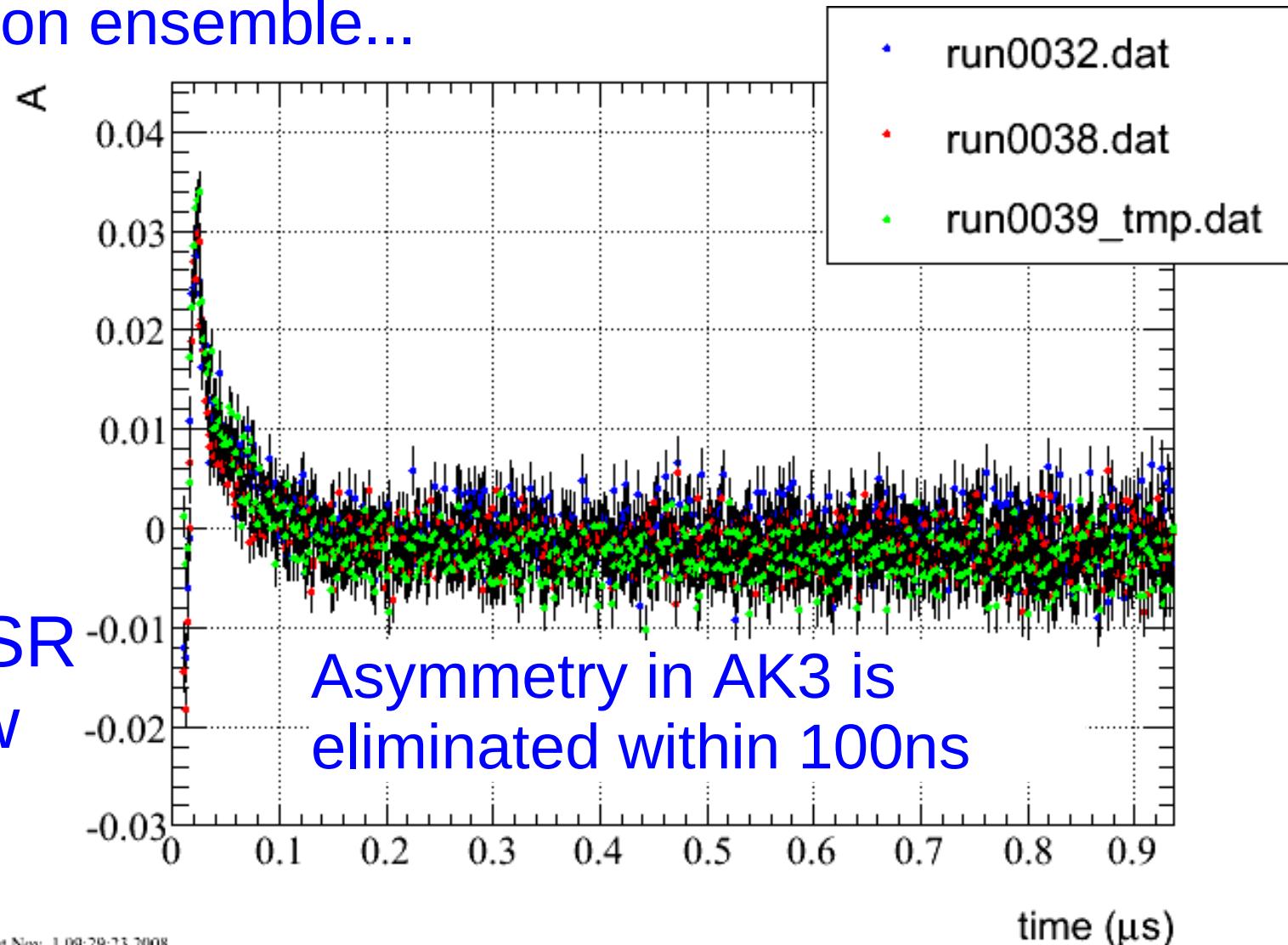
Detector A'



