

# P7 Compressed Sensing Image Reconstruction for CASA

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abstract plcaceholder

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# 1 Large Scale Image Reconstruction

Measurement Equation

## 1.1 Measurements

$$\underset{x}{\text{minimize}} \left\| v - F^{-1}x \right\|_2^2 + \lambda \|\Psi x\|_1 \quad (1.1)$$

Image size 32k squared

Visibility size

## 1.2 Python imaging pipeline

## 2 Wide Field of View Imaging

So far the small Field of View inverse problem has been introduced where each antenna pair measures a Visibility of the sky brightness distribution. This leads to the small Field of View measurement equation (2.1). It is identical to the two dimensional Fourier Transform. In practice the Fast Fourier Transform (FFT) is used, since it scales with  $n \log(n)$  instead of  $n^2$  pixels.

$$V(u, v) = \iint x(l, m) e^{2\pi i(ux+vy)} dl dm \quad (2.1)$$

For wide Field of View imaging, two effects break the two dimensional Fourier Transform relationship: Non-coplanar Baselines and the celestial sphere which lead to the measurement equation (2.2). Note that for small Field of View  $1 - x^2 - y^2 \ll 1$ , and (2.2) reduces to the 2d measurement equation (2.1).

$$V(u, v, w) = \iint \frac{X(x, y)}{\sqrt{1 - x^2 - y^2}} e^{2\pi i(ux+vy+w\sqrt{1-x^2-y^2})} dx dy \quad (2.2)$$

Non-coplanar Baselines lead to a third component  $w$  for each Visibility. Figure 1 shows the the  $u$   $v$  and  $w$  coordinate system.  $w$  is essentially the pointing direction of the instrument. The UV-Plane is the projection of the antennas on a plane perpendicular to the pointing direction. Which point in the UV-Plane get sampled and what  $w$  component it has depends on the pointing direction. If the instrument points straight up, the UV-Plane is a tangent to earth's surface, and the  $w$  term compensates for earth's surface curvature. If however the instrument points at the horizon, the projected UV-Plane gets squashed and  $w$  compensates for antennas which lie far behind the UV-Plane. In essence,  $w$  is a phase delay that corrects antenna positions in three dimensions. The wide Field of View measurement equation (2.2) would account for the  $w$  phase delay, but it breaks the the two dimensional Fourier relationship and the FFT cannot be used. The W-Projection [?] algorithm approximates the effect of the  $w$  term restores the two dimensional Fourier relationship.

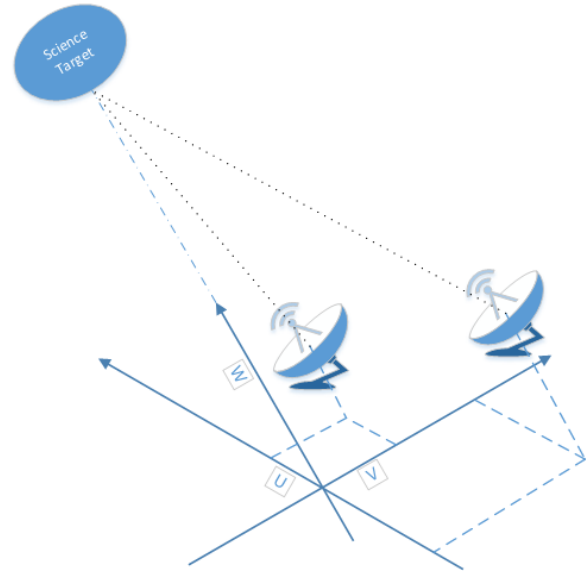


Figure 1: U V and W coordinate space

### 2.1 Interpretation of the W-Term and Spherical Harmonics

A-Projection [? ]

### 2.2 Separation of the W-term

WSClean

### 2.3 Calibration

### 3 Benchmark

## **4 results**

## 5 Conclusion



## References

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## 6 Ehrlichkeitserklärung

Hiermit erkläre ich, dass ich die vorliegende schriftliche Arbeit selbstständig und nur unter Zuhilfenahme der in den Verzeichnissen oder in den Anmerkungen genannten Quellen angefertigt habe. Ich versichere zudem, diese Arbeit nicht bereits anderweitig als Leistungsnachweis verwendet zu haben. Eine Überprüfung der Arbeit auf Plagiate unter Einsatz entsprechender Software darf vorgenommen werden.

Windisch, December 2, 2018

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