Advanced Quantum Mechanics

exercise 1

We define the operators: $H_0:=\frac{p^2}{2m}-\frac{e^2}{r}$, $H':=\sqrt{m^2c^2+p^2c^{2l}}-\frac{e^2}{r}$ additionally: $X:=\frac{p^2}{m^2c^2}$

Now if makes sense to define $f: (-1, \infty) \to \mathbb{R}$, $f(x) := \sqrt{1+x}$ for all $x \in (-1, \infty)$ and expanding f(x) for $x \in (-1, \infty)$ into a taylor series.

taylor series. $=) \quad \int_{k=0}^{\infty} \frac{f(0)}{k!} x^{k} \quad (we assume that this series is absolut convergent for the given values)$

We approximate f up to second order: $f(x) \approx 1 + \frac{x}{2} - \frac{x^2}{8}$ $\Rightarrow \int \left(\frac{\rho^2}{m^2c^2}\right) \approx 1 + \frac{\rho^2}{2m^2c^2} - \frac{\rho^4}{8m^4c^4}$

=>
$$\sqrt{m^2c^4 + p^2c^2} \approx mc^2 + \frac{p^2}{2m} - \frac{p^4}{8m^3c^2}$$

Then we neglect the constant term and define the new approximated Hamiltonian by

$$\widetilde{H} := \frac{p^2}{2m} - \frac{p^4}{8m^3c^2} - \frac{e^2}{r}$$

additionally we define a perturbation: $S:=-\frac{p^4}{8m^3c^2}$

Now we are able to use first-order perturbation theory. We define $\widetilde{H}: \mathbb{L}0.1] \rightarrow \mathbb{L}(H, \mathbb{K})$ where \mathbb{K} is our used hilbert space and $\mathbb{L}(H, \mathbb{K})$ is the set of operators. For all $1 \in \mathbb{L}0.1$ we say $\widetilde{H}(\lambda) := H_0 + \lambda S$

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Turther we define 19(0)): [0,1] -> X, E:[0,1] -> 1R
         F(21 |9(2)) = E(1) |9(2)> for all 2 = [011]
Now we assume that we can expand E and 18(0) into
       |\varphi(\lambda)\rangle = \sum_{k=0}^{\infty} \lambda^{k} |\varphi_{k}\rangle, E(\lambda) = \sum_{k=0}^{\infty} \lambda^{k} E_{k}
      for all de [o,1] with 19h) EXE, EXER for all kello.
Here we only use first-order approximation.
       = > |9(2)| = |9_0| + |1/9_1|
                                                for all de [0,1]
              E(\lambda) \approx E_0 + \lambda E_1 \quad (30), (92) \in \mathcal{H}, E_0, E_1 \in \mathbb{R}
Using the definition of A, 18(1), E we get
   H(2) (40) ~ (Ho+25) (190) + 2 (90)
               = Ho (90) + 1 (Ho (91) + S(90)) + 12 S(91)
≈ E(1) (9(11) ≈ (E,+ 1En) (190) + 1 (91)
                 = Eol 90) + d(Eol91) + E190) + d' E191)
 Comparing the coefficients of It for ke {0.1.23 this results in
        (I) Hold) = Eolyo)
        (II) Ho (9x) + S/90) = Eo/91) + E193)
        (III) E_{\Lambda}(\mathcal{G}_{\Lambda}) = S(\mathcal{G}_{\Lambda}) (this equation has no real
                                    meaning because we did only
                                      a first order approximention)
(I) says that 1907 and Eo are eigenfunction and eigenvolve
  (energy) of the unperturbed system described by Ho
(II) shall be multiplied by (90) to get Ex:
    <9.(H. 19,) + <9.1519) = <9.(E.19,) + <9.(E.19)
                              = Eo (9,191)
    = E. (90191)
    => E = < 9.16(90)
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This means Ex is the first-order energy-shift of a given system perturbed by S.

Now we write 5 in operator form: $S = -\frac{\pi^4}{8mc^2}\Delta^2$

Then we rewrite a in sperical coordinates and use the knowledge that (90) should be the ground state of an hydrogen atom.

Therefore we can amit the derivatives of angles.

$$=\frac{2}{r}\left(-\frac{2}{r^{2}}\partial_{r}+\frac{2}{r}\partial_{r}^{2}\right)+\frac{2}{r}\partial_{r}^{3}+\left(\frac{4}{r^{3}}\partial_{r}+\frac{-2}{r^{2}}\partial_{r}^{2}-\frac{2}{r^{2}}\partial_{r}^{2}+\frac{2}{r}\partial_{r}^{3}\right)+\partial_{r}^{4}$$

The groundstate 190 can be written as $90(r) = Ae^{-ar}$ for constants $A \in \mathbb{R}$ and $a \in (0,\infty)$.

