

Page: 1

## point\_source

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#### Abstract

Calculates the xspec normalization per square arcminute for the Cosmic X-ray background after removal of point sources to some limiting level (min\_flux). ELF is shorthand for the logN-logS.

## 1 Instruments/Modes

	Instrument	Mode	
EPIC		Imaging	

### 2 Use

pipeline processing	no
interactive analysis	yes

# 3 Description

Calculates the xspec normalization per square arcminute for the Cosmic X-ray background after removal of point sources to some limiting level (min\_flux). ELF is shorthand for the logN-logS.

We assume that integral[ELF] from 0 to infinity is less than cxrb\_norm, or that:

 $CXRB = X + integral[ELF]_0^i n finity.$ 

Therefore, for any given "blank sky" observation where the brightest point source has a flux smax, the total X-ray emission in the field would be:

 $X + integral[ELF]_0^s max.$ 

The currently available functions are:

hms: Hasinger, Miyaji, & Schmidt (2005), from ROSAT, XMM, & Chandra):

mushotzky: (????):

cappelluti: Cappelluti et al (2008), from COSMOS:

mateos: Mateos et al (2008), from XMM'

Output: Xspec normalization for power law in units of photons/cm2/s/am2/keV

### Examples::

point-source func=mateos min\_src\_flux=5.e-14 cxrb\_norm=10.6 index=1.40

## 4 Parameters

This section documents the parameters recognized by this task (if any).

Parameter Mand Type Default Constraints	
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func	yes	string	mateos	

Name of logN-logS function.

hms (Hasinger, Miyaji, & Schmidt 2005)

mushotzky (REF TBD)

cappelluti (Cappelluti et al. 2008)

mateos (Mateos et al. 2008)

min_src_flux	yes	real	1.0e-14		
Source flow out off in anglem 2/g					

Source flux cutoff in erg/cm2/s.

cxrb_norm	yes	real	10.6		
Normalization of the accraig V ray hadraround					

Normalization of the cosmic X-ray background.

index	yes	real	1.46	
D1				

Photon power law index.

- 5 Input Files
- 6 Output Files
- 7 Algorithm
- 8 Comments

## References