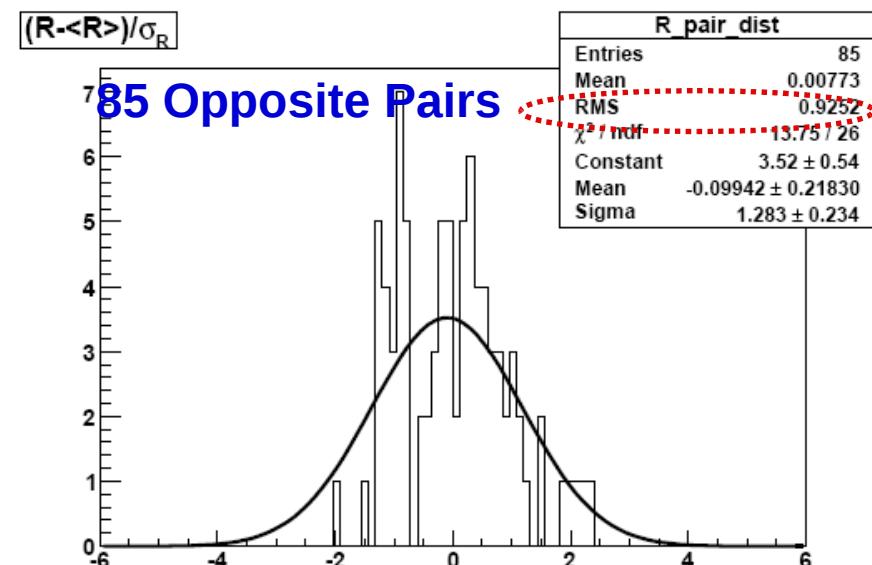
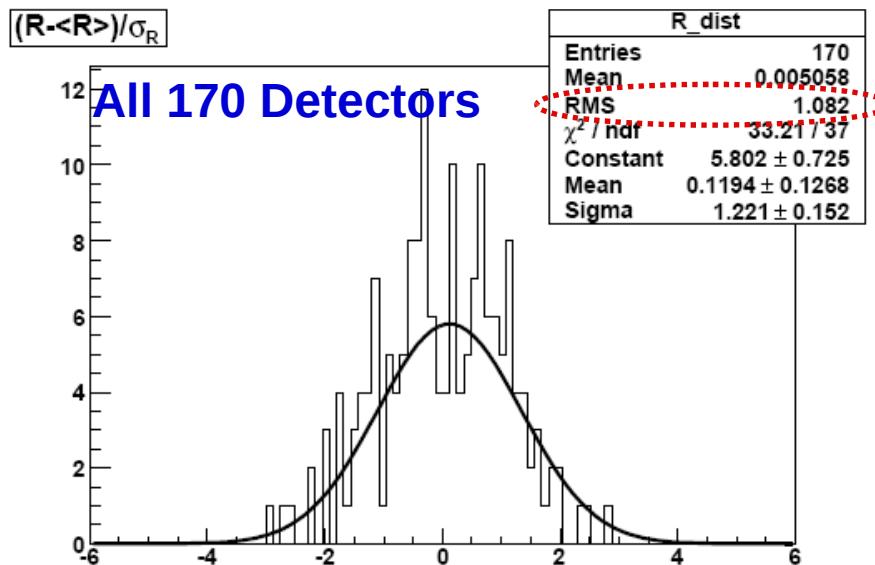
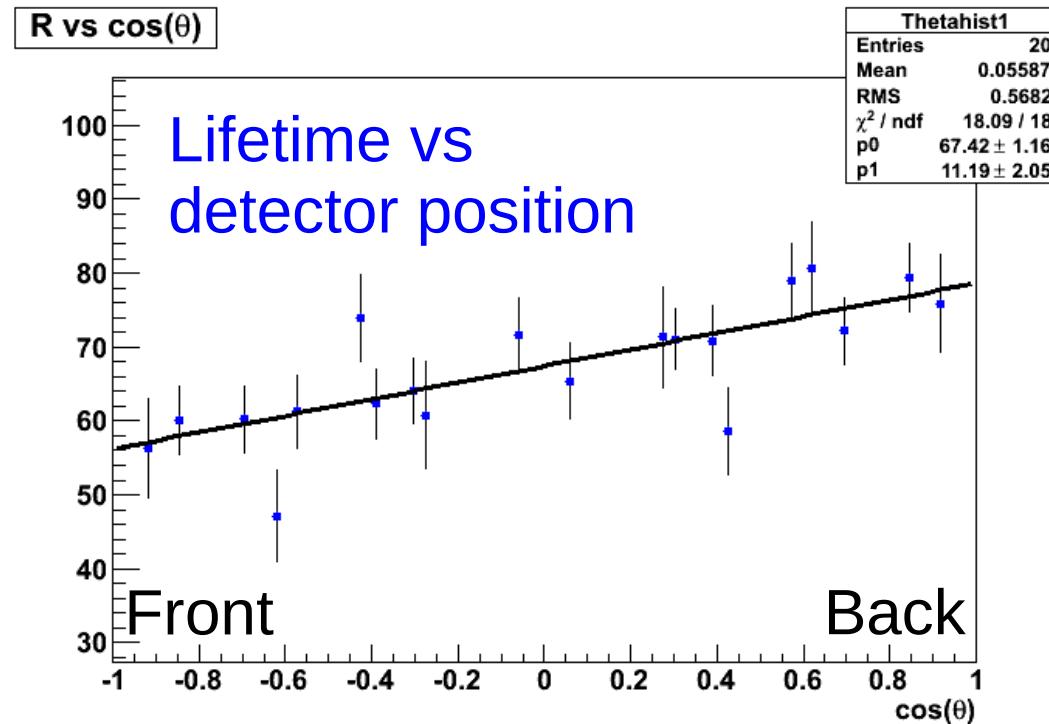
In 2006, we chose a target with high internal magnetic field (Arnokrome III) to minimize the residual polarization

The high internal field should rapidly dephase the incoming muon ensemble...