$$= > E_{1} = \langle 9_{0} | 5 | 9_{0} \rangle = \frac{-h^{4}}{8m^{2}} \int_{0}^{\infty} \int_{0}^{T} \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) \left(J_{r}^{4} + \frac{4}{r} J_{r}^{3} \right) g_{0}(r) r^{2} sin \vartheta \int_{0}^{2\pi} g_{0}^{*}(r) r^{2} sin \vartheta \int_{0}^{2$$

$$= 4\pi A^{2} \cdot \frac{-h^{4}}{8m^{3}c^{2}} \cdot \int_{0}^{\infty} e^{-\alpha r} \left(\partial_{r}^{4} r + \frac{4}{r} \partial_{r}^{3}\right) e^{-\alpha r} \cdot r^{2} dr$$

$$\int_{0}^{\infty} r^{2} e^{-2\alpha r} dr = \frac{2}{(2\alpha)^{3}} = \frac{1}{4\alpha^{3}} \qquad \int_{0}^{\infty} r e^{-2\alpha r} dr = \frac{1}{4\alpha^{2}}$$

$$= > E_1 = \frac{-4\pi A^2 h^4}{8m^3c^2} \left(\frac{\alpha^4}{4a^3} - \frac{4a^3}{4a^2} \right) = \frac{3\pi a A^2 h^4}{8m^3c^2}$$

ABER:
$$A = ?$$
 (-1.0)

Exercise 2 Let (X, <.1.) be a hilbert space. We define the unperturbed Hamiltonian $\mathcal{H}_0:\mathcal{H}\to\mathcal{H}$ as $\mathcal{H}_0:=\frac{\rho^2}{2m}+\frac{m\omega^2}{2}\chi^2=\hbar\omega\left(a^{\dagger}a+\frac{1}{2}\right).$ For a small parameter $\lambda \in (0,\infty)$ we porturb the system with the following quarter potential operator V: H -> H, V:= 1664 (a+a+) 4 for le e C and qat & L(H, H). Therefore the Hountsonian H for the perturbed system can be defined as H: H -> H $H = H_0 + \lambda V$ Determination of 6: Let A, B & L (H, H) be two operators. Then it holds that (A+ cB) (A-cB) = A2-cAB+cBA-c2B2 = A2 - i [A,B] + B2 with [A,B] := AB - BA $=) A^2 + B^2 = (A + iB)(A - iB) + i[A,B]$ Use this proposition for Ho:

 $= \Rightarrow H_0 = \frac{p^2}{2m} + \frac{m\omega^2\chi^2}{2} = \left(\sqrt{\frac{m\omega^2}{2}} \left(\chi + i\sqrt{\frac{1}{2m}}\right) \left(\sqrt{\frac{m\omega^2}{2}}\right) - i\sqrt{\frac{1}{2m}}\right)$ + ([/mw2/X, /1/p]

$$= \frac{m\omega^{2}}{2} \left(x + \frac{ip}{m\omega} \right) \left(x - \frac{ip}{m\omega} \right) + \frac{i\omega}{2} \left[x_{i}p \right]$$

We already know: [x,p] = ih 1 for the one-dimensional case. Additionally it holds that ata = aat-1.

=>
$$\hbar\omega\left(a^{\dagger}a + \frac{1}{2}\right) = \hbar\omega\left(aa^{\dagger} - 1 + \frac{1}{2}\right) = \hbar\omega\left(aa^{\dagger} - \frac{1}{2}\right) = H_0$$

=>
$$\alpha \alpha^{+} = \sqrt{\frac{m\omega}{2\pi}} \left(x + \frac{ip}{m\omega} \right) \cdot \sqrt{\frac{n\omega}{2\pi}} \left(x - \frac{ip}{m\omega} \right)$$

k was given that:
$$16 \frac{6}{4} \times \frac{4}{4} = (ata^{\dagger})^{\frac{4}{4}} = 2 \frac{2}{4} \times \frac{2}{4} = ata^{\dagger}$$

$$a + a^{\dagger} = \sqrt{\frac{2m\omega}{2n}} 2x = \sqrt{\frac{2m\omega}{n}} \times = 2 \frac{1}{2} \sqrt{\frac{2m\omega}{n}} = \sqrt{\frac{m\omega}{2n}}$$

