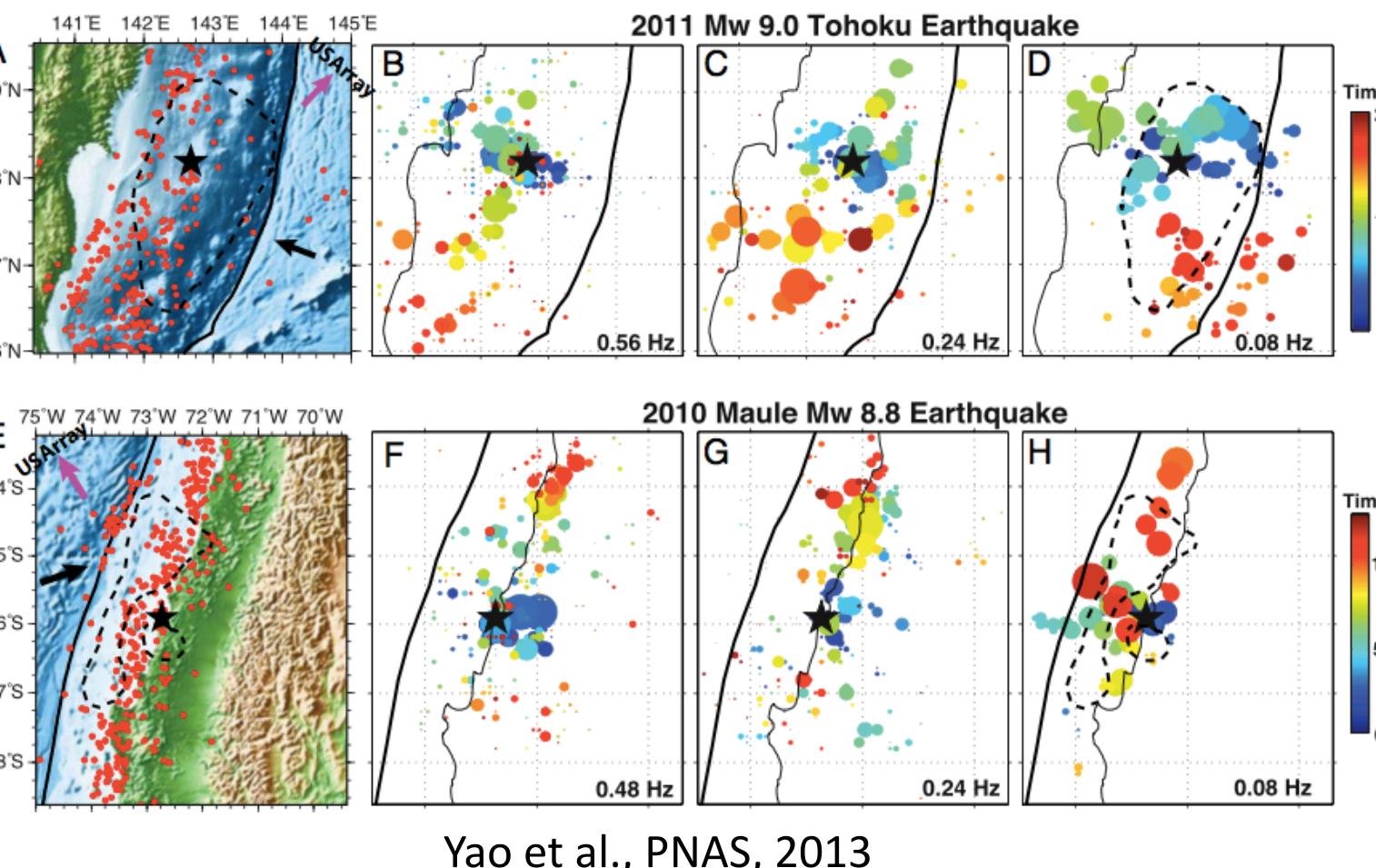


# Improving Distance Metrics in Ground Motion Prediction Equations Based on Seismic Array Back-Projection

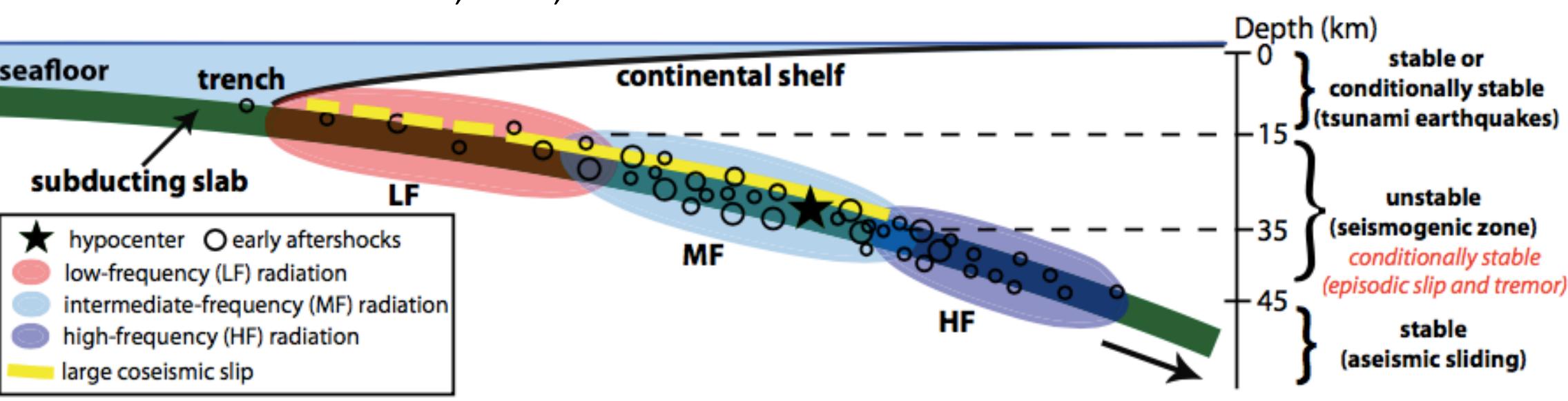
Lingsen Meng<sup>1</sup> ([meng@ess.ucla.edu](mailto:meng@ess.ucla.edu)) and