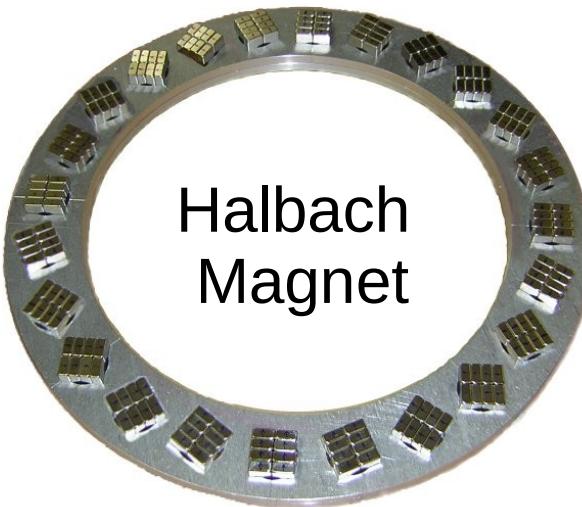
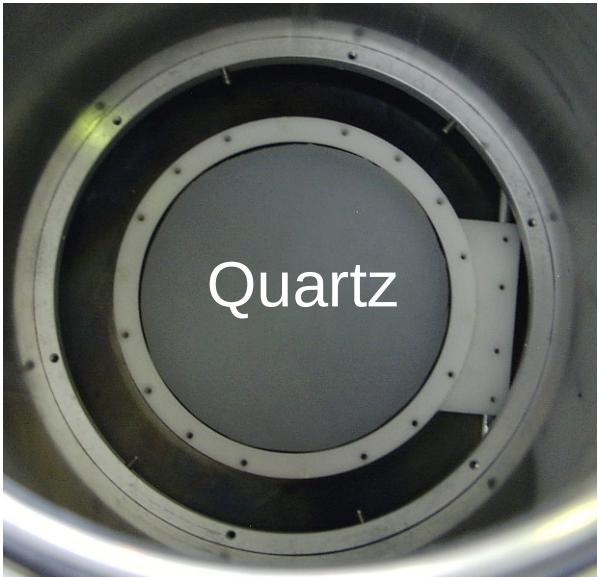
... and dedicated μ SR studies show exactly that.



