Environmental Systematics and the Impact on 21-cm Epoch of Reionization Measurements

by

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A Thesis Presented in Partial Fulfillment of the Requirements for the Degree Bachelor of Science with Honors

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Abstract

1. Introduction

1.1. The Epoch of Reionization

Immediately after the Big Bang, the expanding universe was a hot, but cooling, plasma of fundamental particles. Approximately 380,000 years later, the universe had cooled and expanded enough for electrons to become bound to atomic nuclei, making the universe transparent to light and emitting the cosmic microwave background. Immediately following this period and lasting until ~ 400 million years after the Big Bang was the cosmic Dark Ages, when the universe was quite literally dark—no stars or other sources of light had yet formed. Then, with the formation of the first luminous sources, the intergalactic medium (IGM), which had been predominantly neutral until then, was ionized by the energetic radiation from these sources during the Epoch or Reionization (EoR). The EoR ended around a billion years after the Big Bang, and the universe's structure has continued evolving into what we now see today. This history is summarized in Figure 1. I should probably have a reference here.

Furlanetto et al. (2006)

Morales & Wyithe (2010)

Pritchard & Loeb (2012)

1.2. The 21-cm Power Spectrum

1.3. The Hydrogen Epoch of Reionization Array

The Hydrogen Epoch of Reionization Array (HERA; Figure 2) is an experiment to study the periods prior to and during the EoR (from redshifts $z \sim 6-30$) via measurements of the redshifted 21-cm line (DeBoer et al., 2017).

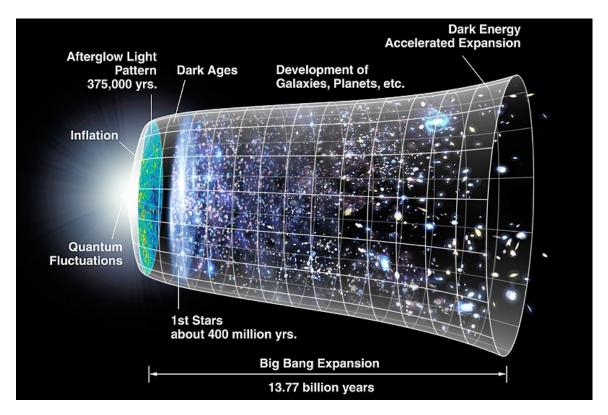


Figure 1. Timeline of the evolution of the universe. Image courtesy of NASA.

1.4. Radio Frequency Interference

2. Methods

- 2.1. RFI Excision Strategies
- 2.2. Calculating the Power Spectrum
 - 2.3. Modelling HERA Data
 - 3. Results
 - 4. Conclusion
 - 5. References

DeBoer, D. R., Parsons, A. R., Aguirre, J. E., et al. 2017, PASP, 129, 045001

Furlanetto, S. R., Oh, S. P., & Briggs, F. H. 2006, PhR, 433, 181



Figure 2. HERA as of late 2017 – early 2018. HERA will observe large-scale structure prior to and during the EoR via the redshifted 21-cm line from the IGM. Image courtesy of the South African Radio Astronomy Observatory.

Morales, M. F., & Wyithe, J. S. B. 2010, ARA&A, 48, 127

Pritchard, J. R., & Loeb, A. 2012, RPPh, 75, 086901