Tian Feng<sup>1</sup>

1. Earth, Planetary and Space Sciences, University of California, Los Angeles, CA 90025, USA.

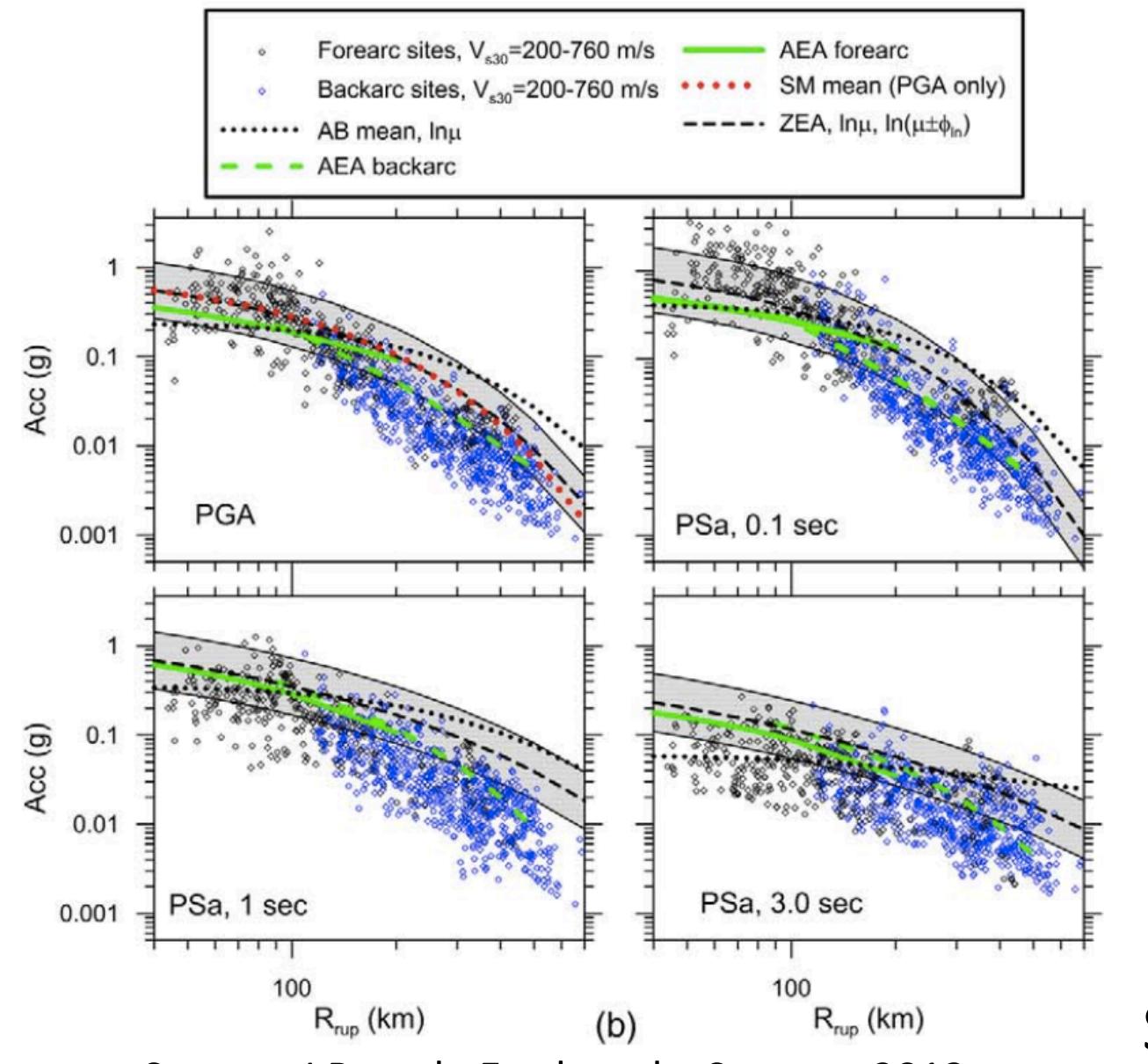
## Frequency-dependent and Depth-varying Seismic Radiation