There's a small longitudinal remnant, but it cancels in the pointwise sums



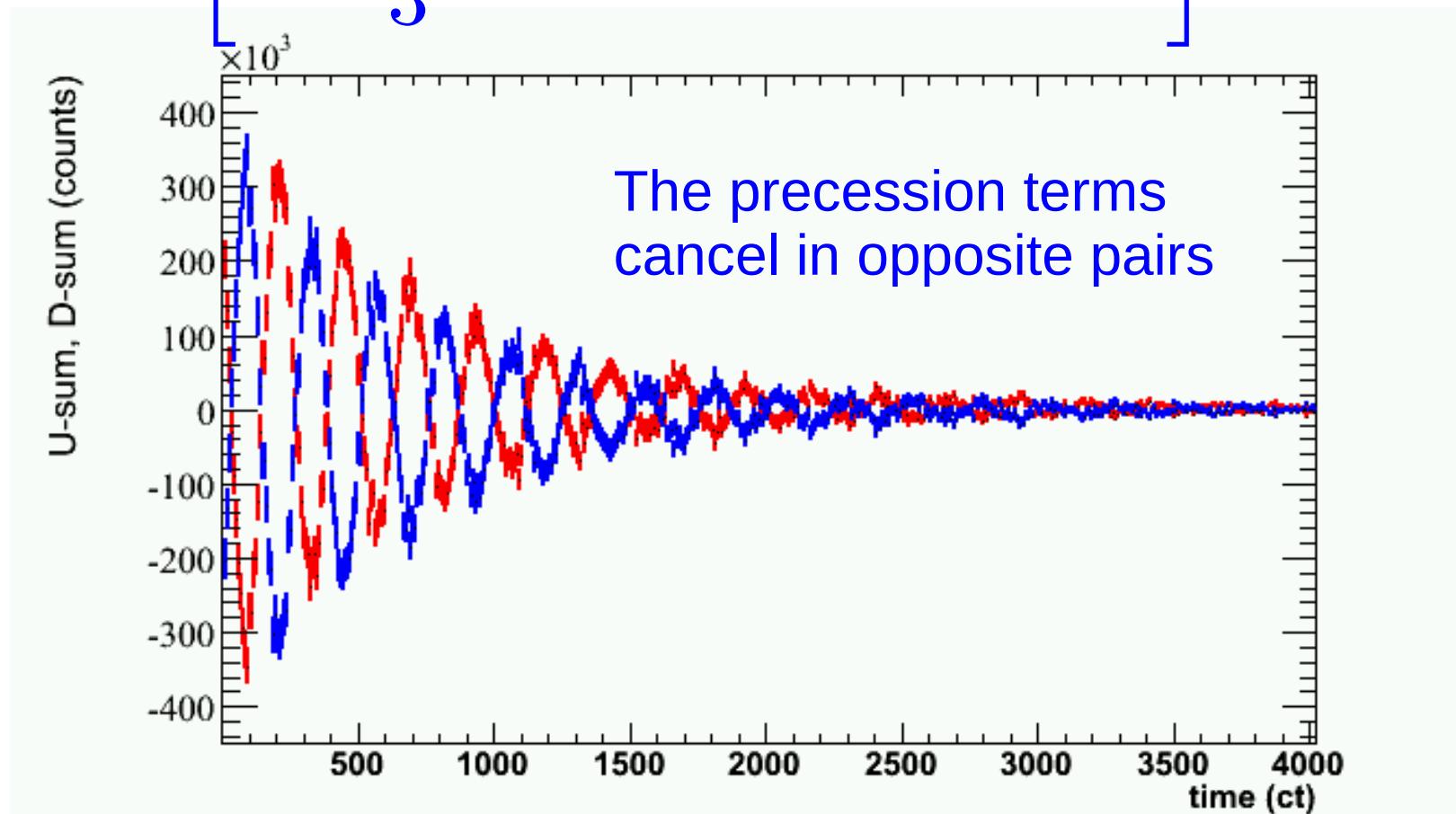
For 2007, we chose a muonium forming target with an externally applied field



- 90% Muonium formation
 - Test of free vs bound lifetime (theory says they're the same)
 - High magnetic moment gives high precession frequency (100x free muons)
- 10% “free muons”
 - We must fit for their precession!

