CS 83: Computer Vision	Winter 2024
Quiz 6 — 02	/22/2024
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## **Credit Statement**

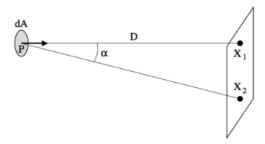
I discussed solution ideas with:

- 1. Ivy (Aiwei) Zhang
- 2. Angelic McPherson

However, all typed work is my own, with reference to class notes.

## Problem 1.

A small Lambertian source dA is centered at P and emits radiance L. The orientation of this patch is the same as that of a plane containing two points,  $X_1$  and  $X_2$ . The point  $X_1$  is the point on this plane that is closest to P, and the distance from P to  $X_1$  is D as shown.



(i) Calculate the solid angle subtended by dA at points  $X_1$  and  $X_2$ .

The solid angle subtended by an area dA at a point x that is a distance r from dA is given by

$$d\omega = \frac{dA\cos\theta}{r^2},$$

where  $\theta$  is the angle between the normal to dA and the line from  $\mathbf{x}$  to dA. In this case;

- 1. Since dA has the same orientation as the plane containing  $X_1$  and  $X_2$ , and we are given that  $X_1$  is the point on this plane that is closest to P, then the normal at dA passes through P. Therefore,  $\theta = 0$  and  $\cos \theta = 1$ . Therefore,  $d\omega_{X_1} = \frac{dA}{D^2}$ .
- **2.** As established in part **1.**, the normal at dA passes through P. Therefore, the angle between the normal to dA and the line from  $X_2$  to dA is  $\alpha$  as shown in the image. Let  $D_2 = ||X_2 P||$ ; then:

$$\cos \alpha = \frac{D}{D_2}$$

$$\implies D_2 = \frac{D}{\cos \alpha}.$$

Therefore;

$$\mathrm{d}\omega_{X_2} = \frac{\mathrm{d}A\cos\alpha}{D_2^2}$$
$$= \frac{\mathrm{d}A\cos\alpha}{D^2/\cos^2\alpha}$$
$$= \frac{\mathrm{d}A\cos^3\alpha}{D^2}.$$

(ii) Calculate the irradiance E incident on the plane at points  $X_1$  and  $X_2$ , and calculate the ratio  $E(X_1)/E(X_2)$ .

The irradiance E incident on a point  ${\bf x}$  is given by

- **1.** The irradiance  $E(X_1)$  incident on  $X_1$  is given by
- **2.** The irradiance  $E(X_2)$  incident on  $X_2$  is given by
- 3. The ratio  $E(X_1)/E(X_2)$  is given by

## Problem 2.

As we discussed in class, the Lambertian and the specular BRDF are the two most commonly used reflectance models in physics-based vision.

(i) For Lambertian surfaces, the BRDF is a constant function of the input and the output directions. For such a material, we often describe the reflectance in terms of its *albedo*, which is given the symbol  $\rho$ . For a Lambertian surface, the BRDF and the albedo are related by  $f_r(\hat{\mathbf{v}}_i, \hat{\mathbf{v}}_r) = \rho/\pi$ . Using conservation of energy, prove that  $0 \le \rho \le 1$ .

First, note that the albedo is a measure of the fraction of incident light that is reflected by the surface. Therefore, it cannot be negatie, so  $0 \le \rho$ . We shall show that  $\rho \le 1$  by using the energy conservation principle in reflectance which stipulates that the energy reflected by a surface cannot exceed the energy incident on the surface. In other words, the integral of the BRDF over the hemisphere must be less than or equal to 1:

$$\int_{H^2} f_r(\hat{\mathbf{v}}_i, \hat{\mathbf{v}}_r) \cos \theta \, d\vec{\omega}_i \le 1.$$

Substituting in the known values gives:

$$\int_{H^2} \frac{\rho}{\pi} \cos \theta \, d\vec{\omega}_i \le 1$$

$$\frac{\rho}{\pi} \int_{H^2} \frac{dA \cos \theta}{r^2} \le 1 \qquad \left( \text{since } d\vec{\omega}_i = \frac{dA}{r^2} \right)$$

$$\frac{\rho}{\pi} \int_{H^2} \frac{r^2 \cos \theta \sin \theta \, d\theta \, d\varphi}{r^2} \le 1 \qquad \left( \text{since } dA = r^2 \sin \theta \, d\theta \, d\varphi \right)$$

$$\frac{\rho}{\pi} \int_{H^2} \cos \theta \sin \theta \, d\theta \, d\varphi \le 1$$

$$\frac{\rho}{\pi} \int_{0}^{2\pi} \int_{0}^{\pi/2} \cos \theta \sin \theta \, d\theta \, d\varphi \le 1$$

$$\frac{\rho}{\pi} \int_{0}^{2\pi} \left[ \frac{\sin^2 \theta}{2} \right]_{0}^{\pi/2} \, d\varphi \le 1 \qquad \left( \text{since } \int \cos \theta \sin \theta \, d\theta = \frac{\sin^2 \theta}{2} \right)$$

$$\frac{\rho}{2\pi} \int_{0}^{2\pi} d\varphi \le 1$$

$$\frac{\rho}{2\pi} \cdot [\varphi]_{0}^{2\pi} \le 1 \qquad \left( \text{since } \int 1 \, d\varphi = \varphi \right)$$

$$\frac{\rho}{2\pi} \cdot 2\pi \le 1$$

$$\rho \le 1.$$

(ii) A specular surface perfectly reflects *radiance* in the *mirror direction*. Concretely, consider a (non-absorbing) specular surface patch with normal  $\hat{\mathbf{n}}$ . For any incident direction  $\hat{\mathbf{v}}_i$ , the mirror direction equals  $\hat{\mathbf{v}}_s = 2(\hat{\mathbf{n}}^T\hat{\mathbf{v}}_i)\hat{\mathbf{n}} - \hat{\mathbf{v}}_i$ , and  $L(\hat{\mathbf{v}}_s) = L(\hat{\mathbf{v}}_i)$ . Given this property, derive an expression for the specular BRDF.

The BRDF of an ideal specular reflection is a Dirac delta function as follows:

$$f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) = F_r(\vec{\omega}_i) \frac{\delta(\vec{\omega}_i - R(\vec{\omega}_r, \hat{\mathbf{n}}))}{\cos \theta_i}$$

$$= F_r(\vec{\omega}_i) \frac{\delta(\vec{\omega}_i - R(\vec{\omega}_r, \hat{\mathbf{n}}))}{\cos \theta_i}$$

$$= \frac{\delta(\hat{\mathbf{v}}_i - \hat{\mathbf{v}}_s)}{\hat{\mathbf{v}}_i \cdot \hat{\mathbf{n}}}$$

$$= \frac{\delta(\hat{\mathbf{v}}_i - 2(\hat{\mathbf{n}}^{\mathsf{T}} \hat{\mathbf{v}}_i) \hat{\mathbf{n}} + \hat{\mathbf{v}}_i)}{\hat{\mathbf{v}}_i \cdot \hat{\mathbf{n}}}$$

$$= \frac{\delta(2\hat{\mathbf{v}}_i - 2(\hat{\mathbf{n}}^{\mathsf{T}} \hat{\mathbf{v}}_i) \hat{\mathbf{n}})}{\hat{\mathbf{v}}_i \cdot \hat{\mathbf{n}}}$$