Recent advances in earthquake source imaging of major subduction earthquakes highlight the frequency-dependent and depth-varying seismic radiations at the plate interfaces. Low-frequency energy mainly emanated from the shallower portion of the megathrusts while dominant high-frequency energy often radiates from the deeper portion of the megathrust (Meng et al., 2011; Lay et al., 2012; Yao et al., 2013).

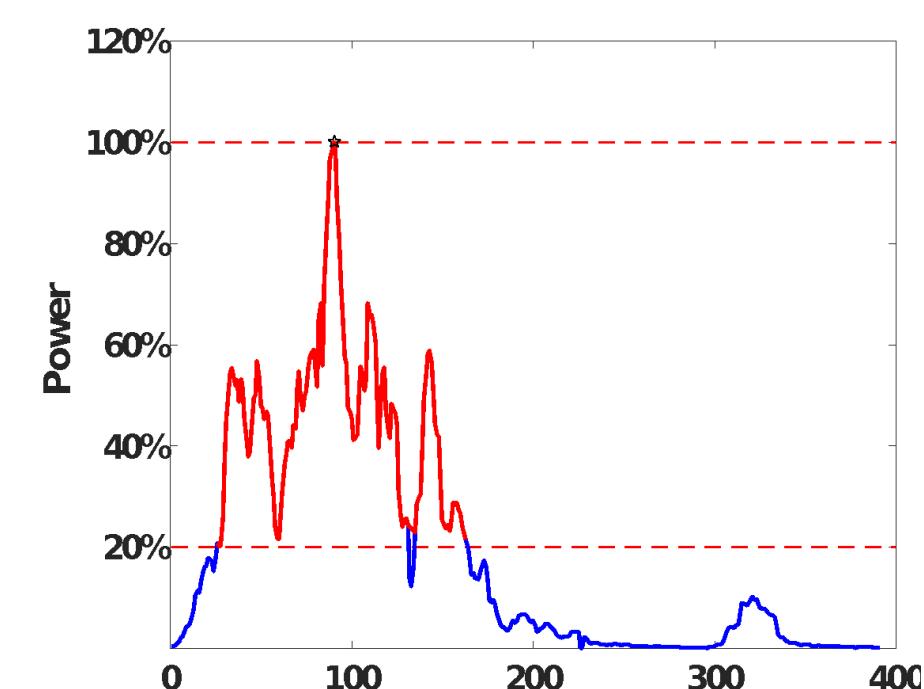


## Distance Metrics of GMPEs

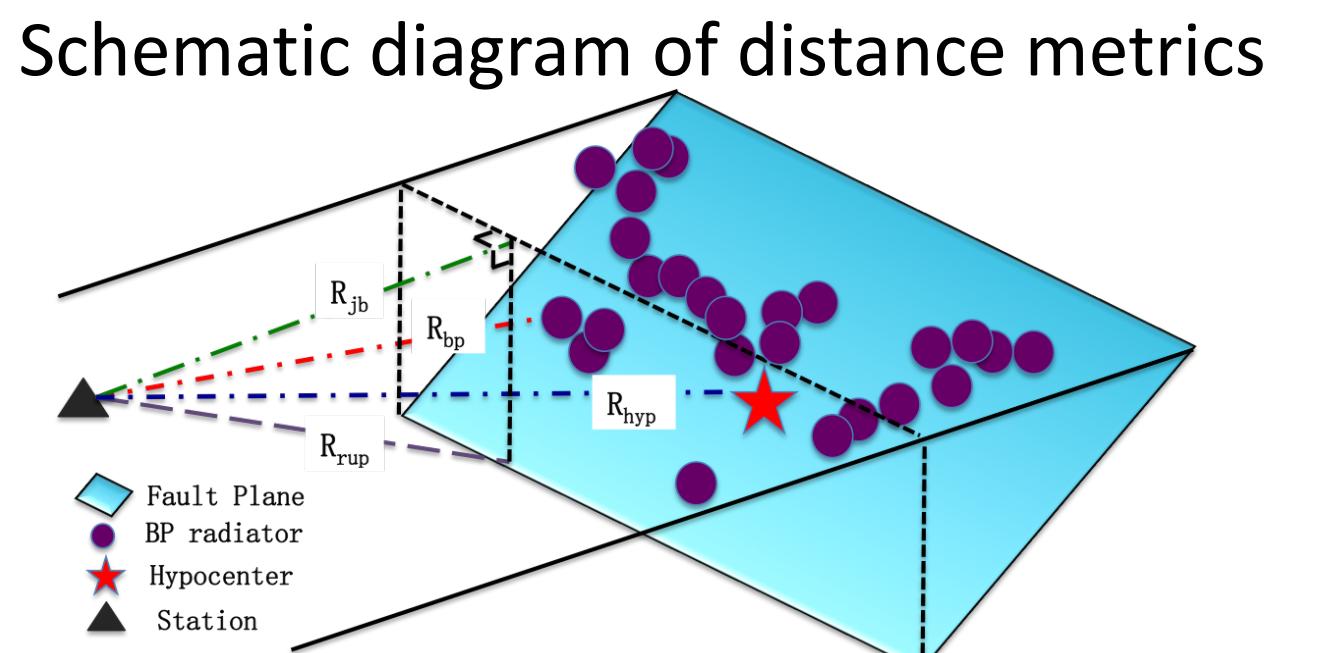


Stewart J P et al., Earthquake Spectra, 2013

The ground-motion misfits are more severe in the high frequency observation (1-10Hz). Since the attenuation relation between distance and ground motion is well established, we hypothesize that the large observed errors can be attributed to a misrepresentation of the source-receiver distance in the high-frequency band. We thus defined a new type of distance measurements, Rbp, the shortest distance from receiver site to the back-projection radiators. Back Projection Radiator (BP radiator) refers to the location where peak