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To do the atmospheric outer boundary cond. calculations, we first need to figure out what the outermost  $\tau, \kappa, T, P, g, R, M, \& L$  values are.

The first block of code in the atmos.F subroutine does this.

$$\text{Start w/ } d\tau = 10^{-3}$$

$$M_{\text{atm}} = M_{\text{max}} - 0.9 M_{\text{max}} \quad (I \text{ think this is what the ATMASS1 \& 2 are about.})$$

The metallicities  $XX \& YY$  are given/supplied as inputs to this calculation, so we don't have to worry  $\&$  determining those fr. scratch.

$$d\tau_{1/2} = 1/2 d\tau = 5 \times 10^{-4}$$

$$I_{\text{tau}23} = 0$$

$$I_{\text{del}M} = 0$$

$$R_{\text{tau}3} = R_{\text{star}}$$

$$z_{\text{tau}3} = M_{\text{max}}$$

$$T_{\text{eff}} = \left[ \frac{L}{4\pi\sigma R_{\text{star}}^2} \right]^{1/4}$$

$\leftarrow L \& R_{\text{star}}$  are also supplied as inputs, here, so we don't have to worry  $\&$  calculating those fr. scratch, either.

$$R_{\text{KO}} = 0$$

$$z_{\text{KO}} = 0$$

$$T_{\text{KO}} = \left[ \frac{3}{4} T_{\text{eff}}^4 \left( \tau + \frac{2}{3} \right) \right]^{1/4} = \left[ T_{\text{eff}}^4 \left( \frac{3}{4} \tau + \frac{1}{2} \right) \right]^{1/2} = T_{\text{out}} \quad \leftarrow \text{in my personal notation, here}$$

$$R_{\text{at}} = R_{\text{star}}$$

$$z_{\text{at}} = z_{\text{M}(N)} = M_{\text{max}}$$

Note in this code,  $z$ -variables typically refer to mass.

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Next, the code claims we "define atmospheric values for  $JK=1$  at  $\frac{1}{2}d\tau$ ." Do that as follows:

$$P_{rad} = c_{rad} \cdot a_{rad} \cdot T_{K0}^4$$

$$P_{out} = \max[2 \cdot P_{rad}, P_{rad} + 10] \leftarrow \text{saying the outermost pressure value in the atmosphere is basically set by the radiation (rather than the gas) pressure, I guess. Maybe these factors } (\times 2, \text{ or } +10) \text{ account in some way for the } \frac{P_{gas}}{P_{rad}} \text{ ratio @ this outermost point?}$$

$$\rho = \rho(T, P, XX, YY, c_{rad}, \text{etc.}) \leftarrow \text{call the eqn. of state function w/ } XX, YY, P_{out} \text{ \& } T_{K0}(T_{out}, \text{basically to find } \rho$$

$$\kappa = \kappa(N, T_{out}) \leftarrow \text{call the opacity funct. to find } \kappa \text{ as a funct. of } T_{out} \text{ \& } \rho.$$

$$GD = \frac{d\tau \cdot G \cdot Z_{at}}{R_{at}^2} = \frac{d\tau \cdot G \cdot M_{star}}{R_{star}^2}$$

$$P_{K0} = P_{out} + \frac{GD}{\kappa} = P_{out} + \frac{d\tau G M_{star}}{R_{star}^2 \kappa} \leftarrow dP = P_{K0} - P_{out} = \frac{d\tau G M}{\kappa R^2}$$

$$\frac{dP}{d\tau} = \frac{GM}{R^2} \cdot \frac{1}{\kappa}$$

↑  
eqn. (5.31)

Assume  $T$  doesn't change in this first, outer  $d\tau$  step (?). Find  $\rho$  at this new  $\tau$  value as a funct. of  $T_{out}$ ,  $P_{K0}$ ,  $XX$ ,  $YY$ .

$$\rho_{new} = \rho(T_{out}, P_{K0}, XX, YY).$$

$$\kappa_{new} = \kappa(T_{out}, \rho_{new})$$

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Still w/in the "define atmos. values for  $TK=1$  @  $\frac{1}{2}dt$ " part of this code, we enter the following ~~for-loop~~ <sup>do-loop</sup> ~~(for-loop?)~~.

while (itr = 1, < 30, itr++):

$$G0 = (PK0 - P_{out})K_{new} - GD \leftarrow = dP \cdot K_{new} - \frac{dEGM_{star}}{R_{star}^2}$$