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Deformination of terrus Pa and Pz:
   (a+a+) = [(a+a+)2]2 = [aa + aa+ ata + ata+]
    = \left[ aa + (N+1) + N + atat \right]^2 = \left[ aa + atat + 2N+1 \right]^2
    = aaaa + aaatat + aa (2N11)
    + atataa + atatatat + atat (2N+1)
    + (2N+1) aa + (2N+1) a tat + (2N+1) (2N+1)
   = a4 + a+4 + [aa (2N+1) + (2N+1)aa]
                    + [a+a+(2N+1)+ (2N+1)a+a+]
                   + atataa + aaatat + (2N+1)(2N+1)
(I)
      [a_iN] = aN - Na = aa^{\dagger}a - Na = (aa^{\dagger} - N)a
              = (N+1-N)a = a
         => aN = Na + [a,N] = (N+1) a
         => Na = aN - [a,N] = a(N-1)
(II) [a^{\dagger}, N] = a^{\dagger}N - Na^{\dagger} = a^{\dagger}N - a^{\dagger}aa^{\dagger} = a^{\dagger}(N - aa^{\dagger}) = -a^{\dagger}
        => a^{\dagger}N = Na^{\dagger} + La^{\dagger}N? = (N-1)a^{\dagger}
        => Nut = atN - [at, N] = a^{\dagger}(N+\Lambda)
   =) ataa = at(N)a \stackrel{(E)}{=} (N-1)ata = (N-1)N = N^2-N
   =) aaa^{\dagger}a^{\dagger} = a(N+1)a^{\dagger} \stackrel{(I)}{=} (N+2)aa^{\dagger} = (N+2)(N+1) = N^2 + N + 2N + 2
   (2N+1)(2N+1) = 4N^2 + 2N + 2N + 1
  => a^{\dagger}a^{\dagger}aa + aaa^{\dagger}a^{\dagger} + (2N+1)(2N+1) = 6N^2 + 6N + 3 = 3(2N^2+2N+1)
   aa(2N+1) \stackrel{(I)}{=} a[2(N+1)+1]a = a(2N+3)a
  (2N+1)aa = a[2(N-1)+1]a = a(2N-1)a
    => aa(2N+1) + (2N+1)aa = a(4N+2)a = a[2(2N+1)]a
   => atu+ (2N+1) + (2N+1)atu+ (I) a+[(2N-1) + (2N+3)]a+= a+[2(2N+1)]a+
 => A = a4+ a+4+ a[2(2N+1)]a + a+[2(2N+1)]a+ + 3(2N2+2N+1)
  -> P_{\lambda}(N) = 3(2N^{2}+2N+1), P_{\lambda}(N) = 2(2N+1)
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Computation of energy shift in first-order perturbation theory: We already know: $a(9n) = \sqrt{n7(9n-n)}$ for all $n \in \mathbb{N}$, a(9n) = 0at Isn > = Vn+11 Isn+x > for all neNo with Isn' Ed for all ne No as the eigenstates of Ho. ({19n7 (ne No) forms an orthonormal basis of H) Additionally we know from exercise $\Lambda: E_n^{(1)} = \langle g_n^{(0)} | V | g_n^{(0)} \rangle$ where $|g_n^{(0)}\rangle := |g_n\rangle$ for all $n \in N_0$ and $E_n^{(1)}$ is the first-order energy shift of the state with quantum number nello in perburbation => $E_n^{(4)} = \frac{1}{16E^q} \left\langle \mathcal{G}_n^{(\omega)} / \Delta \left| \mathcal{G}_n^{(\omega)} \right\rangle = \frac{4\hbar^2}{16m^2\omega^2} \left\langle \mathcal{G}_n \left| \Delta \left| \mathcal{G}_n \right\rangle = \hbar\omega \cdot \frac{\hbar}{4m^2\omega^2} \right\rangle \left\langle \mathcal{G}_n \left| \Delta \left| \mathcal{G}_n \right\rangle \right\rangle$ = tw. th (g(N) 19n) + aP(N) a 19n) + a+P(N) a+ 19n) + a 4 (9n) + a + 4 (9n)) for all ne No $a^{+4}|S_n\rangle = \sqrt{\frac{(n!4)!}{n!}}|S_{n+4}\rangle, \quad a^{4}|S_n\rangle = \begin{cases} 0 &: n < 4\\ \sqrt{\frac{n!}{(n-4)!}}|S_{n-4}\rangle &: else \end{cases}$ for all $a^{\dagger}P_{2}(N) a^{\dagger}(\mathcal{G}_{h}) = P_{2}(n+1) \sqrt{\frac{(n+2)!}{n!}} |\mathcal{G}_{n+2}\rangle$ neNo. here A(N) $\alpha P_{2}(N) \alpha |9_{u}\rangle = \begin{cases} 0 & : n < 2 \\ P_{2}(n-1) \sqrt{\frac{n!}{(n-2)!}} |19_{n-2}\rangle & : else \end{cases}$ and P2(N) mean operators and Pala), $P_{\lambda}(N)(\mathcal{S}_n) = P_{\lambda}(n)(\mathcal{S}_n)$ P2(n) mean polynoms UM No with the summe $= \sum E_n^{(1)} = \hbar \omega \cdot \frac{\hbar}{4m^2 \omega^3} P_{\lambda}(n)$ for all $ne N_0$ (use $\langle 9_n | 9_m \rangle = 8_{nm}$ coephienes for all nime Wo) => first-order perturbation energy levels:

=) first-order perturbation energy levels: (with $E_n^{(o)}$ as energy levels of unperturbed system for all $n \in \mathbb{N}_o$)

= $\hbar \omega \left(\frac{1}{2} + n + \frac{\hbar \lambda}{4m^2 \omega^3} 3(2n^2 + 2n + 1) \right)$ for all $n \in \mathbb{N}_o$

AUF 2: 6.0/6.0