energy is released. To avoid the effect of coda waves, we consider that only the significant radiators (with BP power larger than 20% of the maximum) are responsible for the generation of strong ground motions.



Threshold of the significant Back Projection Radiators



## Events & Fault slip models selected

Event	Time	Institute	Mw	Location	Depth (km)	Fault model
Tohoku	2011/03/11	Iris	9.1	38.2963N 142.498E	19.7	Wei(2011)
		USGS		38.297N 142.493E	29	
		JMA	9.0	38.10N 142.86E	24	
		Chu	4.9	38.19N 142.68E	21	
Tokachi	2003/09/26	Iris	8.3	41.815N 142.872E	33	Koketsu (2004)
		USGS		41.815N 143.91E	27	
		JMA	8.0	41.78N 144.07E	42	
		Offshore Honshu (Foothock of Tohoku Eq)		38.4403N 142.9803E	26.2	Hayes (NEIC, offshore Honshu 2011)
Off Miyagi	2005/08/16	Iris	7.2	38.33N 142.28E	8	
		USGS		38.2013N 142.1143E	37.8	Shao and Ji (UCSB, Honshu 2005)
		JMA		38.33N 142.03E	36	
				38.15N 142.28E	42	
Iwate-Miyagi Nairiku	2008/06/13	Iris	6.9	38.34N 140.73E	11.6	Cultura et al. (2013)
		USGS		39.03N 140.88E	7.8	
		JMA	7.2	39.03N 140.88E	8	

Original Boore's GMPE equation ( Boore and Atkinson, 2007):

