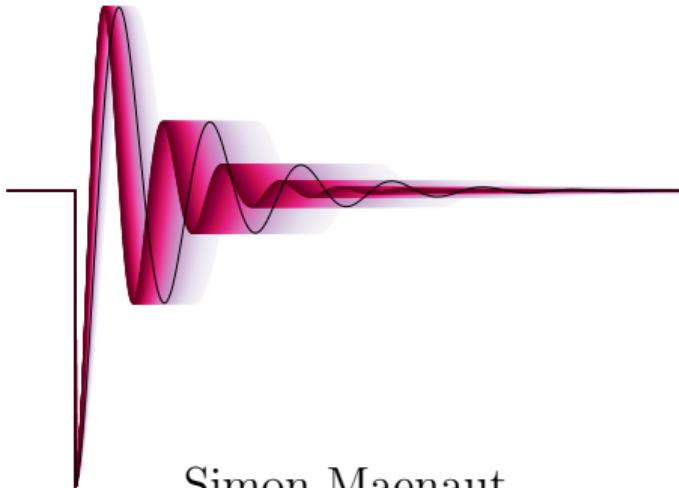


# pyRing for LISA

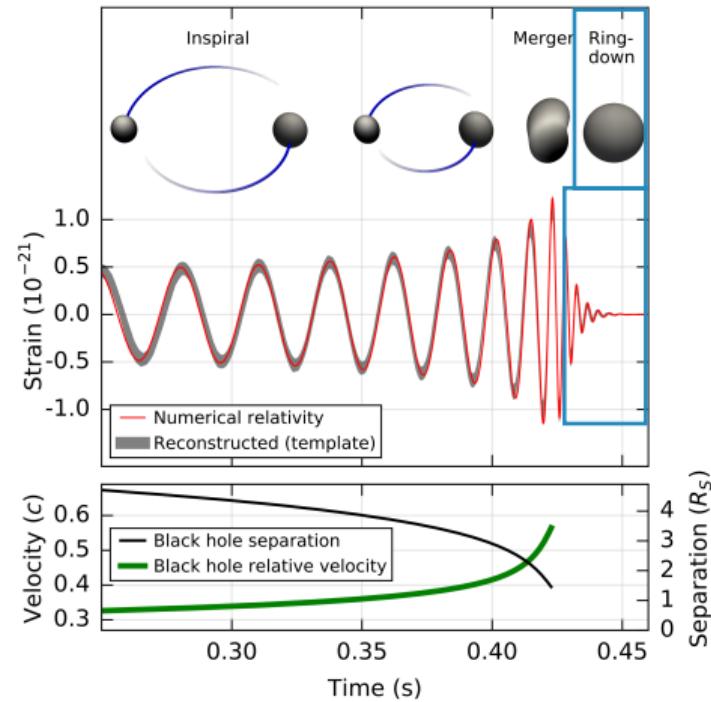


Simon Maenaut

LISA Fundamental Physics Working Group

# BINARY BLACK HOLES COALESCENCES

- Three main phases of the coalescence:
  - **Inspiral:** quasi-adiabatic evolution (PN theory + resummation, EOB)
  - **Plunge-merger:** highly dynamical (NR)
  - **Ringdown:** remnant approaches equilibrium.  
Damped **normal-modes** emission  
**(perturbation theory + NR)**



# SCIENCE GOALS

- **Testing** the emission:
  - Remnant compact object **nature**;  
*Are we really observing black holes?*
  - General Relativity predictions for **spectral emission**;  
*Is General Relativity a correct description of gravity at high curvatures?*
  - Black Hole **Uniqueness Theorems**;  
*Do non-extremal black holes have additional hairs?*
  - Possible **quantum horizon** effects and classical **BH thermodynamics**.  
*Is our classical description of black holes valid?*

