Numerical Relativity: Simulating the Evolution of Binary Black Holes

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Introduction

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Example:

➤ SXS: Gravitational Lensing of GW150914 → Web Version

What is Numerical Relativity?

science about numerical solutions to the Einstein field equations

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Example:

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What is Numerical Relativity?

science about numerical solutions to the Einstein field equations

Why?

prediction and explanation of cosmic phenomena and relativistic instabilities



Background

Einstein Field Equations

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$$

Einstein Field Equations

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$$

Vacuum ($T_{\mu\nu}=0$):

$$R_{\mu\nu} = 0$$

Analytical Solutions

Schwarzschild Metric:

$$(g_{\mu\nu})(t,r,\vartheta,\varphi) = \operatorname{diag}\left(-\left(1-\frac{2M}{r}\right), \left(1-\frac{2M}{r}\right)^{-1}, r^2, r^2 \sin^2\vartheta\right)$$

Other Solutions:

- Kerr metric
- Kerr-Newman metric

Geodesics

$$\ddot{x}^i + \Gamma^i_{mn} \dot{x}^m \dot{x}^n = 0$$

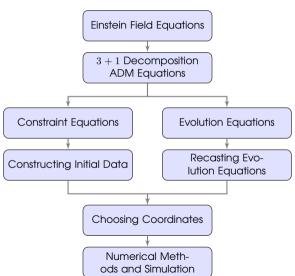
freely-falling test particles and light move along geodesic curves

Examples:

- ► Newton versus Schwarzschild trajectories
- ▶ Kerr-Newman-Orbit

The Basic Idea

The Basic Idea



The Details

3+1 Decomposition

3+1 Decomposition: The ADM Equations

Constraint Equations:

$$R + K^2 - K_{ij}K^{ij} = 16\pi\rho$$
$$D_j \left(K^{ij} - \gamma^{ij}K\right) = 8\pi S^i$$

$$\rho \coloneqq n_a n_b T^{ab} , \qquad S^i \coloneqq -\gamma^{ij} n^a T_{aj}$$

3+1 Decomposition: The ADM Equations

Evolution Equations:

$$\partial_t \gamma_{ij} = -2\alpha K_{ij} + D_i \beta_j + D_j \beta_i$$

$$\partial_t K_{ij} = \alpha \left(R_{ij} - 2K_{ik} K_j^k + K_{ij} \right) - D_i D_j \alpha - 8\pi \alpha \left[S_{ij} - \frac{1}{2} \gamma_{ij} \left(S - \rho \right) \right]$$

$$+ \beta^k D_k K_{ij} + K_{ik} D_j \beta^k + K_{kj} D_i \beta^k$$

$$S_{ij} := \gamma_{ia}\gamma_{jb}T^{ab}$$
, $S := \gamma^{ij}S_{ij}$

Constructing Initial Data

Constructing Initial Data: CTT Decomposition

Solve Constraint Equations for W^i and ψ :

$$\left(\bar{\Delta}_L W\right)^i - \frac{2}{3} \psi^6 \bar{\gamma}^{ij} \bar{D}_j K = 8\pi \psi^{10} S^i$$
$$8\bar{D}^2 \psi - \psi \bar{R} - \frac{2}{3} \psi^5 K^2 + \psi^{-7} \bar{A}_{ij} \bar{A}^{ij} = -16\pi \psi^5 \rho$$

$$\bar{A}^{ij} = \bar{A}^{ij}_{\mathrm{TT}} + \left(\bar{L}W\right)^{ij}$$

Constructing Initial Data: CTT Decomposition

Construct Physical Solution:

$$\gamma_{ij} = \psi^4 \bar{\gamma}_{ij}$$

$$K_{ij} = A_{ij} + \frac{1}{3} \gamma_{ij} K$$

$$= \psi^{-2} \bar{A}_{ij} + \frac{1}{3} \gamma_{ij} K$$

Recasting The Evolution Equations

Recasting the Evolution Equations: The BSSN Formalism

Idea:

$$\gamma_{ij} = e^{4\varphi} \bar{\gamma}_{ij}$$

$$K_{ij} = e^{4\varphi} \tilde{A}_{ij} + \frac{1}{3} \gamma_{ij} K$$

$$\bar{\Gamma}^i := \bar{\gamma}^{jk} \bar{\Gamma}^i_{jk} = -\partial_j \bar{\gamma}^{ij}$$

Choosing Coordinates

Choosing Coordinates: The Lapse and Shift

Geodesic Slicing:

$$\alpha = 1 \; , \qquad \beta^i = 0$$

Choosing Coordinates: The Lapse and Shift

Geodesic Slicing:

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Maximal Slicing:

$$K = 0 = \partial_t K$$
 \Longrightarrow $D^2 \alpha = \alpha \left[K_{ij} K^{ij} + 4\pi \left(\rho + S \right) \right]$

Numerical Methods

Numerical Methods: Classification of PDE

$$A\partial_1^2 \varphi + 2B\partial_1 \partial_2 \varphi + C\partial_2^2 \varphi = \rho$$

Elliptic:

$$AC - B^2 > 0$$

Parabolic:

$$AC - B^2 = 0$$

Hyperbolic:

$$AC - B^2 < 0$$

Numerical Methods: Finite Difference Method

Examples of Stencils:

$$\left[\partial_1 \varphi\right]_{i,j} = \frac{[\varphi]_{i+1,j} - [\varphi]_{i-1,j}}{2h} + \mathcal{O}\left(h^2\right)$$

$$\left[\partial_1^2\varphi\right]_{i,j} = \frac{-\left[\varphi\right]_{i-2,j} + 16\left[\varphi\right]_{i-1,j} - 30\left[\varphi\right]_{i,j} + 16\left[\varphi\right]_{i+1,j} - \left[\varphi\right]_{i+2,j}}{12h^2} + \mathcal{O}\left(h^4\right)$$

Results

Visualizing the results:

- ➤ SXS: BBH head-on collision ➤ Web Version
- ➤ SXS: Highly precessing BBH ➤ Web Version

Going Further:

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