~~then~~

$$PK1 = 1.001(PK0)$$

$$P_{K1} = S(T_{out}, PK1, XX, YY)$$

$$K_{K1} = K(P_{K1}, T_{out})$$

$$G1 = (PK1 - P_{out})K_{K1} - GD \rightarrow = (1.001 - 1)PK0 \cdot K_{K1} - \frac{dEGM_{star}}{R_{at}^2}$$

$$\begin{aligned} \frac{dG}{dP} &= \frac{G0 - G1}{PK0 - PK1} = \frac{(PK0 - P_{out})K_{new} - GD - (PK1 - P_{out})K_{K1} + GD}{PK0(1 - 1.001)} \\ &= \frac{PK0 \cdot K_{new} - P_{K1}K_{K1} + P_{out}(K_{K1} - K_{new})}{-0.001(PK0)} \end{aligned}$$

$$dP = -\frac{G0}{dG/dP}$$

$$dP = \max(dP, -\frac{1}{2}PK0)$$

$$dP = \min(dP, 0.9PK0)$$

$$dP = \max(dP, 0.8(P_{rad} - PK0))$$

← is  $P_{rad} - PK0 = P_{gas}$ ?  
why is  $\frac{4}{5}P_{gas}$  an important quantity, here?

if itr > 11,  $dP = \frac{1}{2}dP$ . ← more forceful nudge towards convergence, I guess?

$$PK0 = PK0 + dP$$

$$P_{new} = S(PK0, T_{out}, XX, YY)$$

$$K_{new} = K(P_{new}, T_{out})$$

$$\text{if } \left| \frac{dP}{P_{K0} + P_{K1}} \right| < 10^{-5} \text{ } \leftarrow$$

then exit this loop! You've found your outer boundary conds.

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After you find your outer boundary ~~contour~~ values for  $\kappa, \rho, P, T, R, M, \& L$ , then you go into a 4th order Runge-Kutta integration in terms of increasing values of  $\tau$ . I'll ~~do~~ leave my dissection of that part of the code for later, b/c it's conceptually straightforward.

For now, I want to focus of figuring out exactly what's happening on the part of the code I just unpacked, particularly w/in the for-loop. What does it mean for those outer  $P, T, \rho, K$  values to be "converged", physically?

~~$P_{out}, T_{out}, R_{star}, M_{star}, G_{out}, K_{out}, G_0$~~

towards the  
center of  
the star

$\uparrow$

~~$P_{ko}, T_{out}, R_{star}(?), M_{star}(?), G_{ko}, K_{ko}, G_0$~~

$= P_{out} + \frac{GD}{K_{out}}$

$\downarrow \frac{dG}{dR} \rightarrow \downarrow dP$

~~$P_{kl}, T_{out}, R_{star}(?), M_{star}(?), G_{kl}, K_{kl}, G_1$~~

Adjust the PKD "location" (value) using the dP value determined in this step.

if  $\left| \frac{dP}{PK_0 + dP + PK_1} \right| < 10^{-5}$ , then "convergence!"

$$\overline{P} = \frac{PK_0 + PK_1}{2}$$

$$\frac{1}{\bar{P}} = \frac{2}{PK_0 + PK_1}$$

⑤

What is the purpose of this iterative PK0/PK1 process? My guess is that it's trying to bring the PK0 & PK1 values closer together somehow. I also suspect that this process brings the PK0 & PK1 values closer inline w/ the Pout values, although I'm less sure about that. (Need to modify the atmos.F subroutine to output values as it's doing these iterations to check how the PK1 & PK0 values change wrt each other & wrt the initial Pout value.)

$M_{\text{star}}$  and  $R_{\text{star}}$  (and  $T_{\text{out}}$ ) don't change throughout this process.  $g$ ,  $P$ ,  $\kappa$ , &  $GD/G0/G1$  do, however. Since  $GD/G0/G1$  basically depend on  $g$ , though, that is why those values change. The amount of mass enclosed in the outermost cell depends on  $g$  (since neither  $R_{\text{star}}$  nor  $M_{\text{star}}$  change - this is all my guess, though), which is why the grav. accel. values at the 3 locations change.

What you're really doing in this analysis, I think, is determining the density in the outermost cell, since that sets both the opacity ( $\kappa$ ) and thus the optical depth ( $\tau$ ). The density determination depends on the grav. accel. ( $GD/G0/G1$ ), which in turn depends on the pressure ( $PK0, PK1, P_{\text{out}}(?)$ ) x thru the  $\frac{dP}{d\tau}$  relation.

Next step: modify the atmos.F subroutine to output info  $\tau$ ,  $P_{\text{out}}$ ,  $PK0$ ,  $PK1$ , densities, opacities, etc. on each pass thru the iteration loop to check/see precisely how those values evolve towards convergence or an answer or... something.