# MISSING PHYSICS?

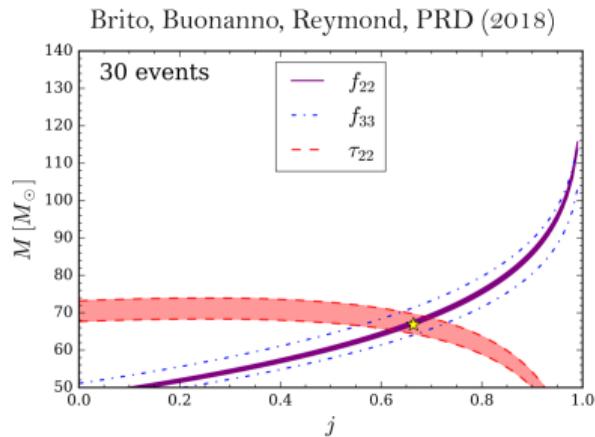
- Why merger-ringdown GW observations compared to **other experiments?**
  - **Largest curvature** regime experimentally accessible;
  - Highly **dynamical** regime;
  - Ringdown searches for new physics in LVK band expected to be **clean**
  - **Gauge** the contribution of **missing physics** for precision tests:
  - In LISA?

# RINGDOWN: QUASI-NORMAL MODES SOLUTIONS

- In terms of gravitational wave **multipoles**:

$$h_+ - i h_\times = -\frac{M}{r} \sum_{l,m,n} \mathcal{A}_{lmn} S(\theta, \phi) e^{i\omega_{lmn} t} e^{-t/\tau_{lmn}}$$

- Perturbation theory predicts spectrum.  
In Kerr,  $\{\omega_{lmn}(M, a), \tau_{lmn}(M, a)\}$  (“*final state conjecture*”)
- Measure  $\{\omega_1 = \omega_{220}, \tau_1 = \tau_{220}\} \rightarrow \{M, a\}$   
Measure  $\{\omega_2 = \omega_{330}\}$  (now predicted)  $\implies$  test **General Relativity**
- “Universal” prediction: “easy” to add **beyond-GR** effects



## WARNING AND GOAL

- Spectroscopy from binary mergers, **systematics**-plagued measurements:
  - “**Ringdown start time**” uncertainty (ill-defined)
  - **Dynamical evolution** of the system (e.g. mass&spin change with time)
  - Affect **measurement precision**, but also **physical interpretation**
  - Difference between:
    - **Fitting** post-peak **waveform** with bunch of damped-sinusoids
    - **Extract physical** vibrational black hole **spectrum**

# WHAT IS PYRING?

- Python package for **Black Hole** ringdown analysis, model comparison and parameter estimation.
  - Can handle interferometric, simulated or numerical relativity data.
  - Natively incorporates several analytical ringdown templates.
  - Supports parametrised tests of **General Relativity** and a variety of quasi-normal modes spectrum predictions for emissions alternative to the Kerr solution.
- Internally reviewed for scientific usage by the **LIGO-Virgo-Kagra** collaboration.
- Routinely used to produce catalogs of BH ringdown properties and related tests of GR by the LVK collaboration.

# pyRing TECHNICAL DETAILS

- Fully time domain formulation, both likelihood and waveform models
- Relies on gwpy for data access and LAL for all the antenna patterns
- Core computational operations in cython, everything else in python

Code and documentation

- LVK repository: [git.ligo.org/lscsoft/pyring](https://git.ligo.org/lscsoft/pyring)
- Full documentation: [lscsoft.docs.ligo.org/pyring](https://lscsoft.docs.ligo.org/pyring)
- Test of models vs NR: [github.com/GCArullo/bayRing](https://github.com/GCArullo/bayRing)

Carullo+ 1902.07527, Isi+ 1905.00869, LIGO & Virgo 2010.14529

## SCIENCE OUTPUT

- Additional studies that made use of this software:
  - Impact of noise on overtone detection in GW150914: [arXiv:2201.00822](#);
  - GWTC-3 Testing GR LVK catalog: [arXiv:2112.06861](#);
  - GW190521 waveforms consistency and astrophysical implications: [arXiv:2112.06856](#);
  - Models and constraints of ringdown in the presence of U(1) black hole charges: [arXiv:2109.13961](#);
  - Constraints on braneworld gravity: [arXiv:2106.05558](#);
  - Verification of the Bekenstein-Hod bound: [arXiv:2103.06167](#);
  - Constraints on alternative theories of gravity using the ParSpec formalism: [arXiv:2102.05939](#);
  - Investigations and observational constraints on the area quantisation hypothesis: [arXiv:2011.03816](#);
  - GWTC-2 Testing GR LVK catalog: [arXiv:2010.14529](#);
  - GW190521 discovery: [arXiv:2009.01075](#) and physics implications: [arXiv:2009.01190](#);
  - Spectroscopy of Rotating Black Holes Pierced by Cosmic Strings: [arXiv:2002.01695](#).
  - Probing the Purely Ingoing Nature of the Black-hole Event Horizon [arXiv:1912.07058](#);

## pyRing RINGDOWN MODELS

- Simple superposition of damped sinusoids
- Models with predicted complex frequency for Kerr
- Models with Kerr spectrum and amplitude predictions
- Tidal Effective One Body post-merger (TEOBPM) model
- Many other models...

