Sky Viewer Investigations for LSST EPO

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Abstract. The Large Synoptic Survey Telescope (LSST) is an 8-m optical ground-based telescope being constructed on Cerro Pachon in Chile. LSST will survey half the sky every few nights in six optical bands. LSST Education and Public Outreach (EPO) is investigating requirements and technologies to create a next generation sky viewer. We discuss and survey current sky viewer technologies and identify new features required for an LSST Skyviewer.

1. Introduction

LSST Education and Public Outreach (EPO) reports on the early investigations of using and implementing a sky viewer across its software systems. We contracted with Theresa Neil: Strategy + Design (TNSD)¹ to develop a proof of concept sky viewer to test with a variety of intended users of the EPO Portal in order to understand and achieve the needs of our sky viewer audiences. We researched available open-source sky viewer software to identify software engineer requirements and features for our implementation (Jacoby et al. 2017; Emmons et al. 2017). Over the summer we prototyped an implementation of this using Aladin and Aladin-lite (Boch & Fernique 2014) and evaluated other available tools.

2. Requirements

The sky viewer has three software engineering requirements.

- 1. Build on an existing sustainable open-source project.
- 2. Simple to extend and integrate with existing tools and libraries.
- 3. Usable in multiple contexts (e.g. mobile, web, science notebooks, planetarium).

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3. Implementation

Aladin v10.009 and Hipsgen were used to create a prototype survey and sky viewer with HSC Public Data Release 1 UDEEP COSMOS data (2 sq. deg.) (Aihara et al. 2017). This was reprocessed by the LSST DM team using the LSST DM pipeline (Jurić et al. 2015). Hierarchical Progressive Survey (HiPS, Fernique et al. 2015) creation took 16 hours on a single node using 8 vCPUs and 16 GB of RAM to process 127 GB of calibrated exposure FITS images across three bands (i, r, g) into a 16.7 GB color PNG image HiPS (HiPS Norder=3-12). To create multicolor HiPS individual monocolor HiPS must first be created. From these monocolor HiPS a multicolor HiPS is then composed. This method increases linearly with disk usage and processing time for each band being processed. This is in contrast to other survey creation methods which do not require the creation of intermediate monocolor surveys.

4. Comparison

Aladin and Aladin Lite use HEALPix (Calabretta & Roukema 2007). HEALPix is supported in multiple libraries in Python 3. Python 3 has seen extensive use at LSST (Jenness, T. 2018) and is used in our Jupyter (Pérez & Granger 2007) science notebook investigations. HiPS and MOC (Fernique et al. 2015), which use HEALPix, are IVOA standards we plan to support. Finally, HEALPix is acceptable for use with the LSST science platform (Dubois-Felsmann et al. 2017) and our formal education teams.

Montage (Berriman & Good 2017) was evaluated to create Tessellated Octahedral Adaptive Subdivision Transform (TOAST) for the World Wide Telescope (WWT) (Goodman et al. 2012) and Leaflet based surveys. Montage supports HEALPix FITS files as input data but it does not create HiPS. WWT, TOAST and Montage were not chosen for two reasons. First, Montage does not create HiPS. Second, TOAST has no viable stand-alone viewer and is only available as part of WWT as a monolithic application. This makes it difficult to customize and provide a sky viewer that is usable in multiple contexts. LSST EPO is interested in WWT support for data cubes, 3D and virtual reality. We compare three types of sky viewers in table 1: Aladin and Aladin

	Aladin & Aladin-lite	Leaflet based sky viewers	World Wide Telescope (WWT)
Projection	HEALPix	Mercator	TOAST
Tile Manager	HiPS	Leaflet	WWT
Tile Creation	Aladin	ImageMagick, custom	Montage
Examples	ESA, CDS, ² DES, ³ LIGO ⁴	DECaLS, ⁵ CFHTLS, ⁶ HSC-SSP, ⁷ NASA, SDSS ⁸	WWT web client

Table 1. Types of sky viewers

Lite, World Wide Telescope (WWT) and Leaflet based sky viewers.

