

5/12/12

①

local shear

$$A(r) = \frac{r}{2} \frac{d\Omega}{dr}$$

local vorticity

$$B(r) = \Omega(r) + A(r) = \Omega(r) + \frac{r}{2} \frac{d\Omega}{dr}$$

$$\kappa^2(r) = \frac{1}{r^3} \frac{d}{dr} [r^2 \Omega(r)]^2 \stackrel{?}{=} 4 B(r) \Omega(r)$$

$$4 B(r) \cdot \Omega(r) = 4 \Omega^2(r) + 2 \Omega r \frac{d\Omega}{dr}$$

$$\frac{d}{dr} [r^4 \Omega^2] = 4 r^3 \Omega^2 + r^4 (2 \Omega) \frac{d\Omega}{dr}$$

$$\frac{1}{r^3} \cdot \frac{d}{dr} [r^2 \Omega(r)]^2 = 4 \Omega^2 + 2 r \Omega \frac{d\Omega}{dr}$$

$$\kappa^2(r) = \frac{1}{r^3} \frac{d}{dr} [r^2 \Omega(r)]^2 \stackrel{\checkmark}{=} 4 B(r) \Omega(r)$$

↑
epicyclic frequency

↑
how the heck is that derived?

Different types of resonances:

Lindblad ← in a Keplerian disk, $\kappa < \Omega$, so how are these different

Corotation ← is the moon in one of these? Earth?

mean motion ← these don't seem to amplify orbital eccentricity, unlike the ~~tidal~~ Lindblad resonances. Not entirely sure why, though...

fr. mean motion resonances?

$$L = 4\pi\sigma R^2 T_{\text{eff}}^4$$

$$T_{\text{eff}}^4 = \left[\frac{L}{4\pi\sigma R^2} \right]^{1/4}$$

$$T_{\text{KO}} = \left[T_{\text{eff}}^4 \left(\frac{3}{4} \kappa_{\text{H}}^{1/2} + \frac{1}{2} \right) \right]^{1/4} = T_{\text{eff}} \left[\frac{3}{4} \kappa_{\text{H}}^{1/2} + \frac{1}{2} \right]^{1/4}$$

$$\cancel{P_{\text{rad}} = \cancel{C_{\text{rad}}} \cdot A_{\text{rad}} \cdot T_{\text{KO}}^4}$$

$$P_{\text{rad}} = \frac{4\sigma}{30} T^4$$

I think we want opacity table #66? But check the input conds you were feeding the code. F & polytropic calcs preceding it: check your input metallicities...