See [pyRing documentation](#) for more details.

# A ZOO OF POSSIBILITIES

- All models mentioned before, available
- Together with many more models/options:

- **Quadratic modes**

London+, 1404.3197, Cheung+, 2208.07374, Mitman+, 2208.07380,  
Baibhav+, 2302.03050, Cheung+, 2310.04489

- Tails

Carullo-De Amicis, 2310.12968

- Black hole charges

Carullo+, 2109.13961, Gu+, 2310.10447

- **Parameterised GR deviations** (agnostic/Parspec)

LVK, 2112.06861  
Carullo, 2102.05939

- Beyond-GR corrections (EsGB, dCS, EFTs, ...)

Silva+, 2205.05132

- Braneworld, area-quantisation, ...

Mishra+, 2311.03556, Laghi+, 2011.03816

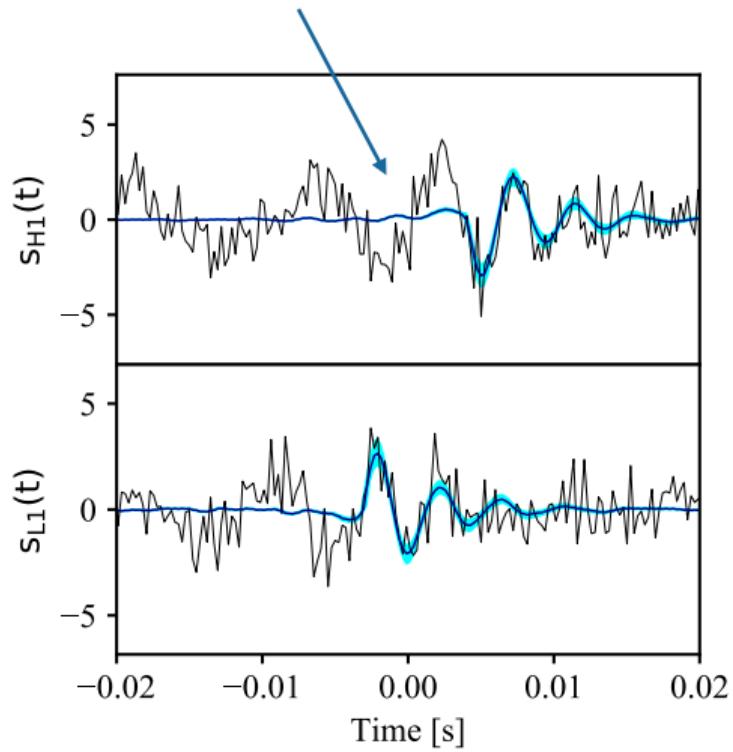
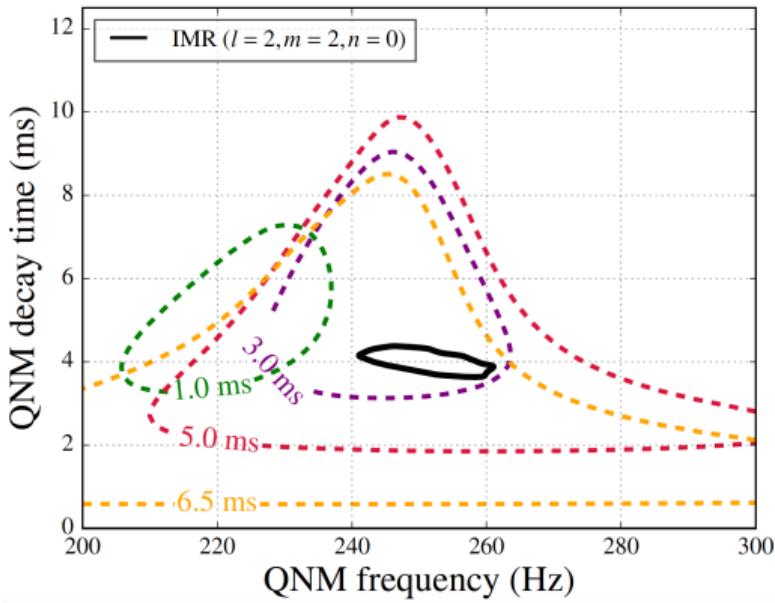
- Older models (Kamaretsos+)

Kamaretsos,+ 1406.3201

- ...