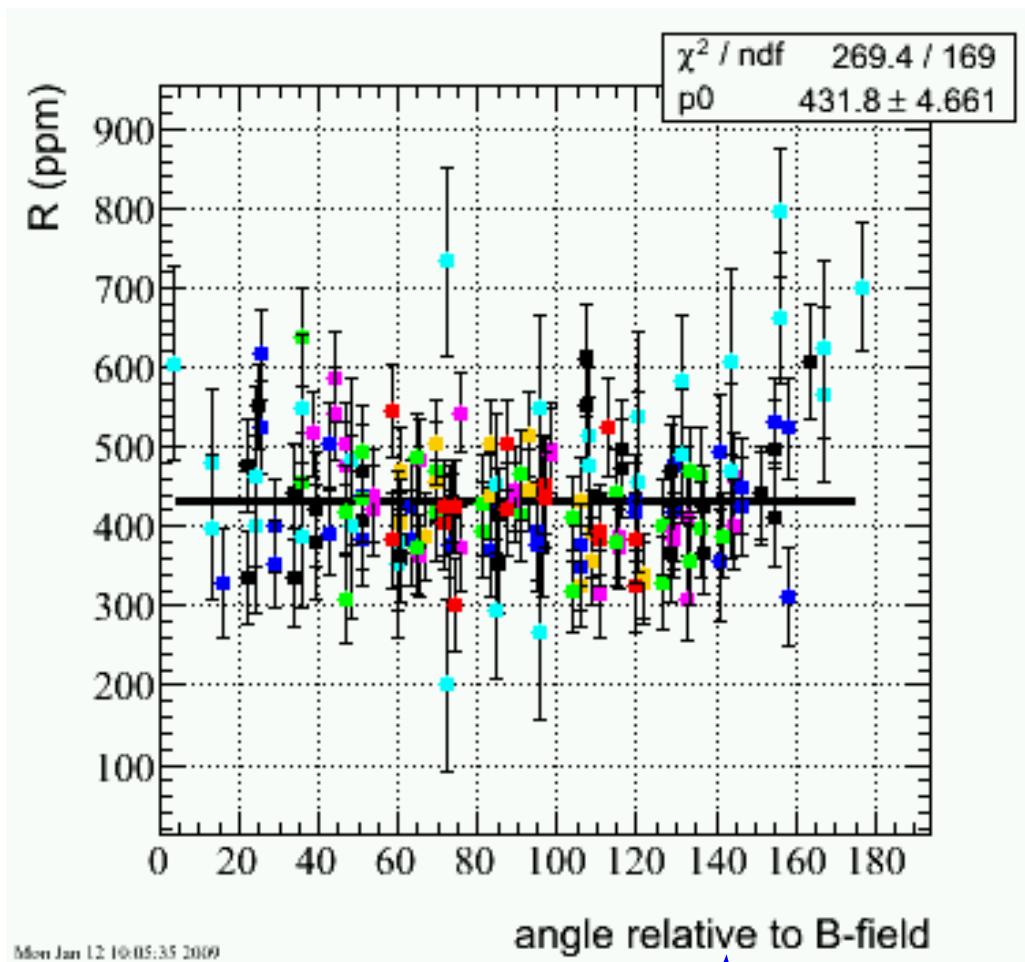
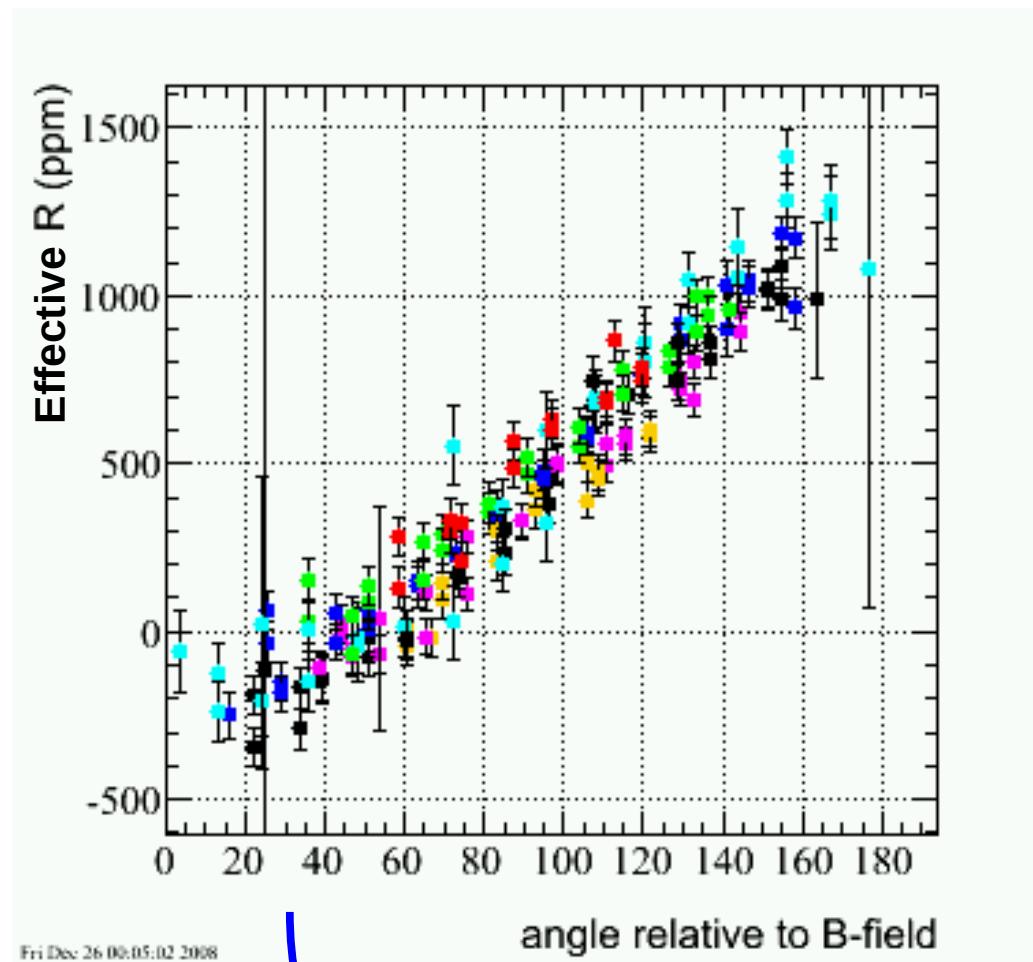
For 2007, we add the precession term directly to the individual detector's fit functions...

$$F(t) = N \left[1 + \frac{1}{3} P_2 \sin(\omega t + \phi) e^{-t/T_2} \right] e^{-t/\tau_{\text{eff}}} + B$$