5. User Interface

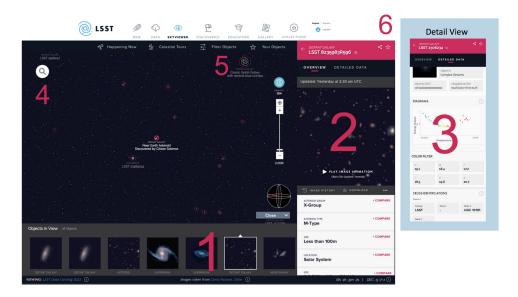


Figure 1. Planned design mockup features for the LSST Skyviewer. Features are labeled as Gallery(1), Object slider(2), Detail slider(3), Search(4), Filter(5) and Collections(6).

EPO contracted design work and user experience testing from TNSD. Through user experience testing TNSD identified weaknesses with existing sky viewer design and iteratively improved the design. Users found sky viewer use was not intuitive and simple. This was resolved by adding a gallery and search. Users wanted information about the object being viewed. This was resolved by adding the object and detail sliders. Finally users wanted a way to filter and save what they viewed. This was resolved by adding filters and collections. The figure 1 is the result.

²Centre de DonnÃes astronomiques de Strasbourg

³Dark Energy Survey

⁴Laser Interferometer Gravitational-Wave Observatory

⁵DECam Legacy Survey

⁶Canada-France-Hawaii Telescope Legacy Survey

⁷Hyper Suprime-Cam Subaru Strategic Program

⁸Sloan Digital Sky Survey

6. Conclusion

EPO plans to implement the LSST Skyviewer as modular open-source software components. These components will build on IVOA standards and should be reusable for different audiences through configuration and extension. Components will be used by our portal and science notebook systems. Since our science notebook system will extend the LSST science platform, our components will be available to LSST scientists.

We plan to use Aladin Lite using HiPS. Improvements in either Hipsgen or Montage to render HiPS more efficiently are necessary. Additionally, WWT using TOAST may still be considered as an alternative if there is a stand-alone viewer.

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References

Aihara, H., et al. 2017, ArXiv e-prints. 1702.08449

Berriman, G. B., & Good, J. C. 2017, PASP, 129, 058006. 1702.02593

Boch, T., & Fernique, P. 2014, in Astronomical Data Analysis Software and Systems XXIII, edited by N. Manset, & P. Forshay, vol. 485 of ASP Conf. Ser., 277

Calabretta, M. R., & Roukema, B. F. 2007, MNRAS, 381, 865

Dubois-Felsmann, G., et al. 2017, LSST Science Platform, Tech. rep., Large Synoptic Survey Telescope. URL https://ldm-542.lsst.io

Emmons, B., et al. 2017, LSST EPO Design, Tech. rep., Large Synoptic Survey Telescope. URL https://docushare.lsst.org/docushare/dsweb/Get/LEP-31

Fernique, P., et al. 2015, A&A, 578, A114. 1505.02291

Goodman, A., et al. 2012, in Astronomical Data Analysis Software and Systems XXI, edited by P. Ballester, D. Egret, & N. P. F. Lorente, vol. 461 of ASP Conf. Ser., 267. 1201.1285

Jacoby, S., et al. 2017, LSST Education and Public Outreach Requirements, Tech. rep., Large Synoptic Survey Telescope. URL https://docushare.lsst.org/docushare/dsweb/Get/LSE-89

Jenness, T. 2018, in ADASS XXVII, edited by TBD (San Francisco: ASP), vol. TBD of ASP Conf. Ser., TBD

Jurić, M., et al. 2015, ArXiv e-prints. 1512.07914

Pérez, F., & Granger, B. E. 2007, Computing in Science and Engineering, 9, 21. URL http://ipython.org