# GW150914: THE DAY WE SAW A BLACK HOLE RINGING

- Frequency Domain: ringdown abrupt start contaminates inference when FFT-ing (Gibbs phenomena)

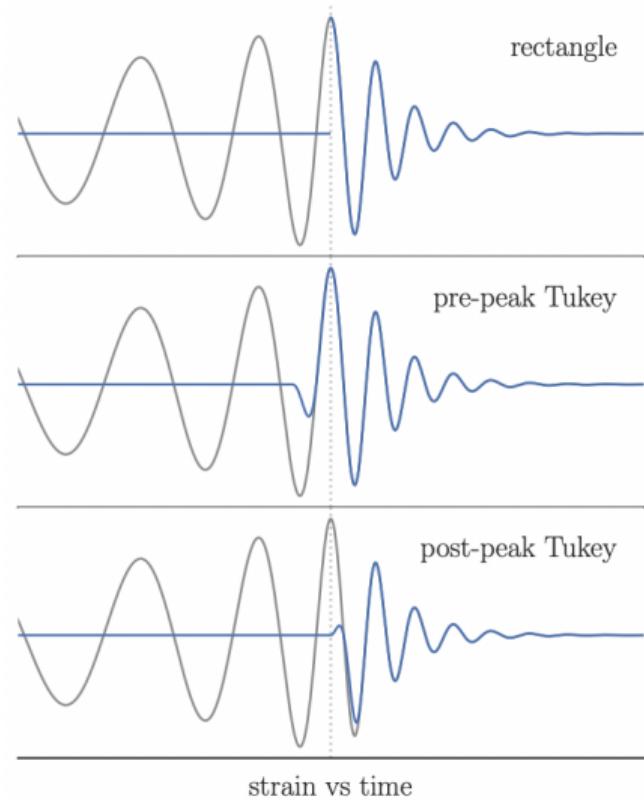


# BLACK HOLE SPECTROSCOPY

- Cure: window

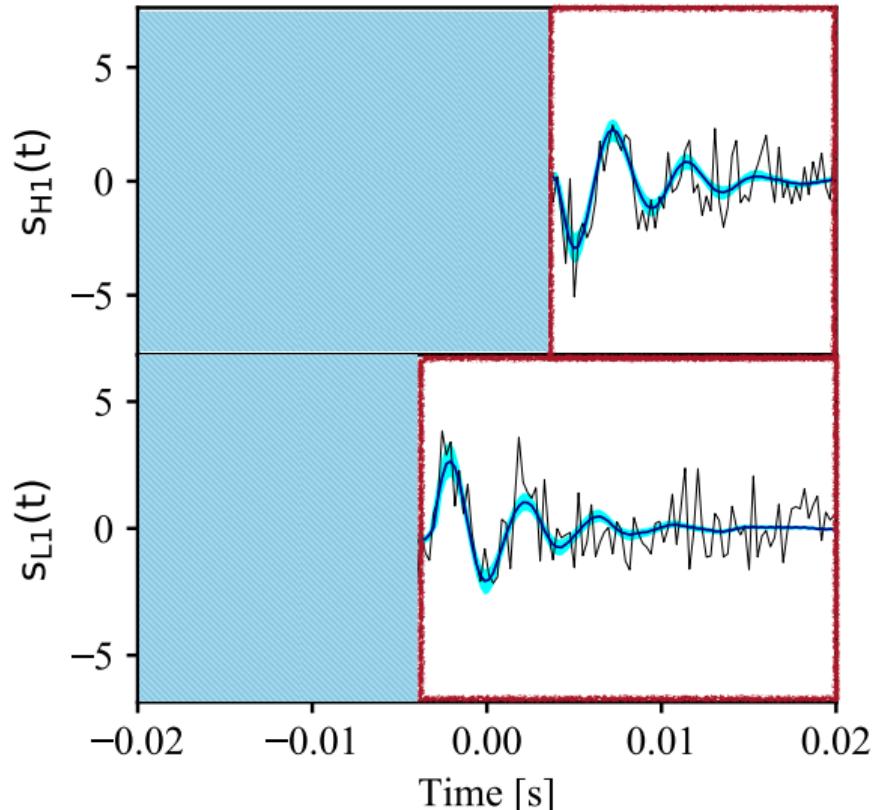
Carullo+, 1805.04760

- **Pro:** simple, FD analysis
- **Cons:** lose SNR, need optimise window parameters



# BLACK HOLE SPECTROSCOPY

- Time-Domain approach:
  - **Isolate** ringdown portion
  - **Avoid** spurious frequencies from abrupt  $t_{\text{start}}$ , stay in time-domain (no FFT)