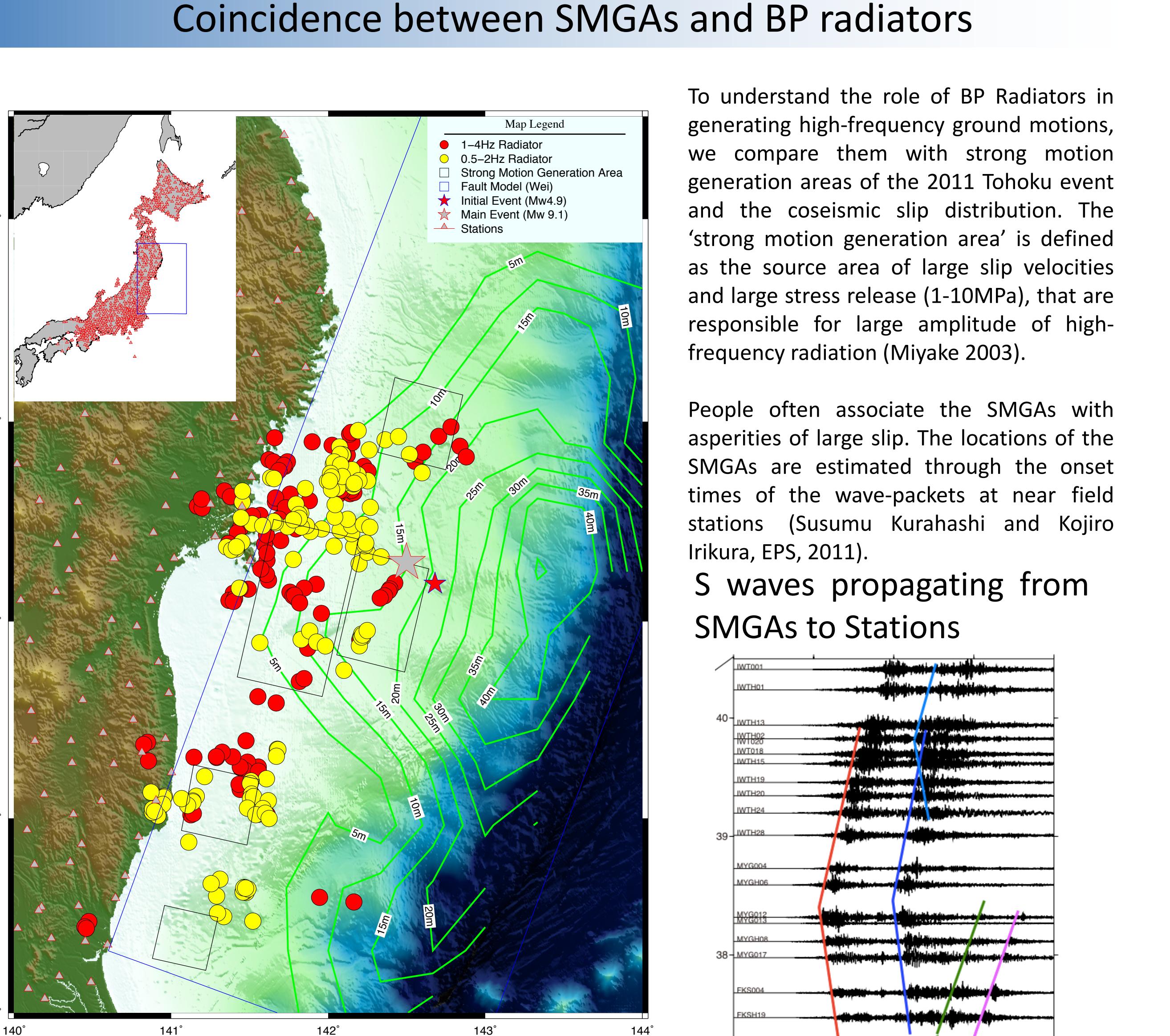
$$\ln Y = F_M(M) + F_D(R, M) + F_S(Vs30, R, M)$$

Where Y can be PSA, PGA and PGV,  $F_M$ ,  $F_D$  and  $F_S$  represent the magnitude scaling, source-receiver distance, and site amplification factor. After transferring the  $F_S$  term from GMPE to left hand side, the observed accelerations after site effect correction is  $PSA' = \frac{PSA}{e^{F_S}}$  and the GMPE becomes:

$$\ln Y = F_M(M) + F_D(R, M)$$

The residual between observation  $PSA'$  and prediction  $\exp(Y)$  describes the accuracy of different measurements.

## Coincidence between SMGAs and BP radiators



To understand the role of BP Radiators in generating high-frequency ground motions, we compare them with strong motion generation areas of the 2011 Tohoku event and the coseismic slip distribution. The 'strong motion generation area' is defined as the source area of large slip velocities and large stress release (1-10MPa), that are responsible for large amplitude of high-frequency radiation (Miyake 2003).

People often associate the SMGAs with asperities of large slip. The locations of the SMGAs are estimated through the onset times of the wave-packets at near field stations (Susumu Kurahashi and Kojiro Irikura, EPS, 2011).

## S waves propagating from SMGAs to Stations