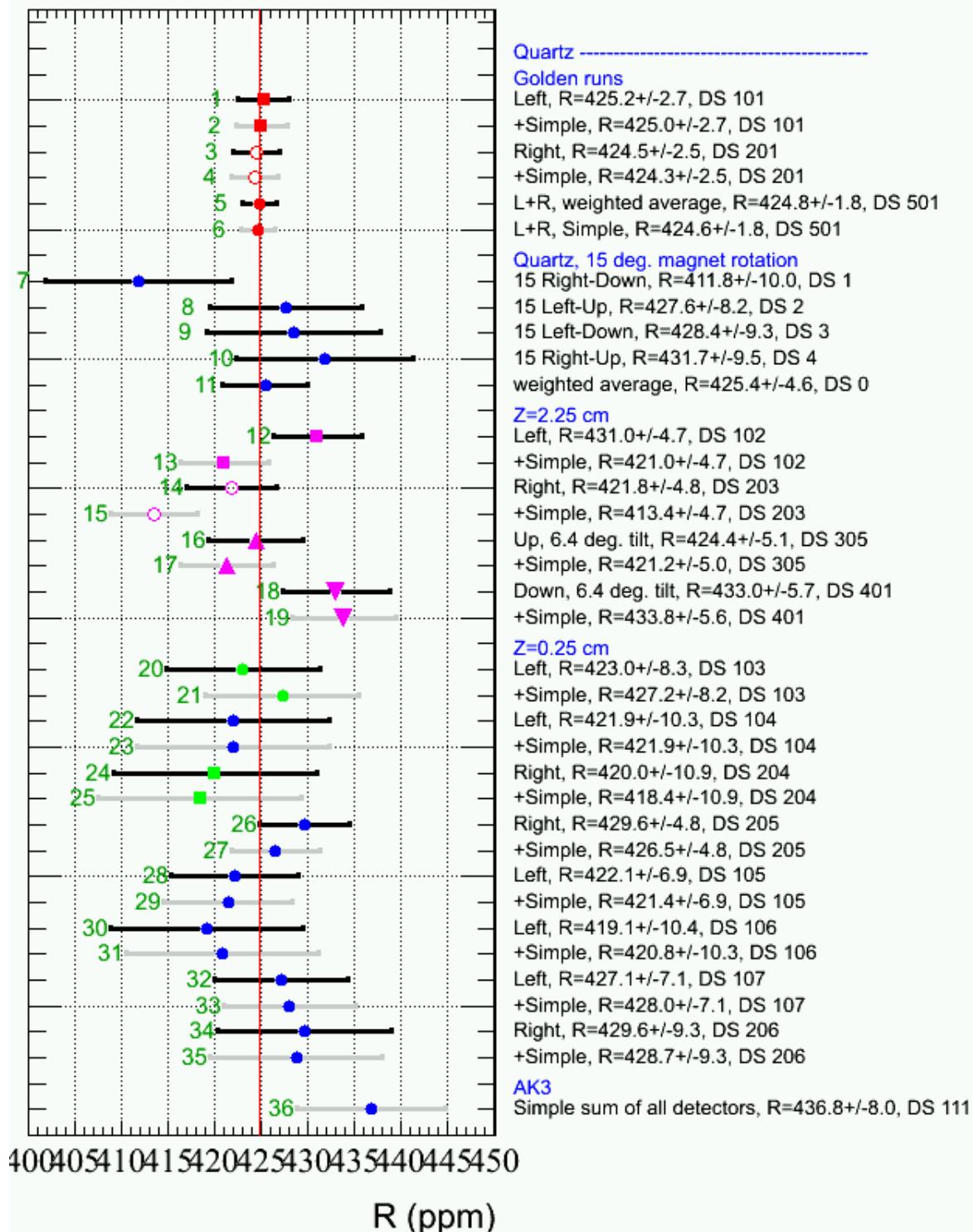


In 2006 (AK3) there is no measurable precession signal, so there's nothing to fit!

...while the effects of the residual longitudinal polarization is measured in the ensemble of all detector fits.