White noise, no-merger: Del Pozzo, Nagar, 1606.03952

Coloured noise: Carullo, Del Pozzo, Veitch, 1902.07527

Full solution: Isi, Farr, 2107.05609

# TIME DOMAIN BLACK HOLE SPECTROSCOPY

- No longer **diagonal** likelihood

$$\log p(d | \theta, I) \propto - \sum_j \frac{|d(f_j) - h(f_j; \theta)|^2}{S(f_j)}$$



- Keep full correlation **matrix** (Toeplitz)

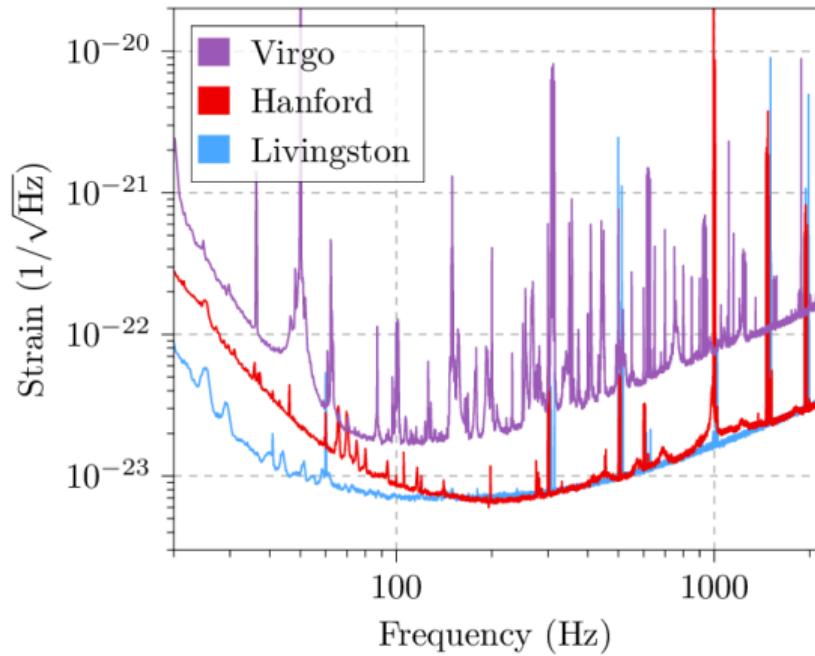
$$\ln P(d | s) = -\frac{1}{2} \sum_{i,j=0}^{N-1} (d_i - s_i) C_{ij}^{-1} (d_j - s_j)$$