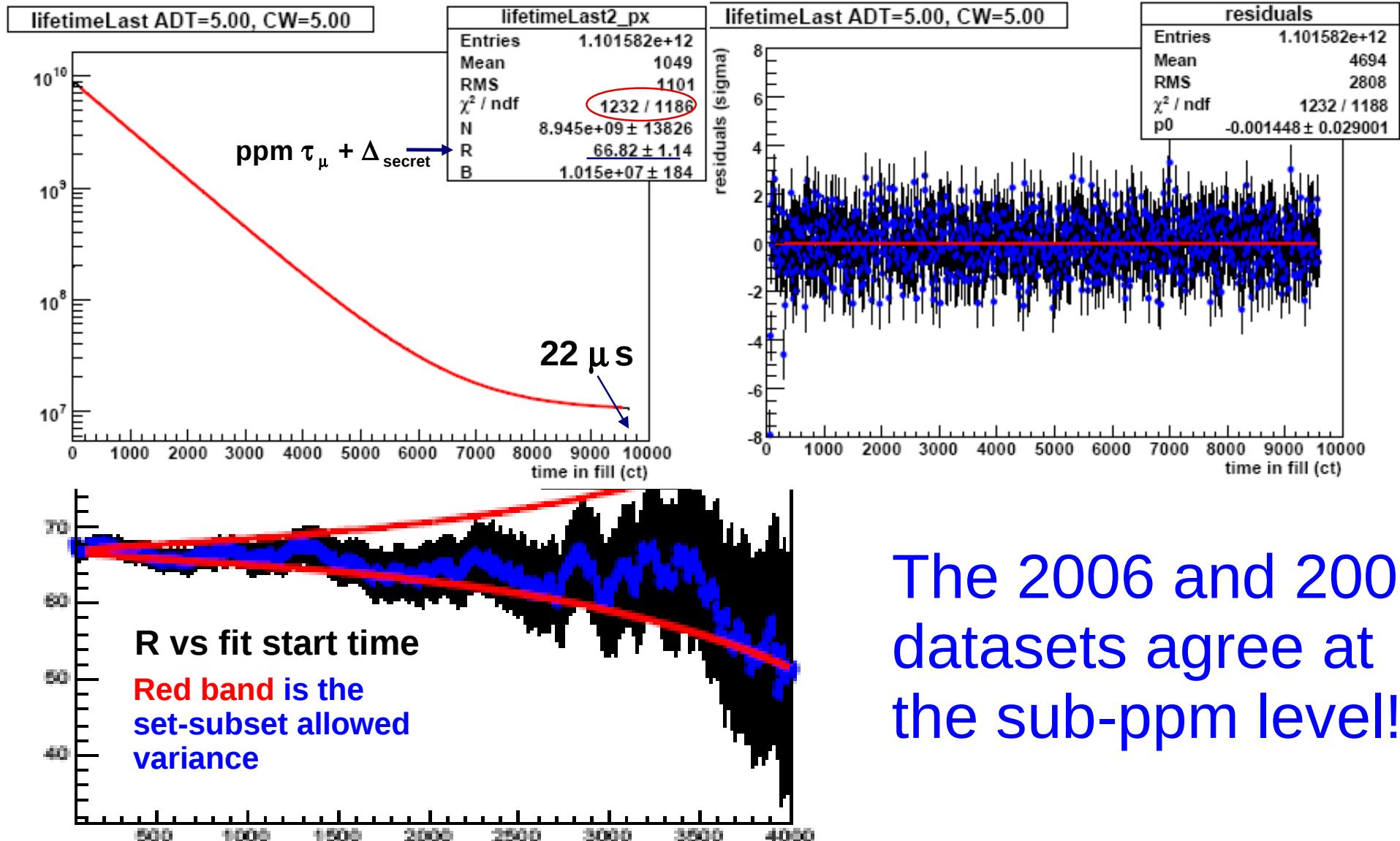


$$\Lambda_\mu = \Lambda_{\text{eff}} - \frac{1}{3} \Lambda_1 \vec{P}_1 \cdot \hat{\vec{r}}_D$$



Consistency of this procedure has been tested against many run conditions, including some truly extreme examples with very large residual longitudinal effects

Fits to all 2006/2007 pileup corrected data passes many consistency tests, including structureless residuals and fit start time scans



Our 2004 result was strongly statistics limited

Statistics: 1.8×10^{10} muons (9.6 ppm)

Source	Size (ppm)
Extinction stability	3.5
Dead time correction	2.0
TDC response	1.0
Gain stability	1.8
Errant muon stops	2.0
Duplicate words (+ 1 ppm shift)	1.0
Queuing loss	0.7
Multiple hit timing shifts	0.8
Total	5.2

$\tau_\mu = 2.197013(21)(11) \mu\text{s}$ (11 ppm)

$G_F = 1.166371(6) \times 10^{-5} \text{ GeV}^{-2}$ (5 ppm)

Our final results will improve on our 2004 results by an order of magnitude

Preliminary

Effect	2006	2007	Comment
Kicker extinction	0.2	0.07	Final
Errant muon stops	~0	~0	In progress; small
Gain stability vs time-in-fill	0.04	0.04	Based on MPV ¹ of data vs time
Gain stability vs time-after-pulse ²	<0.2	<0.2	MPVs in next fill & laser studies
Timing stability vs time-in-fill	0.014	0.014	Final; laser studies
Timing stability vs time-after-pulse	~0	~0	Final; laser studies
Electronic pedestal fluctuation ³	~0.2	~0.2	In progress; upper limit
Pileup correction	~0.3	~0.3	In progress; studies to be done
Residual polarization	~0	~0.2	Incomplete cancellation (quartz)
Total Systematic (DRAFT !)	~0.4	~0.4	Highly correlated for 2006/2007
Total Statistical	1.14	1.7	

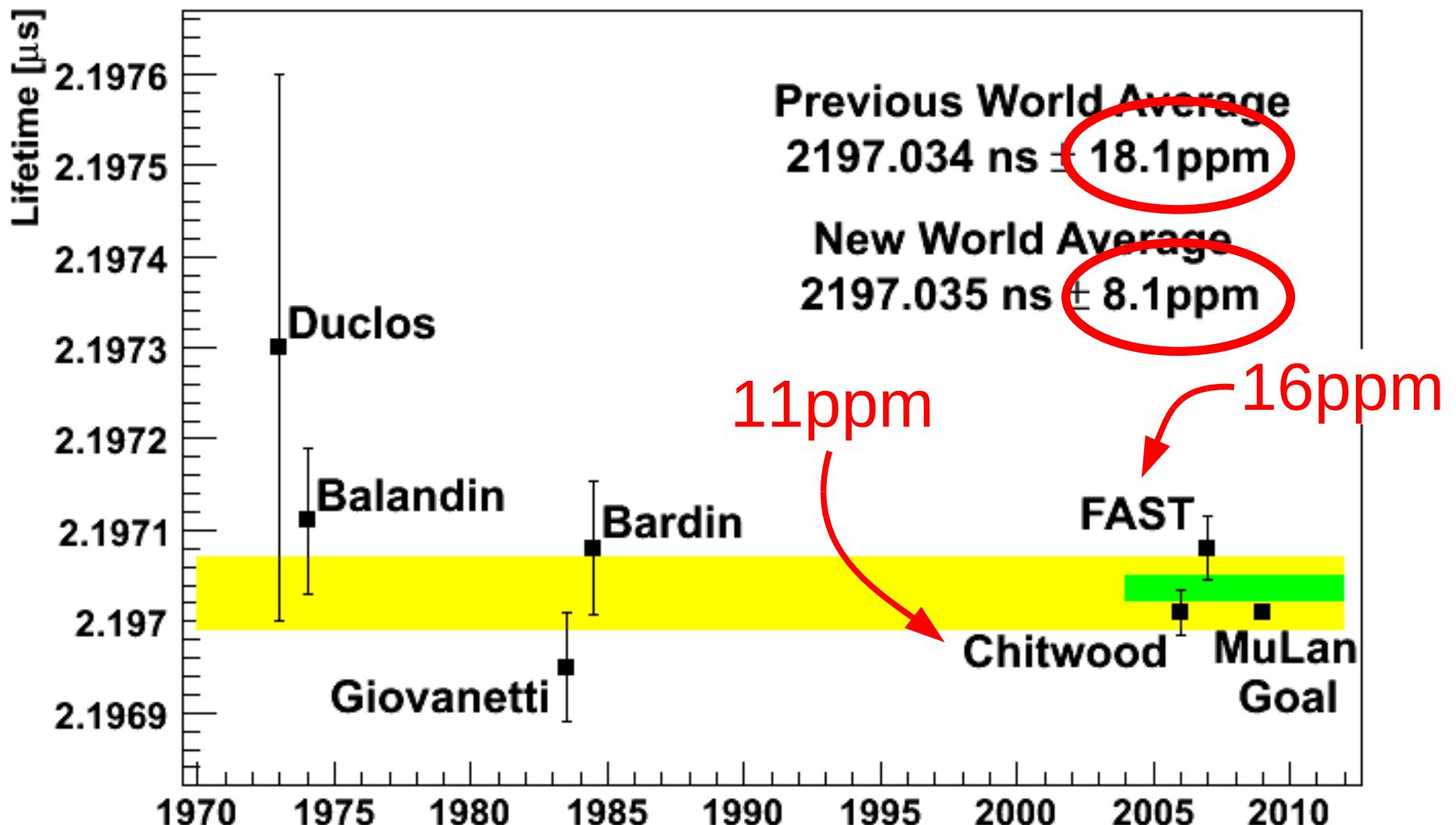
¹Most probable value of energy deposition

²Time-after-pulse is to the “next” pulse following a hit

³Coherent effect, measured in lab tests and easily inserted into fit function

Combined (roughly): 0.95 ppm (statistical) & ~0.5 ppm (systematic)

The current world average lifetime is driven by two measurements, but will (very, very shortly!) be eclipsed by our final result



Chitwood, et.al. Phys. Rev. Lett. 99, 032001 (2007)

FAST Collaboration, Phys.Lett.B663:172-180,2008

Toward 1ppm ...