$$C_{ij} = \rho(|i - j|)$$

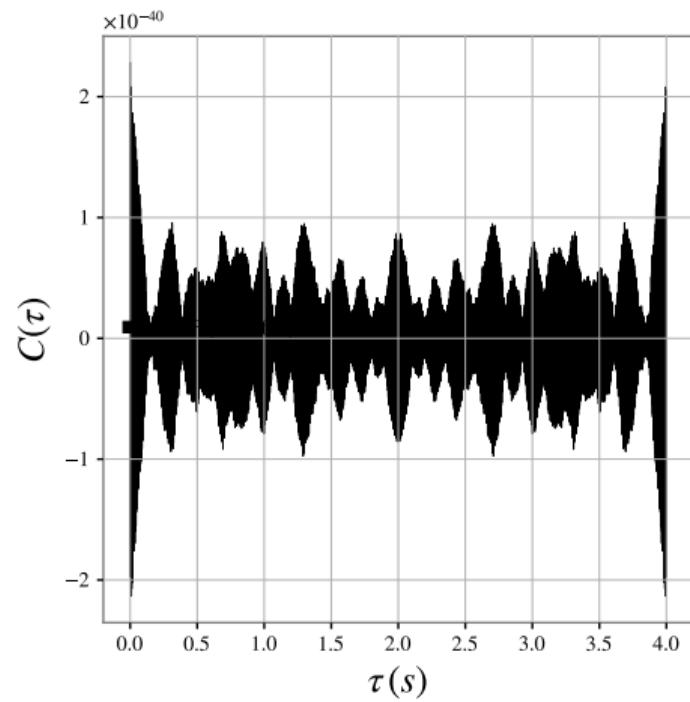
$$\hat{\rho}(k) = \frac{1}{N_\rho} \sum_{i=0}^{N_\rho-1} n_i n_{i+k}$$

# TIME DOMAIN BLACK HOLE SPECTROSCOPY

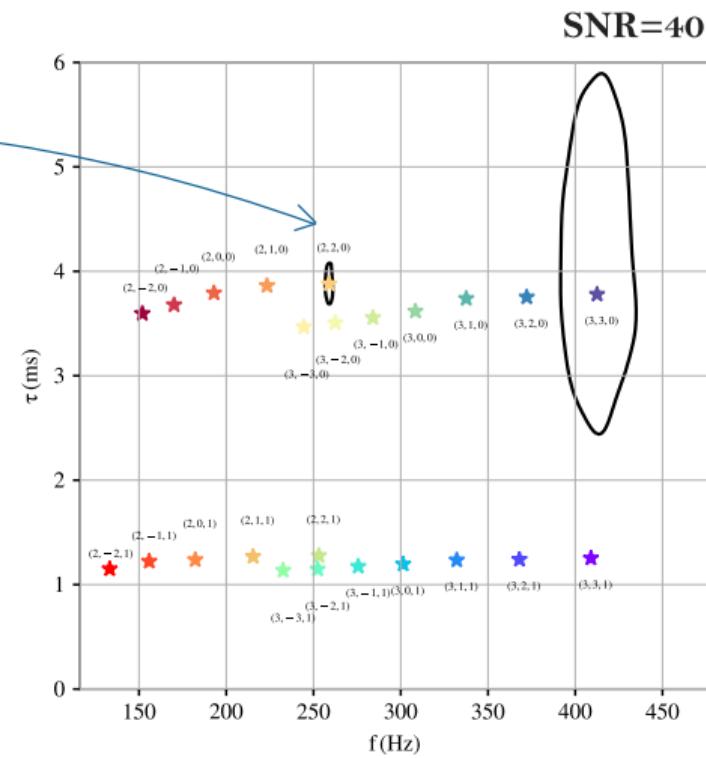
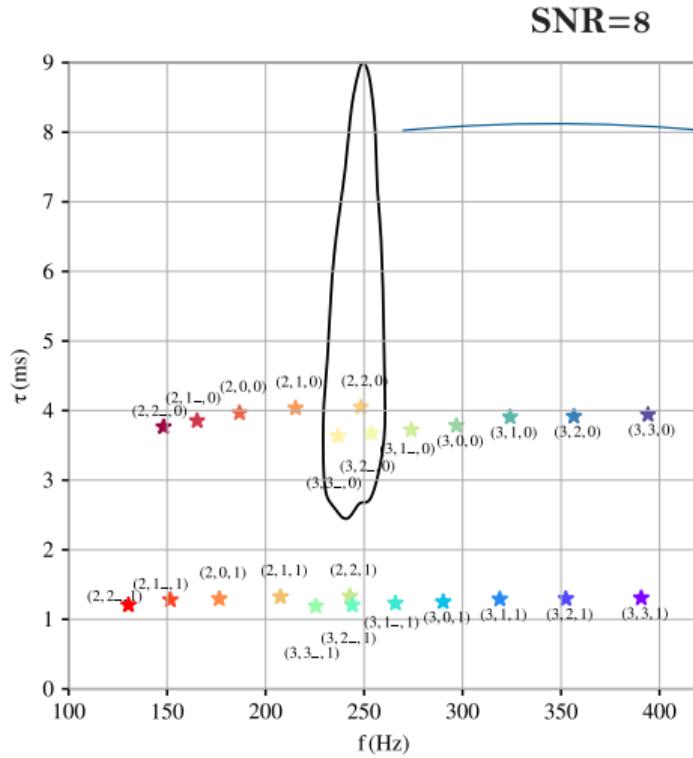
- The familiar:



- Becomes:



# FUTURE BLACK HOLE SPECTROSCOPY



## OPEN PROBLEMS

- Precession Finch-Moore, 2102.07794  
Hamilton-London-Hannam, 2301.06558
  - Eccentricity Carullo+, 2309.07228
  - Impact of memory/BMS frame Zertuche+, 2110.15922
  - High-precision multi-mode fits Baibhav+, 2302.03050  
Cheung+, 2310.04489  
Zertuche+, 2110.15922
  - Quadratic modes London+, 1404.3197  
Cheung+, 2208.07374  
Mitman+, 2208.07380  
Baibhav+, 2302.03050  
Cheung+, 2310.04489

# LISA specifics

## Work in progress

1. Completely different detector response and reference frame
  2. Indirect measurement of the strain (TDI variables X, Y, Z or A, E, T)
  3. Calculation of LISA specific auto-covariance and cross-covariance functions
- 
- LISA Wiki: [Ringdown Tests Of The No-hair Theorem With LISA](#)
  - Gitlab: [gitlab.in2p3.fr/simon.maenaut/pyring](https://gitlab.in2p3.fr/simon.maenaut/pyring)
  - Zulip: [fpwg-projects.zulipchat.com](https://fpwg-projects.zulipchat.com)

# LISA challenges

1. LISA response (adiabatic approximation) from LISA GW Response ✓
2. Calculate TDI variables quickly from strain (linear interpolation) ✓
3. Compute ACF and CCF from LISA PSD and CSD for TDI variables
4. Implement multi-channel likelihood calculation for a single detector
5. Validate parameter-estimation, especially for high SNR signals

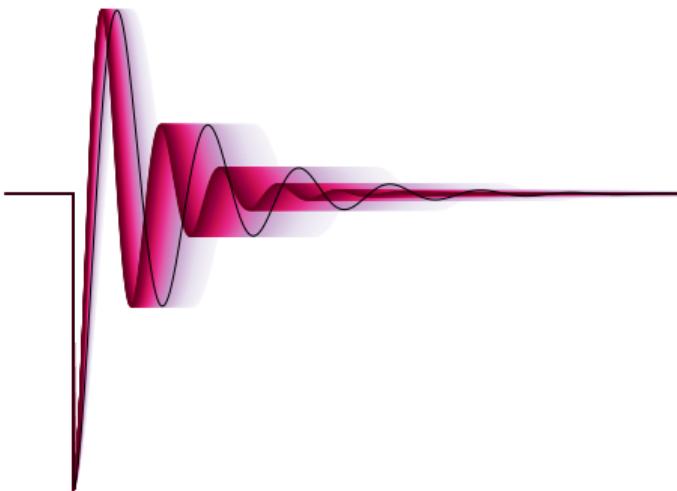
$$C(p, q) = R\left(\frac{p - q}{f_s}\right) = \int_{-\frac{f_s}{2}}^{+\frac{f_s}{2}} S_n(f) e^{2j\pi f \frac{p-q}{f_s}} df$$

# Future outlook

Including ongoing LISA FPWG Ringdown projects

- Apply pyRing on latest LISA data challenge (high SNR) SMBHBs mergers
- Validation with injections of waveforms in LISA noise simulations
- Study the sensitivity to potential beyond GR ringdown effects
- Open questions on ringdown (see earlier) not specific to LISA
- Various ideas to study this dynamic strong gravity regime...

Thank you for your attention



# ALTERNATIVE FORMALISMS

- TD not the only option
- Stay in FD, filter-out pre-merger (“in-painting”)  
Equivalent to TD, more expensive (efficient) for short (long) segments

Capano+, 2105.05238  
Isi, Farr, 2107.05609

- Stay in FD, model agnostically pre-merger (“marginalise”), e.g.  
superposition of wavelets

Finch-Moore, 2205.07809

# FUTURE BLACK HOLE SPECTROSCOPY

